

Radio Polarimetry

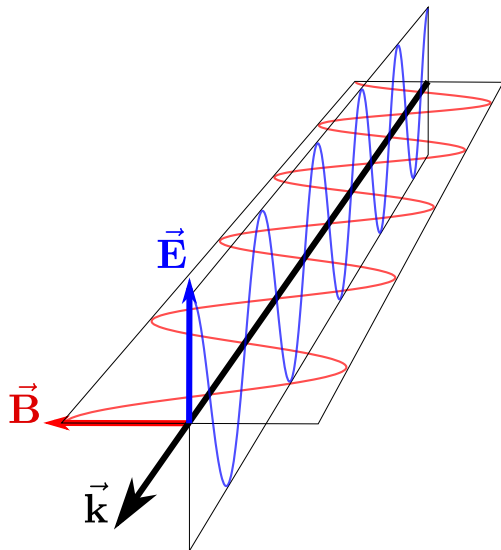
Michiel Brentjens

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ASTRON, Dwingeloo, The Netherlands

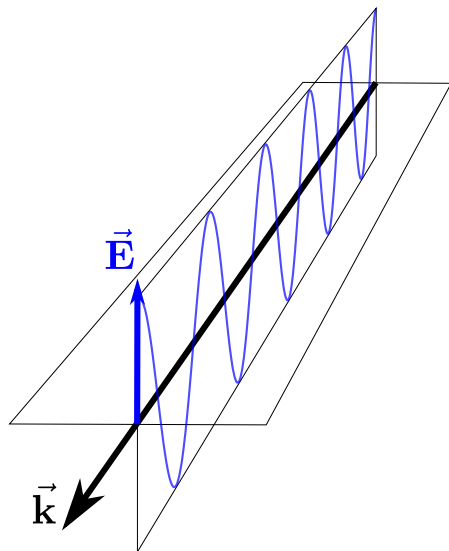
September 9, 2013

- Born & Wolf *Principles of optics*
- Thompson, Moran & Swenson *Interferometry and Synthesis in Radio Astronomy*
- Taylor, Carilli & Perley *Synthesis Imaging in Radio Astronomy II*
- Bracewell *The Fourier Transform & Its Applications*
- Hamaker, Bregman & Sault *Understanding radio polarimetry: paper I*(1996)
- Sault, Hamaker & Bregman *paper II*(1996)
- Hamaker & Bregman *paper III* (1996)
- Hamaker *paper IV* (2000)
- Hamaker *paper V* (2006)
- Brentjens & de Bruyn *Faraday rotation measure synthesis* (2005)

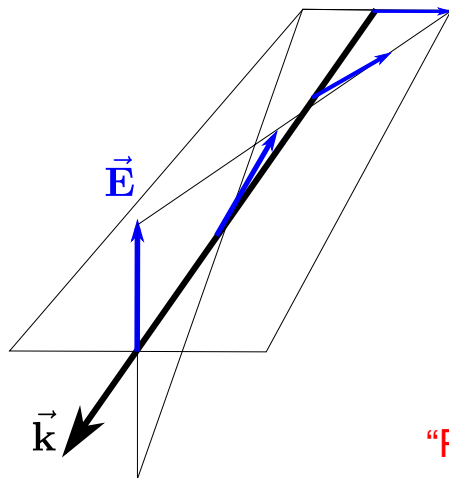
- 1 EM wave physics
- 2 Polarized EM-waves
- 3 Interferometric polarimetry
- 4 Messy reality
- 5 An example



- **Vector** phenomenon
- From Maxwell's equations:
 $\hat{\mathbf{k}} = \hat{\mathbf{E}} \times \hat{\mathbf{B}}$
- We know \mathbf{k} (yesterday)
- Measure either \mathbf{E} or \mathbf{B}



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- \mathbf{E} is easier (yesterday)

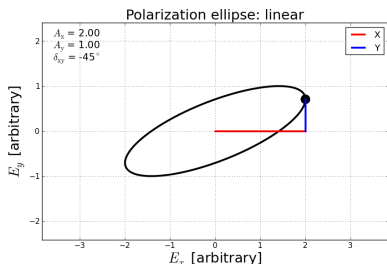


- **Vector** phenomenon
- From Maxwell's equations:
 $\hat{\mathbf{k}} = \hat{\mathbf{E}} \times \hat{\mathbf{B}}$
- We know \mathbf{k} (yesterday)
- Measure either \mathbf{E} or \mathbf{B}
- \mathbf{E} is easier (yesterday)
- But:
- E_x and E_y **not equal**
- \mathbf{E} may **rotate** as function of \mathbf{x} and t .
- \mathbf{E} traces **ellipse**

“Polarization”

- 1 EM wave physics
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Geometry



Viewing from antenna towards source, watching orientation and length of \mathbf{E} vector on a plane at a fixed location in space.

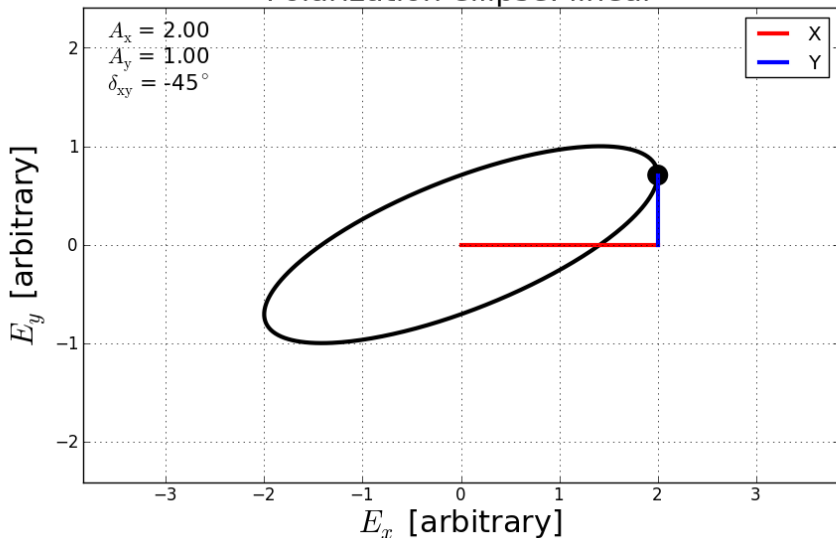
$$\mathbf{E} = E_x \hat{\mathbf{e}}_x + E_y \hat{\mathbf{e}}_y$$

$$E_x = A_x \cos(2\pi\nu t + \delta_x)$$

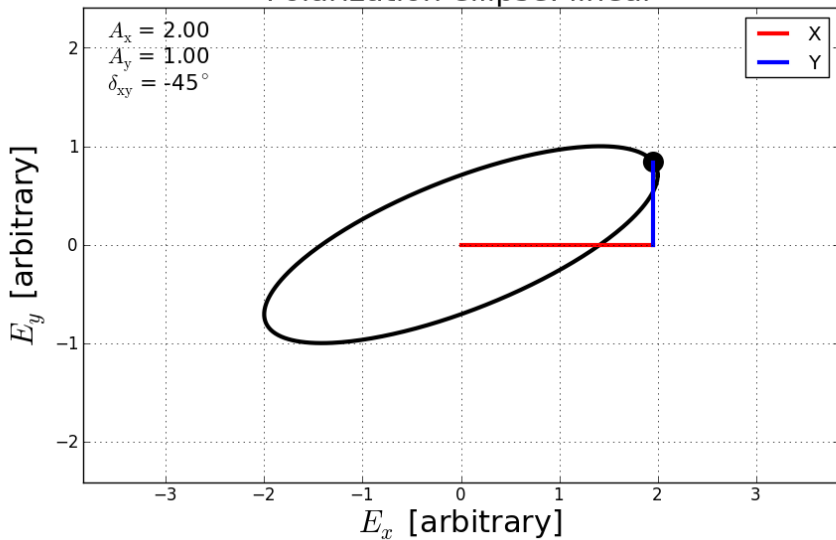
$$E_y = A_y \cos(2\pi\nu t + \delta_y)$$

- $A_x = x$ -amplitude
- $A_y = y$ -amplitude
- $\delta_{xy} = \delta_y - \delta_x$
- $\delta_{xy} =$ measure of ellipticity
- $\delta_{xy} > 0$: CW rotation \Rightarrow LEP
- $\delta_{xy} = 0$: linear polarization
- $\delta_{xy} < 0$: CCW rotation \Rightarrow REP

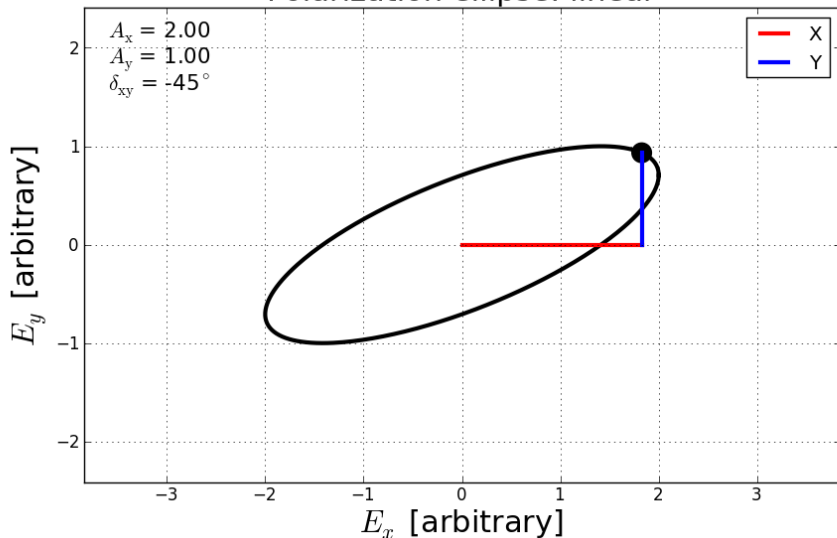
Polarization ellipse: linear



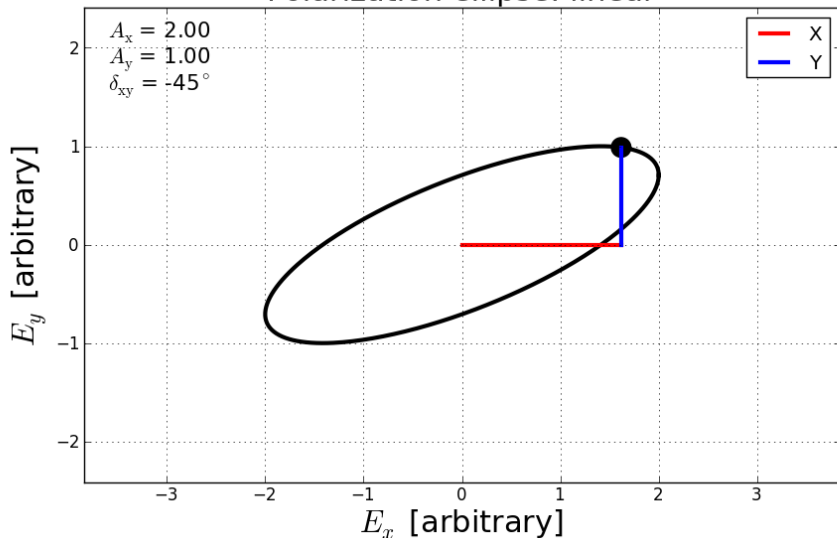
Polarization ellipse: linear



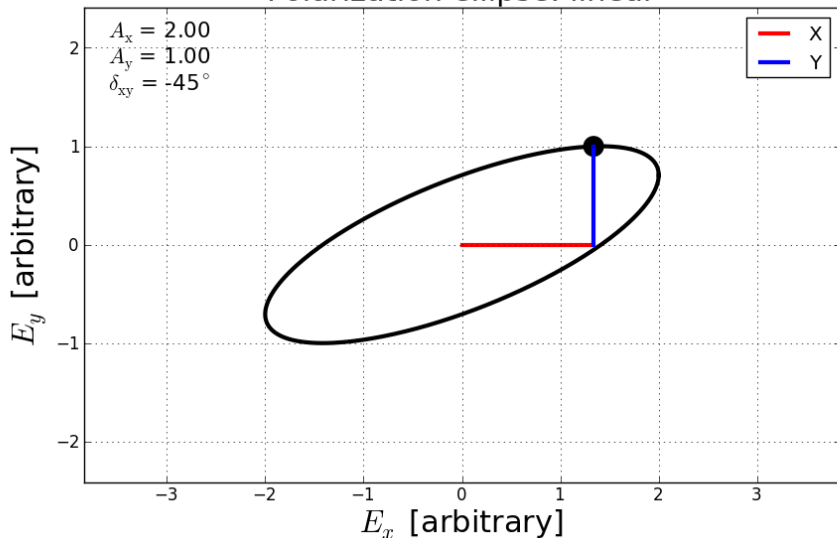
Polarization ellipse: linear



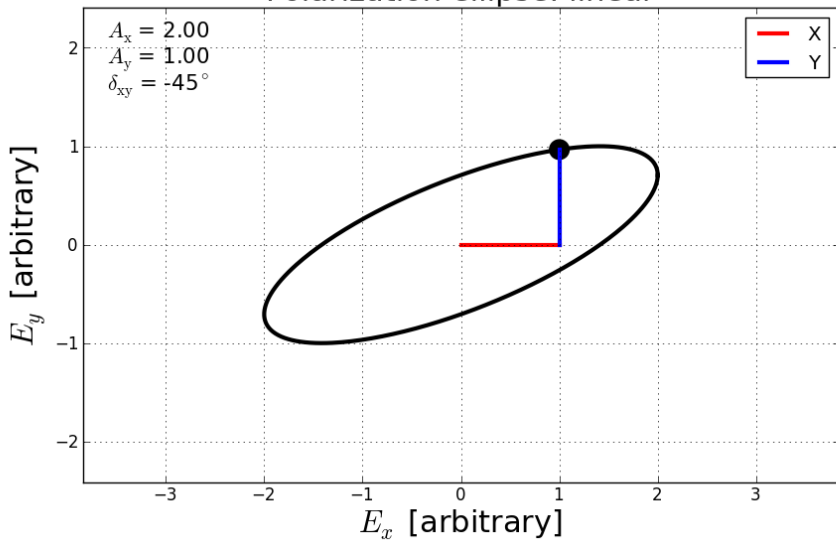
Polarization ellipse: linear



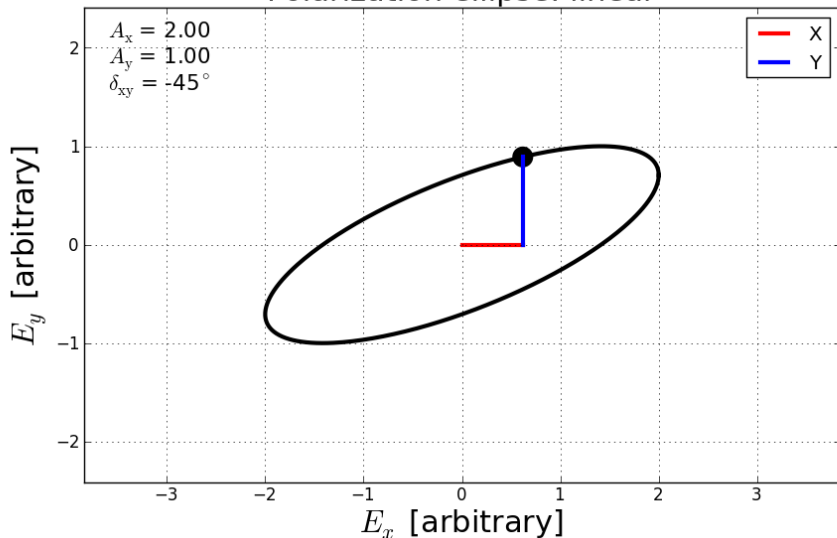
Polarization ellipse: linear



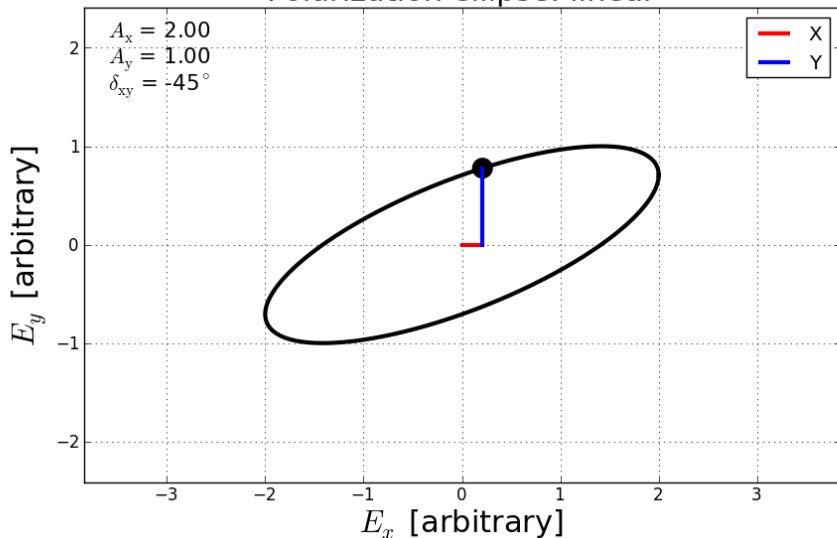
Polarization ellipse: linear



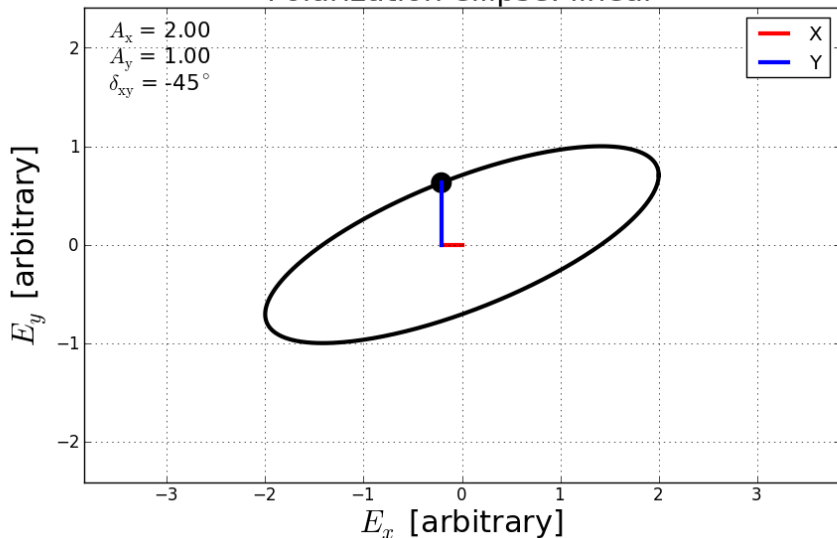
Polarization ellipse: linear



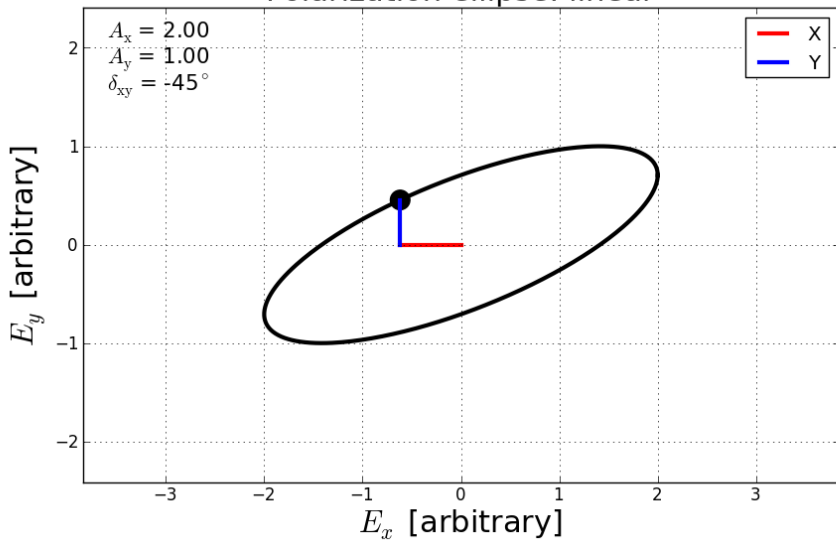
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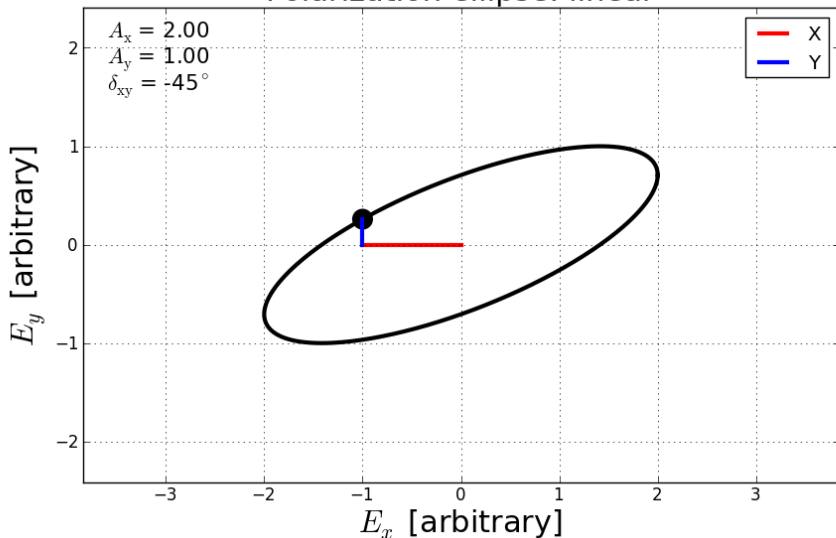
Polarization ellipse: linear



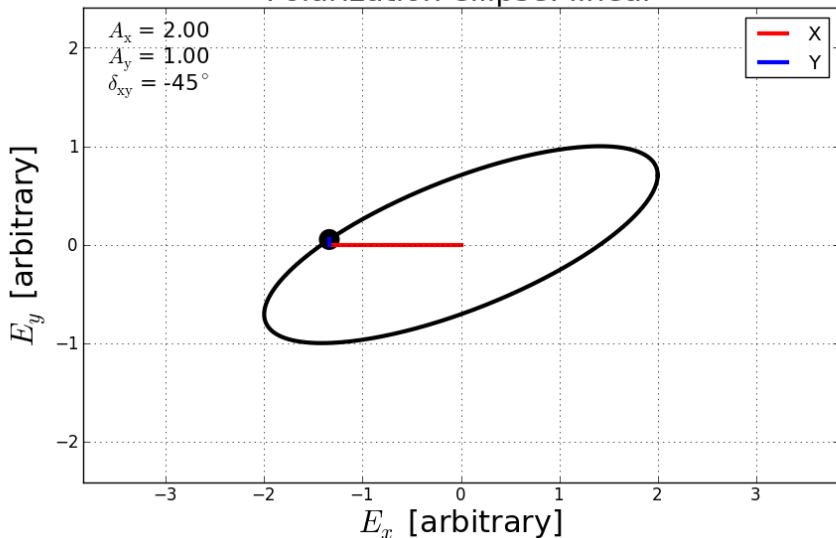
Polarization ellipse: linear



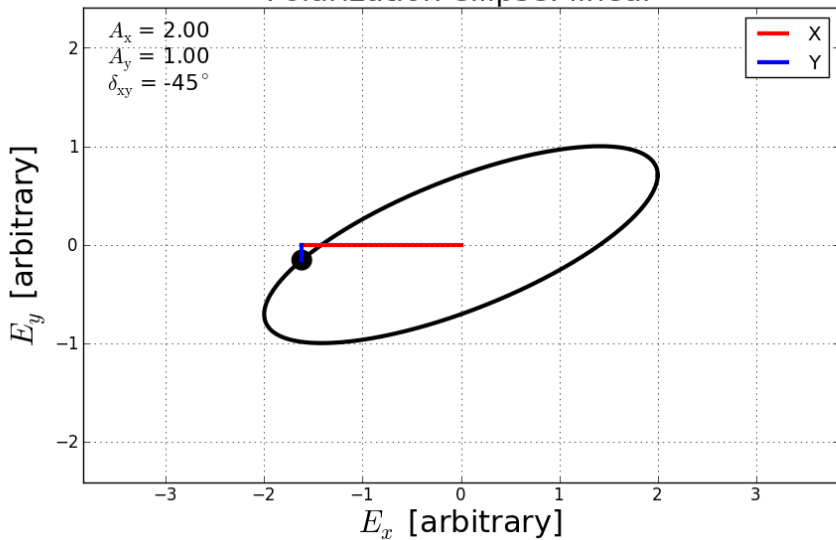
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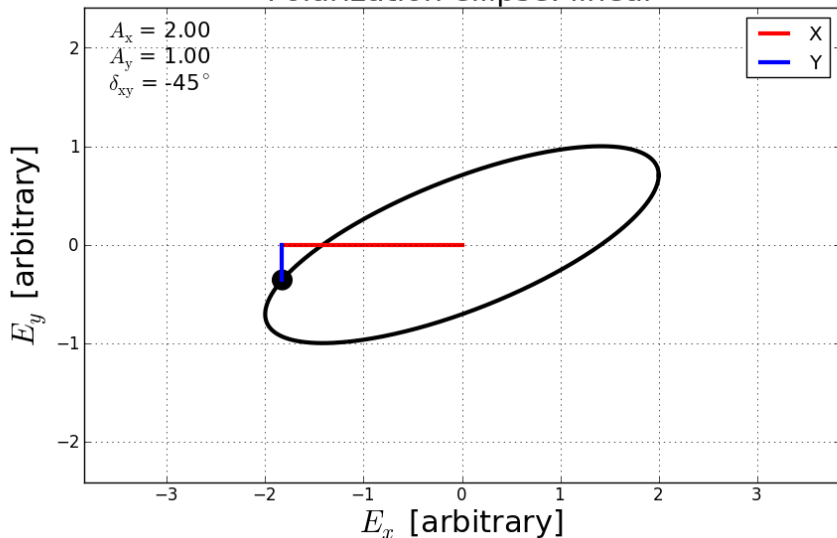
Polarization ellipse: linear



