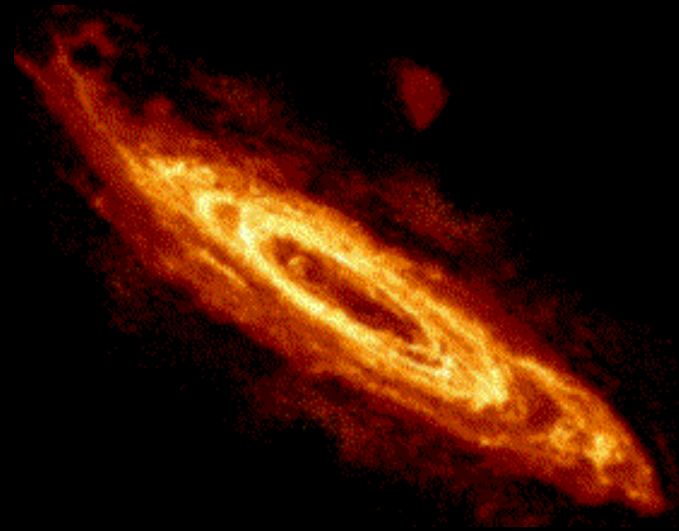


Digital Processing for LOFAR & Embrace

André W. Gunst

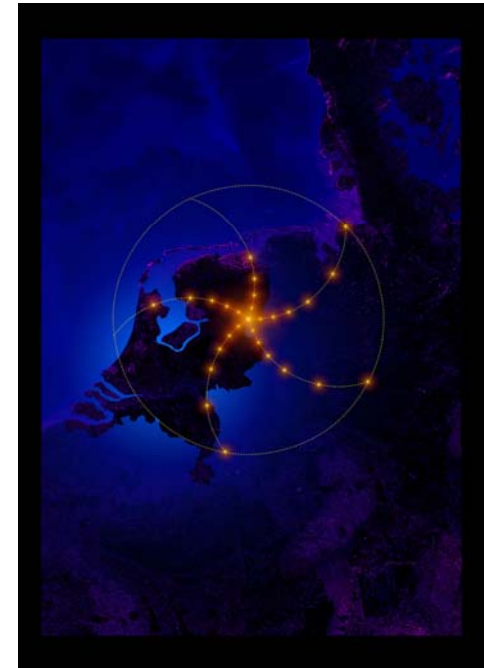


- Many small antenna elements
 - ☞ No mechanical parts necessary
 - ☞ Mass production
 - ☞ Electronic beam steering possible
 - ☞ Multi-beam capabilities

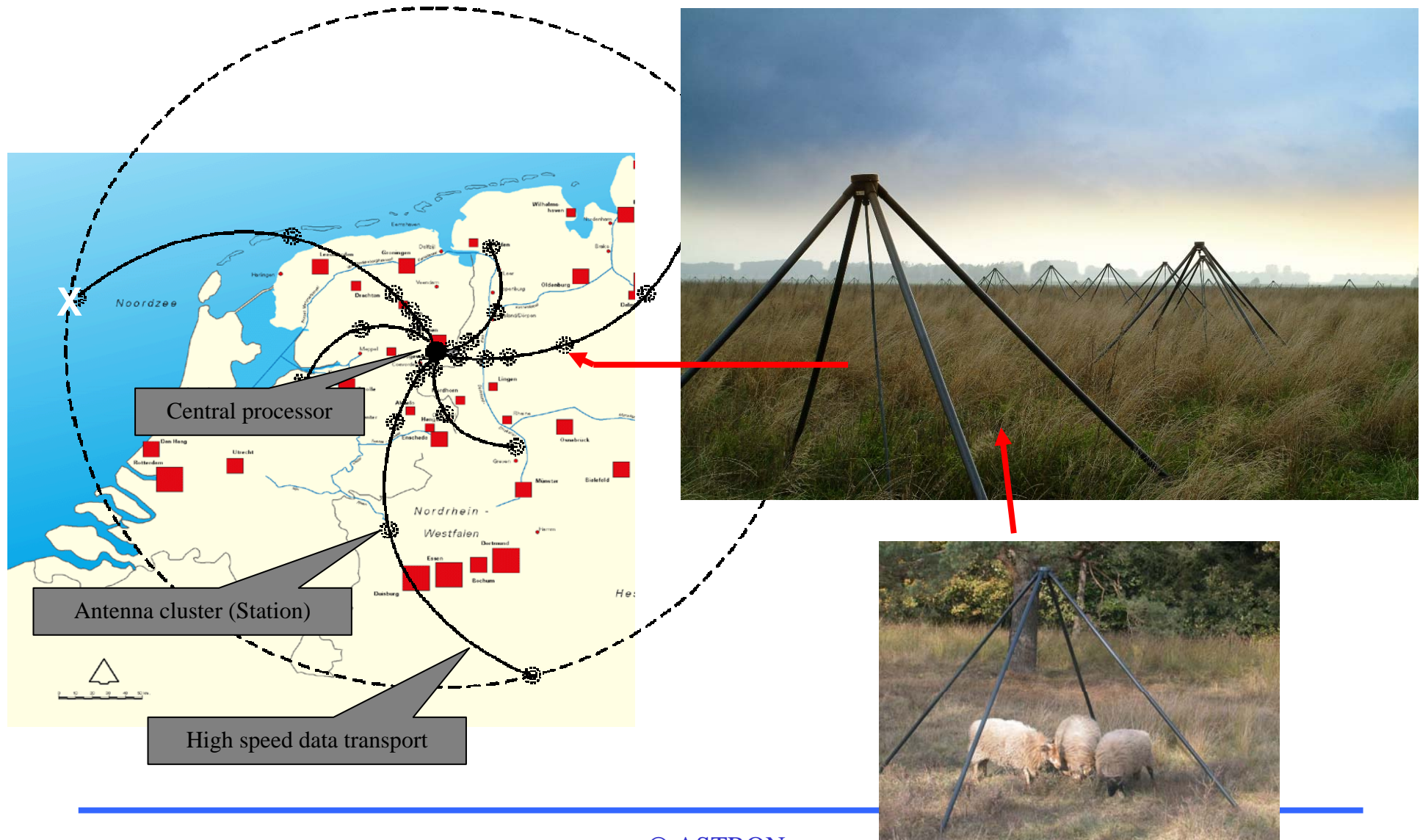


LOW Frequency Array characteristics

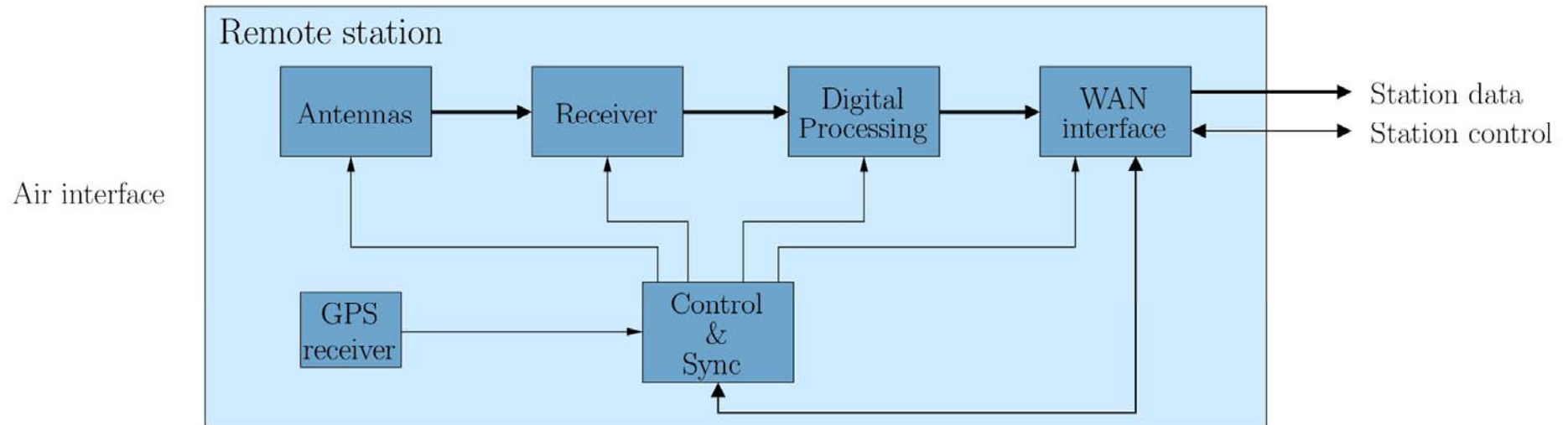
- Phased array antennas
- Diameter of about 100 km
- Low frequency window
- Order 10^4 antennas
- RFI robust
- Large data network
- Central distributed super computer
- Flexible/reconfigurable



