Xarray and Aperture Array Optimisations

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- AA Configuration design space
- Xarray phased array simulator
- Optimisation Strategy (Thibault Clavier)
- Simulation Efforts
- Future work

Overview

Sensitivity

 Depends on diameter, number of elements, and their configuration

Beam-width (Calibration)

 We can increase this by either reducing station size or using a tapering function but at the cost of A/T

Side lobes (Noise suppression)

- We can reduce this by tapering and irregular configurations like GRS
- Filling Factor (one used figure of merit)

Configuration design space

Xarray - Fast station-level phased array simulator/optimiser based on an irregular 2D-FFT implementation

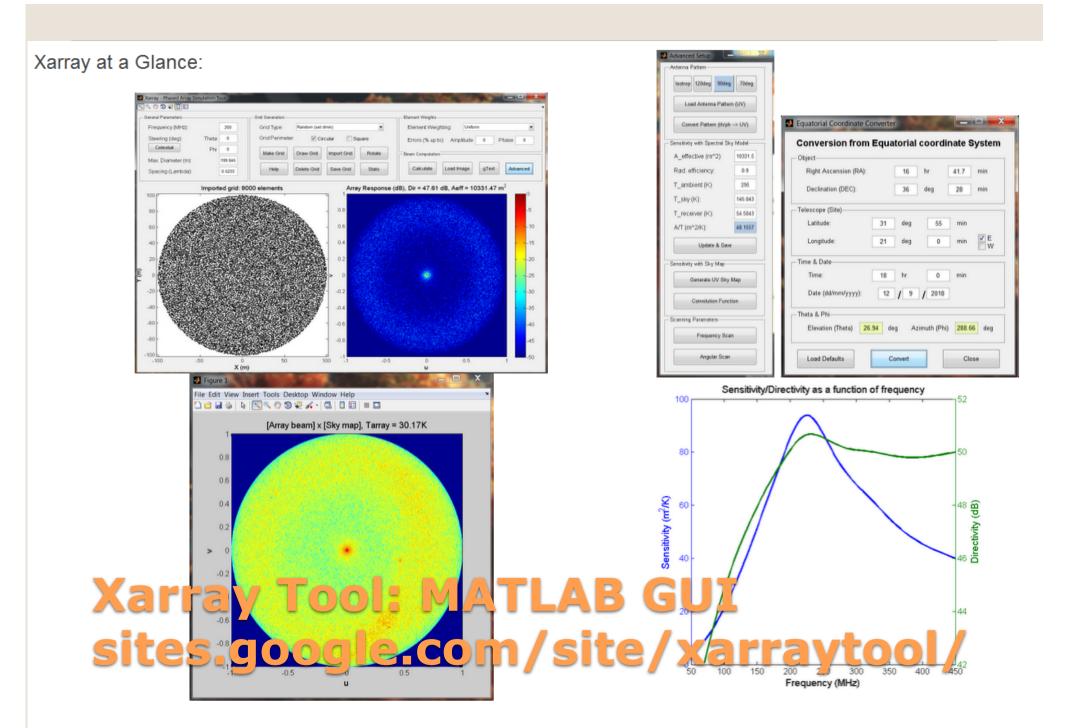
Xarray - Phased Array Simulation Tool				Advanced Setup
N <				Antenna Pattern
General Parameters	Grid Generation	Element Weights		Isotrop 120deg 90deg 70deg
Frequency (MHz): 100	Grid Type: Random (Fast Gen)	Element Weighting: Uniform	_	Isotrop Izoteg Sodeg Todeg
Steering (deg): Theta 20.15	Grid Perimeter: 🔽 Circular 🔲 Squa	errors (% up to): Amplitude P	hase	Load Antenna Pattern (UV)
Celestial Phi 111.73	Make Grid Draw Grid Import Grid	Rotate Beam Computation		Convert Pattern (th/ph> UV)
Max. Diameter (m): 84.581	Help Delete Grid Save Grid	Stats Calculate Load Image gText	Advanced	Sensitivity with Spectral Sky Model
Spacing (Lambda): 0.45			Advanced	A_effective (m^2): 5513.22
Random Grid (3): N=2035, dmin(avg)=1.456m, Diam=84.6m, FFg=0.83 Power (dB), Dir=38.52 dB, Aeff=5091.50 m ² , FF=0.86, Ω=2.09°x2.18°				Rad. efficiency: 0.9
40 -				T_ambient (K) 295
	0.8		5	T_sky (K): 957.78
30 -			10	T_receiver (K): 125.778
20 -				A/T (m ² /K): 5.41957
	0.4		15	Update & Save
10 -			<mark></mark> -20	Sensitivity with Sky Map
Ê 0-	(@) 0.2 (@) uis (@) uis (@) uis		25	Generate UV Sky Map
			-20	
-10 -	-0.2		30	Convolution Function
-20 -	-0.4		35	Scanning Parameters
-20	-0.6			Frequency Scan
-30 -	-0.0		40	Angular Scan
	-0.8		<mark>-</mark> -45	
-40			-50	Fix Angular Resolution
-40 -30 -20 -10 0 X (m	10 20 30 40	-0.5 0 0.5 u = sin(θ)cos(φ)	1	NFFT Pixel size (^2): 2048