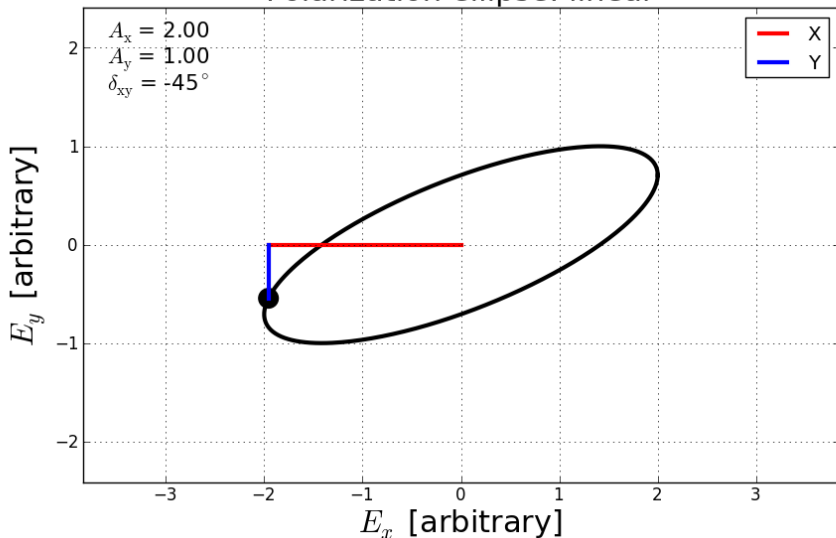
Polarization ellipse: linear



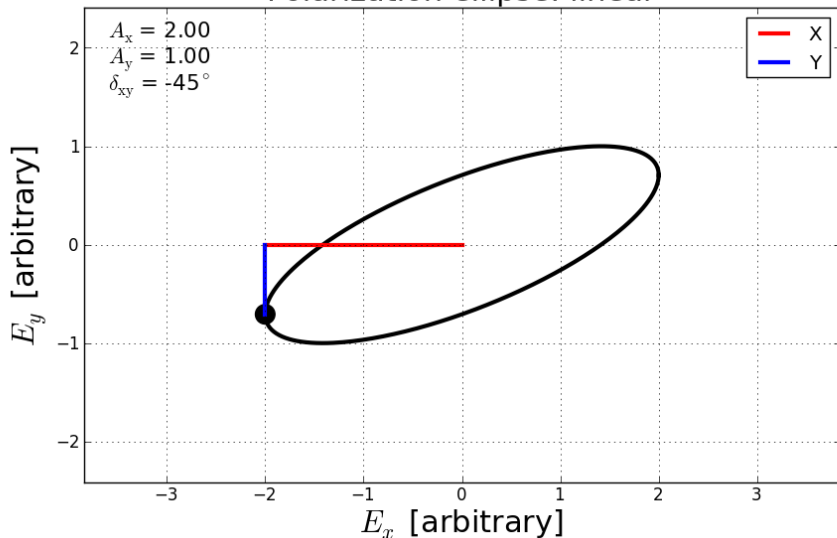
Polarization ellipse: linear



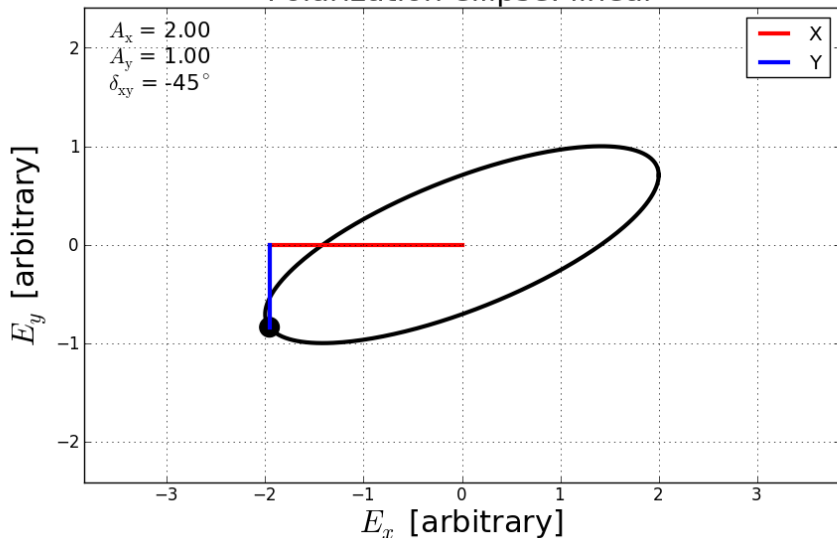
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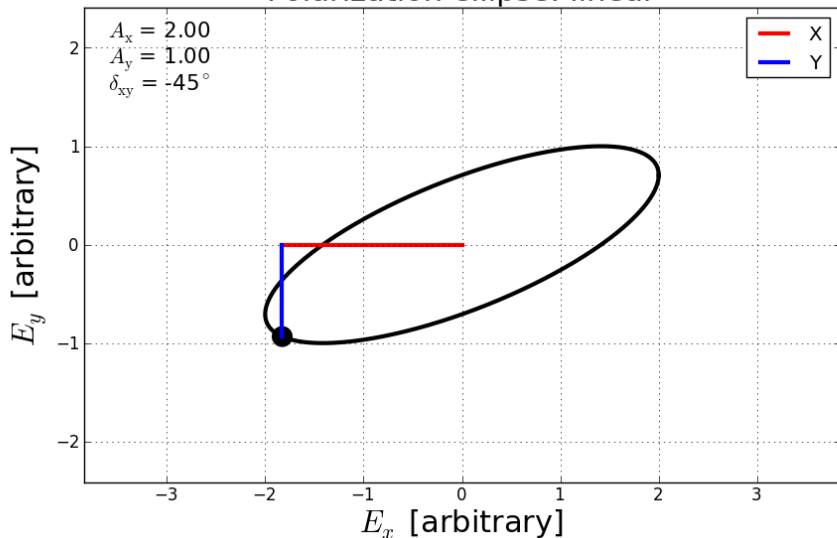
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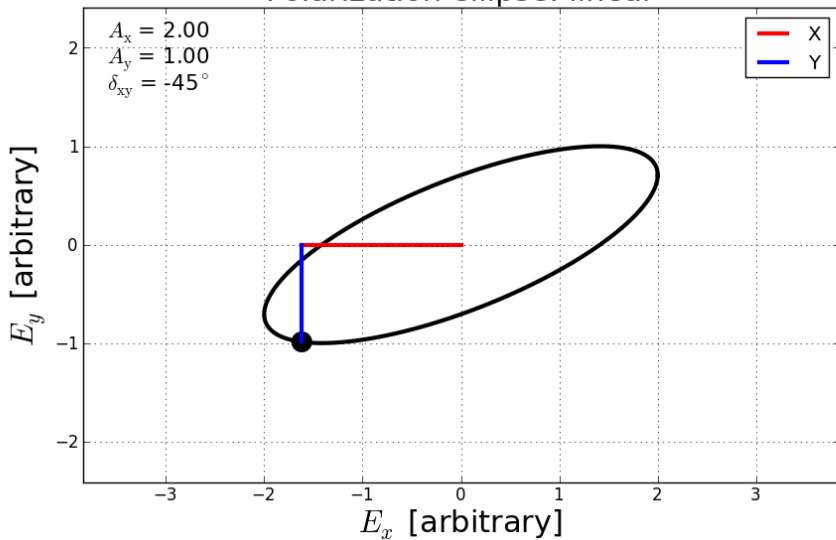
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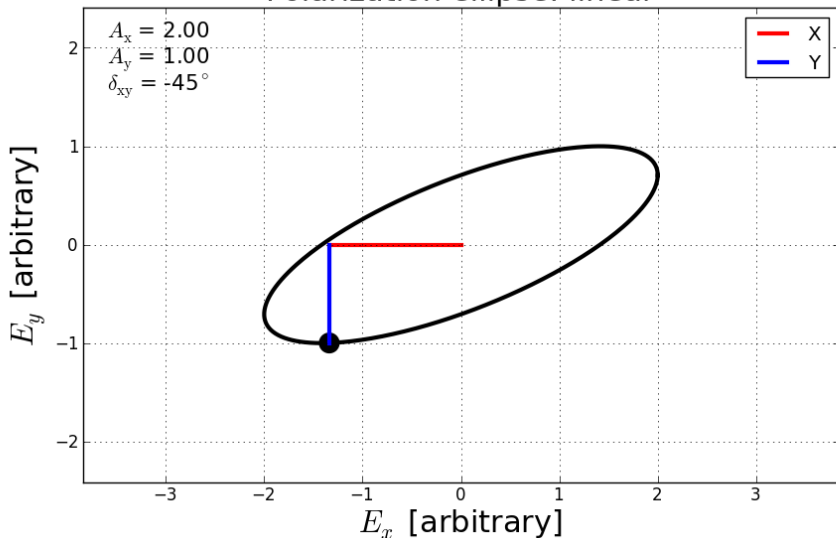
Polarization ellipse: linear



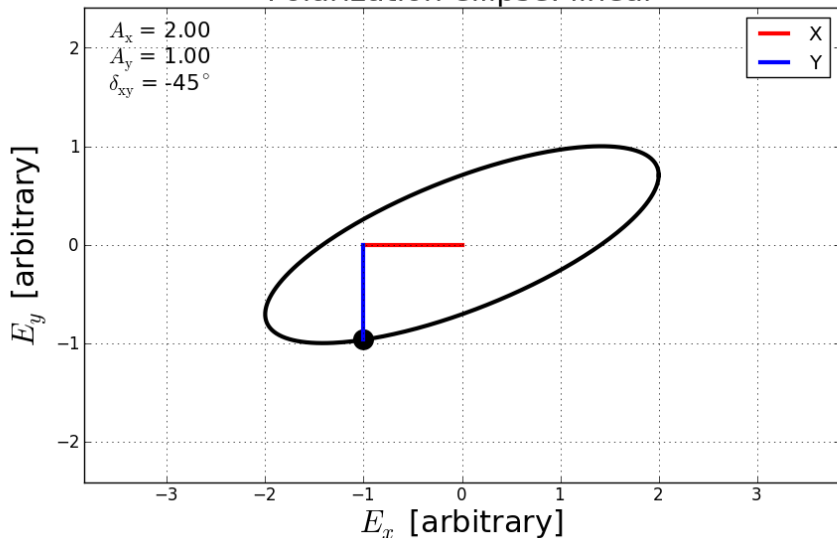
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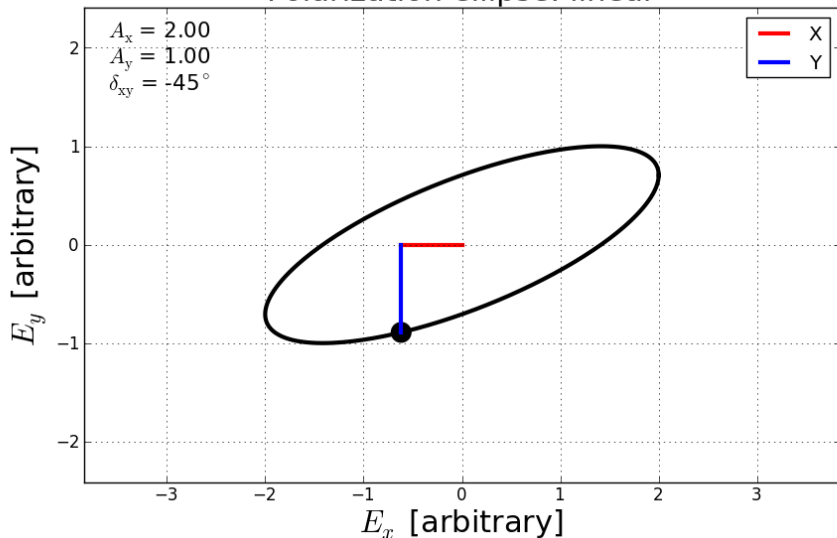
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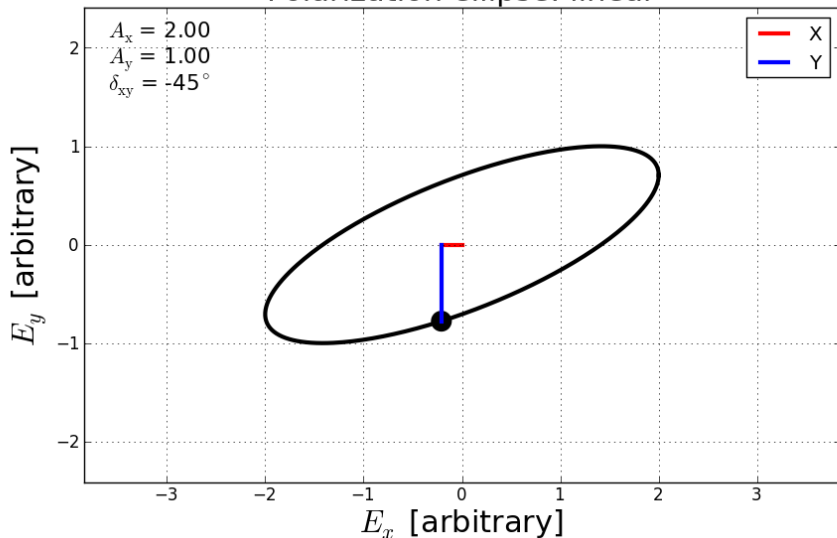
Polarization ellipse: linear



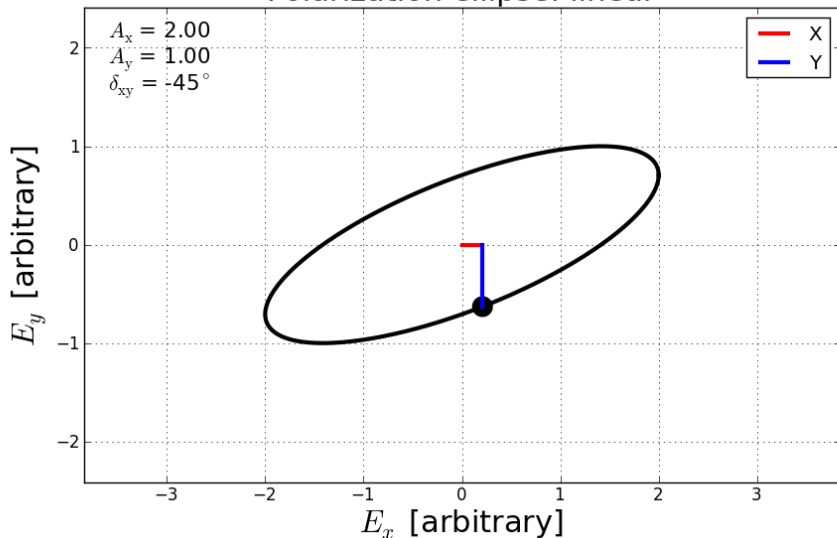
Polarization ellipse: linear



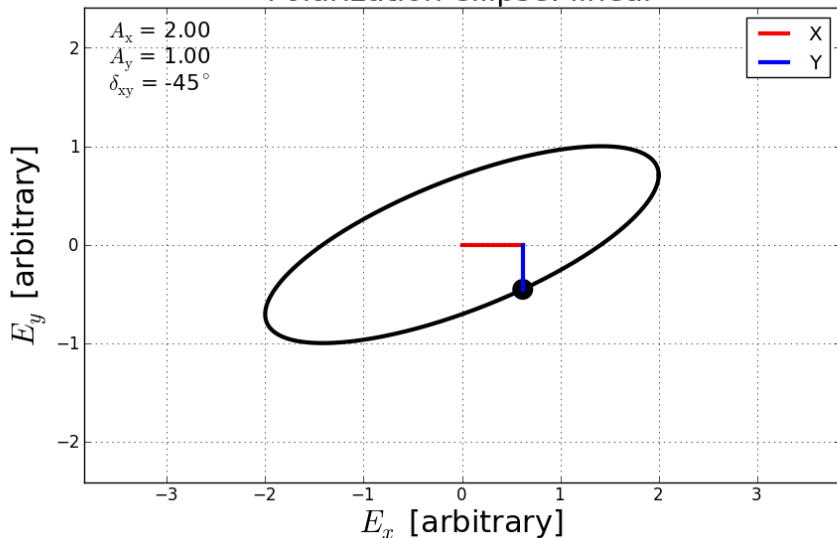
Polarization ellipse: linear



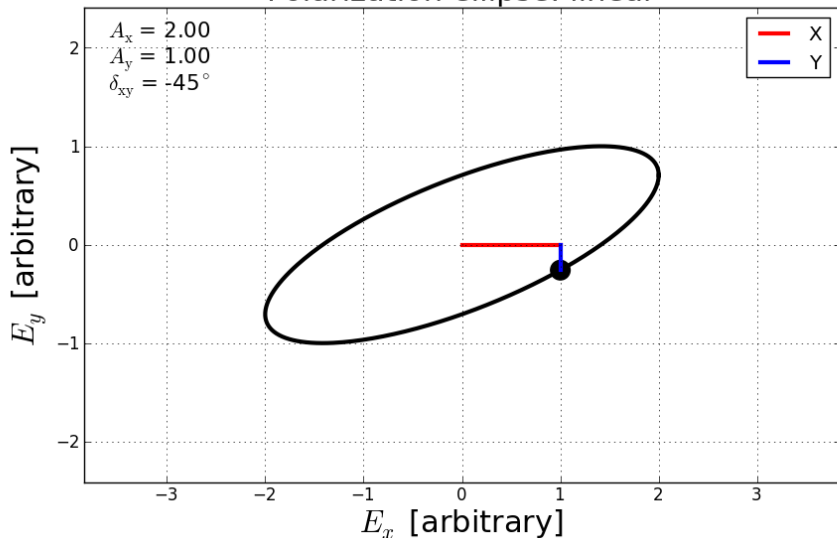
Polarization ellipse: linear



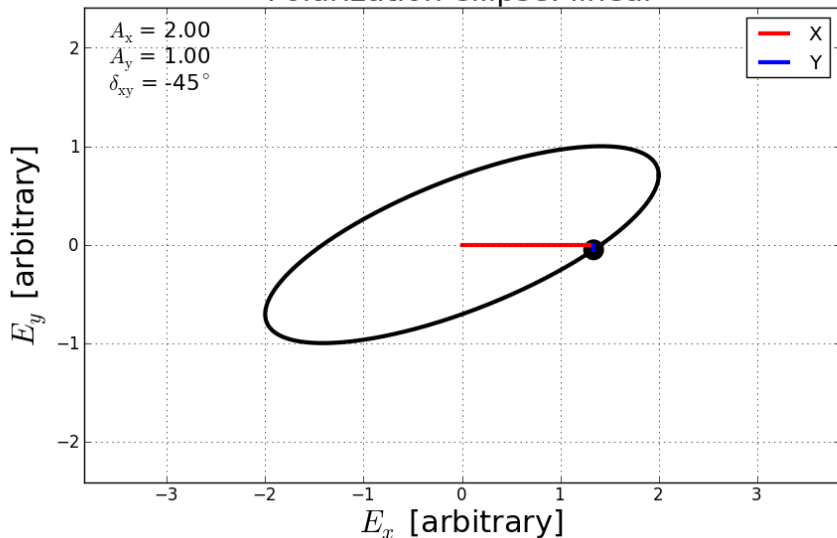
Polarization ellipse: linear



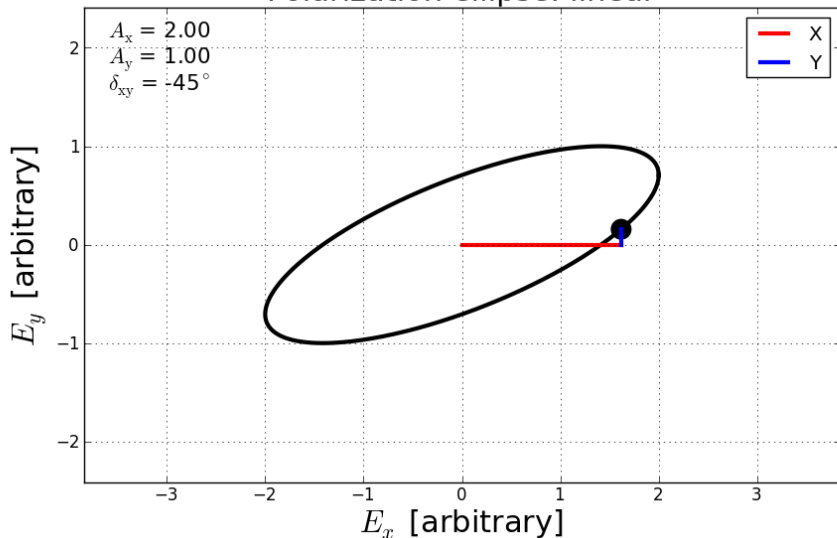
Polarization ellipse: linear



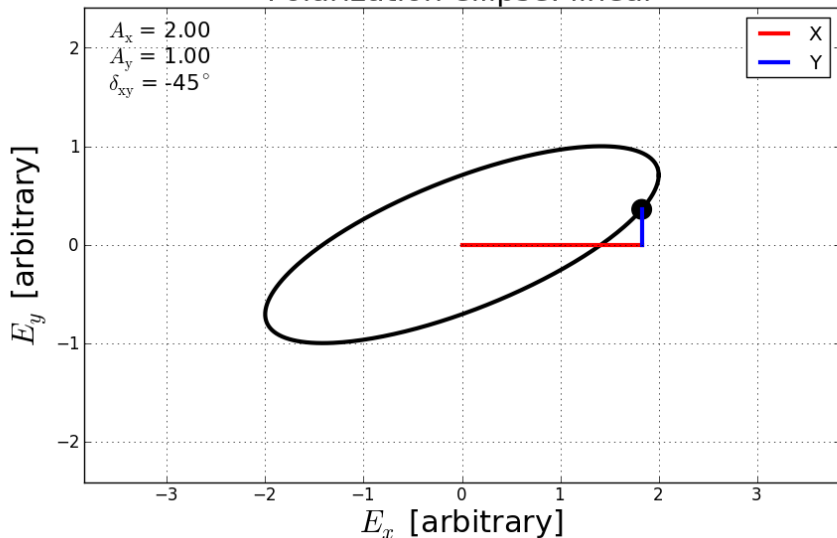
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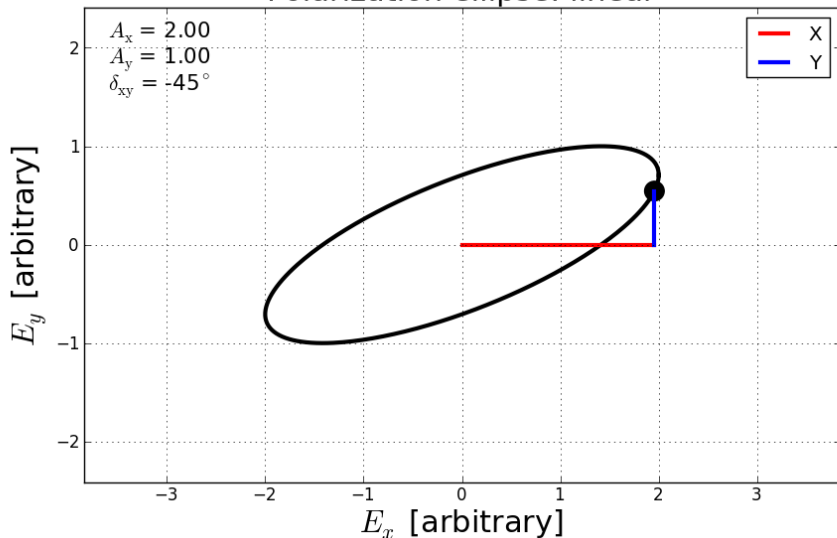
Polarization ellipse: linear



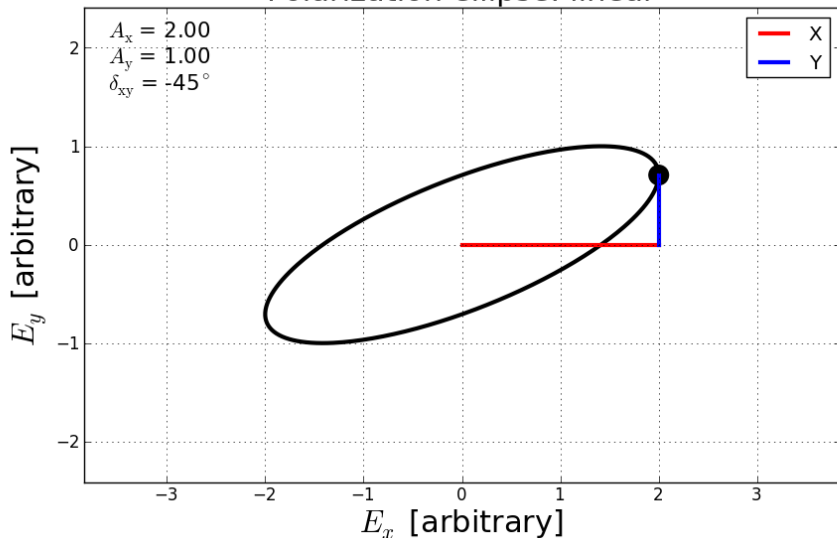
Polarization ellipse: linear



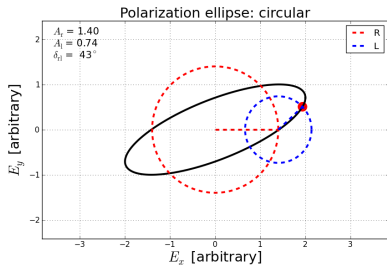
Polarization ellipse: linear



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Geometry



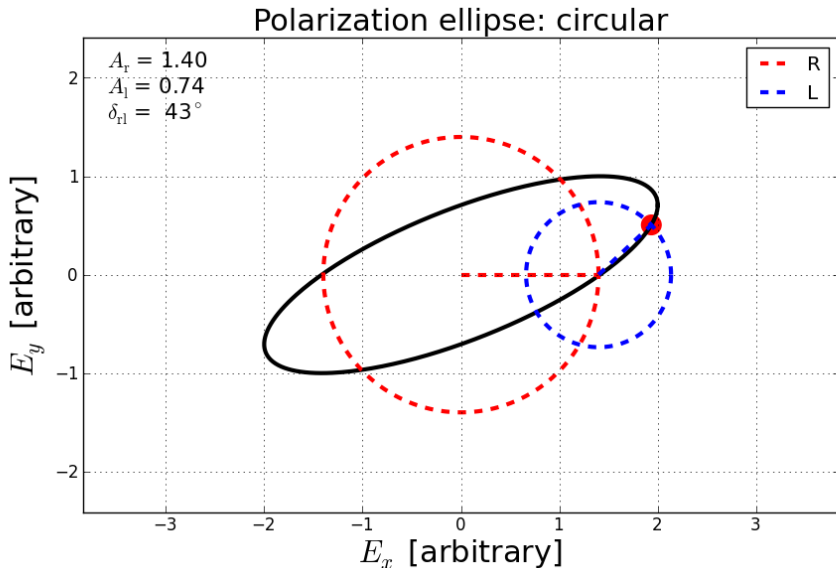
Viewing from antenna towards source, watching orientation and length of \mathbf{E} vector on a plane at a fixed location in space.