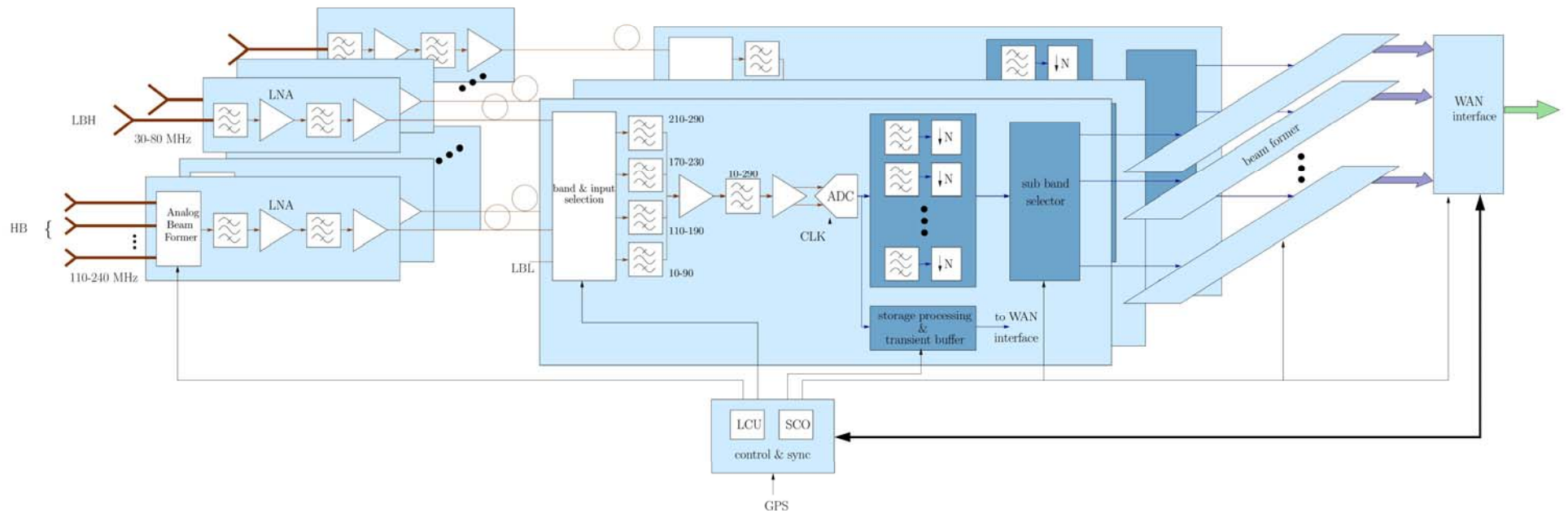
LOFAR Configuration



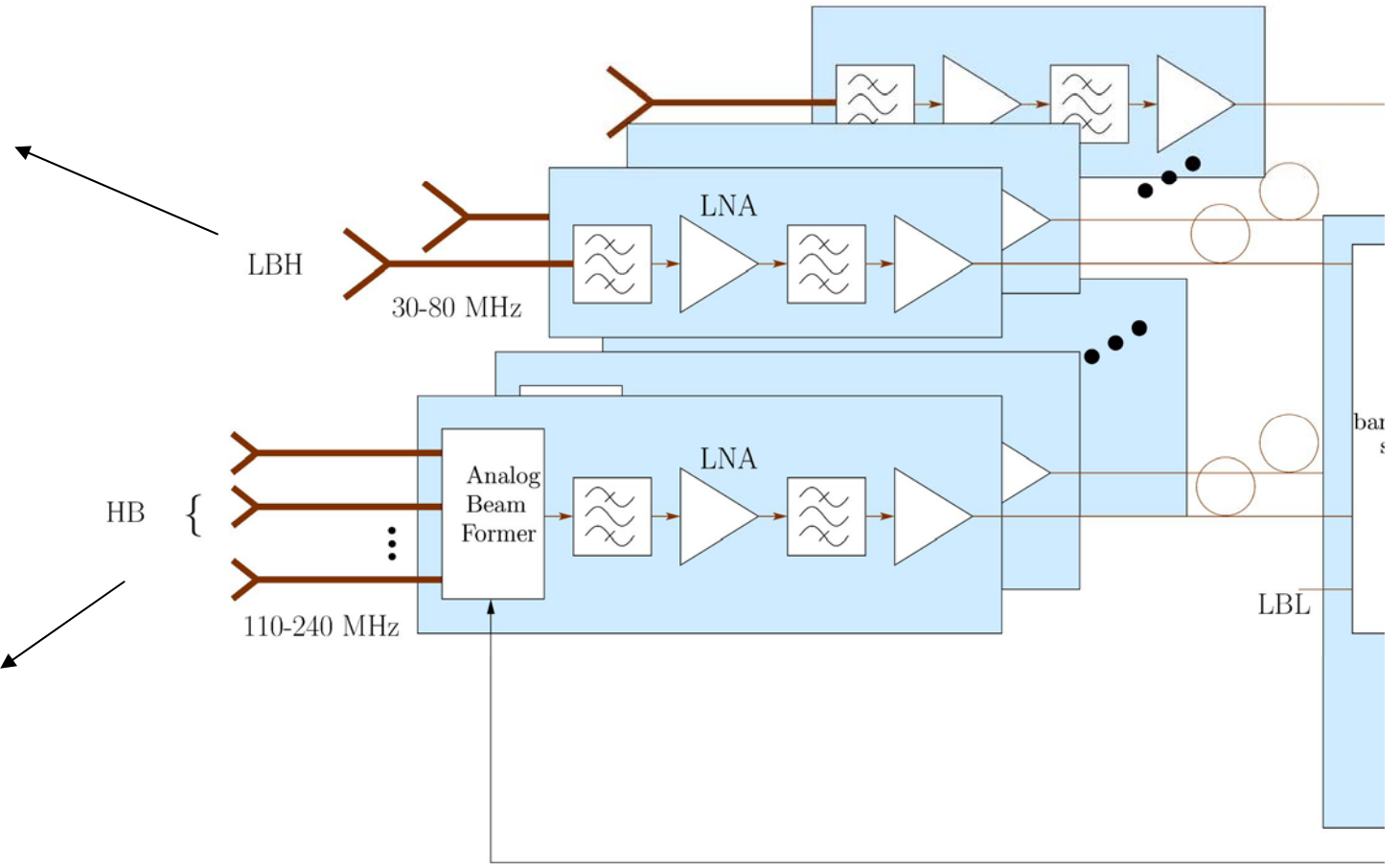
Remote Station Architecture

