

Relative numbers of Radio Loud vs Radio Quiet AGN

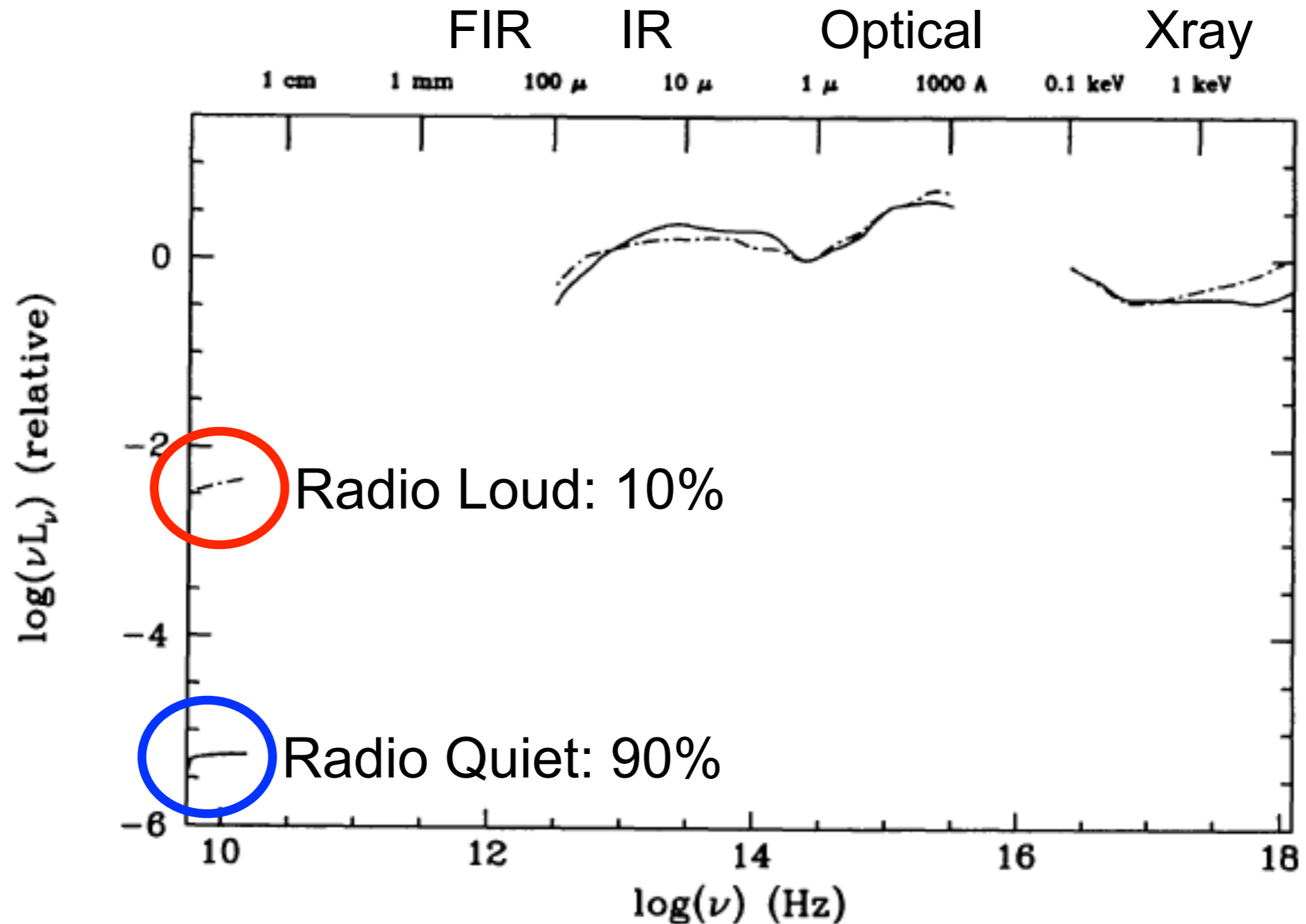
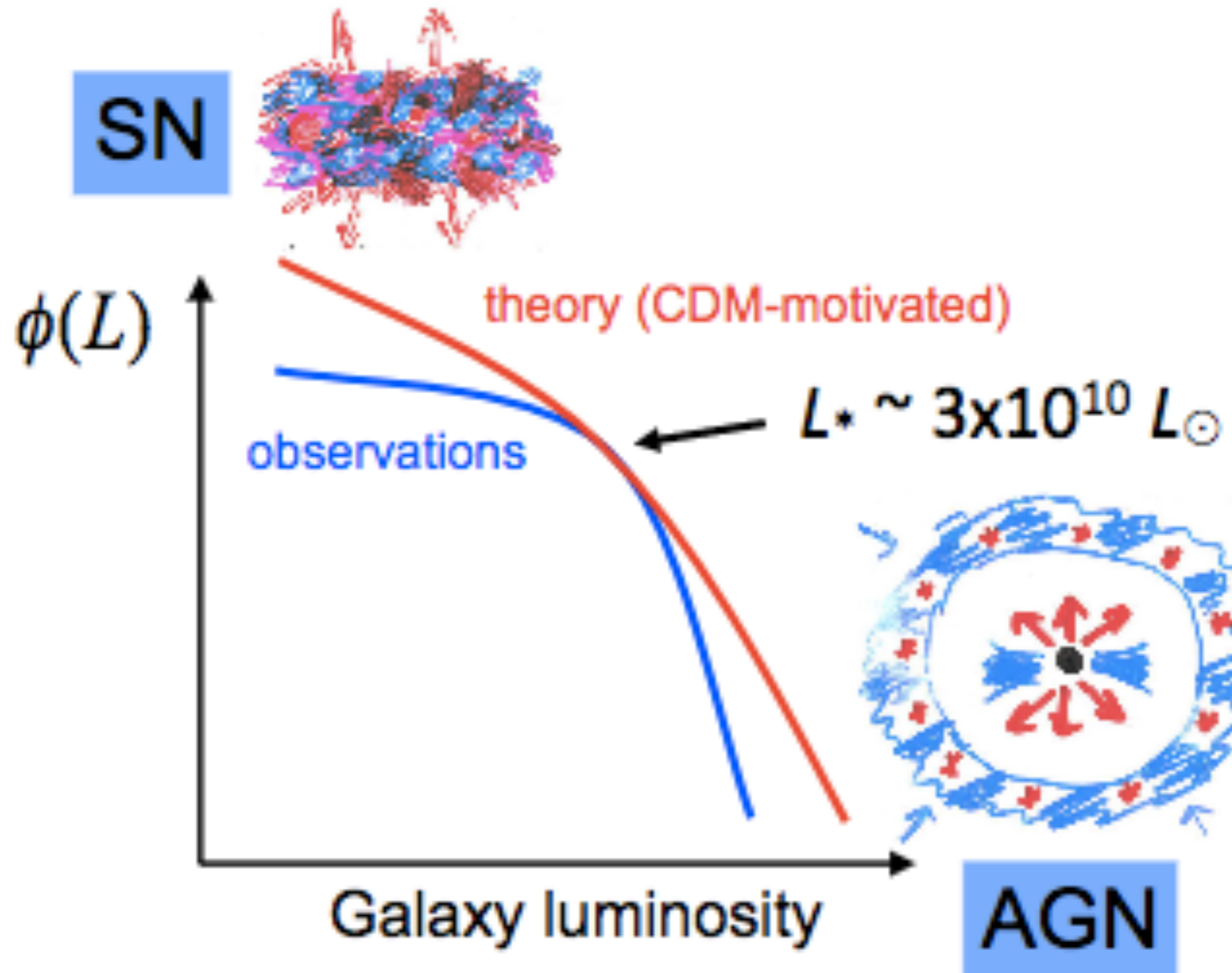


FIG. 10.—The mean quasar energy distribution, normalized at 1.25 μm , for radio-loud (*dashed line*) and radio-quiet (*solid line*) quasars. Spectral regions where few or no data are available are omitted. The radio-loud distribution has a rising X-ray spectrum in this plot, while the radio-quiet X-ray spectrum is horizontal.

Relative importance of radio feedback vs radiative feedback



Discussion - Radio quiet and weak radio sources

1. Radio loud vs radio quiet

- Role of BH spin and mass for RL/RQ
- Role of mergers/environment for BH spin
- Can RL/RQ help understand FRI/FRII?

2. Feedback without jets

- Do you need a jet to influence large scales? No - EELRs
- Do you need a jet to physically move material? No - BALs
- What about winds? Shocks?

3. Host galaxy properties

- Different hosts for different types
- LERGs with no optical indication of AGN activity

4. Jet feedback in sources with low/intermediate radio power

- How significant are the effects of jet-induced feedback in intermediate radio power sources?
- What are the optical morphologies of the hosts of intermediate power sources, and how are such sources triggered?