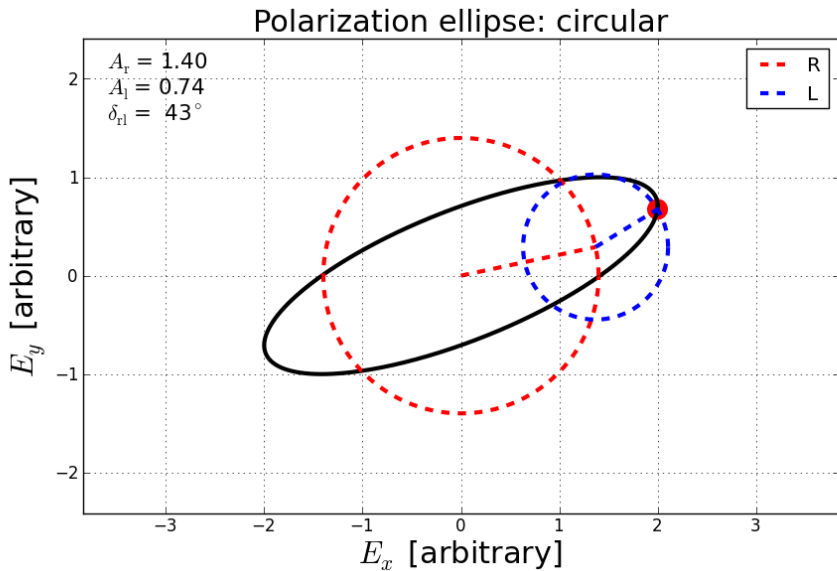
$$\mathbf{E} = A_r \hat{\mathbf{e}}_r + A_l \hat{\mathbf{e}}_l$$

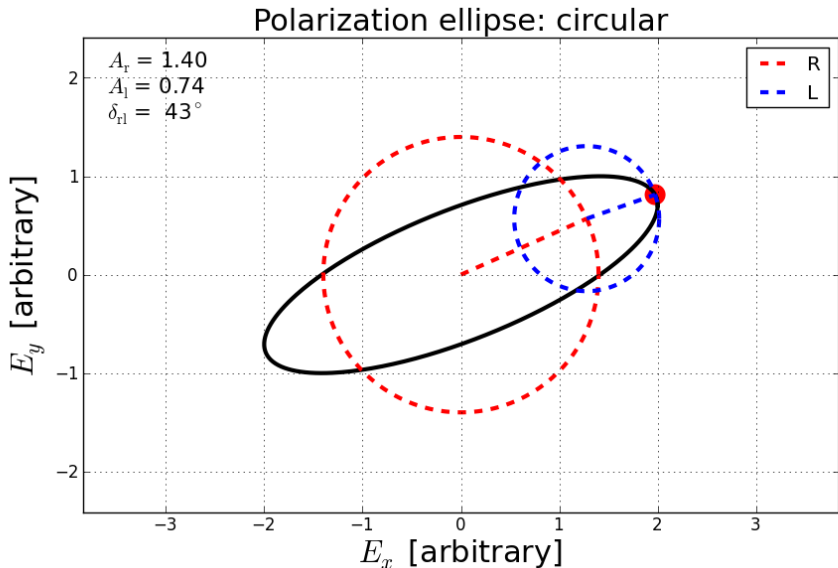
$$\hat{\mathbf{e}}_r = \begin{pmatrix} \cos(2\pi\nu t + \delta_r) \\ \sin(2\pi\nu t + \delta_r) \end{pmatrix}$$

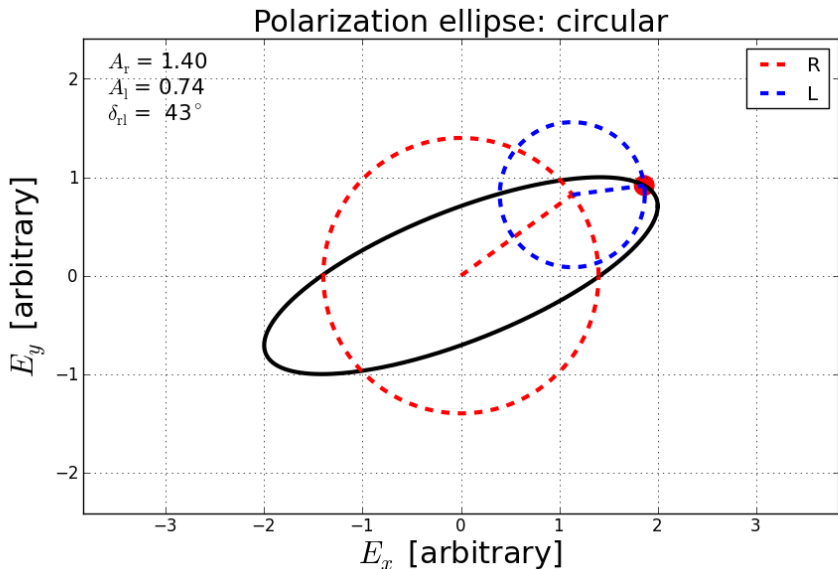
$$\hat{\mathbf{e}}_l = \begin{pmatrix} \cos(2\pi\nu t + \delta_l) \\ -\sin(2\pi\nu t + \delta_l) \end{pmatrix}$$

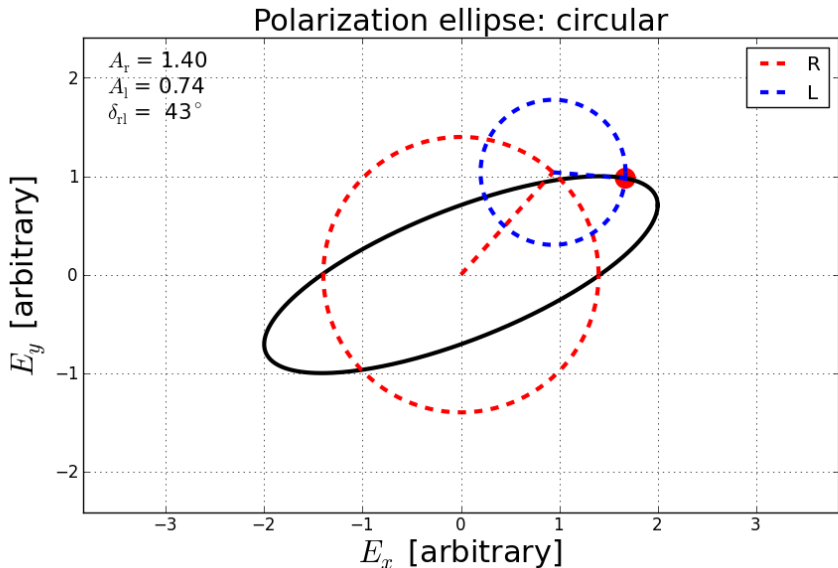
- $A_r + A_l =$ semi-major axis
- $\|A_r - A_l\| =$ semi-minor axis
- $\delta_{rl} = \delta_r - \delta_l$
- $\delta_{rl} =$ orientation of major axis
- $\delta_{rl} > 0$: MA rotated CCW
- $\delta_{rl} = 0$: MA along x-axis
- $\delta_{rl} < 0$: MA rotated CW