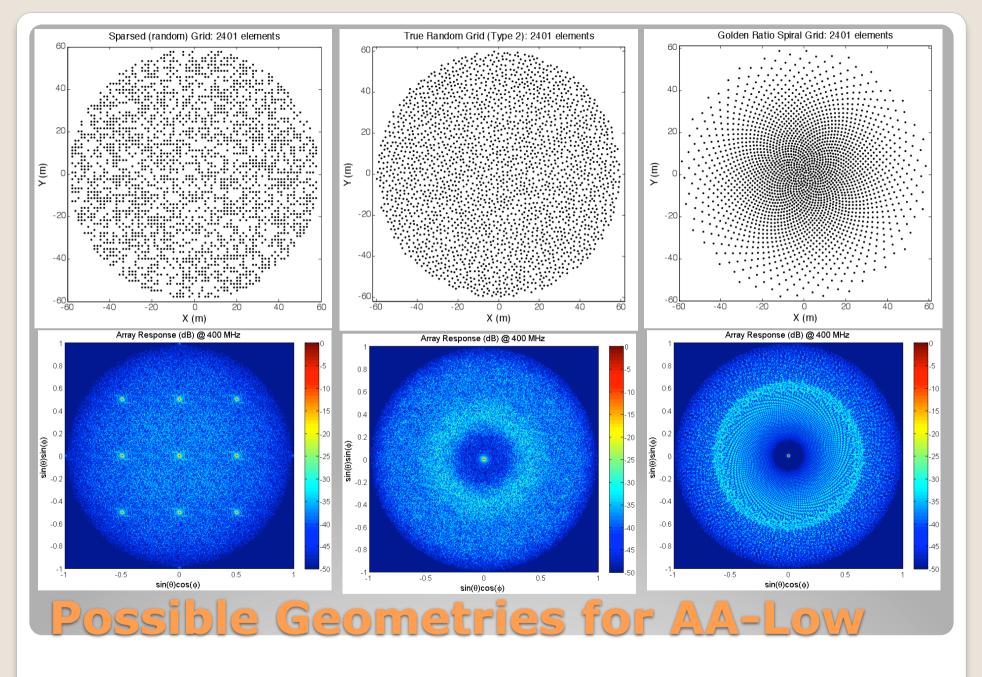
- General parameters: design frequency, steering (spherical or equatorial coordinates), diameter, inter-element spacing
- Type of arrays: regular/triangular, thinned, sparse, concentric rings, random (3 types), golden ratio spirals, spatial tapering
- Grid features: import any, draw/edit, rotation (gradual or full)
- Element weighting: Gaussian, Dolph-Chebyshev, Taylor, percentage errors (amplitude and/or phase)
- Element type: import any embedded pattern, hemispherical, 120°, 90°, 70°