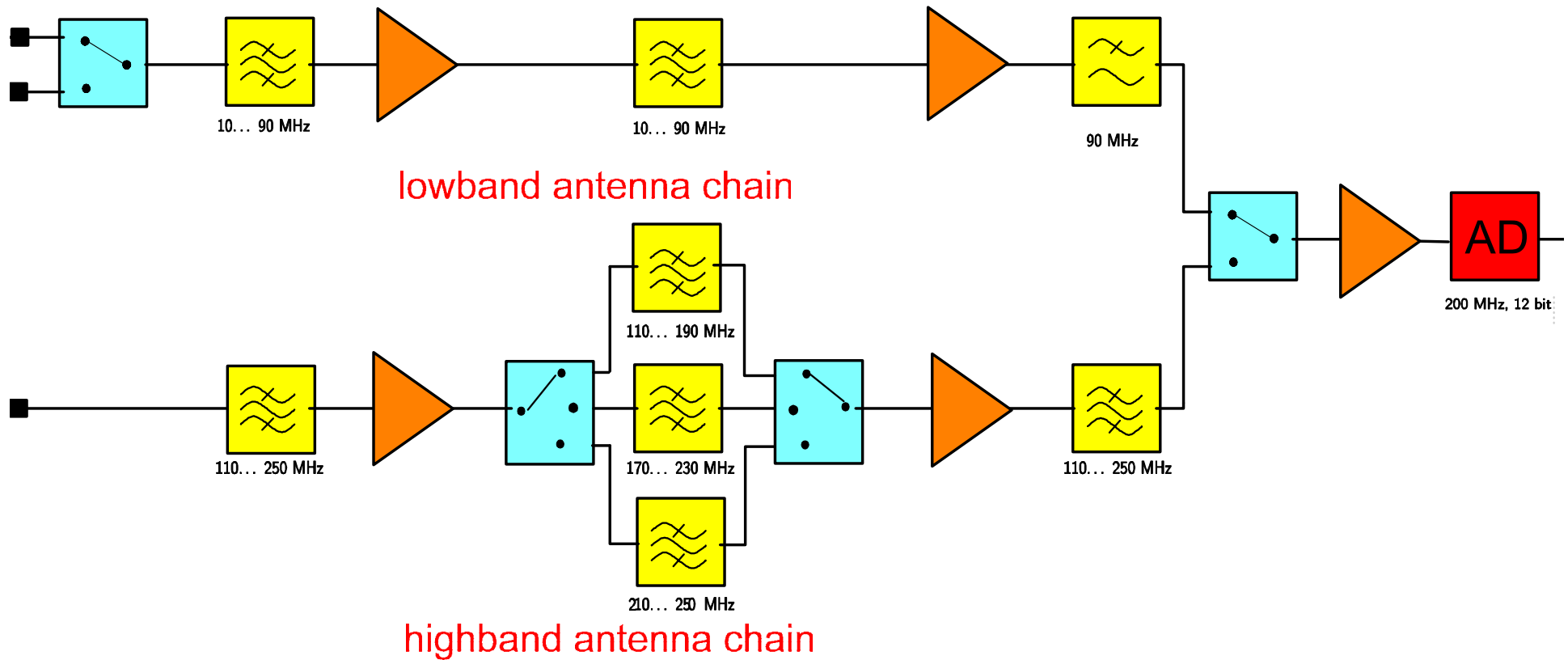


Remote Station Architecture (detailed)

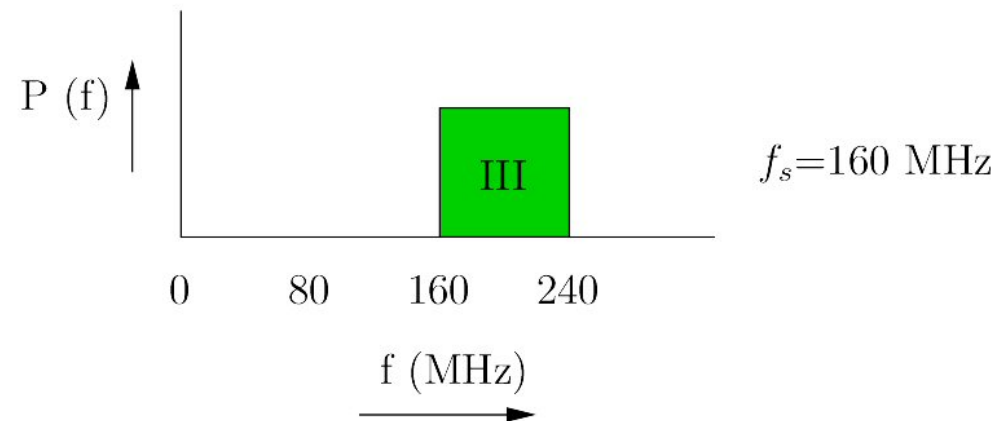
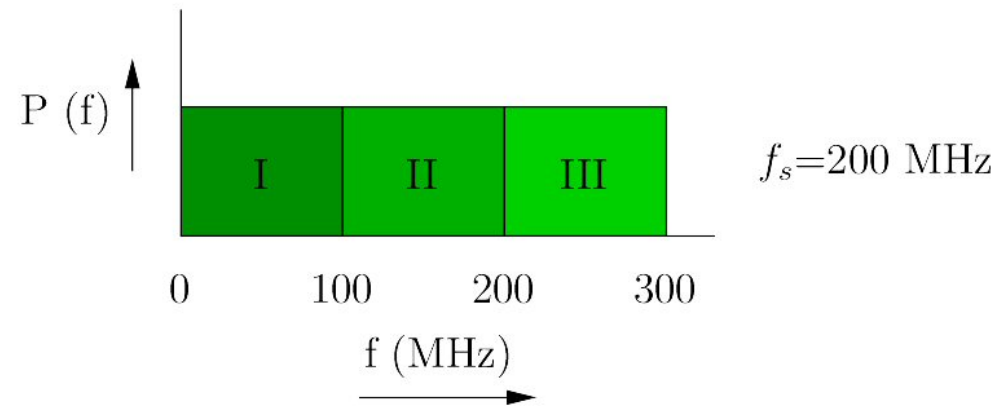