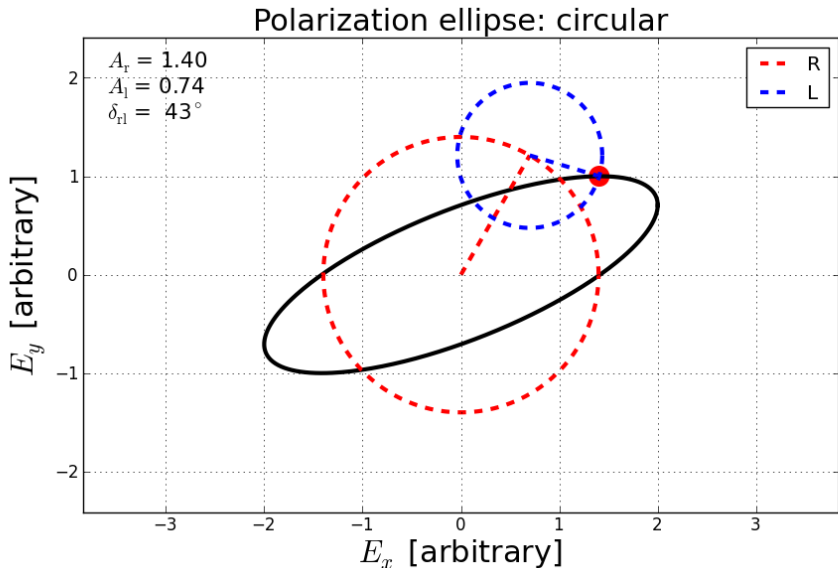


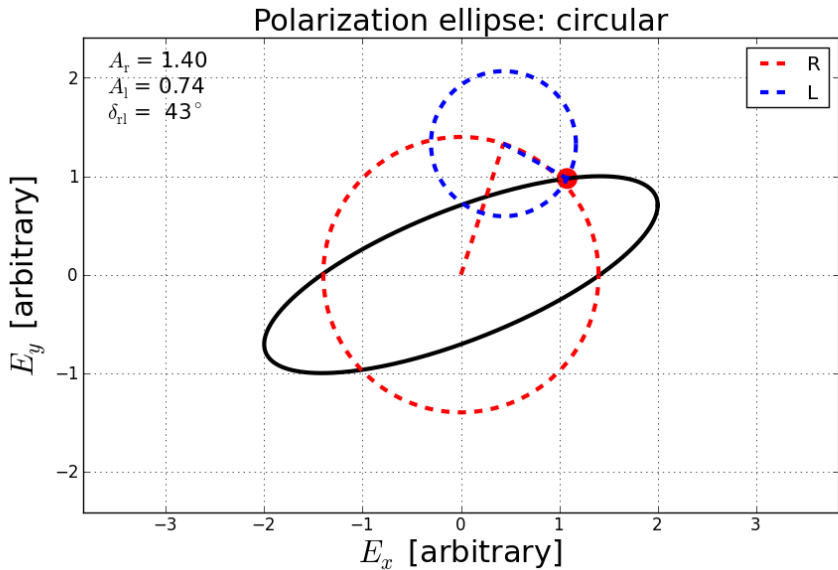


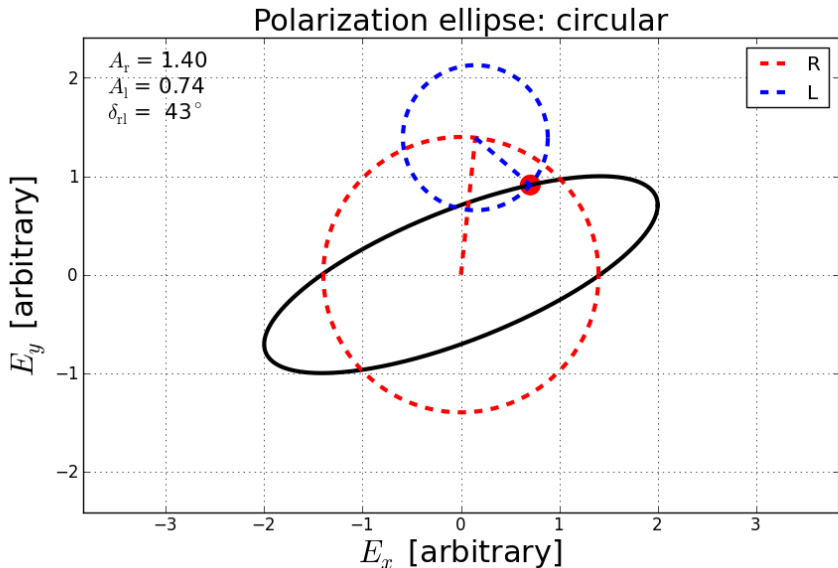


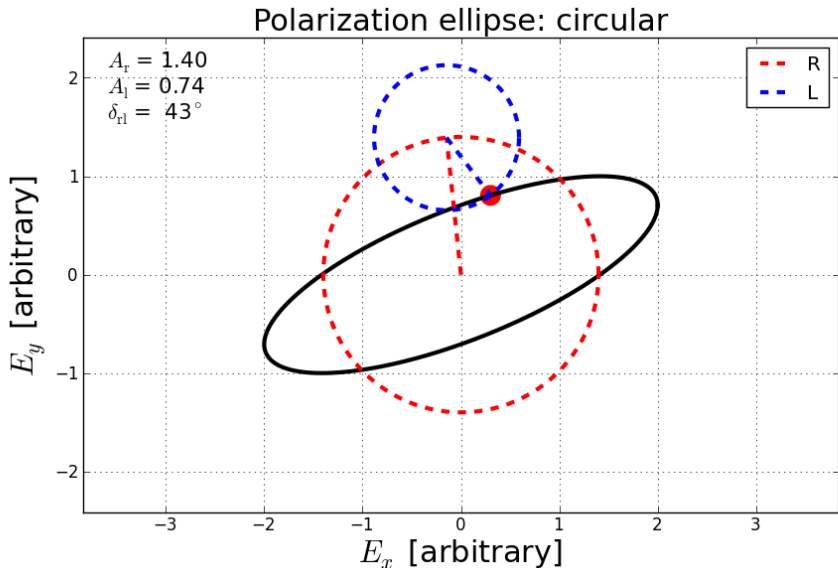


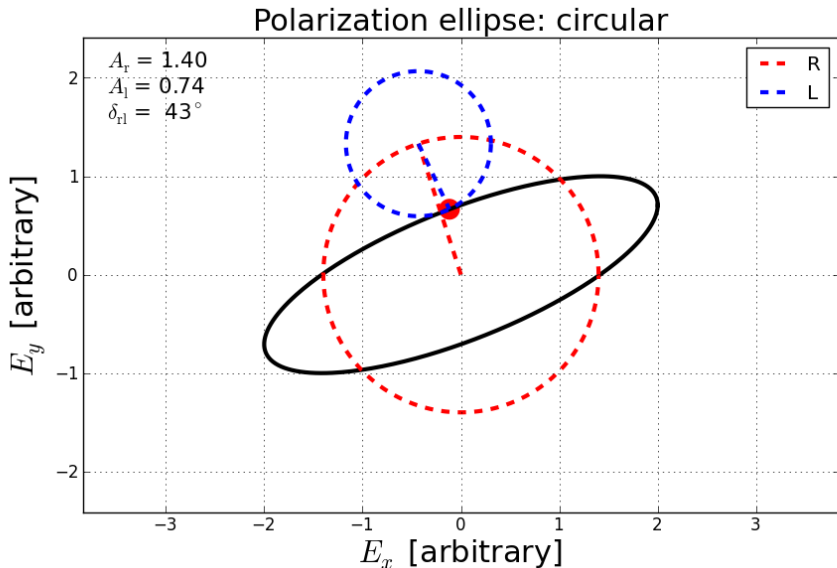


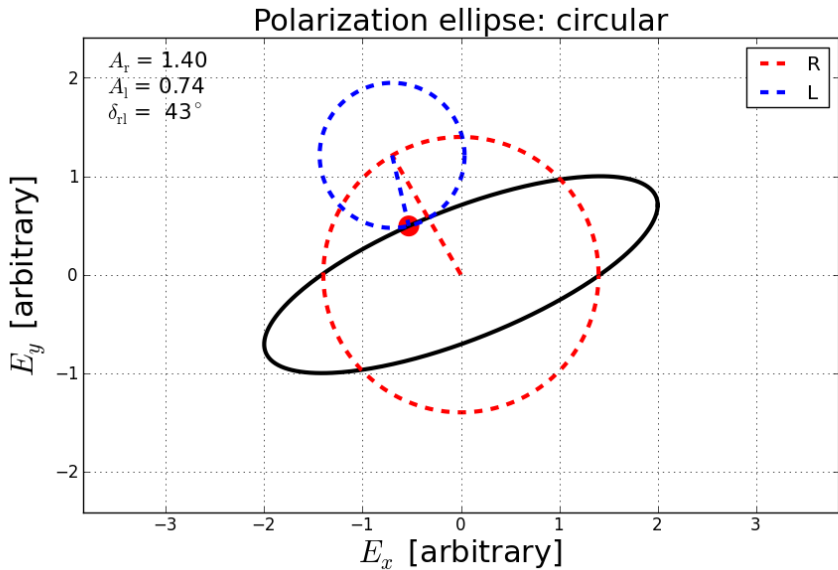


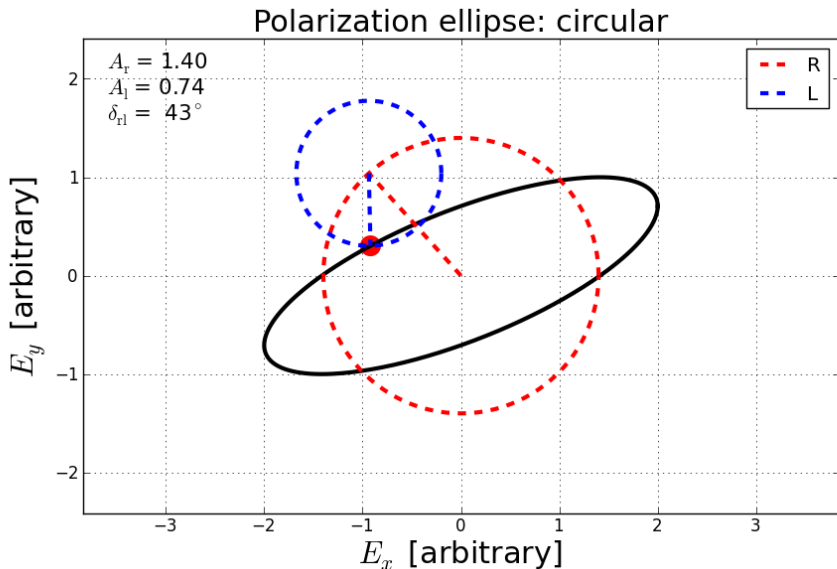


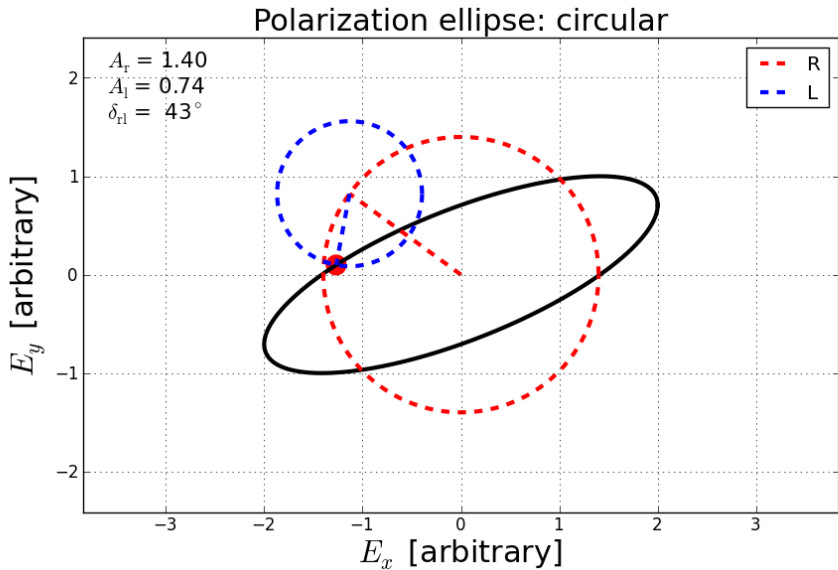


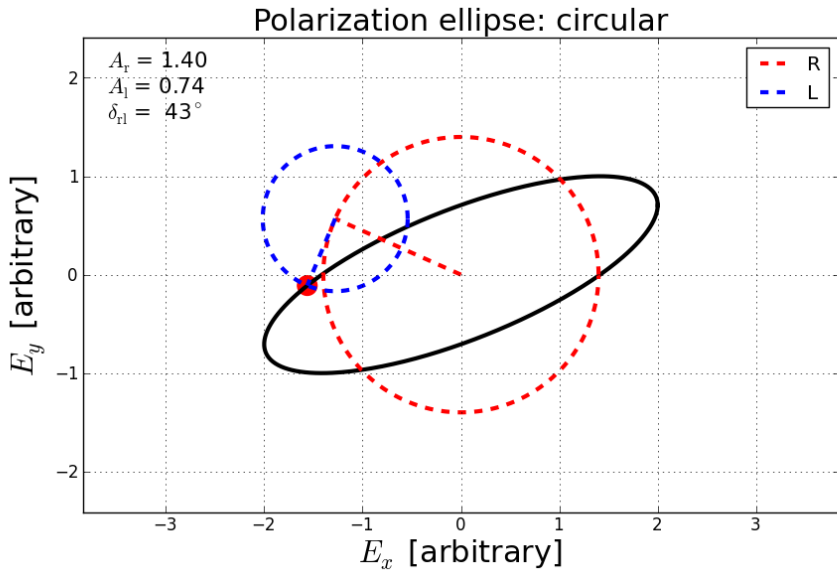


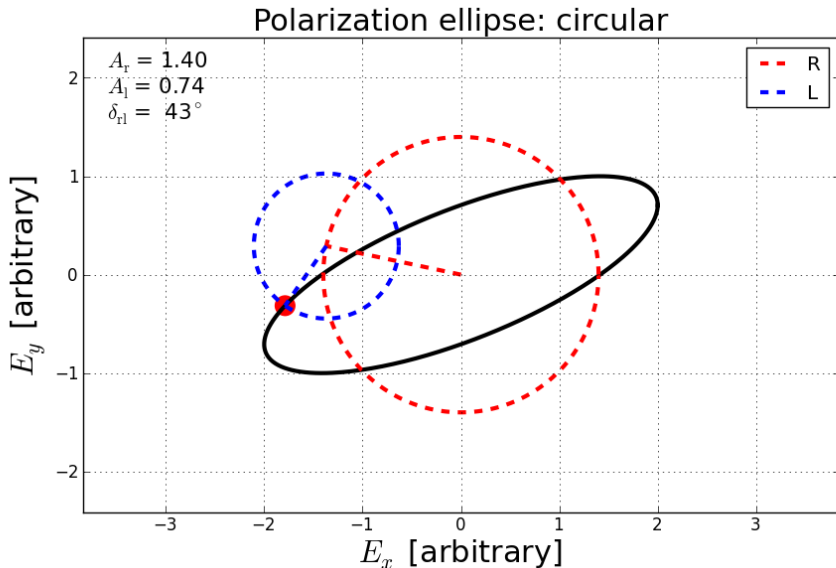


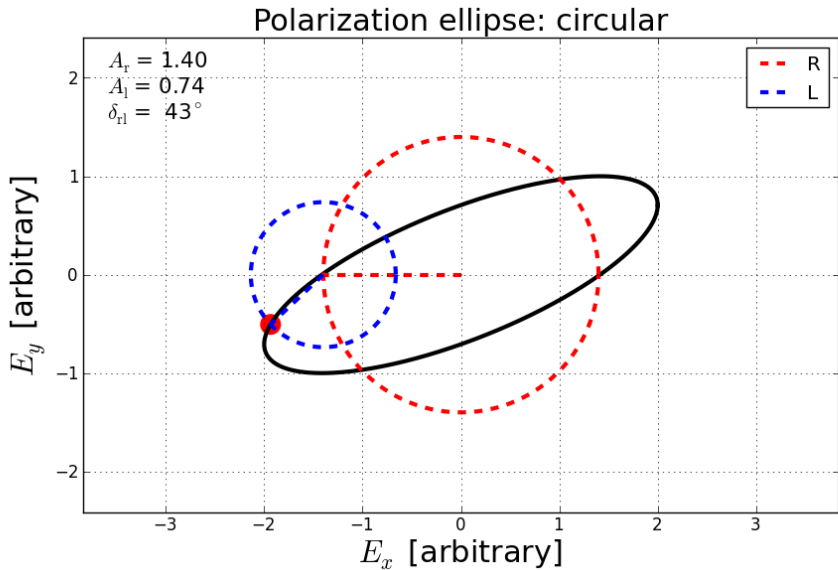


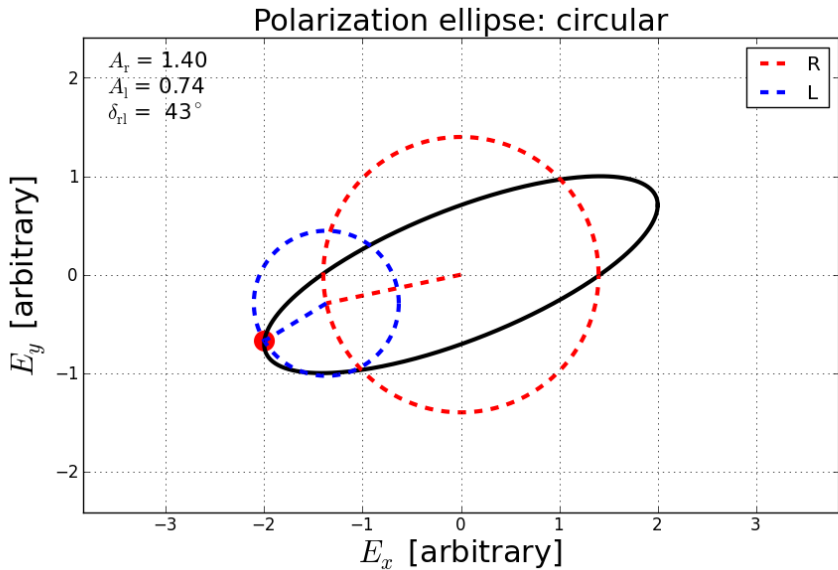


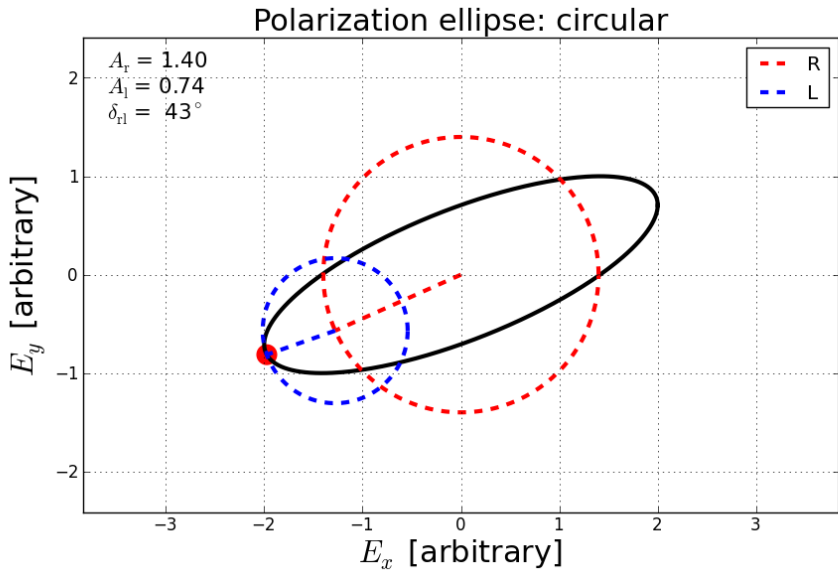


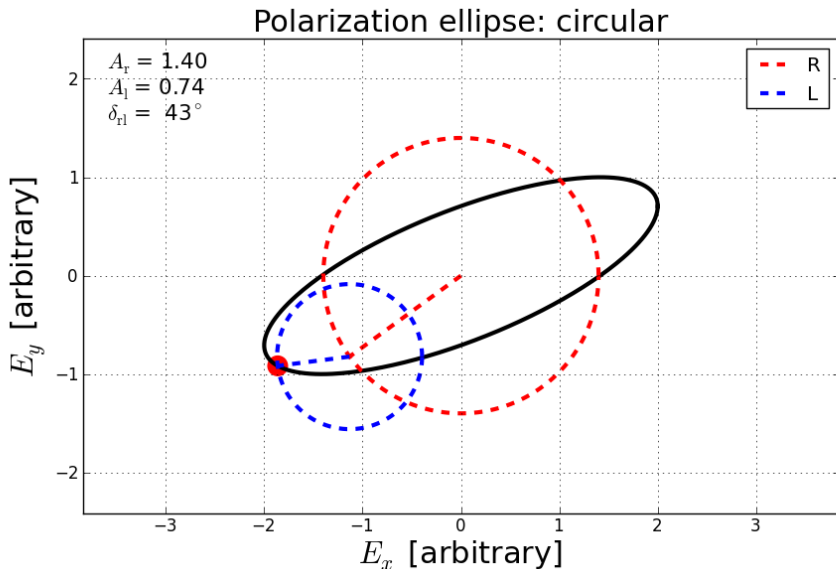


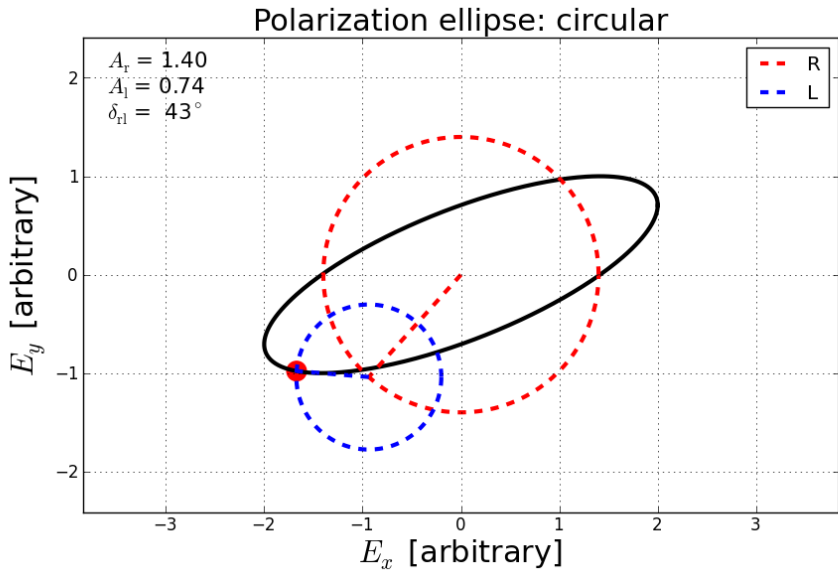


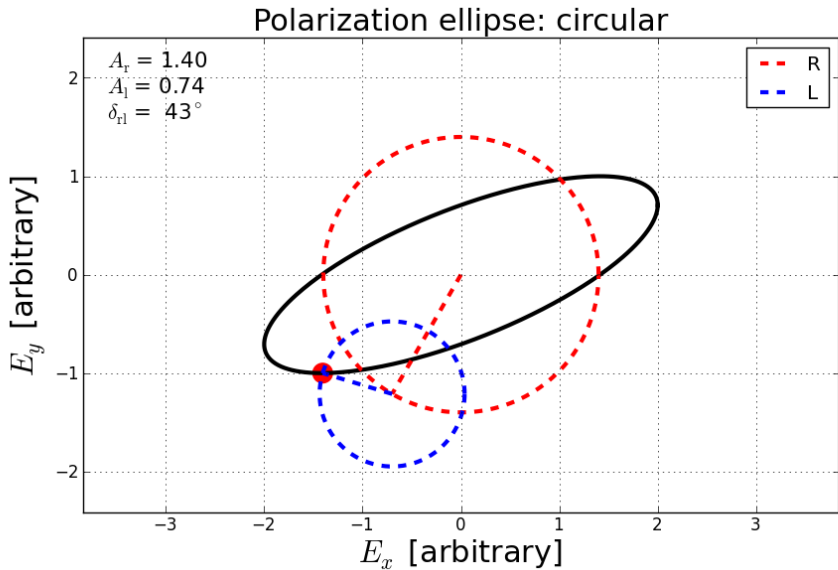


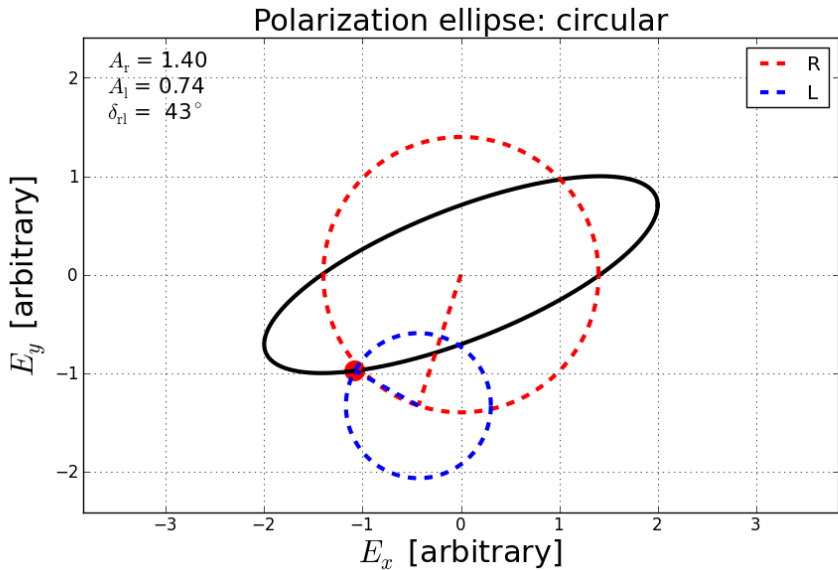


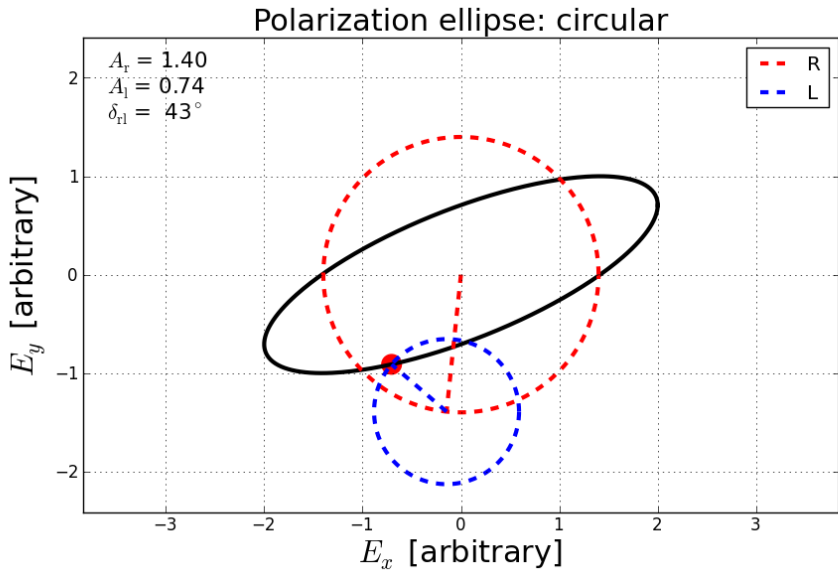


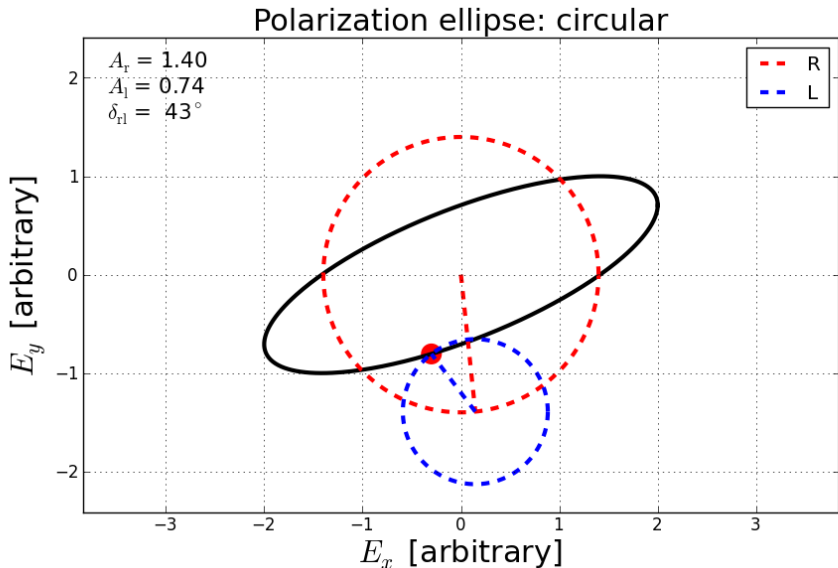


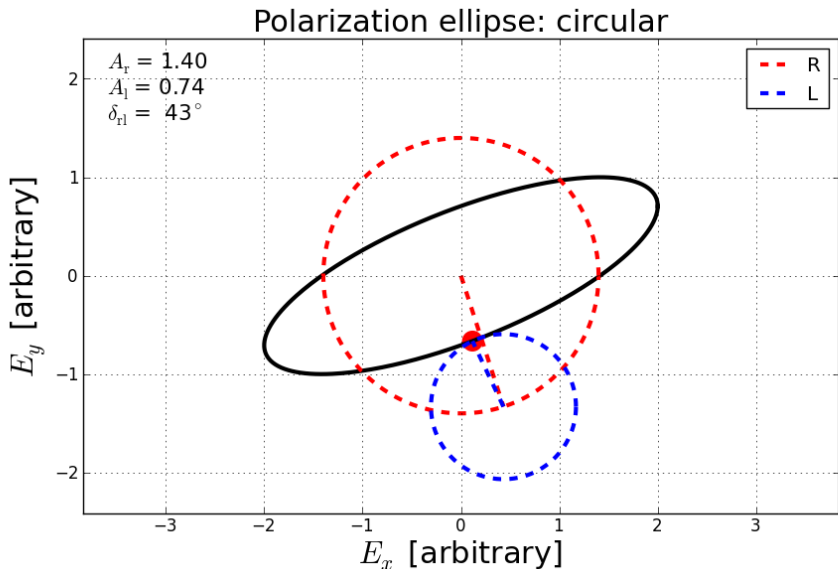


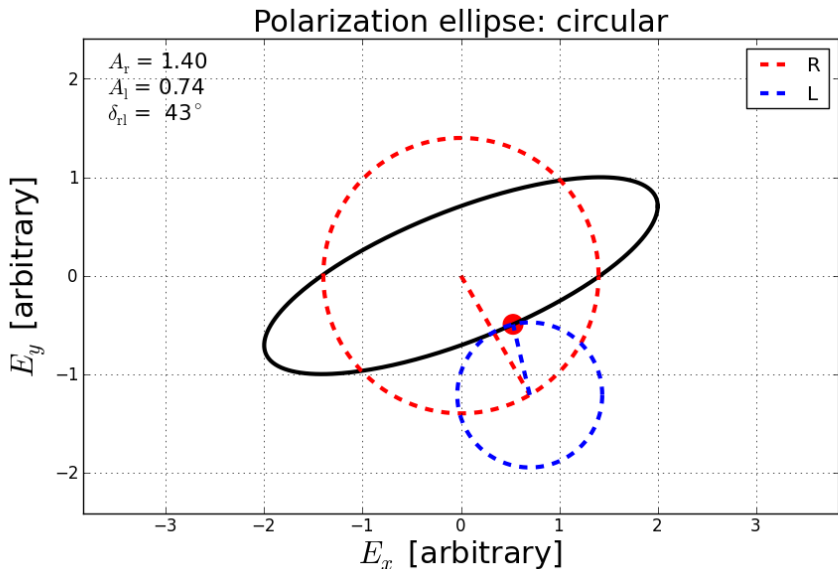


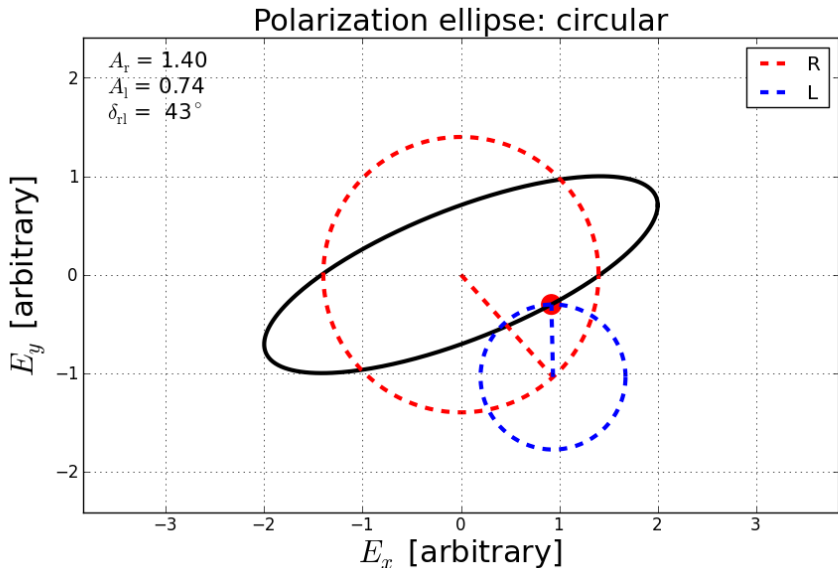


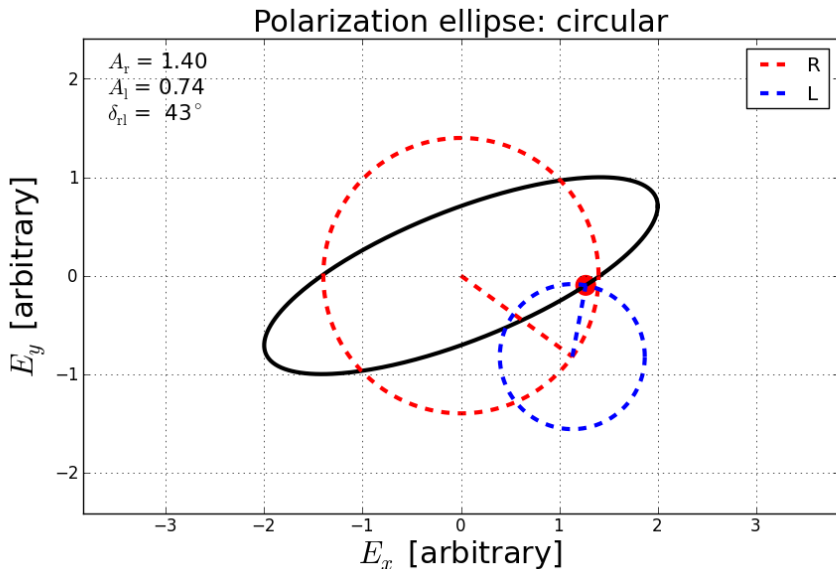


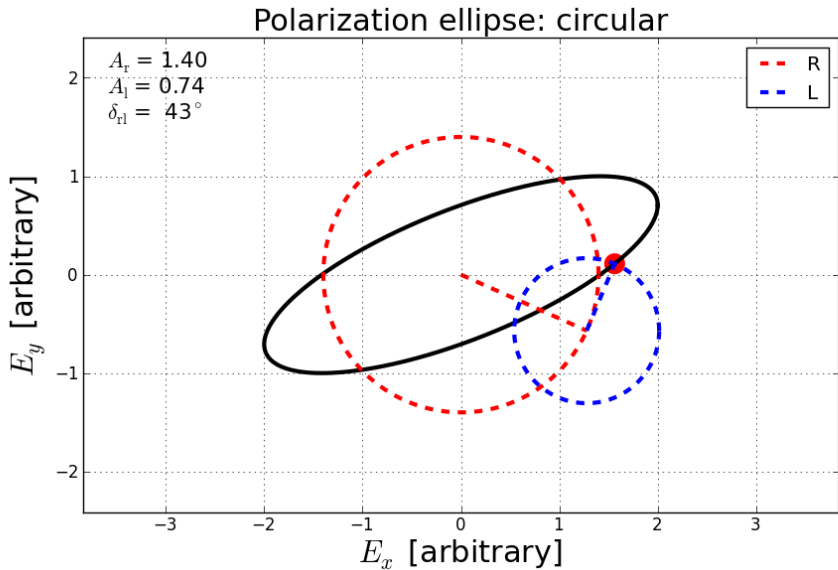


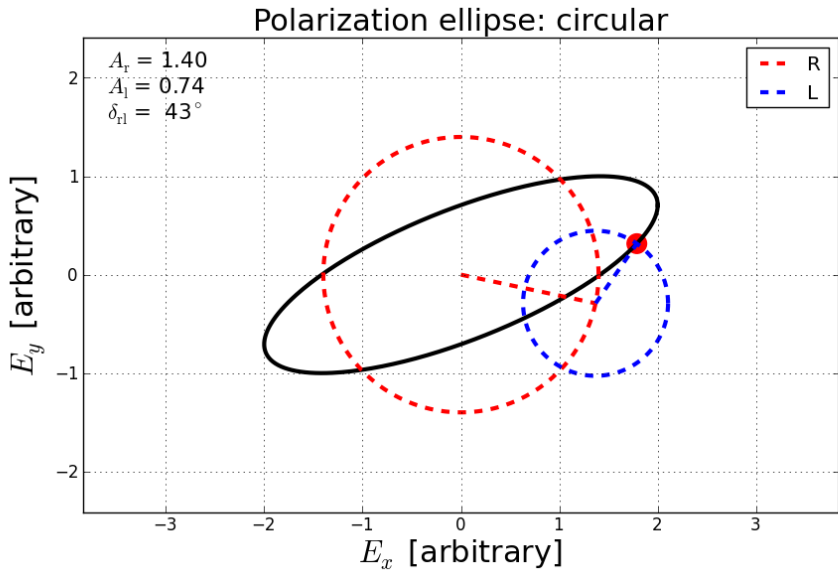


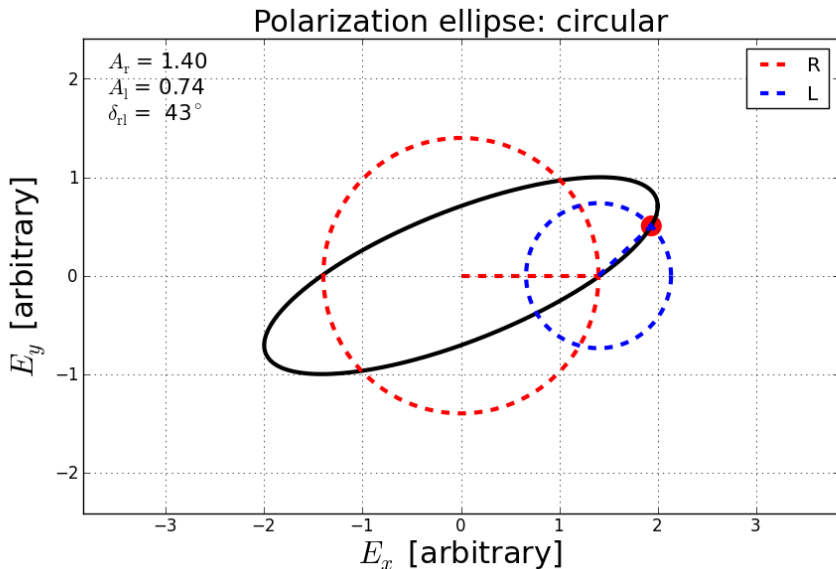










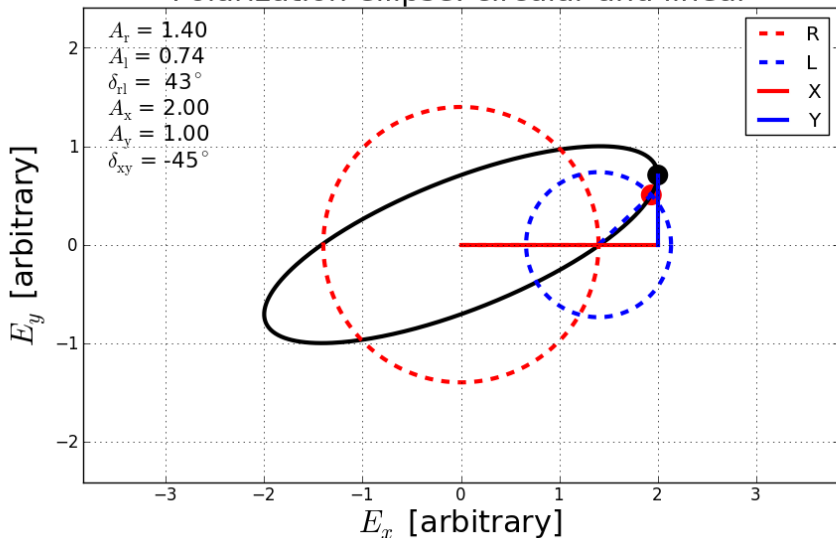


$$A_r = \frac{1}{2} \sqrt{A_x^2 + A_y^2 - 2A_x A_y \sin \delta_{xy}}$$

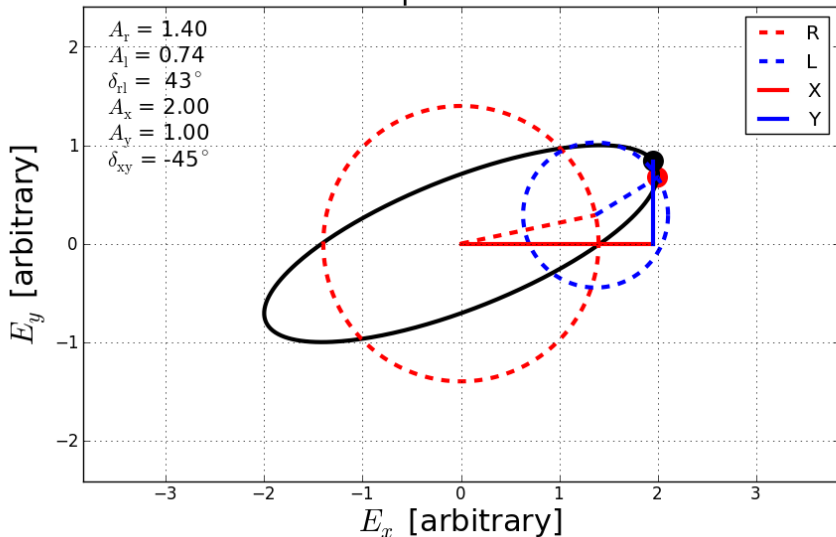
$$A_l = \frac{1}{2} \sqrt{A_x^2 + A_y^2 + 2A_x A_y \sin \delta_{xy}}$$

$$\tan \delta_{rl} = \frac{2A_x A_y \cos \delta_{xy}}{A_x^2 - A_y^2}$$

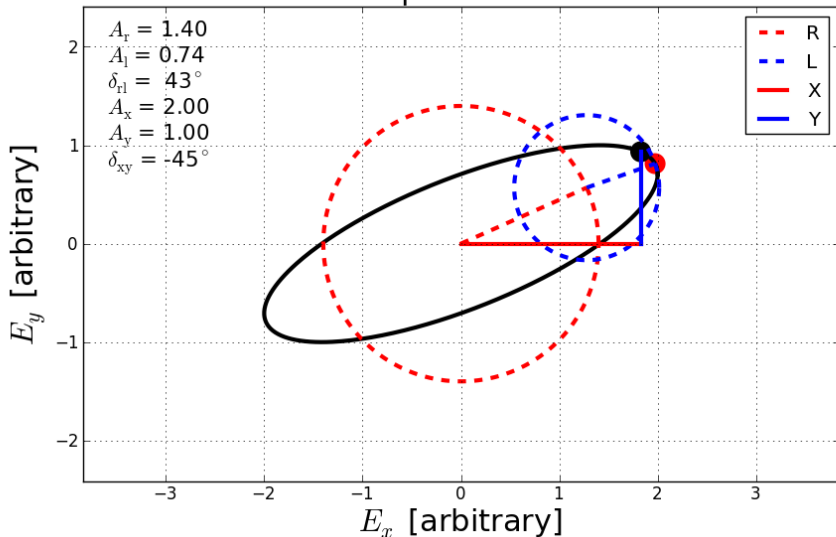
Polarization ellipse: circular and linear



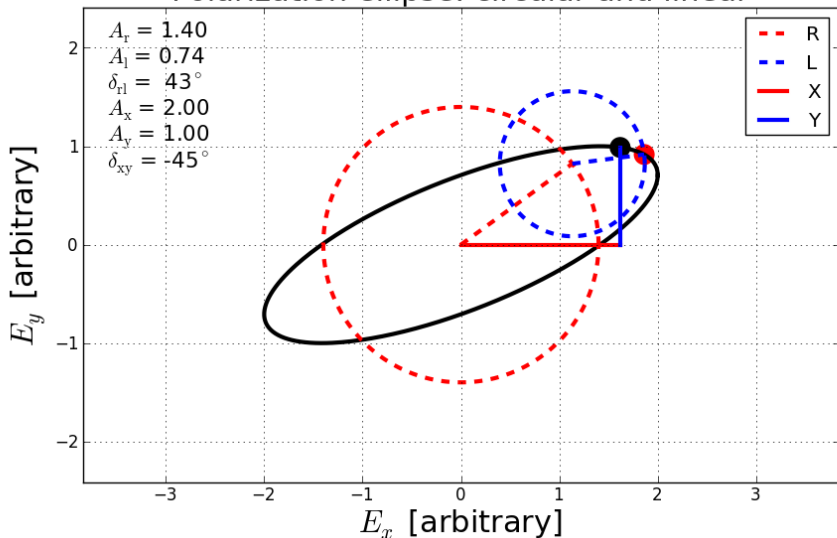
Polarization ellipse: circular and linear



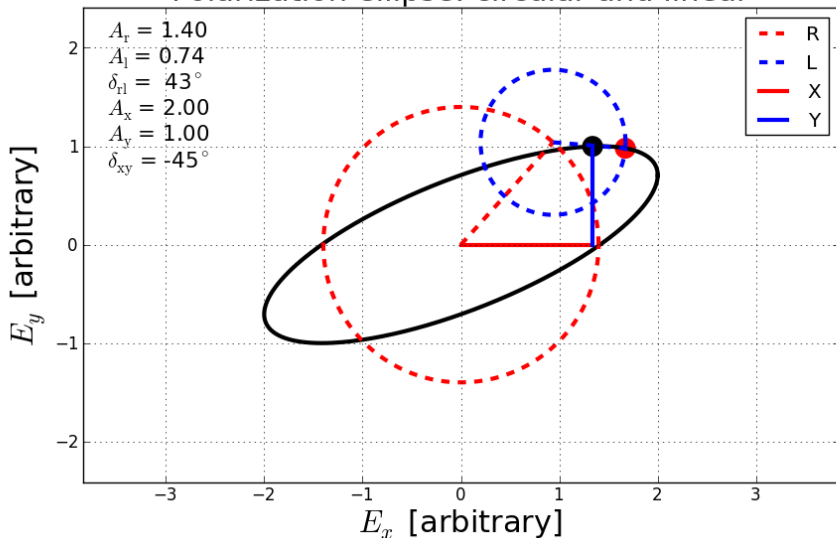
Polarization ellipse: circular and linear