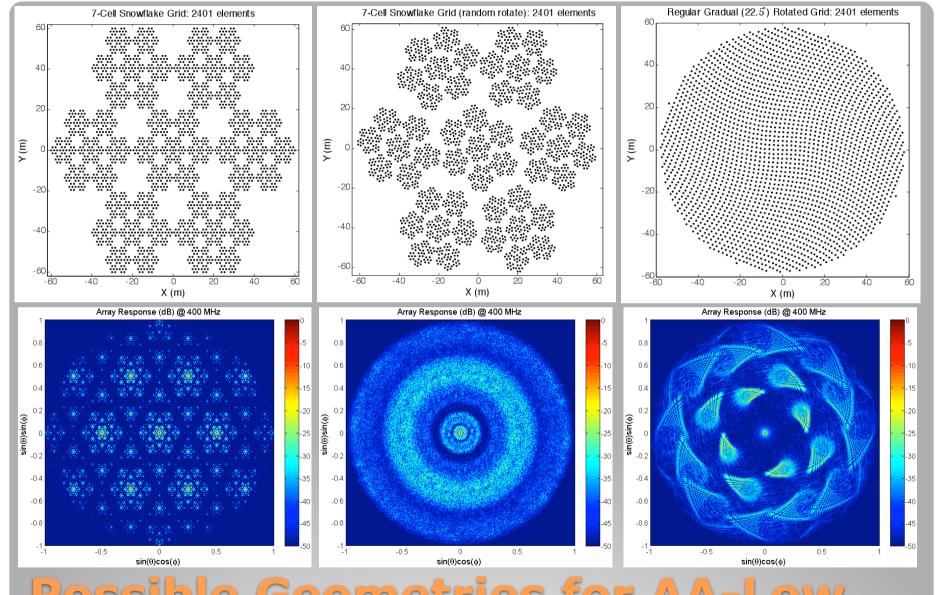
Xarray input parameters

- Scanning in frequency or angle assuming flat or realistic sky brightness distribution (scaled version of Haslam 408MHz map)
- Beam output in FITS format
- Stored parameters: Directivity/Aeff, Tsys, A/T, filling factor, beamwidths, max and mean sidelobe level
- Statistics of XY configuration
- Optimisation parameters (coming soon)
- Beam modelling parameters (future work)