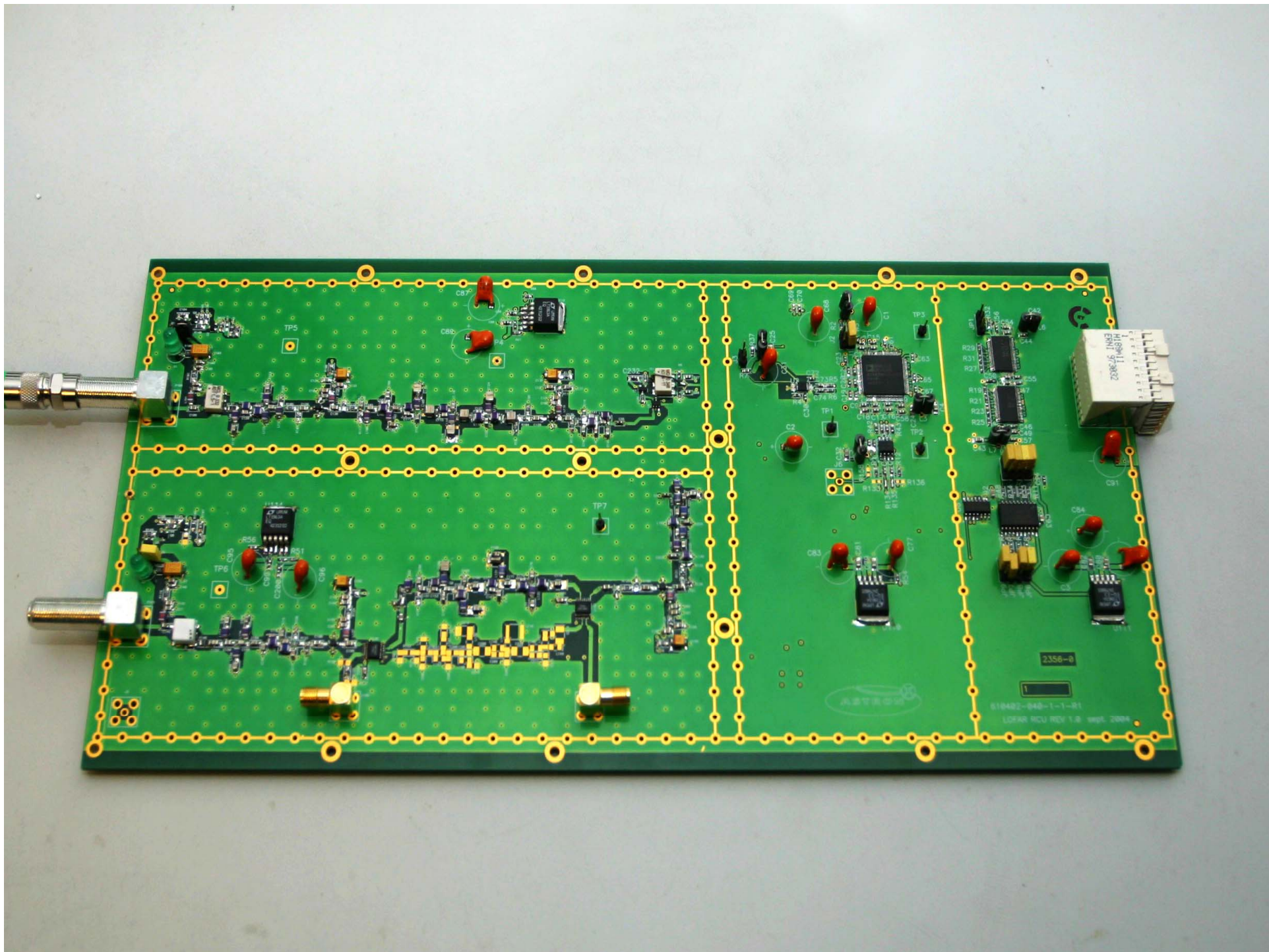
Antennas

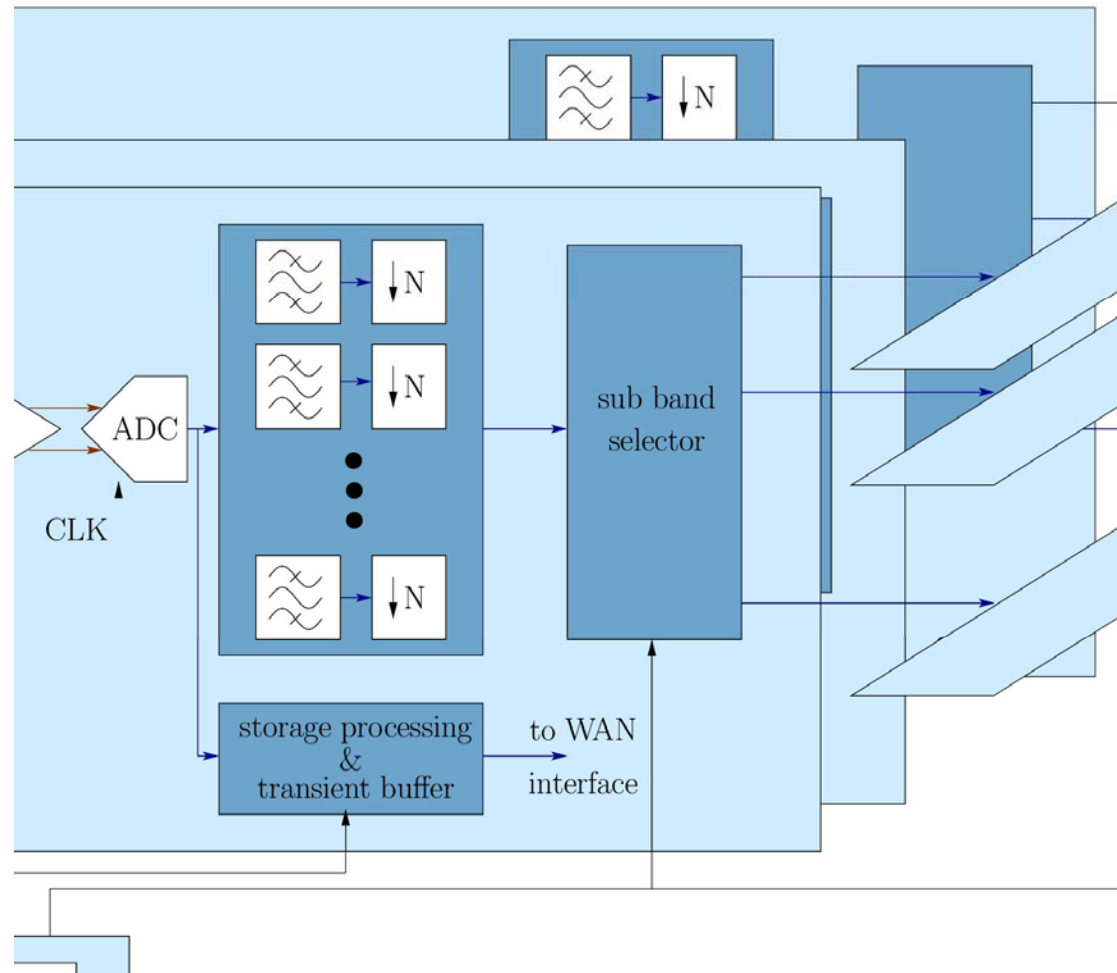


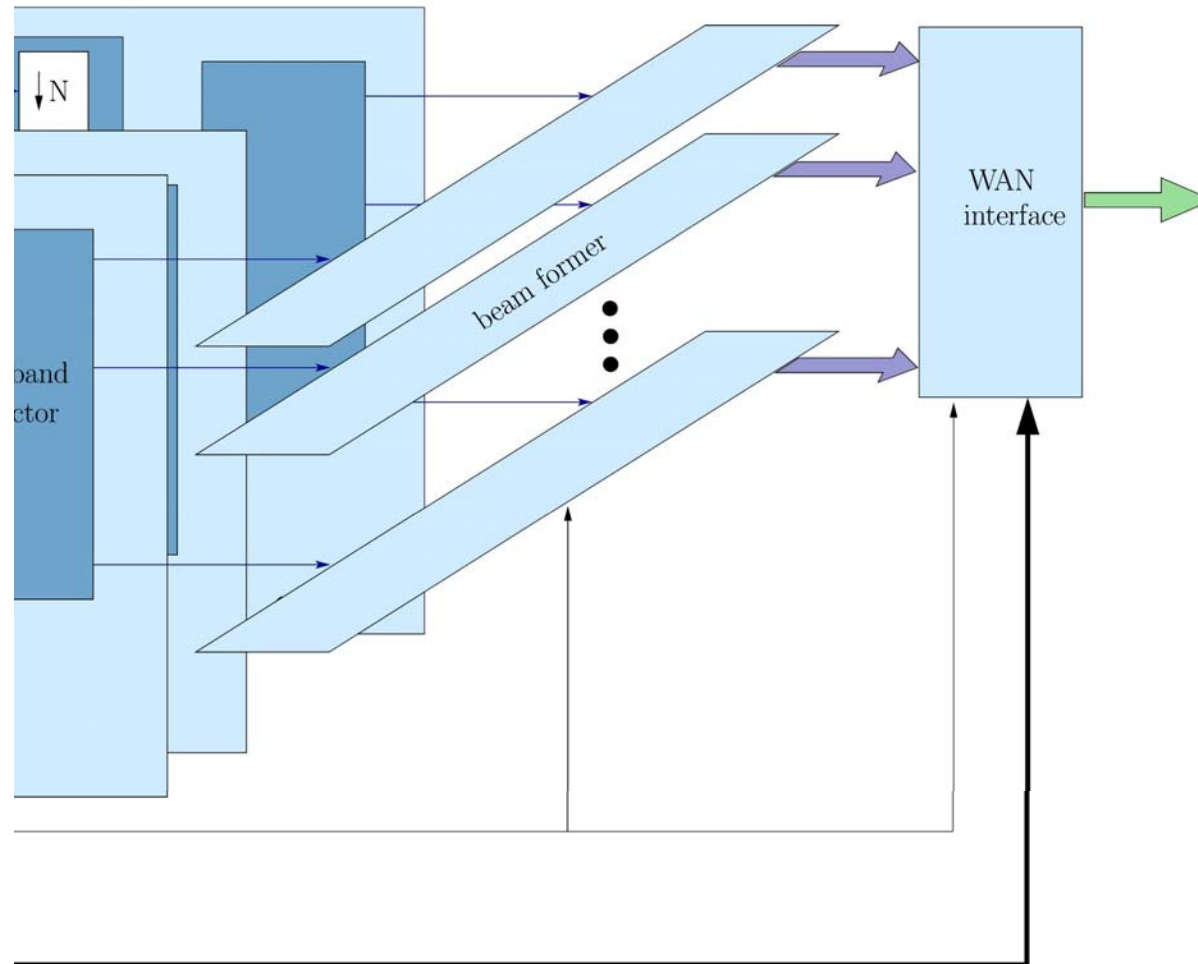


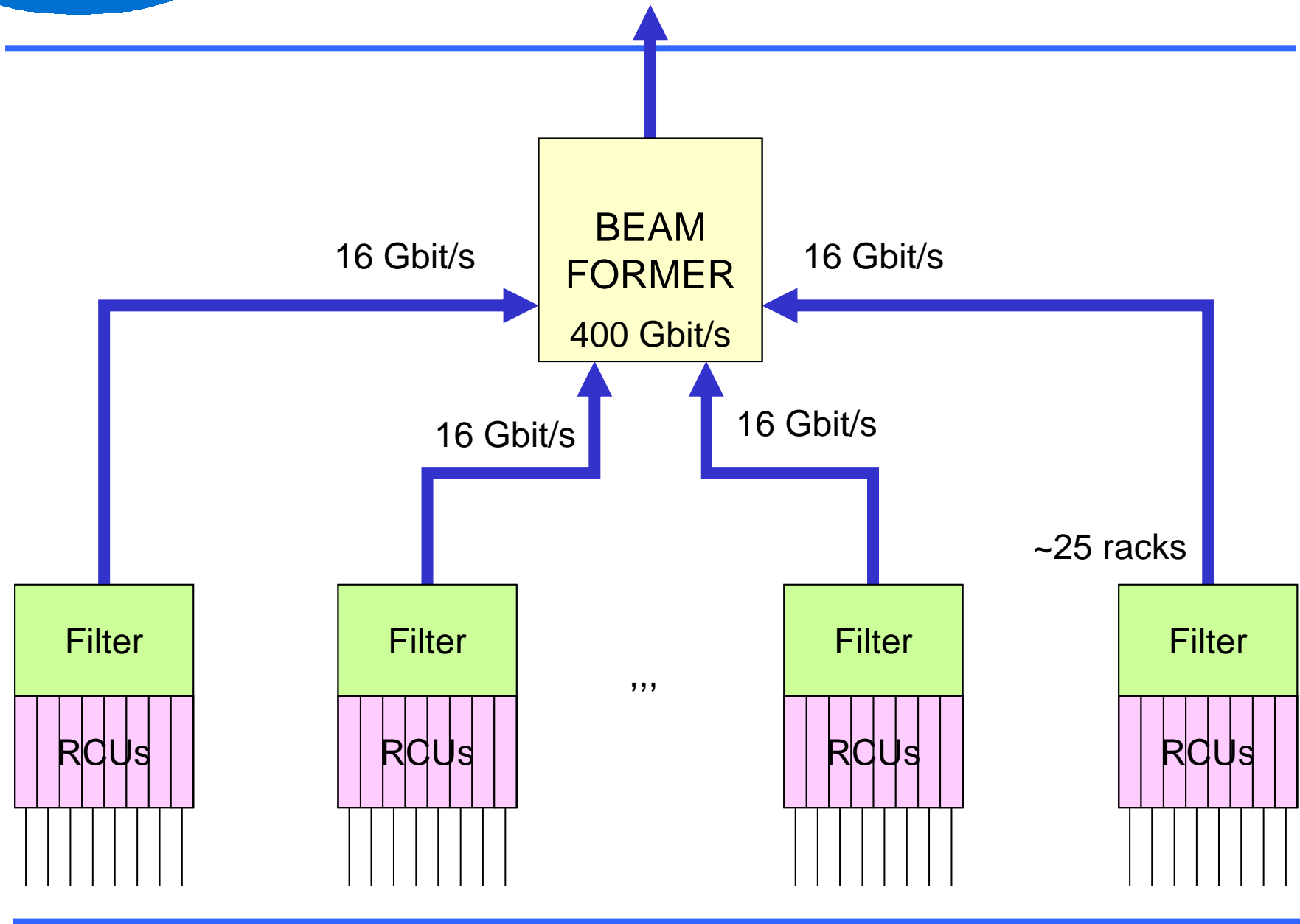
Direct Conversion Architecture