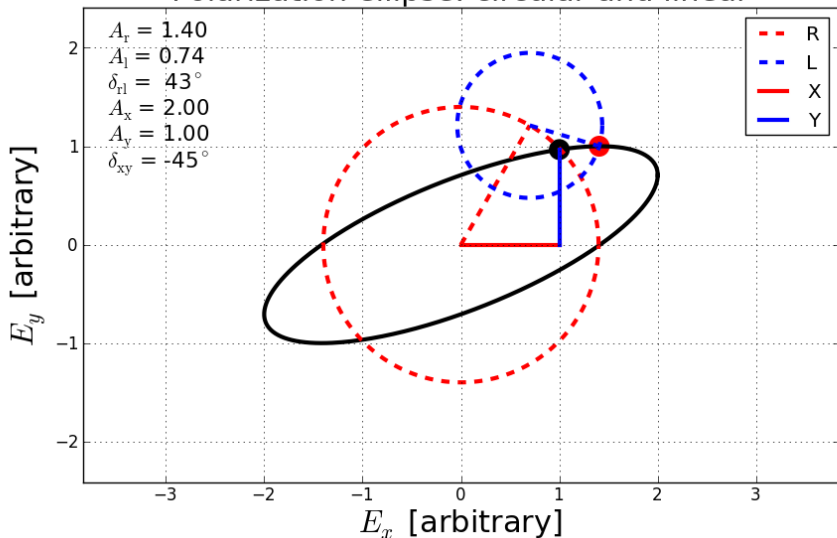
Polarization ellipse: circular and linear



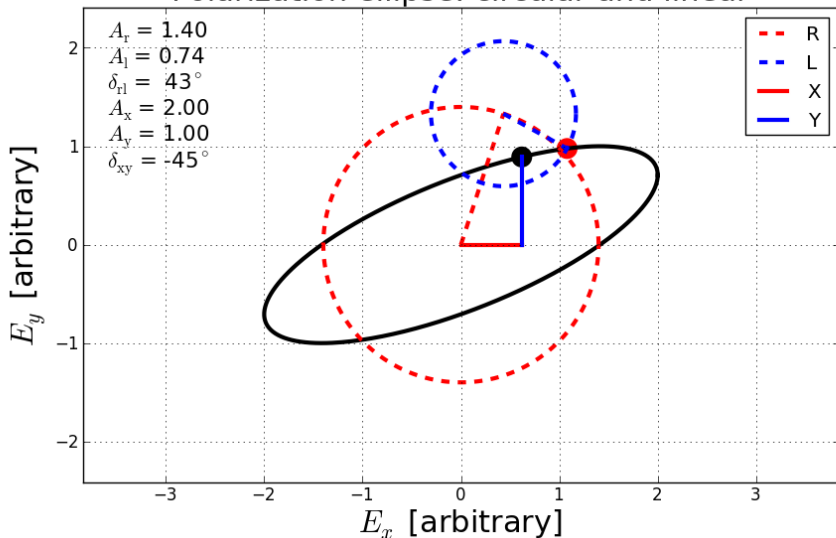
Polarization ellipse: circular and linear



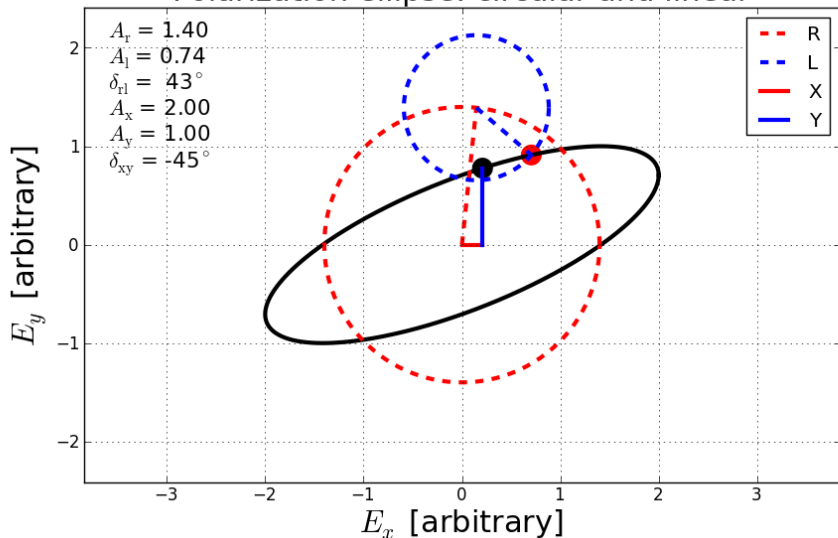
Polarization ellipse: circular and linear



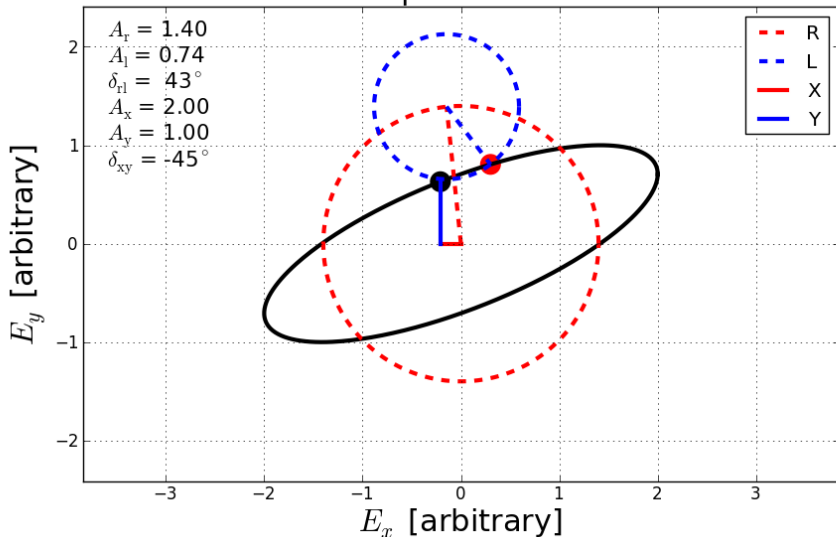
Polarization ellipse: circular and linear



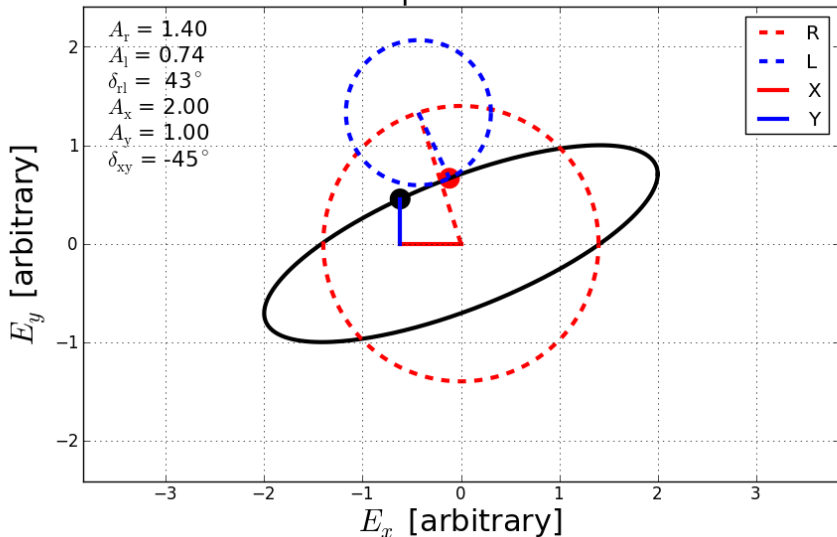
Polarization ellipse: circular and linear



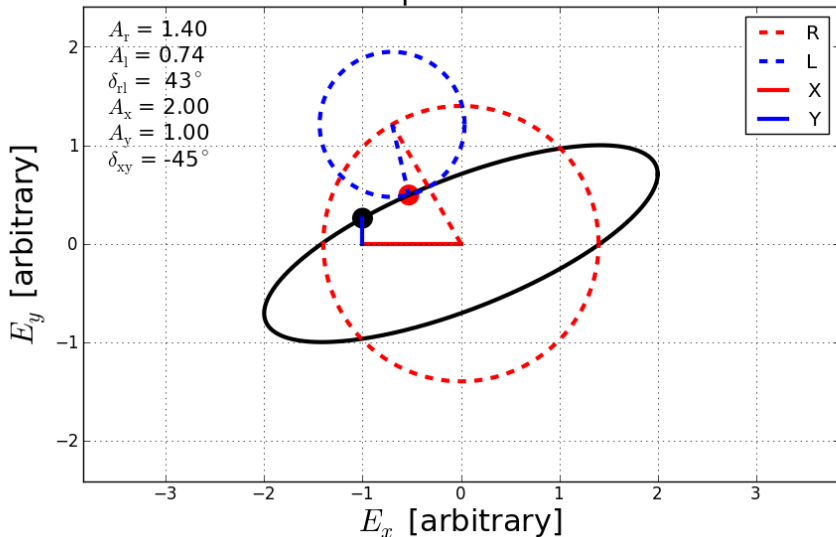
Polarization ellipse: circular and linear



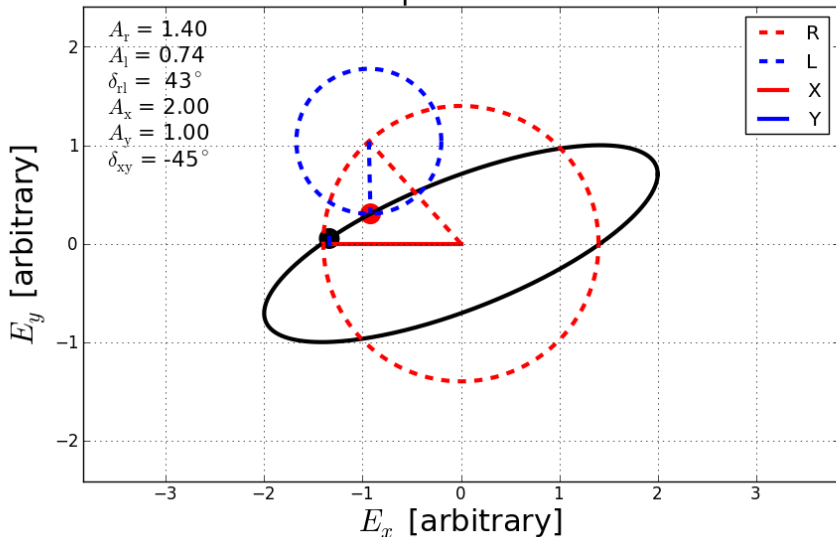
Polarization ellipse: circular and linear



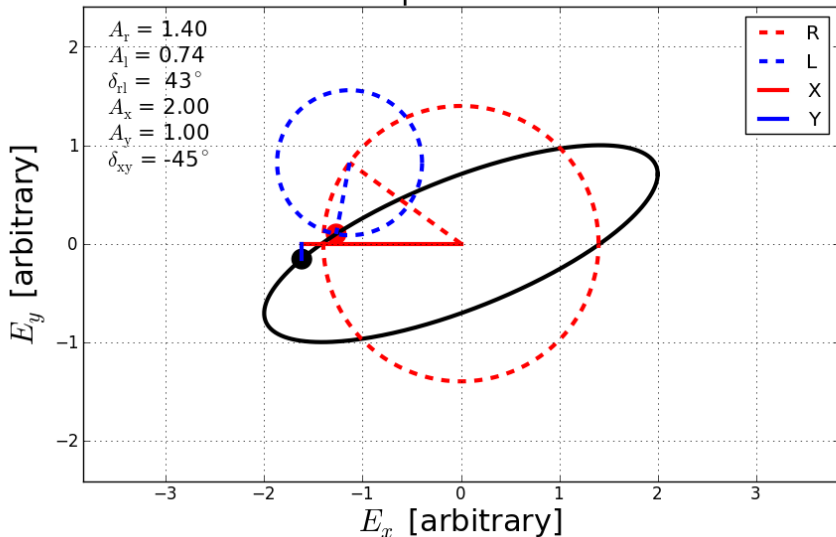
Polarization ellipse: circular and linear



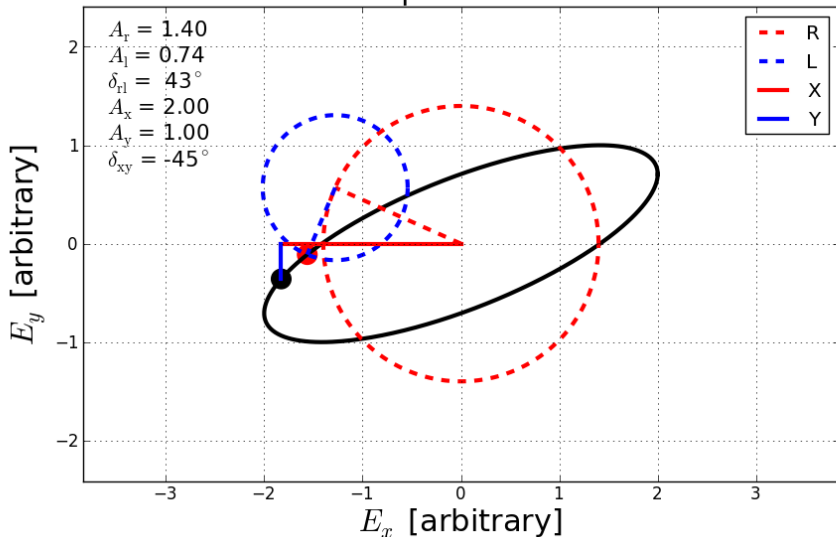
Polarization ellipse: circular and linear



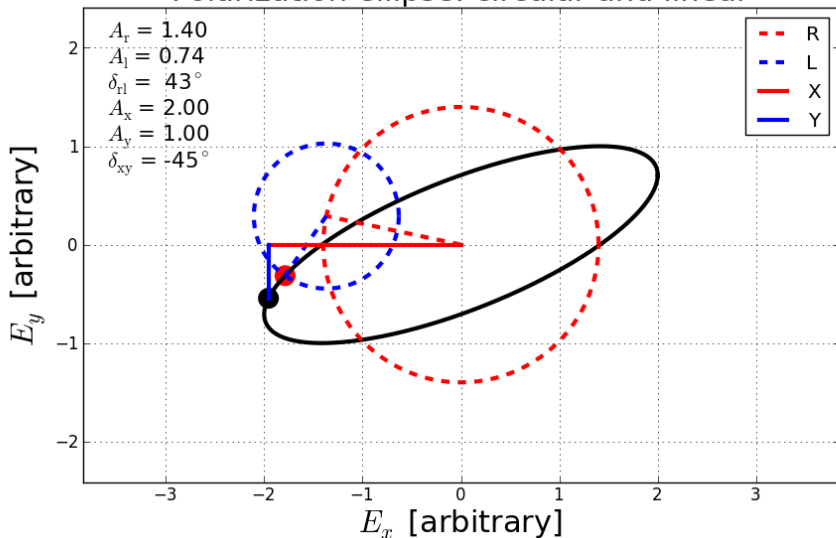
Polarization ellipse: circular and linear



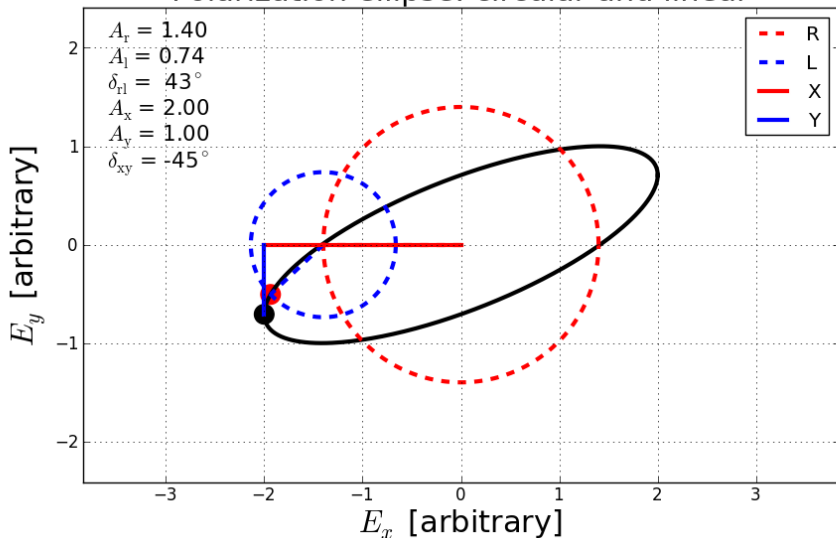
Polarization ellipse: circular and linear



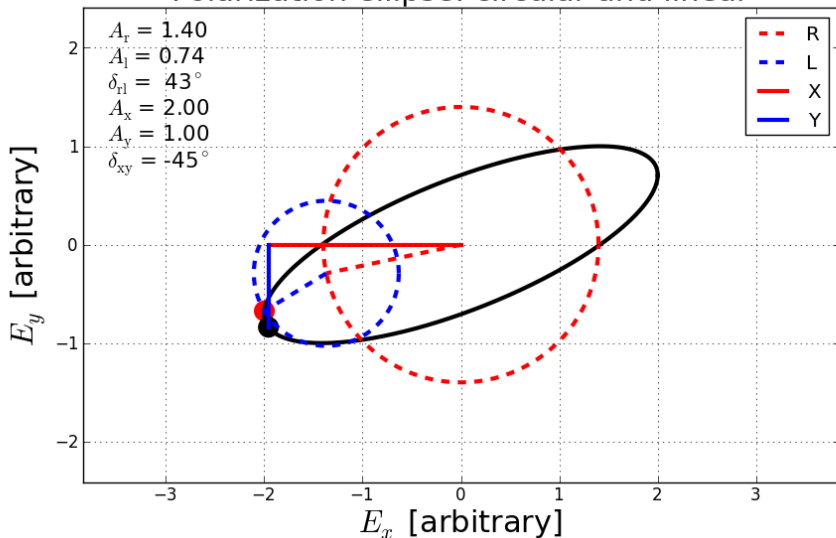
Polarization ellipse: circular and linear



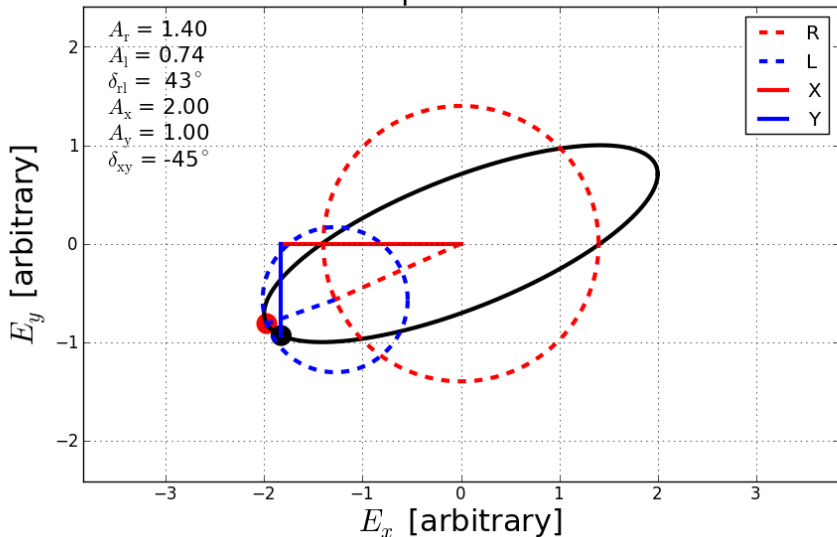
Polarization ellipse: circular and linear



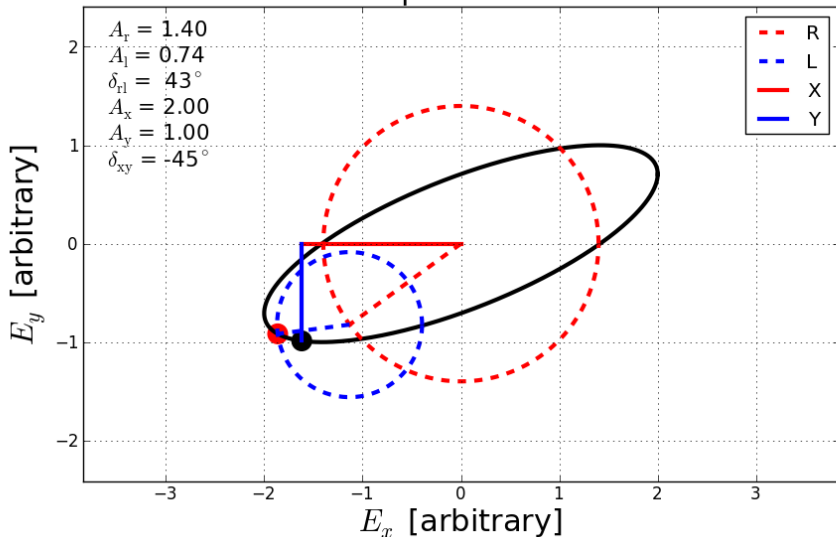
Polarization ellipse: circular and linear



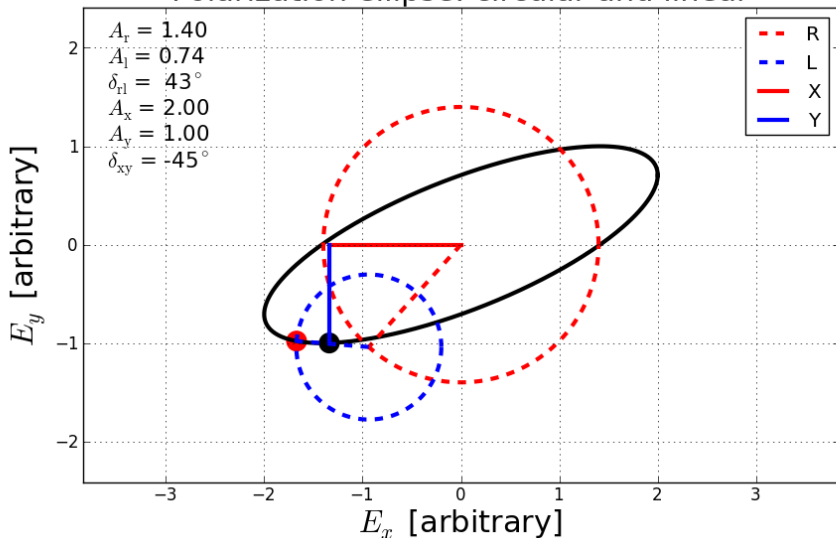
Polarization ellipse: circular and linear



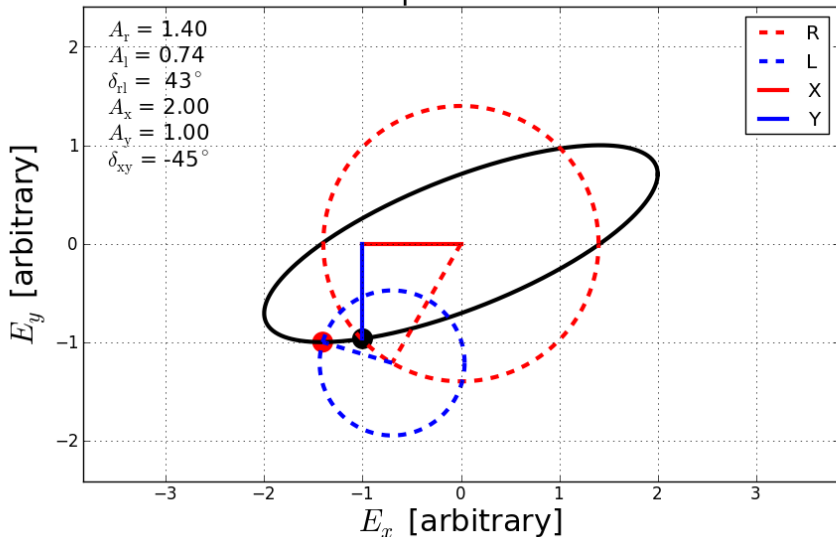
Polarization ellipse: circular and linear



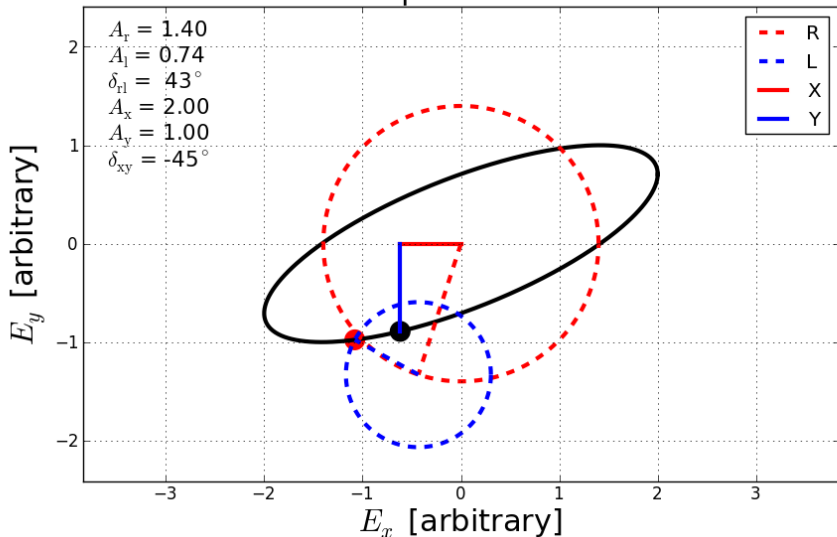
Polarization ellipse: circular and linear