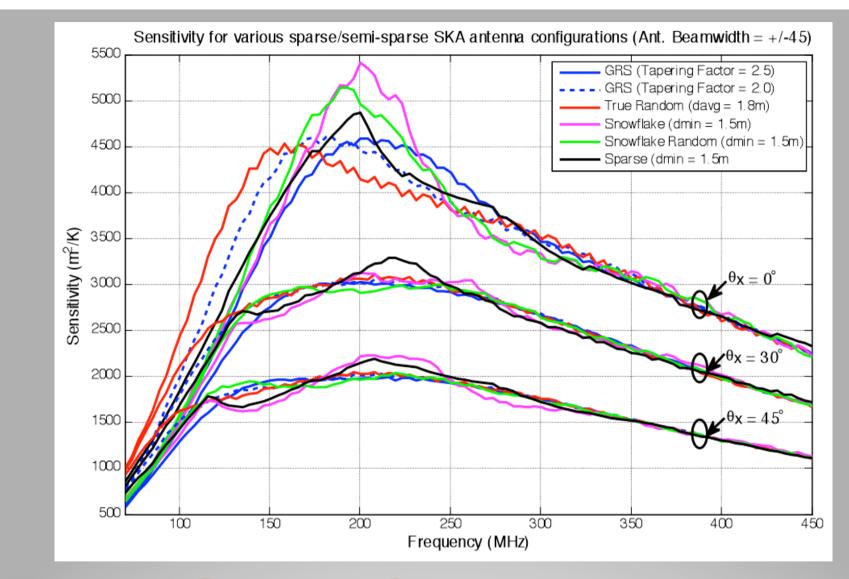
Xarray output parameters





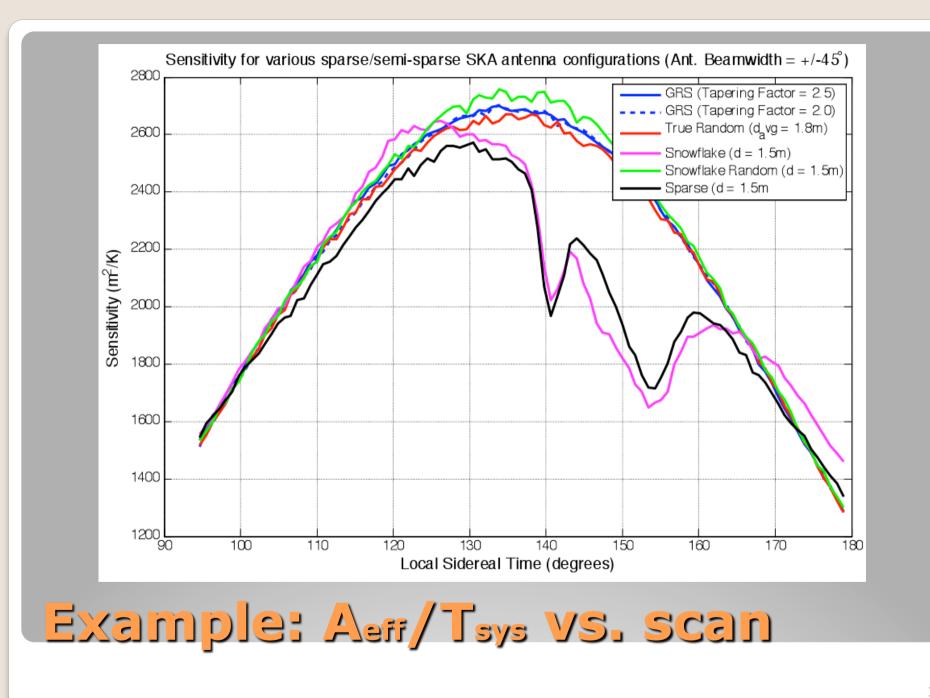






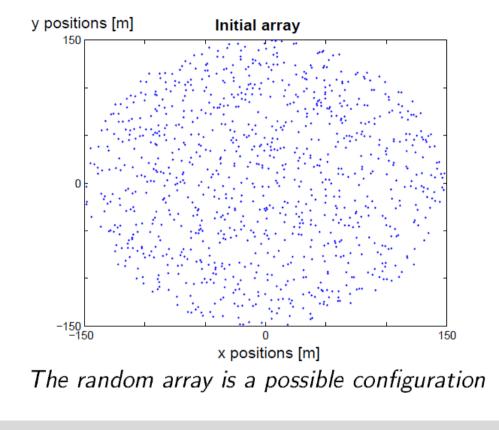
Example: Aeff/Tsys

10



AA element optimisation by Thibault Clavier

The Master's Thesis focusses on the design of one station



Question How to modify the positioning of a given array to improve its performance ?

The objective is to minimize a weighting of the side lobes

Constraints

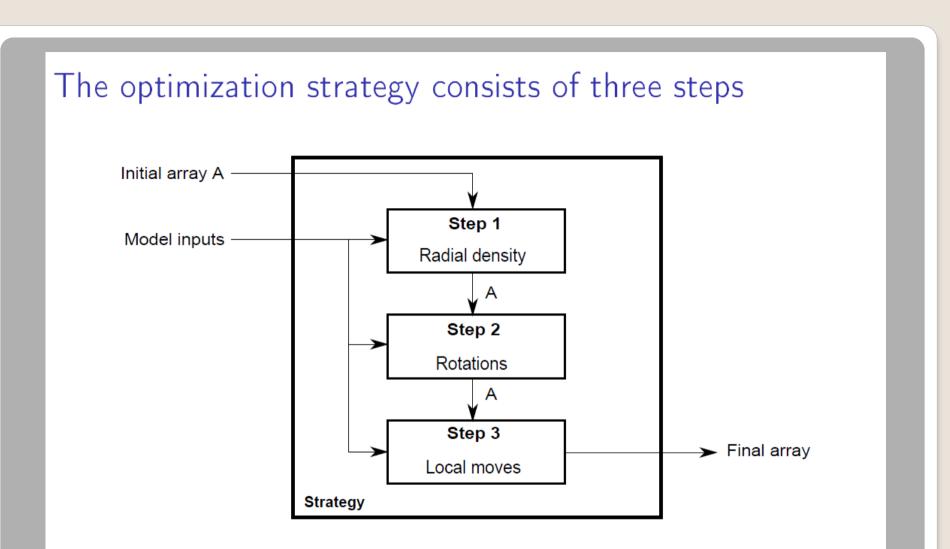
Main beam width is fixed to w_{mb} Minimal distance between antennas is δ