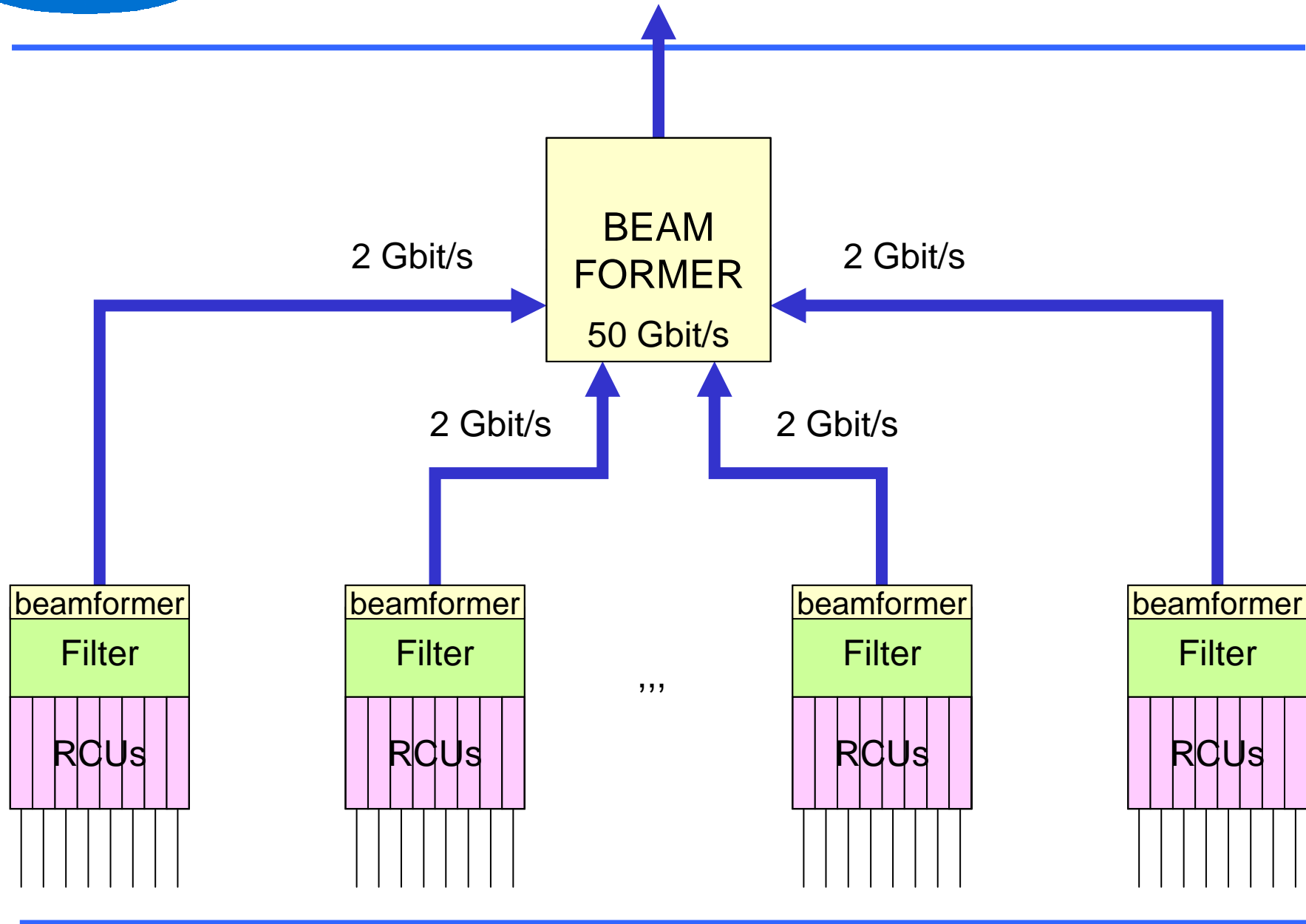


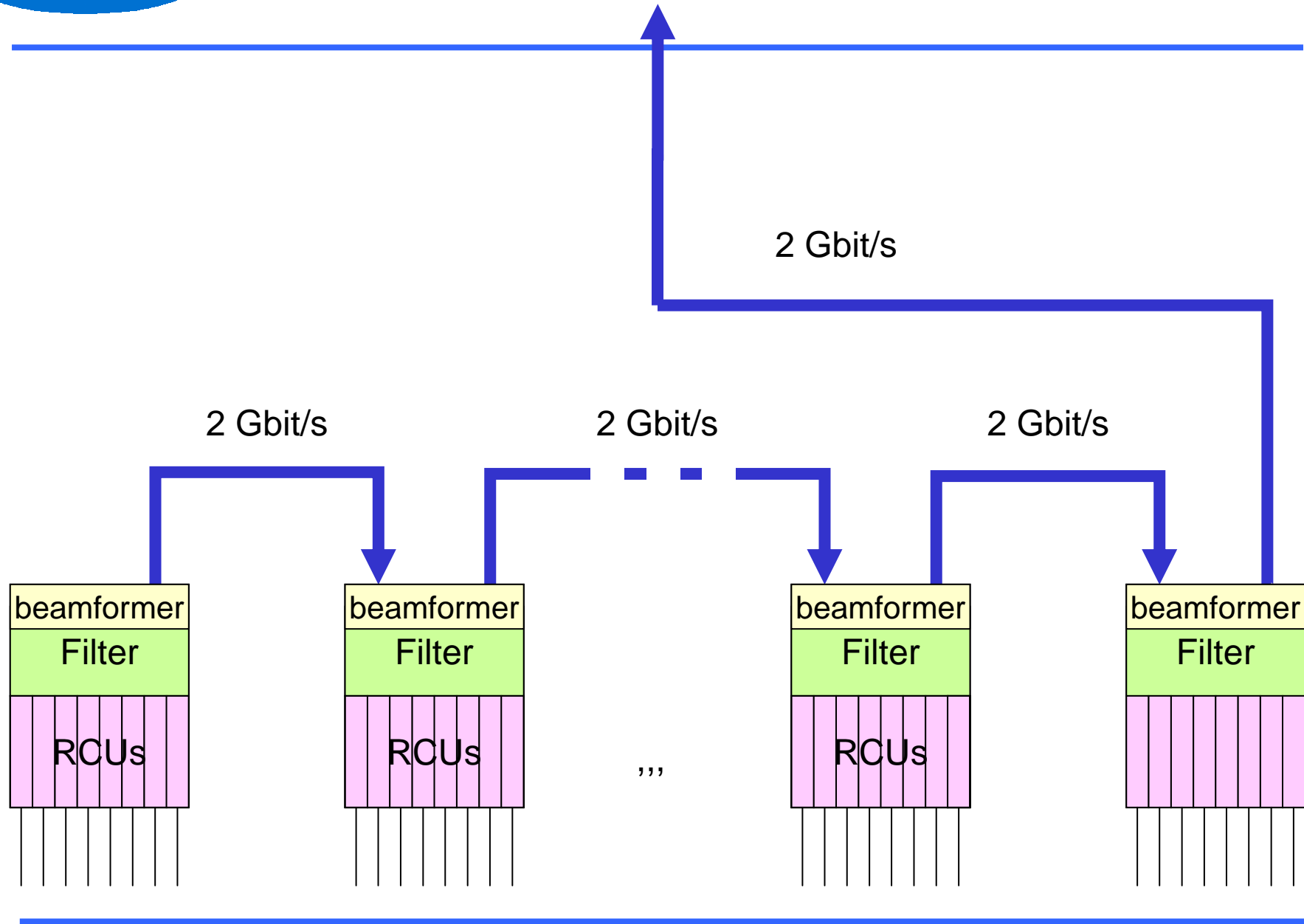


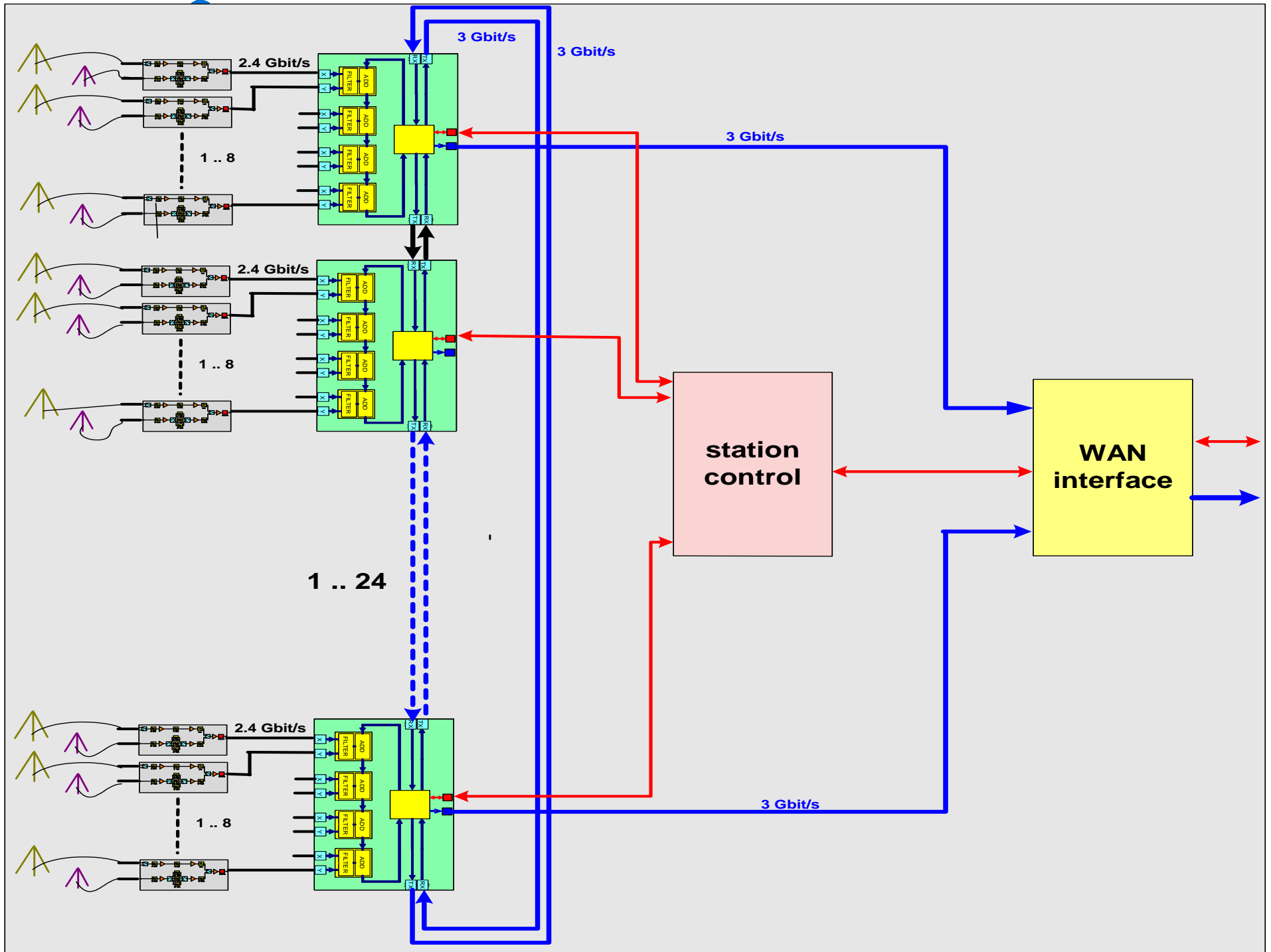




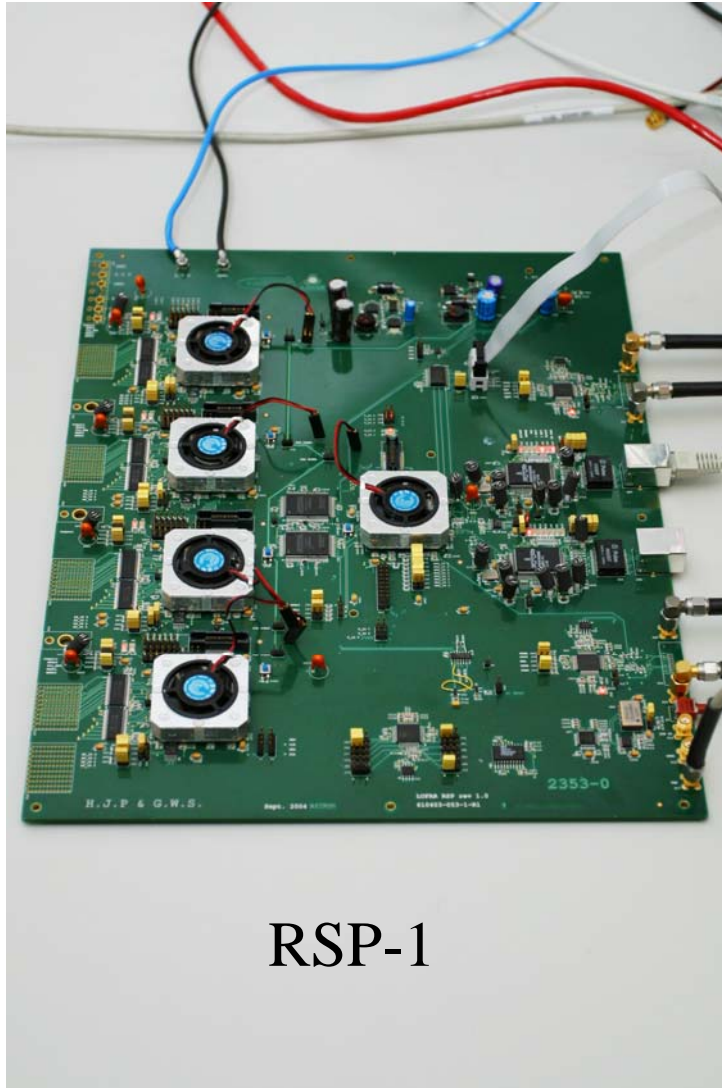