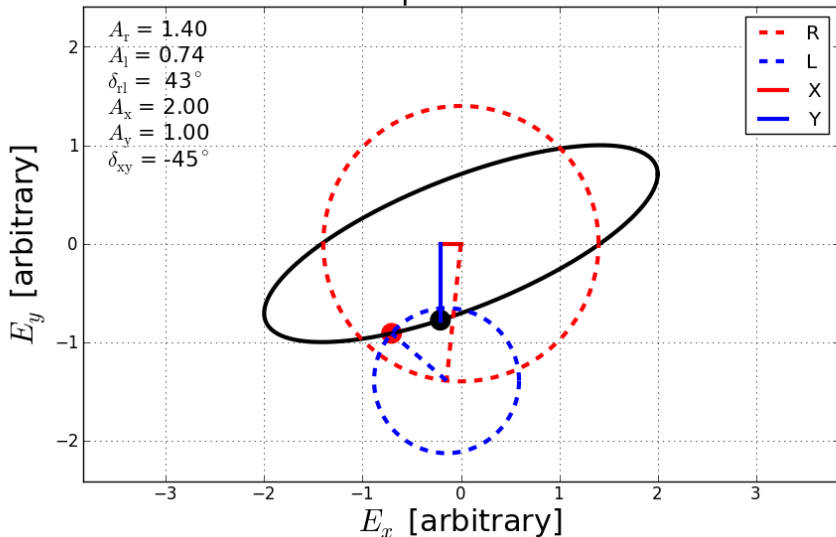
Polarization ellipse: circular and linear



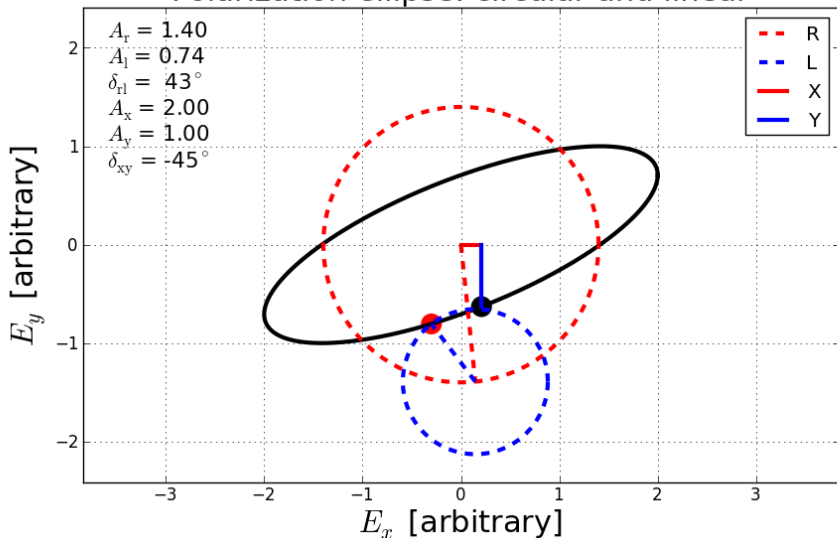
Polarization ellipse: circular and linear



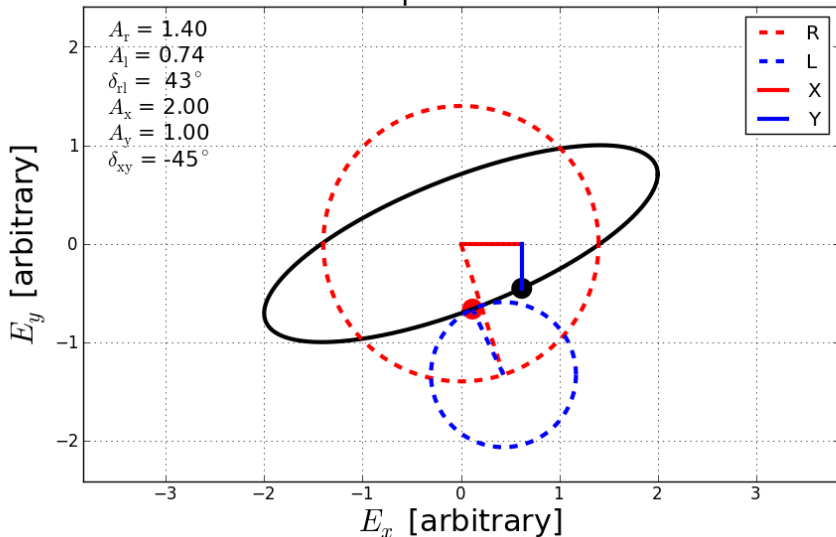
Polarization ellipse: circular and linear



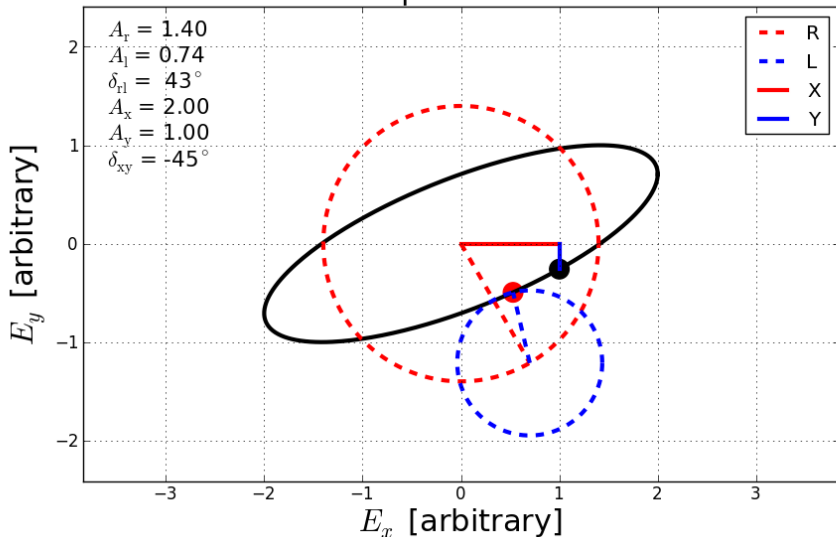
Polarization ellipse: circular and linear



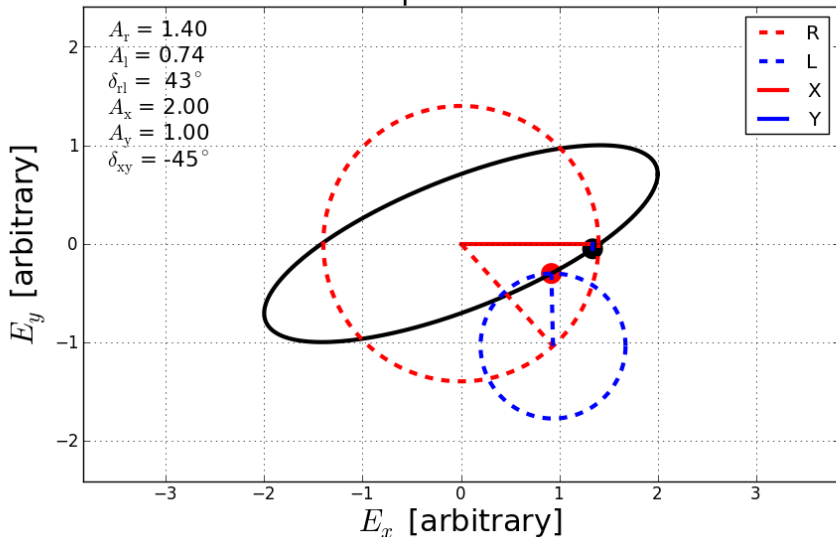
Polarization ellipse: circular and linear



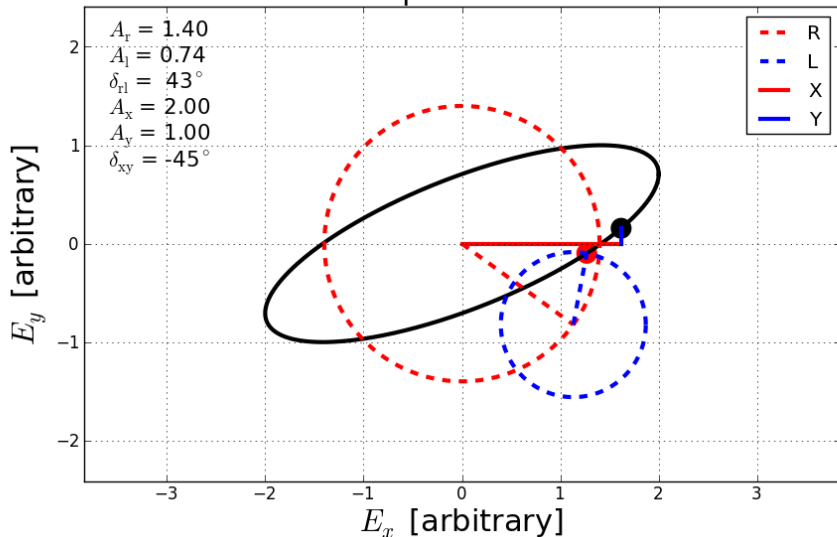
Polarization ellipse: circular and linear



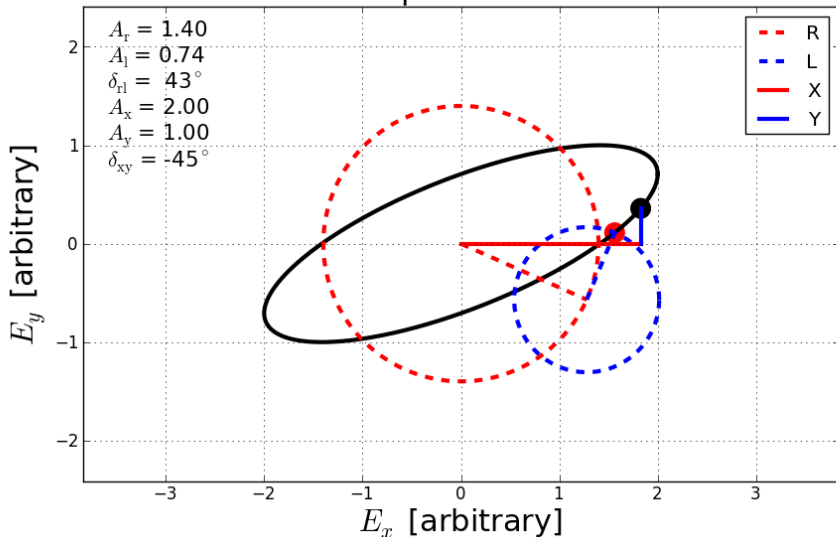
Polarization ellipse: circular and linear



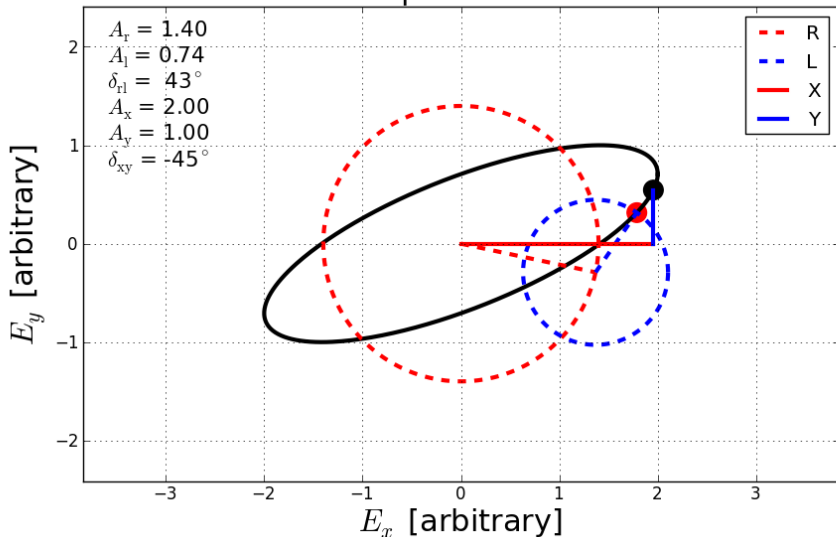
Polarization ellipse: circular and linear



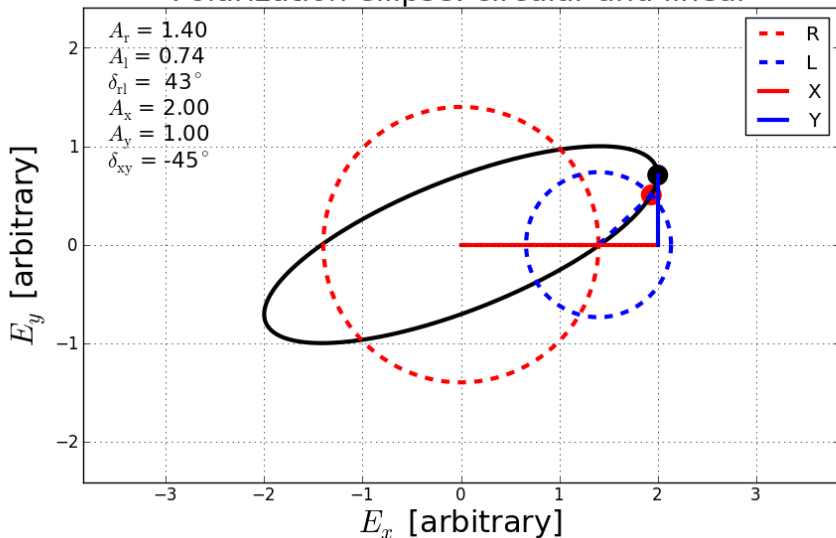
Polarization ellipse: circular and linear



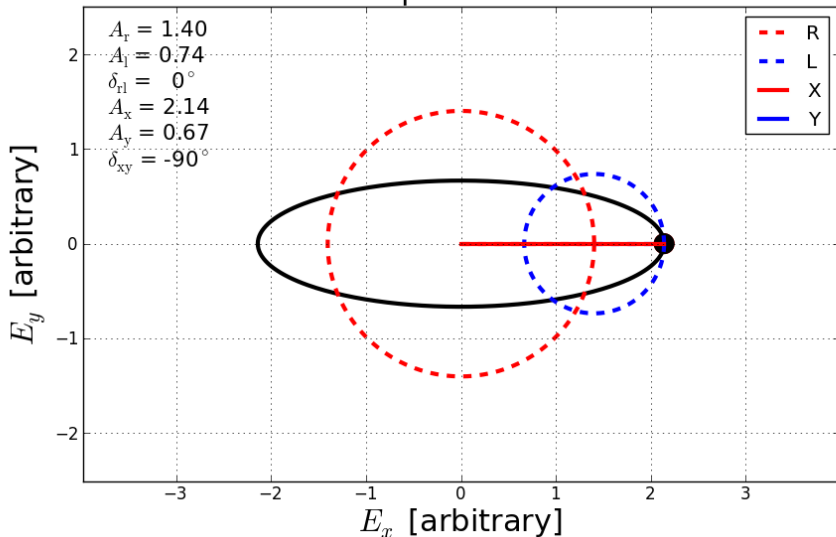
Polarization ellipse: circular and linear



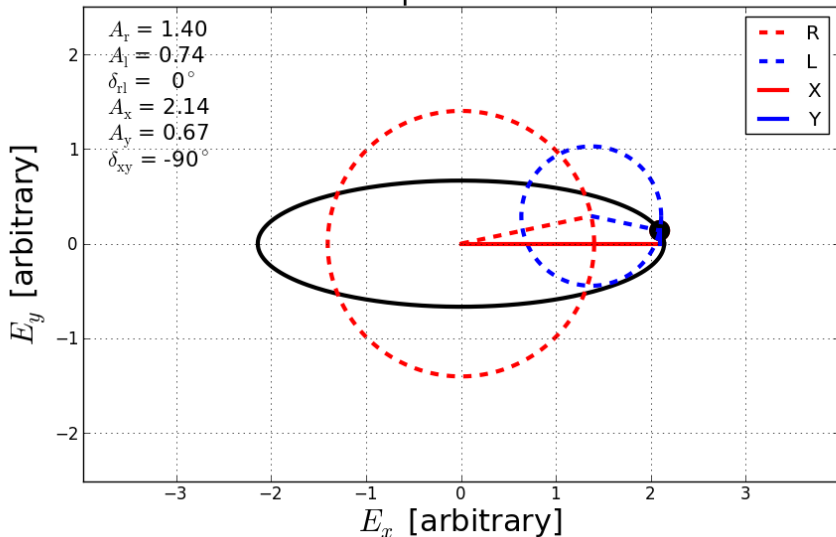
Polarization ellipse: circular and linear



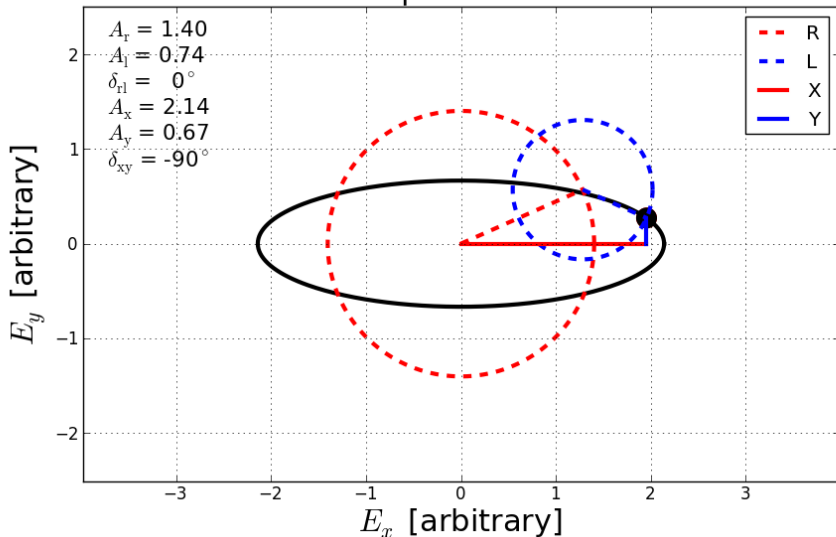
Polarization ellipse: circular and linear



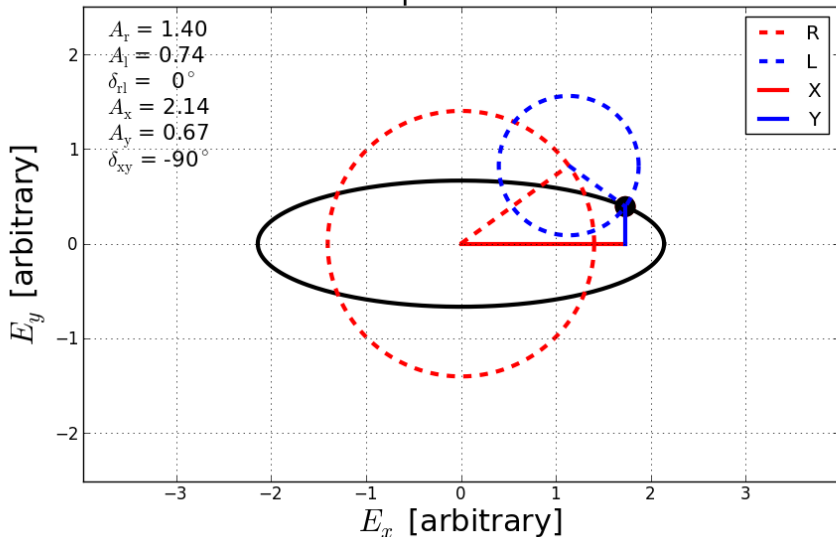
Polarization ellipse: circular and linear

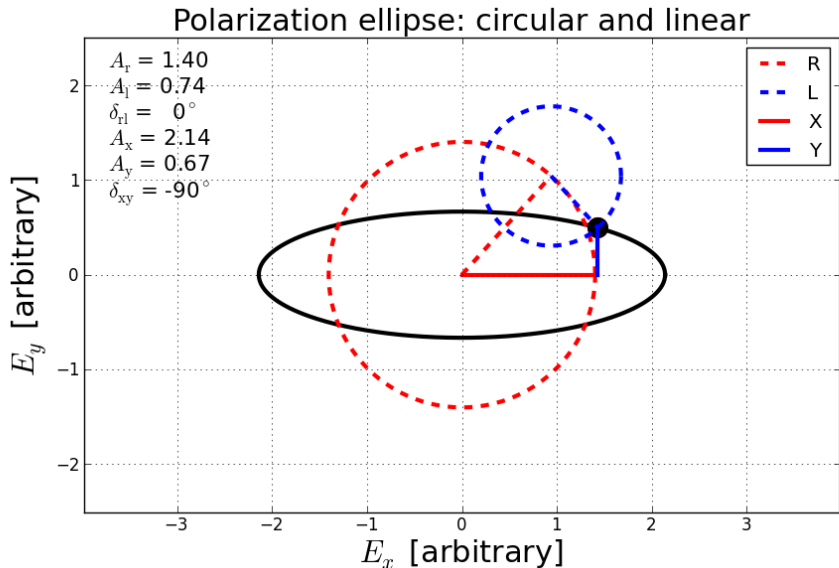


Polarization ellipse: circular and linear

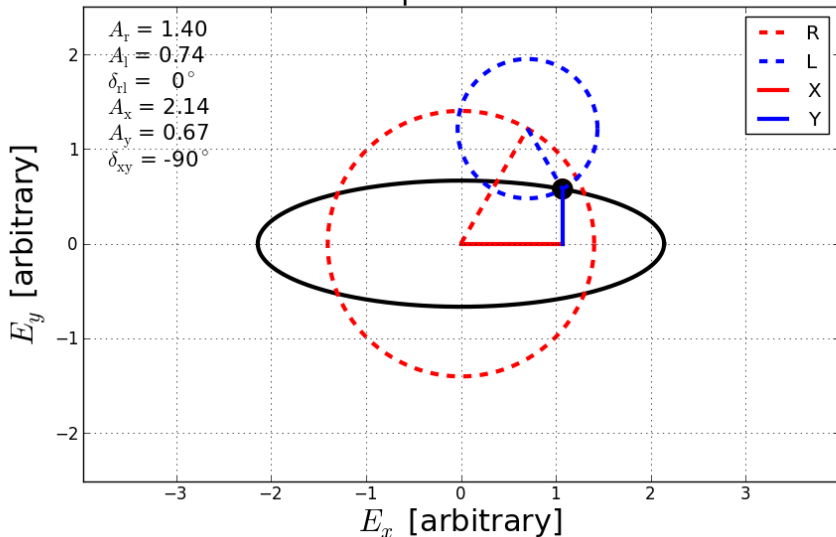


Polarization ellipse: circular and linear

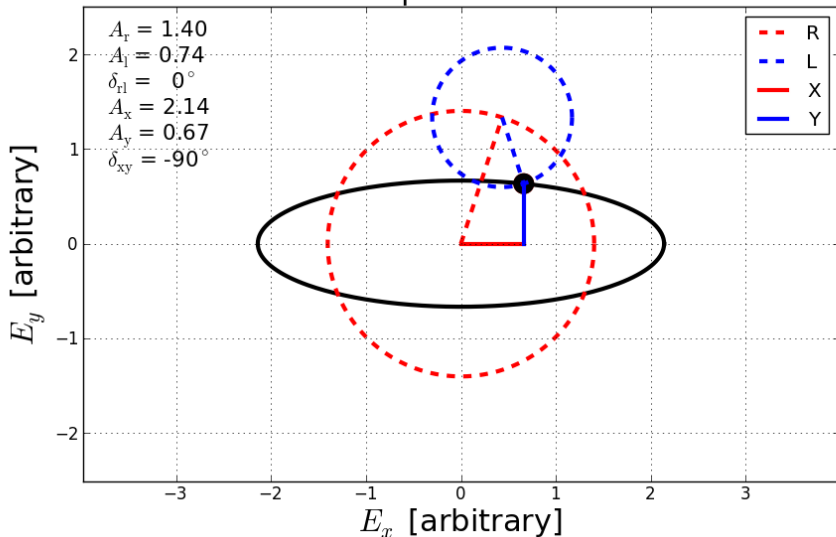




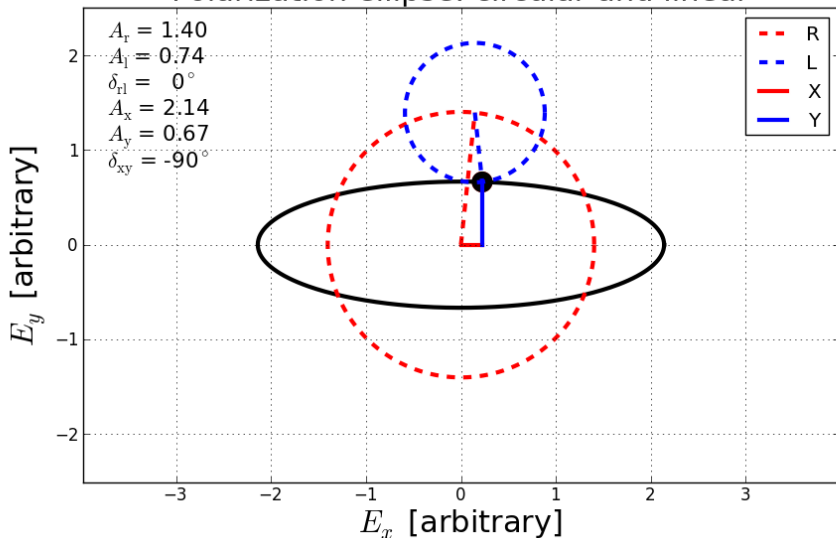
Polarization ellipse: circular and linear



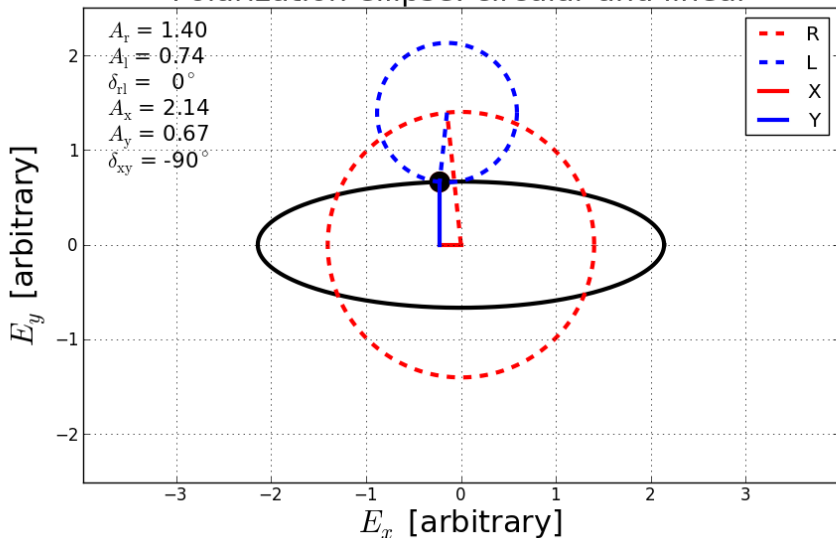
Polarization ellipse: circular and linear