Cost function

$$\min_{x_{\boldsymbol{n}},y_{\boldsymbol{n}}} \left[\int_{\mathcal{U}} \left(W_{\boldsymbol{e}}(u_{\boldsymbol{x}},u_{\boldsymbol{y}}) |R_{\boldsymbol{A}}(u_{\boldsymbol{x}},u_{\boldsymbol{y}})|^2 \right)^p du_{\boldsymbol{x}} du_{\boldsymbol{y}} \right]^{1/p}$$

where p measures the importance of high peaks W_e is a weight function on \mathcal{U}

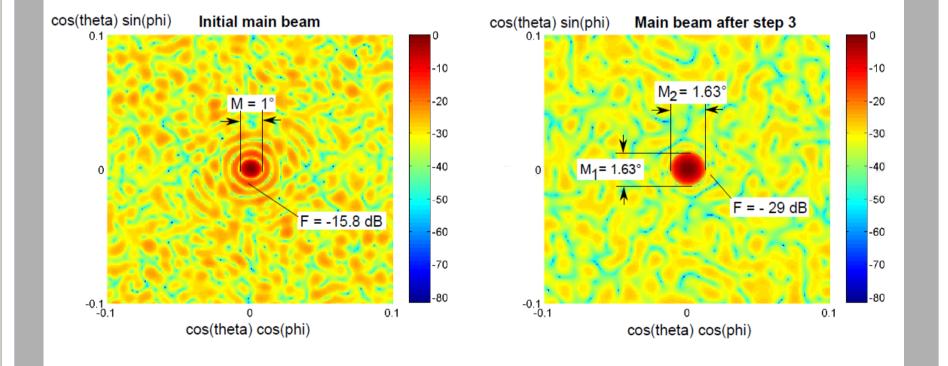
 Use *I_p* norm to optimise for a defined array and interferometric weighting function *W_e*



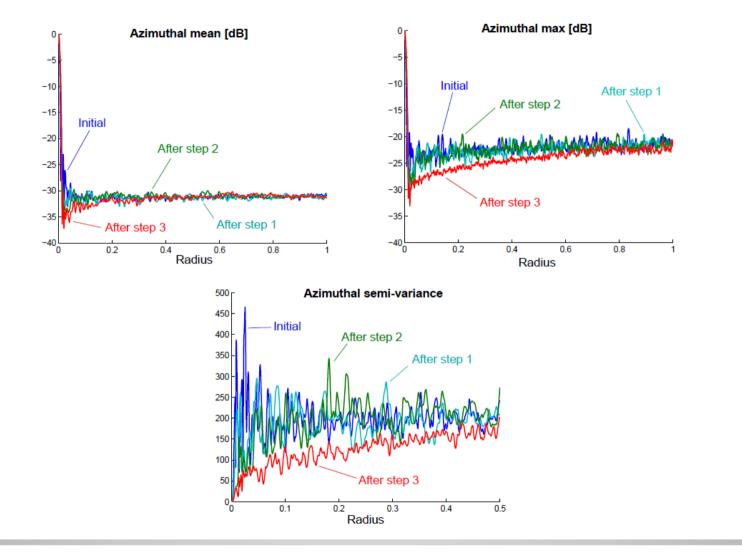
The model inputs are the required main beam width w_{mb} , the minimal separation δ , the parameters p and e

The main beam width is morphed

$$w_{mb} = 1.64^{\circ}$$



The statistics show a marked improvement on the side lobes



- MBF-MOM (element and station level characterisation)
- Xarray (station level)
- AA optimiser (station + interferometric level)
 OSKAR2 (interferometric level)

Still to come...

- Comparison tools to be developed: figures of merit need to be defined and an assessments carried out.
- How does this link with calibration assessment?

AA simulation effort

- What is defined by a "good" station beam
 - Define better metrics!
 - Fix beamwidth, low sidelobe or localised sidelobes through optimisation
 - How calibratible is it?
- High FF core station design: ~850m sea of elements and their practical considerations (how to change elements, paths through elements, physical perimeters and coupling between elements)



- Define a useful programme for station beam optimisation – work with UCL
- Update Xarray to incorporate compact representations of main beam and first few sidelobes.
- Utilise more accurate simulations of the station beam based on MBF approach (collaboration with UCL)
- Use computational framework for SKA simulator (OSKAR2) to evaluate antenna configurations and develop comparison tool (ongoing work with Oxford)

Future work and collaborations

Thank You.