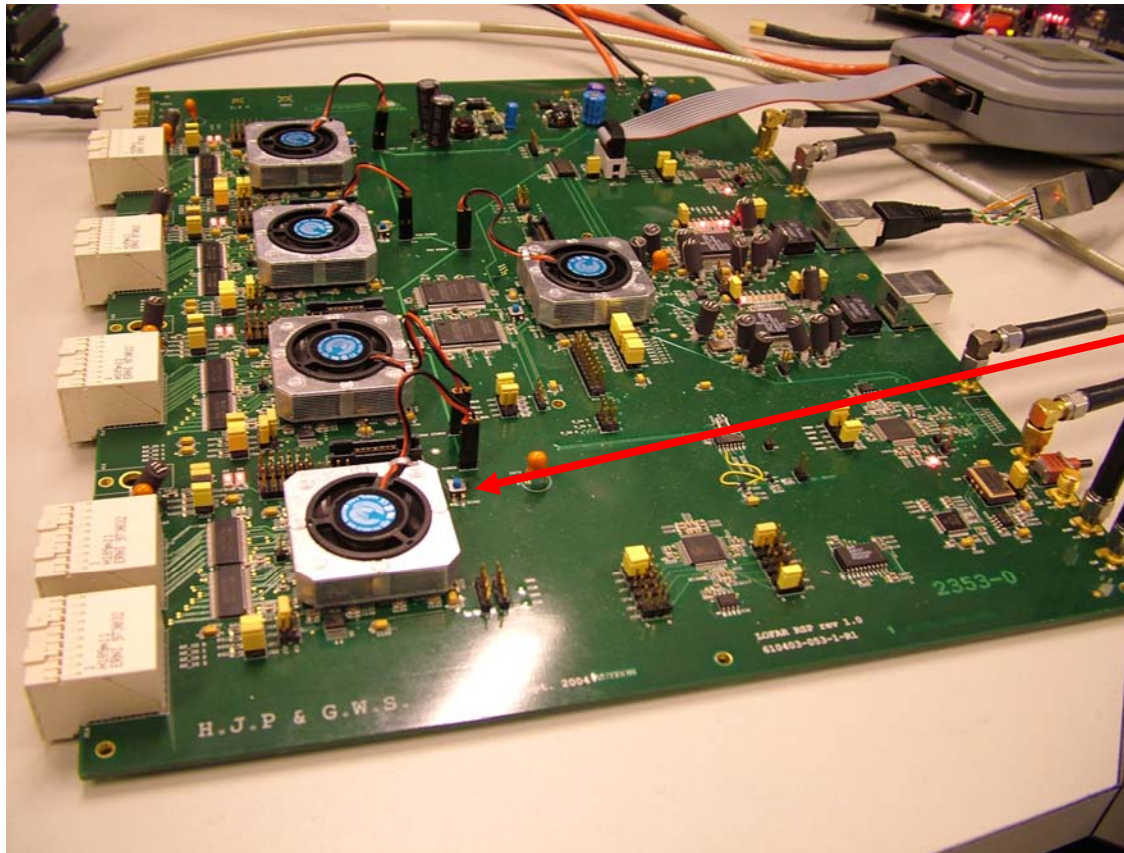
- Splitting the 100 MHz band in 512 subbands
- Selecting subbands
- Digital beamforming with an update rate of 1 s
- Correlation at station level
- Storing data on a trigger



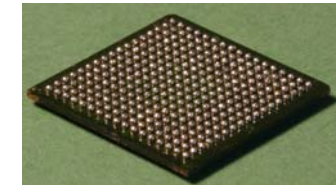
RSP-1

- Board size: 280x366.6 mm
- # layers: 12
- # connections: 6369
- # differential pairs: 439
- input bandwidth: 19.2 Gbps
- max. output bandwidth: 10 Gbps