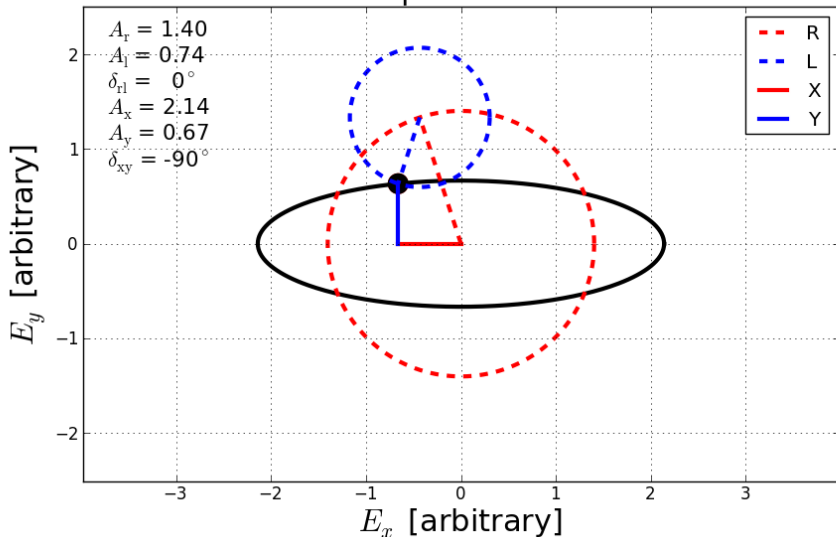
Polarization ellipse: circular and linear



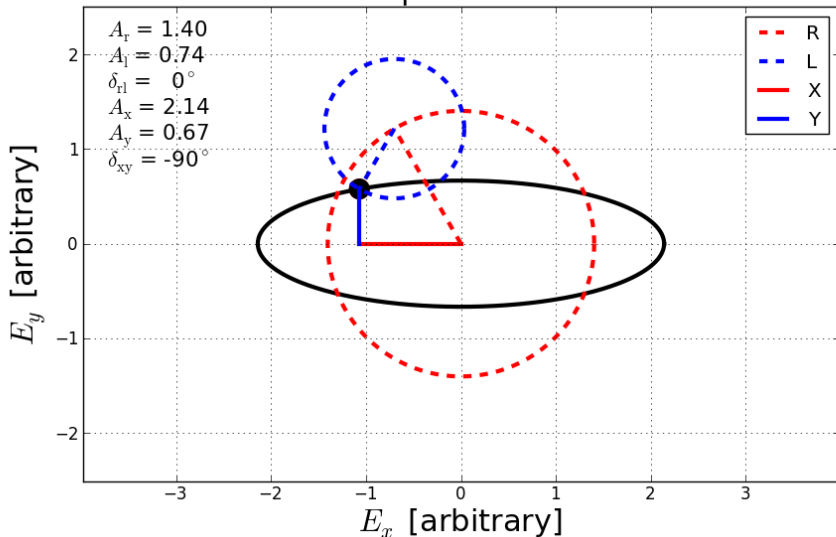
Polarization ellipse: circular and linear



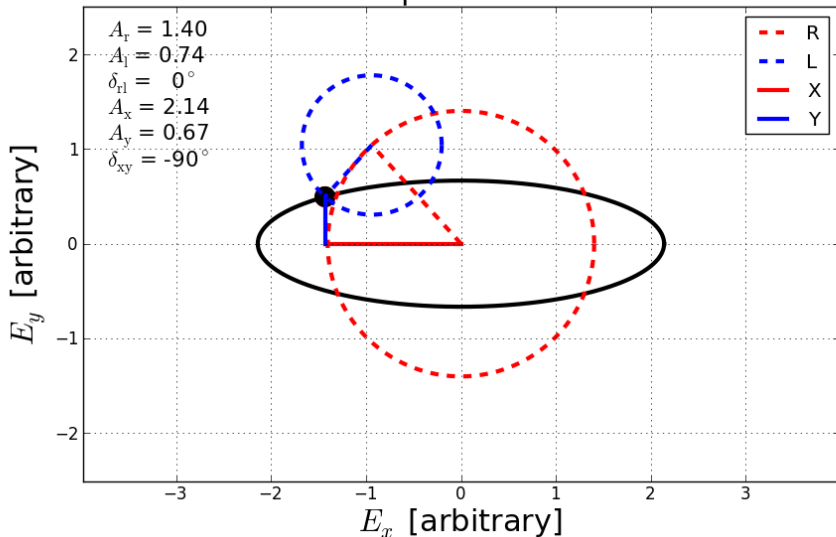
Polarization ellipse: circular and linear



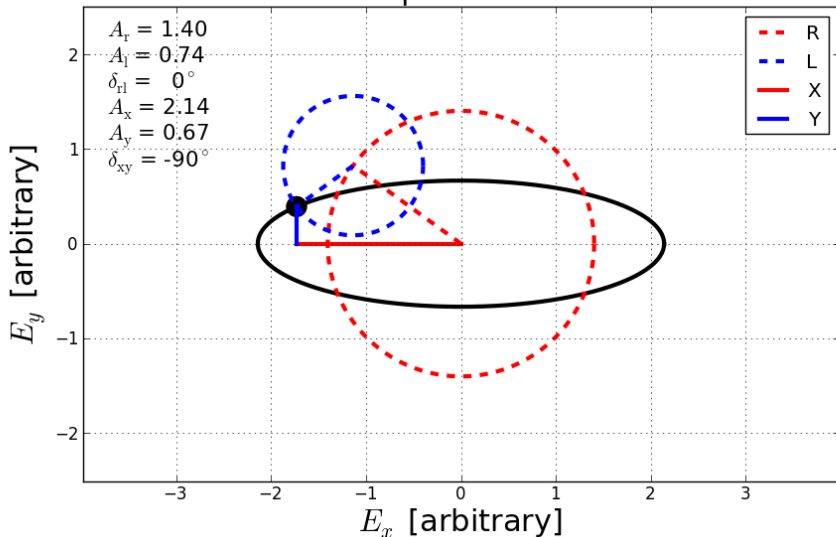
Polarization ellipse: circular and linear



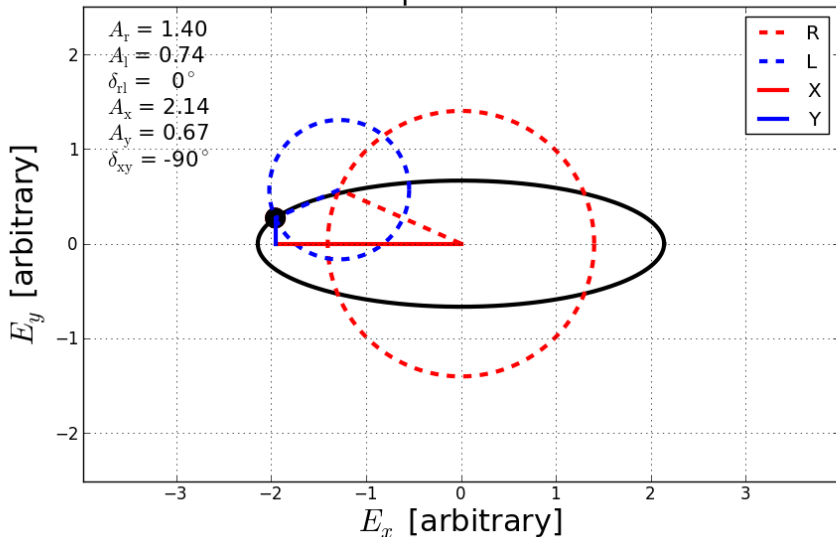
Polarization ellipse: circular and linear



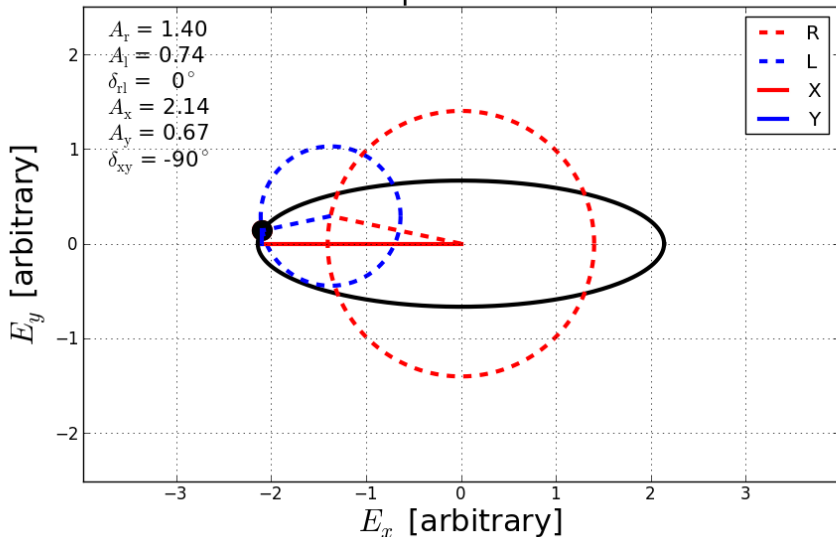
Polarization ellipse: circular and linear



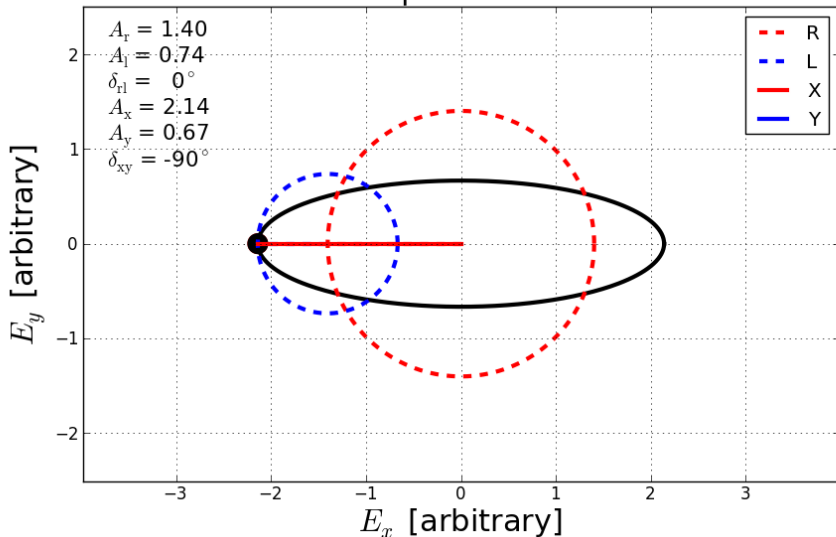
Polarization ellipse: circular and linear



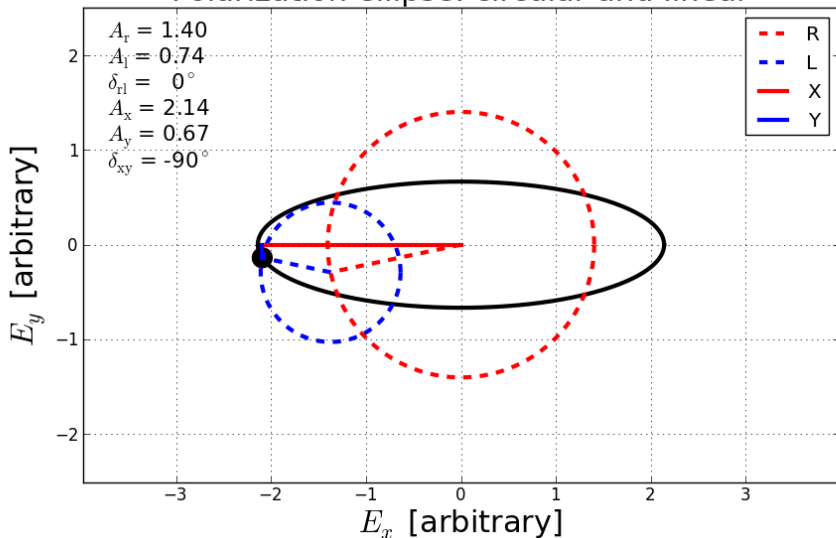
Polarization ellipse: circular and linear



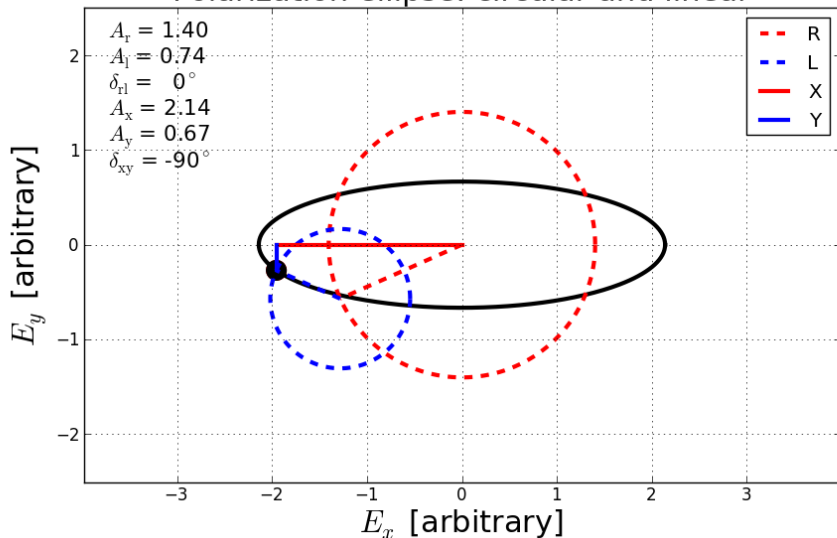
Polarization ellipse: circular and linear



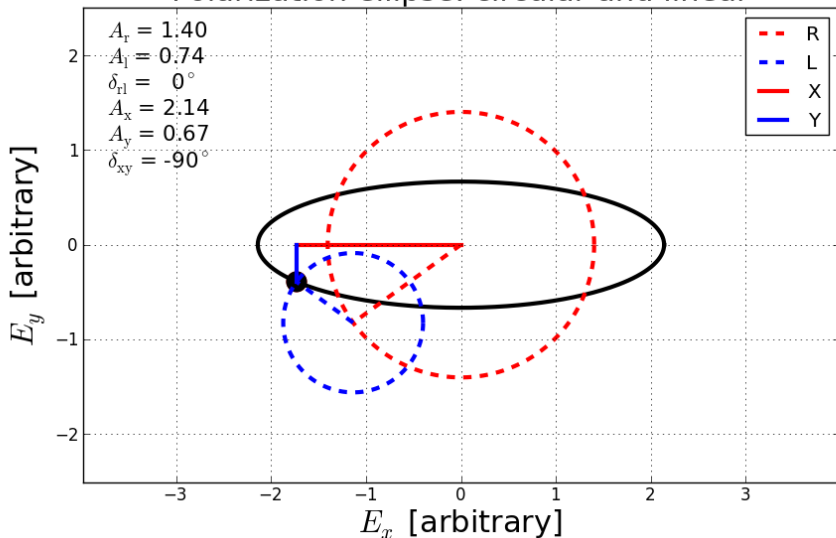
Polarization ellipse: circular and linear



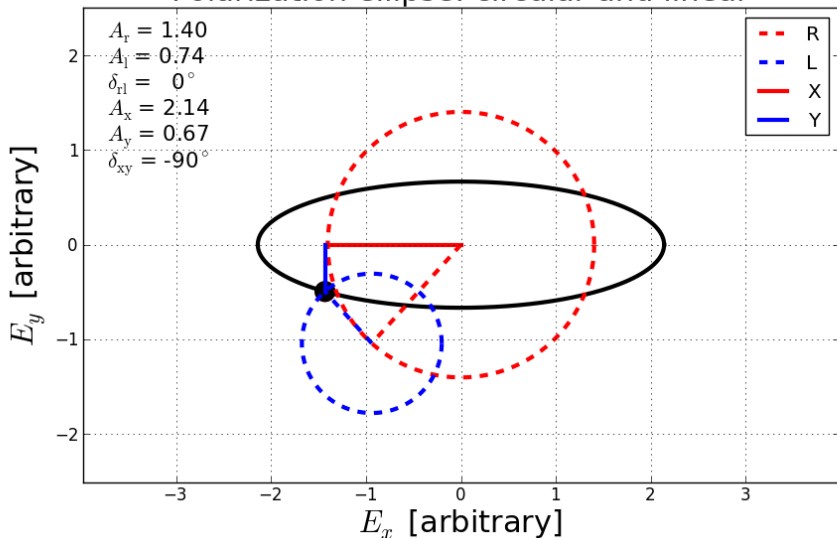
Polarization ellipse: circular and linear



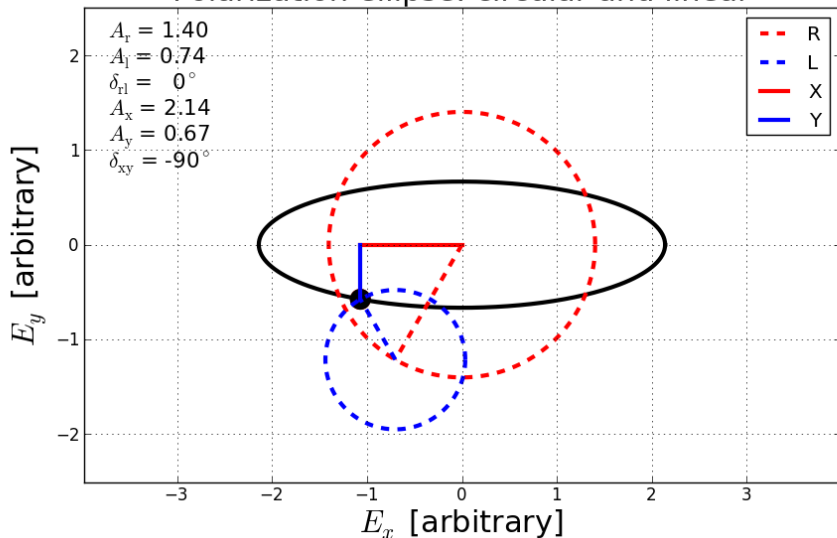
Polarization ellipse: circular and linear



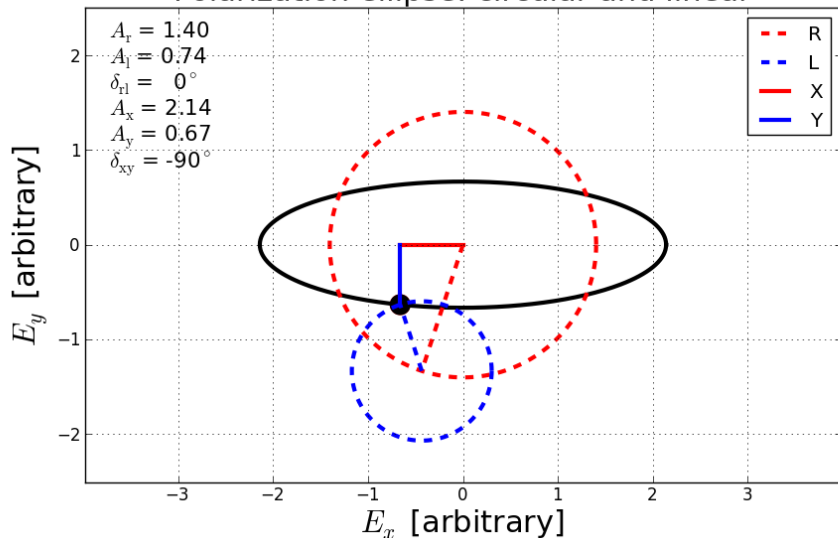
Polarization ellipse: circular and linear



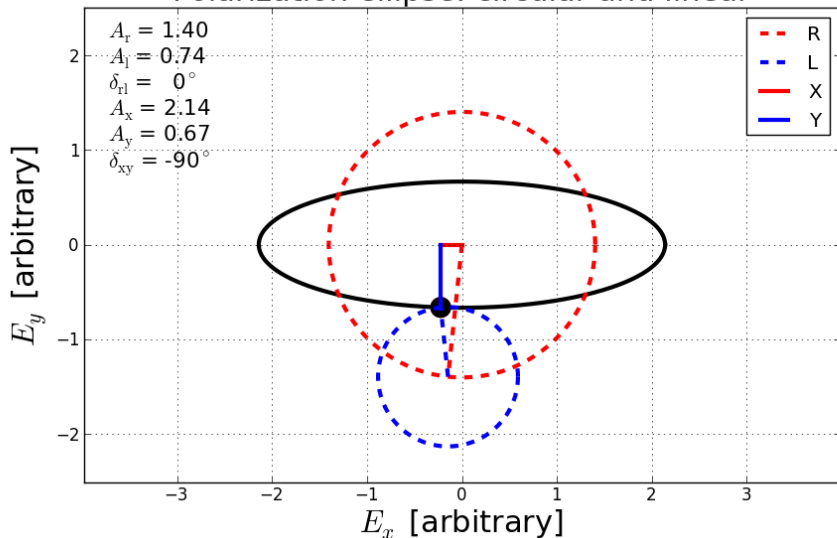
Polarization ellipse: circular and linear



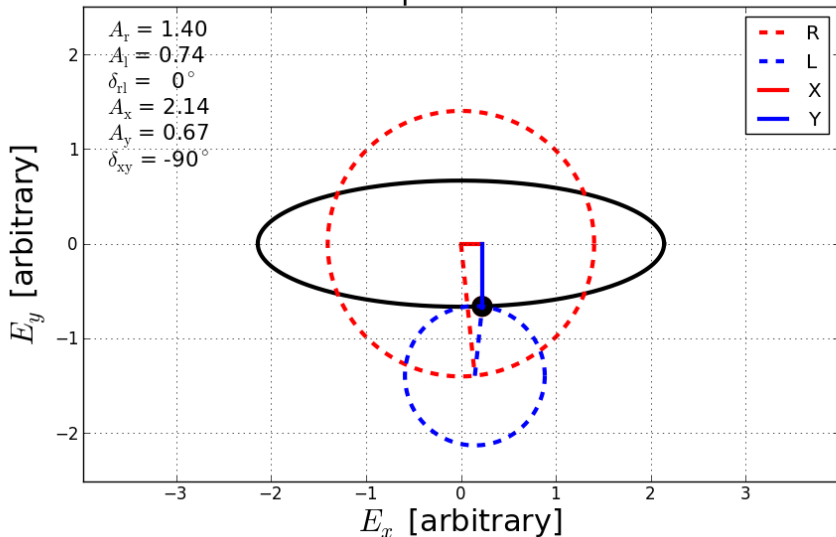
Polarization ellipse: circular and linear



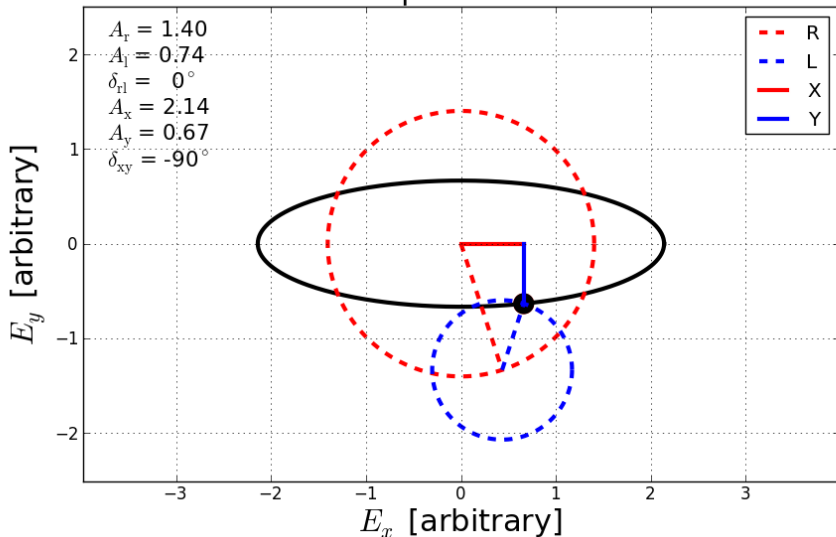
Polarization ellipse: circular and linear



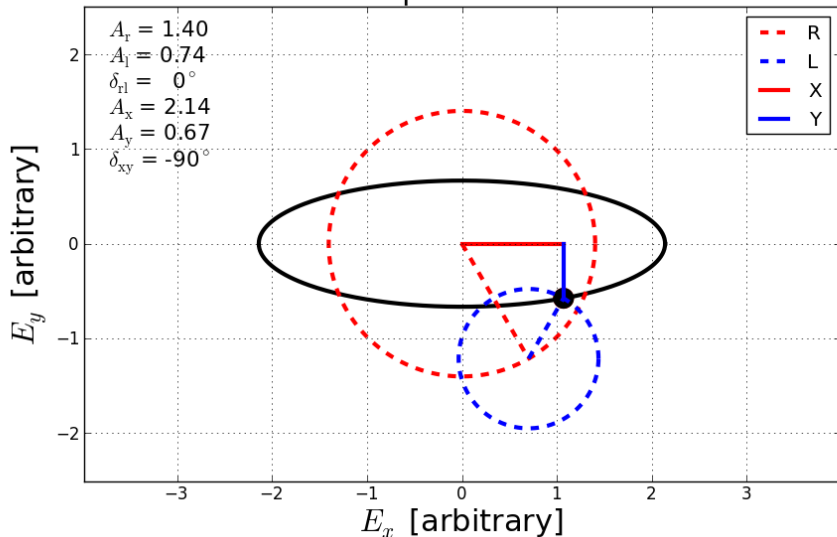
Polarization ellipse: circular and linear



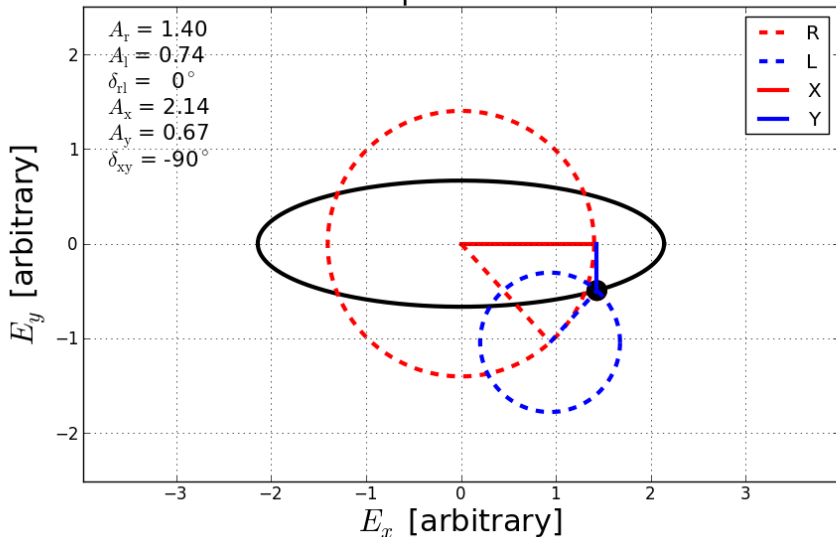
Polarization ellipse: circular and linear



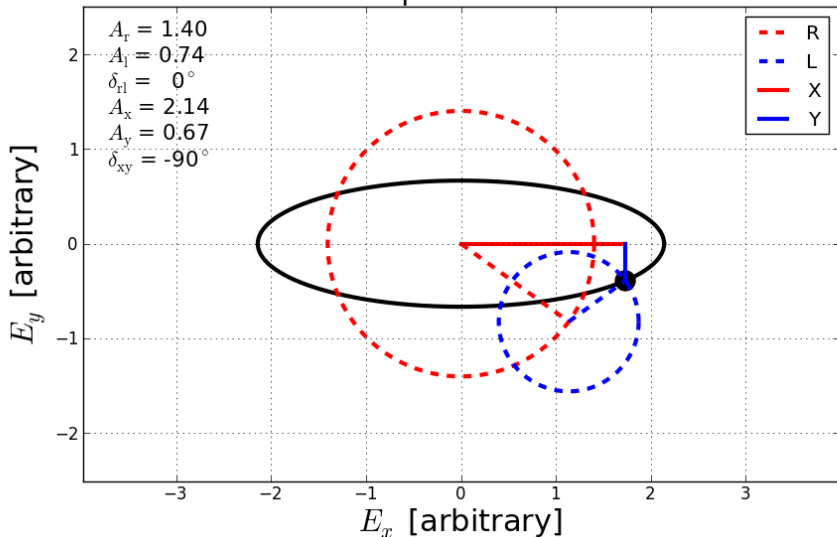
Polarization ellipse: circular and linear

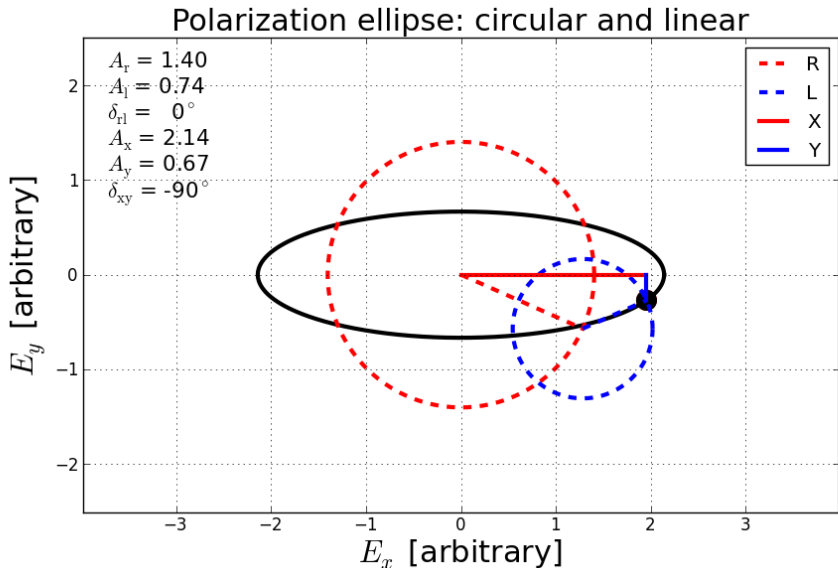


Polarization ellipse: circular and linear

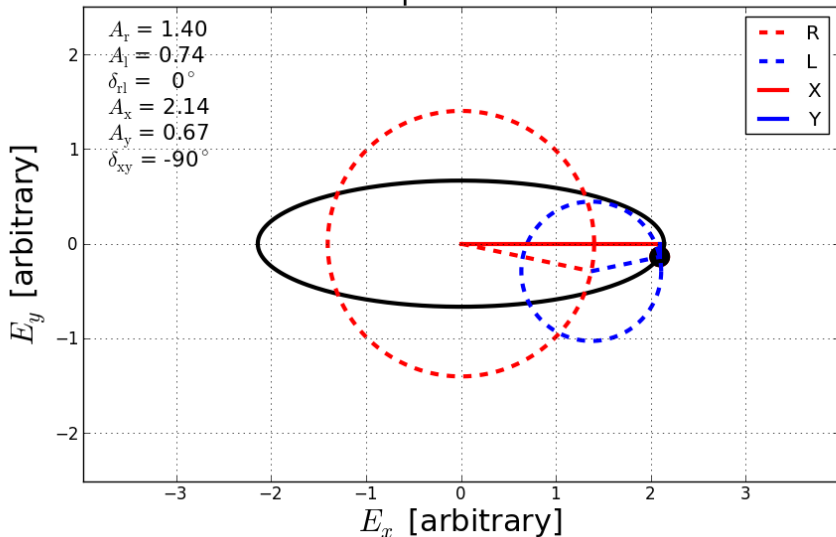


Polarization ellipse: circular and linear

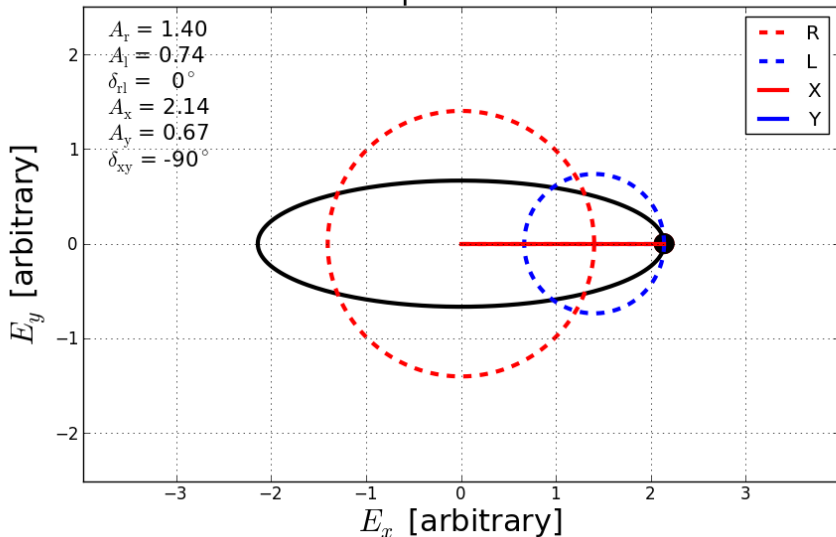




Polarization ellipse: circular and linear



Polarization ellipse: circular and linear



- Three parameters enough
- Same units is convenient
- George Stokes defined four parameters (1856)
- Chandrasekhar introduced them to astronomy (1946)

$$I = A_x^2 + A_y^2$$

$$Q = A_x^2 - A_y^2$$

$$U = 2A_x A_y \cos \delta_{xy}$$

$$V = 2A_x A_y \sin \delta_{xy}$$

$$I = A_r^2 + A_l^2$$

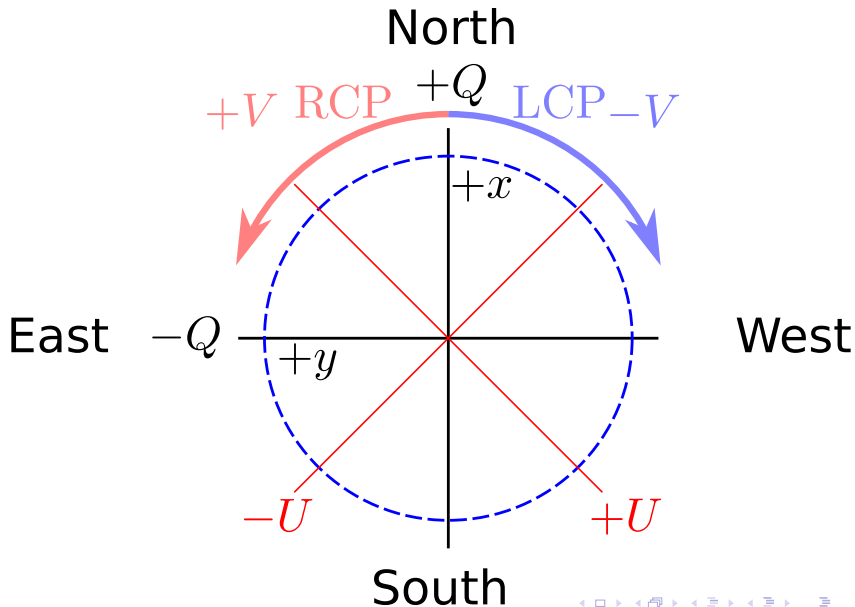
$$Q = 2A_r A_l \cos \delta_{rl}$$

$$U = 2A_r A_l \sin \delta_{rl}$$

$$V = A_r^2 - A_l^2$$

- **Monochromatic** wave 100% polarized:

$$I^2 = Q^2 + U^2 + V^2$$



Quasi-monochromatic approximation

- Monochromatic radiation does not exist
- Finite bandwidth $\Delta\nu$; averaging time $\tau \gg \Delta\nu^{-1}$

$$I = \langle A_x^2 \rangle + \langle A_y^2 \rangle$$

$$Q = \langle A_x^2 \rangle - \langle A_y^2 \rangle$$

$$U = \langle 2A_x A_y \cos \delta_{xy} \rangle$$

$$V = \langle 2A_x A_y \sin \delta_{xy} \rangle$$

$$I = \langle A_r^2 \rangle + \langle A_l^2 \rangle$$

$$Q = \langle 2A_r A_l \cos \delta_{rl} \rangle$$

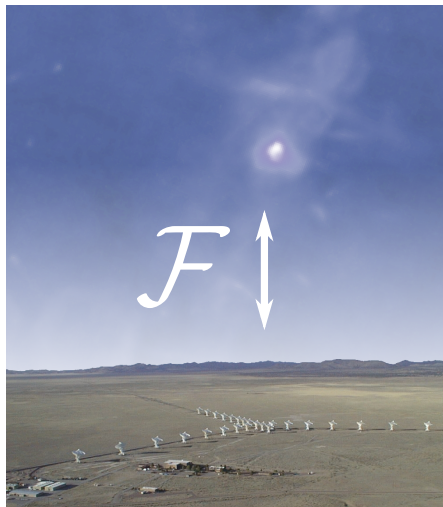
$$U = \langle 2A_r A_l \sin \delta_{rl} \rangle$$

$$V = \langle A_r^2 \rangle - \langle A_l^2 \rangle$$

$$I^2 \geq Q^2 + U^2 + V^2$$

- Fractional linear pol: $p = \sqrt{Q^2 + U^2} / I \leq 1$
- Fractional circular pol: $v = \|V\| / I \leq 1$

- 1 EM wave physics
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- 4 Messy reality
- 5 An example



$$\begin{aligned} \mathcal{I}(u, v) &= \mathcal{F}^+(I(l, m)) \\ \mathcal{Q}(u, v) &= \mathcal{F}^+(Q(l, m)) \\ \mathcal{U}(u, v) &= \mathcal{F}^+(U(l, m)) \\ \mathcal{V}(u, v) &= \mathcal{F}^+(V(l, m)), \end{aligned}$$

where

$$\mathcal{F}^+(f) = \int_{lm} f e^{+2\pi i \nu (ul+vm)/c} dl dm$$

Cartesian

$$E_x = \Re \left\{ A_x e^{2\pi i \nu t} \right\}$$

$$E_y = \Re \left\{ A_y e^{i\delta_{xy}} e^{2\pi i \nu t} \right\}$$

$$I = \langle A_x^2 \rangle + \langle A_y^2 \rangle = \langle E_x E_x^* \rangle + \langle E_y E_y^* \rangle$$

$$Q = \langle A_x^2 \rangle - \langle A_y^2 \rangle = \langle E_x E_x^* \rangle - \langle E_y E_y^* \rangle$$

$$U = \langle 2A_x A_y \cos \delta_{xy} \rangle = \langle E_x E_y^* \rangle + \langle E_y E_x^* \rangle$$

$$V = \langle 2A_x A_y \sin \delta_{xy} \rangle = i \left(\langle E_x E_y^* \rangle - \langle E_y E_x^* \rangle \right)$$

Circular

$$E_r = \Re \left\{ A_r e^{2\pi i \nu t} \right\}$$

$$E_l = \Re \left\{ A_l e^{i\delta_{rl}} e^{2\pi i \nu t} \right\}$$

$$I = \langle A_r^2 \rangle + \langle A_l^2 \rangle$$

$$= \langle E_r E_r^* \rangle + \langle E_l E_l^* \rangle$$

$$Q = \langle 2A_r A_l \cos \delta_{rl} \rangle$$

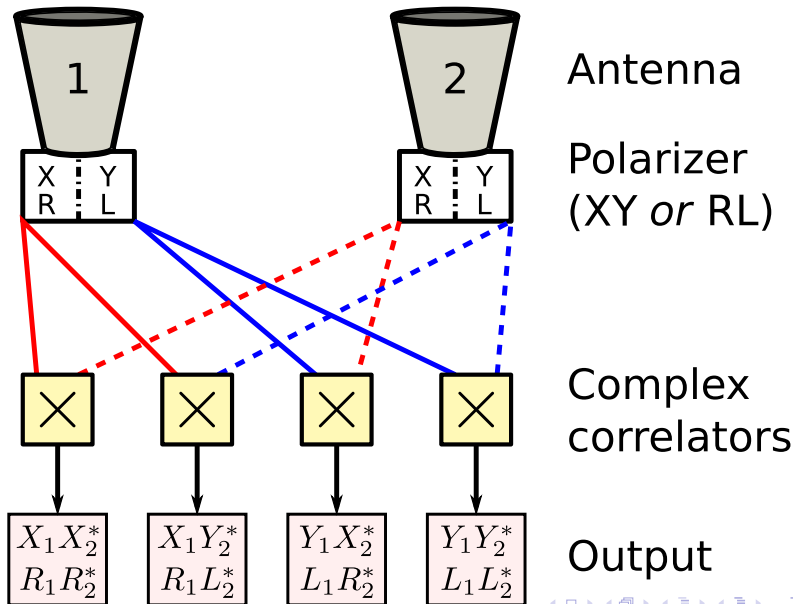
$$= \langle E_r E_l^* \rangle + \langle E_l E_r^* \rangle$$

$$U = \langle 2A_r A_l \sin \delta_{rl} \rangle$$

$$= i(\langle E_r E_l^* \rangle - \langle E_l E_r^* \rangle)$$

$$V = \langle A_r^2 \rangle - \langle A_l^2 \rangle$$

$$= \langle E_r E_r^* \rangle - \langle E_l E_l^* \rangle$$

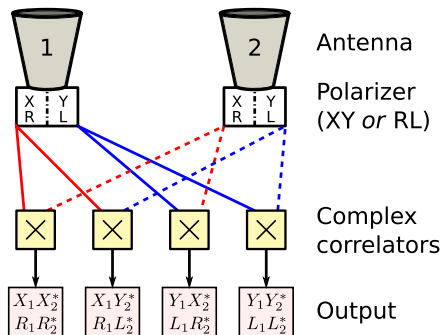


Antenna

Polarizer
(XY or RL)

Complex correlators

Output



- From here on, $\langle \cdot \rangle$ is implied for correlator outputs.

Cartesian

$$\mathcal{I} = x_1 x_2^* + y_1 y_2^*$$

$$\mathcal{Q} = x_1 x_2^* - y_1 y_2^*$$

$$\mathcal{U} = x_1 y_2^* + y_1 x_2^*$$

$$\mathcal{V} = i(x_1 y_2^* - y_1 x_2^*)$$

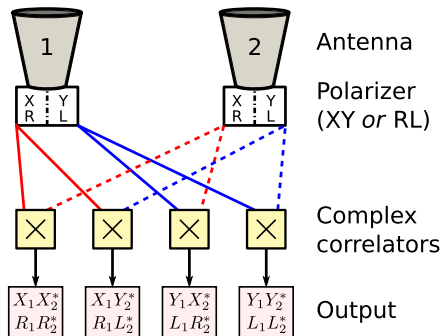
Circular

$$\mathcal{I} = r_1 r_2^* + l_1 l_2^*$$

$$\mathcal{Q} = r_1 l_2^* + l_1 r_2^*$$

$$\mathcal{U} = i(r_1 l_2^* - l_1 r_2^*)$$

$$\mathcal{V} = r_1 r_2^* - l_1 l_2^*$$



From here on, p and q designate either x and y , or r and l .

- Polarizers produce vector:

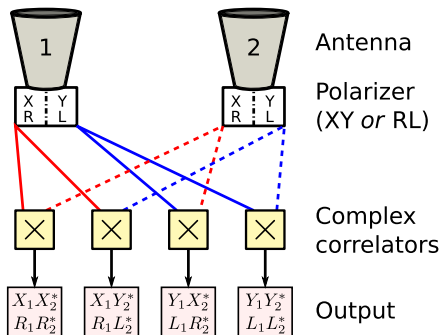
$$\mathbf{e}_i = \begin{pmatrix} p_i \\ q_i \end{pmatrix}$$

- Correlator multiplies:

$$\mathbf{E}_{ij} = \mathbf{e}_i \mathbf{e}_j^\dagger = \begin{pmatrix} p_i \\ q_i \end{pmatrix} \begin{pmatrix} p_j^* & q_j^* \end{pmatrix}$$

$$\mathbf{E}_{ij} = \begin{pmatrix} p_i p_j^* & p_i q_j^* \\ q_i p_j^* & q_i q_j^* \end{pmatrix}$$

- \mathbf{E}_{ij} is the **coherency matrix**



Until now...

- Assumed all systems perfect

From now...

- Assume all systems linear:

$$\mathbf{e}'_i = \mathbf{J}_i \mathbf{e}_i$$

- \mathbf{J}_i (2×2) is **Jones matrix**
- Cross correlation:

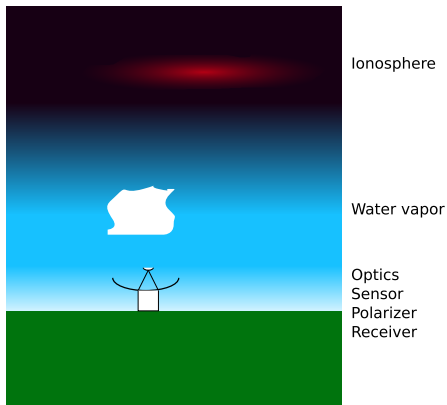
$$\mathbf{E}'_{ij} = \mathbf{e}'_i \mathbf{e}'_j{}^\dagger$$

$$\mathbf{E}'_{ij} = \mathbf{J}_i \mathbf{e}_i (\mathbf{J}_j \mathbf{e}_j)^\dagger$$

$$\mathbf{E}'_{ij} = \mathbf{J}_i \mathbf{e}_i \mathbf{e}_j{}^\dagger \mathbf{J}_j^\dagger$$

$$\mathbf{E}'_{ij} = \mathbf{J}_i \mathbf{E}_{ij} \mathbf{J}_j^\dagger$$

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- The measurement equation:

$$\mathbf{E}'_{ij} = \mathbf{J}_i \mathbf{E}_{ij} \mathbf{J}_j^\dagger$$

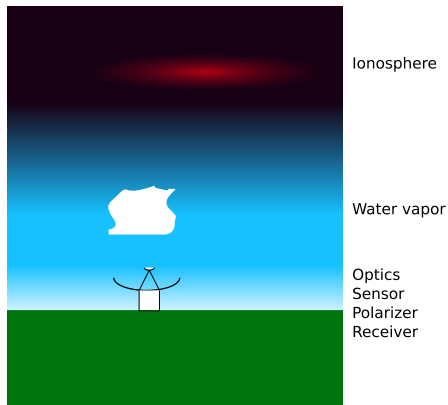
- Invertable!

$$\mathbf{E}_{ij} = \mathbf{J}_i^{-1} \mathbf{E}'_{ij} \mathbf{J}_j^{\dagger-1},$$

- where

$$\mathbf{J} = \text{RPDOWTF} \dots$$

- ... riiiiight...



- Perfect instrument:

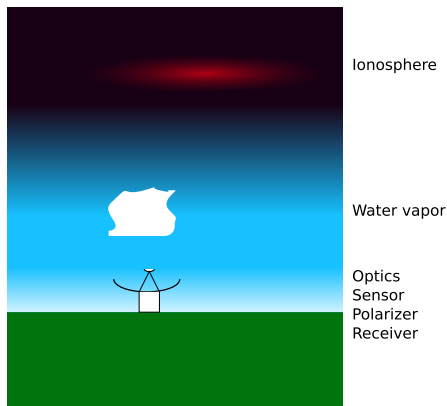
$$\mathbf{J} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

- Ionospheric time delay:

$$\mathbf{J} = \begin{pmatrix} e^{2\pi i\nu\tau} & 0 \\ 0 & e^{2\pi i\nu\tau} \end{pmatrix}$$

- Receiver gain:

$$\mathbf{J} = \begin{pmatrix} g_p & 0 \\ 0 & g_q \end{pmatrix}$$



- Polarization leakage:

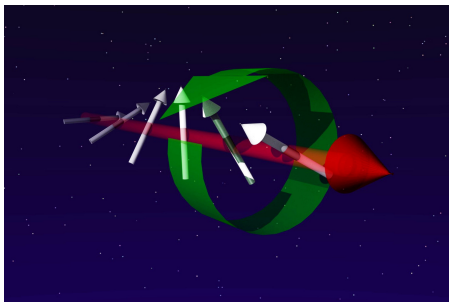
$$\mathbf{J} = \begin{pmatrix} g_p & d_{qp} \\ d_{pq} & g_q \end{pmatrix}$$

- Parallax angle or feed rotation XY:

$$\mathbf{J} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

- Parallax angle or feed rotation RL:

$$\mathbf{J} = \begin{pmatrix} 1 & 0 \\ 0 & e^{-i\theta} \end{pmatrix}$$

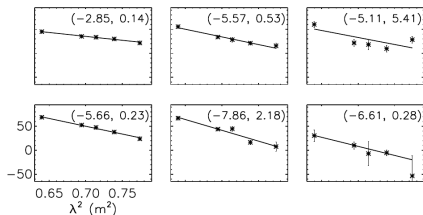


Process

- Modifies polarization state
- Delay between LCP and RCP
- Rotates linear pol angle
- $\Delta\chi = \chi_0 + \phi\lambda^2$

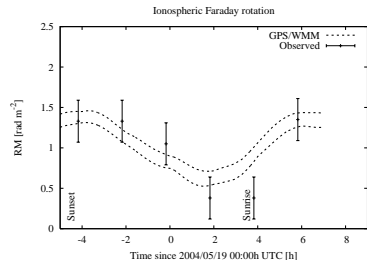
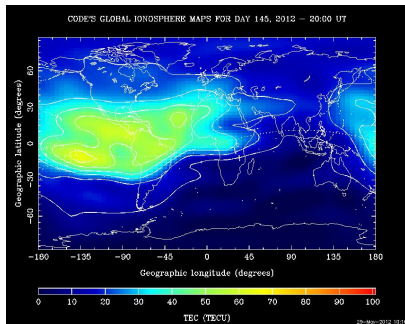
$$\phi = 0.812 \int_{\text{there}}^{\text{here}} n_e \mathbf{B} \cdot d\mathbf{l}$$

λ^2 law *Haverkorn et al. (2001)*



Polarimetry provides

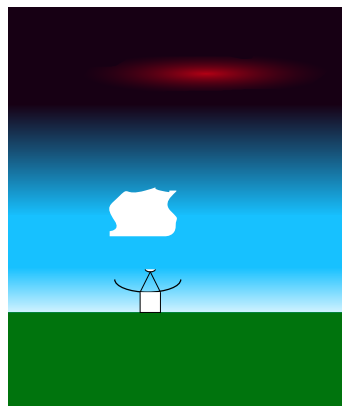
- Source plasma properties
- Intervening plasma properties
- Rare cases: 3D tomography



- Remember: $\Delta\chi = \chi_0 + \phi\lambda^2$
- Faraday depth

$$\phi = 0.812 \int_{\text{there}}^{\text{here}} n_e \mathbf{B} \cdot d\mathbf{l}$$

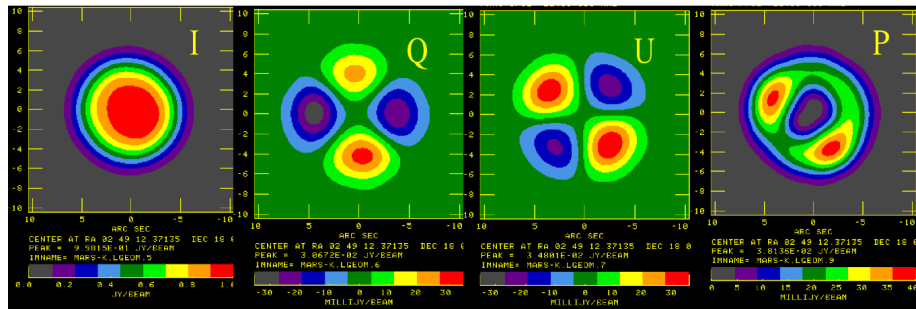
- ionosphere: plasma within Earth's magnetic field
- $\phi \approx -10 - +10 \text{ rad m}^{-2}$
- *Very* significant below 1 GHz
- Use TEC/IGRF models for correction, check with pulsar.



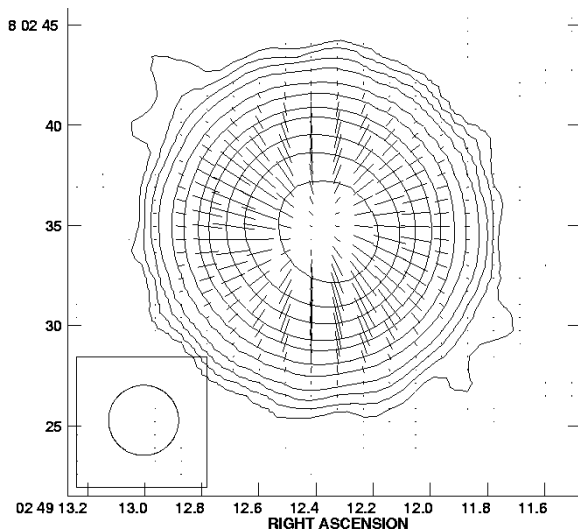
Ionosphere

- $\Delta\chi = \chi_0 + \phi\lambda^2$
- Rotation of linear pol = delay between RCP and LCP
- Antennas see different ionosphere
- Leakage from LL to RR or v.v. after cross correlation
- Rotates $\begin{pmatrix} \mathcal{I} \\ \mathcal{V} \end{pmatrix}$ vector
- Important below 300 MHz at baselines ≥ 20 km

- 1 EM wave physics
- 2 Polarized EM-waves
- 3 Interferometric polarimetry
- 4 Messy reality
- 5 An example**



- Thermal emission
- What do pol vectors look like?
- Why is it even polarized?



Peak contour flux = $9.9738E-01$ JY/BEAM
 Levs = $9.974E-03$ * (-0.250, 0.250, 0.500, 1, 2,
 5, 10, 20, 30, 50, 70, 90)

- Born & Wolf *Principles of optics*
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- Bracewell *The Fourier Transform & Its Applications*
- Hamaker, Bregman & Sault *Understanding radio polarimetry: paper I*(1996)
- Sault, Hamaker & Bregman *paper II*(1996)
- Hamaker & Bregman *paper III* (1996)
- Hamaker *paper IV* (2000)
- Hamaker *paper V* (2006)
- Brentjens & de Bruyn *Faraday rotation measure synthesis* (2005)