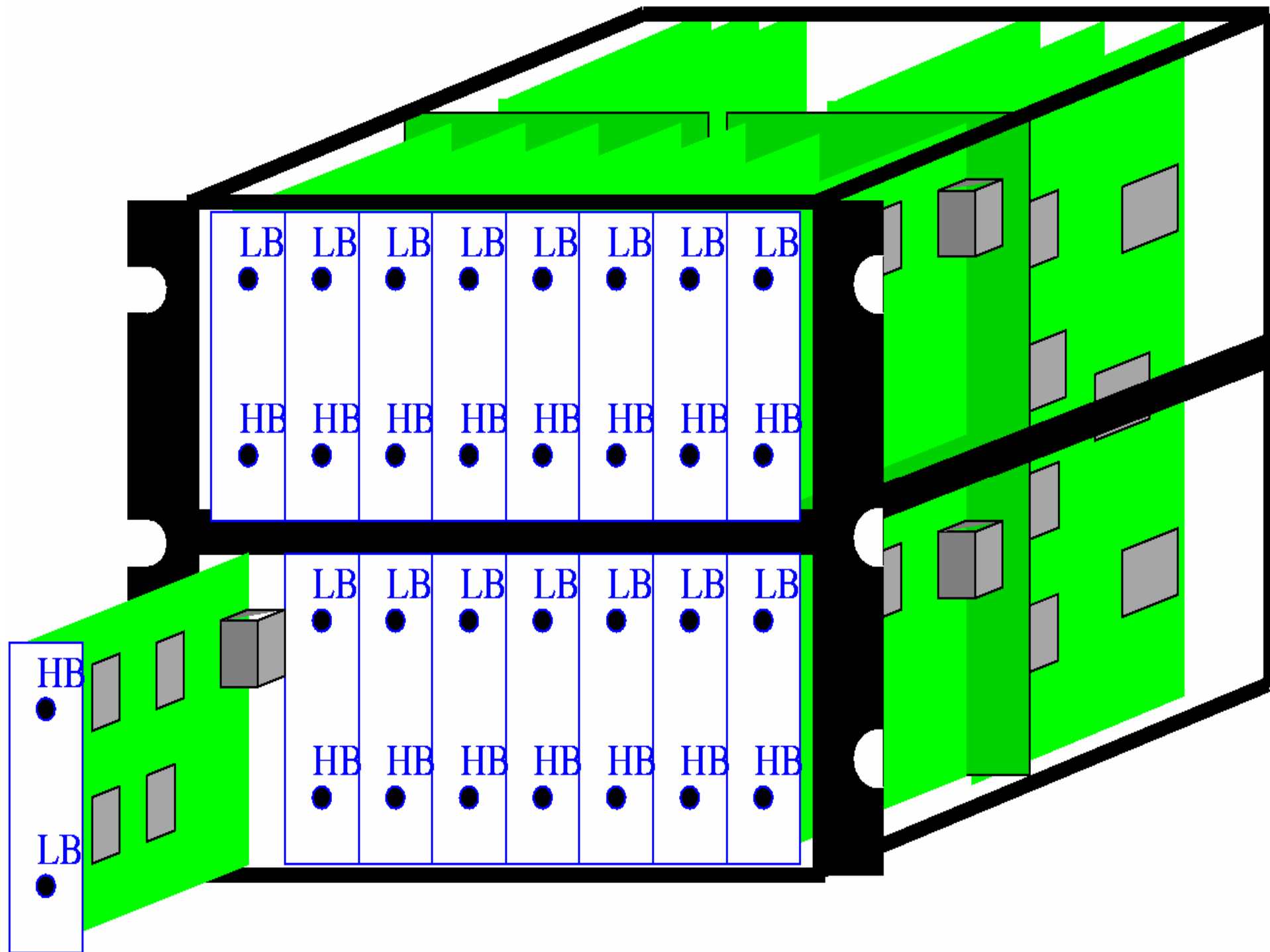
RSP-2 Chip Characteristics



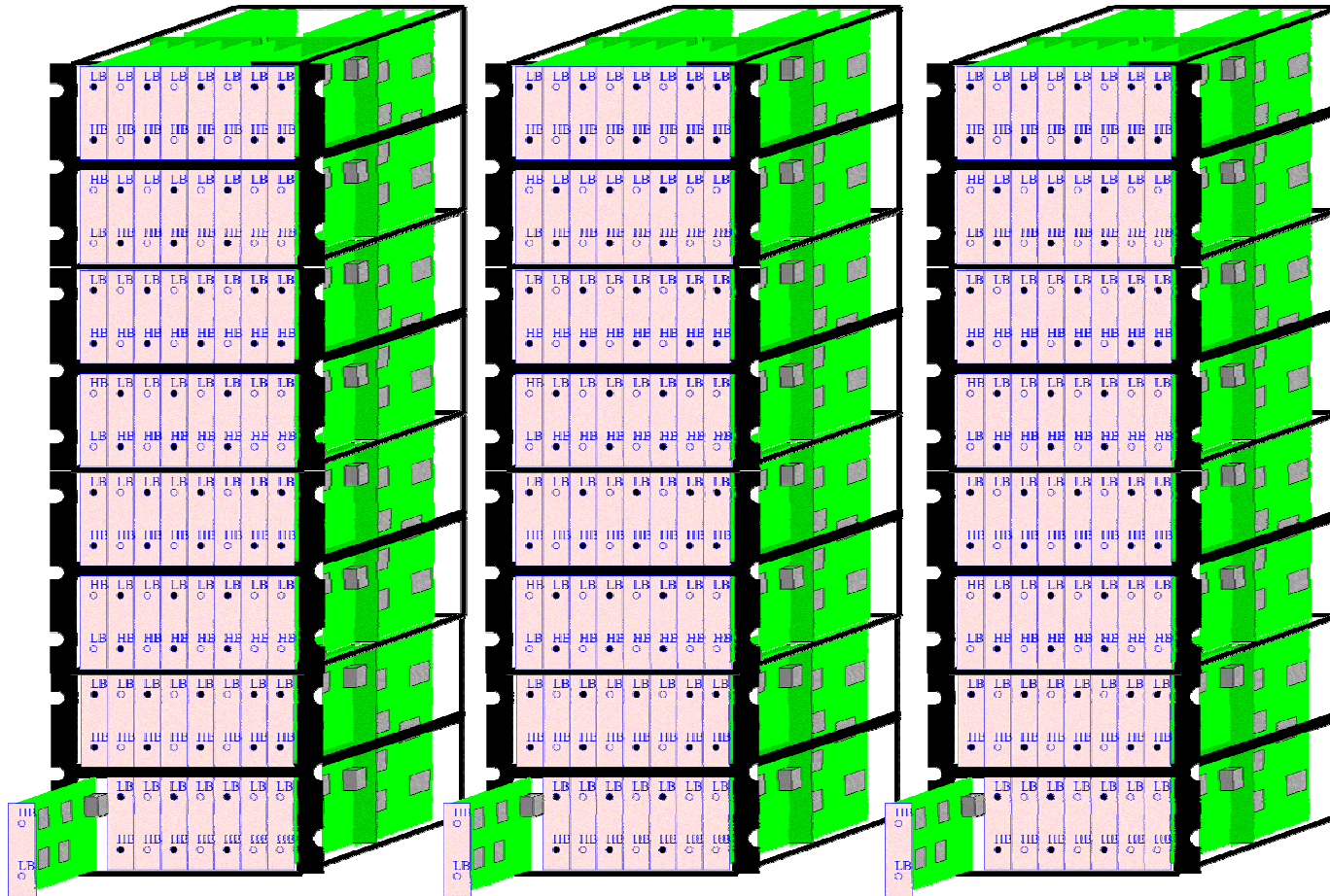
- 90 nm technology
- 192, 18x18 multipliers
- 4.5 Mbit RAM
- 1020 “balls” on chip



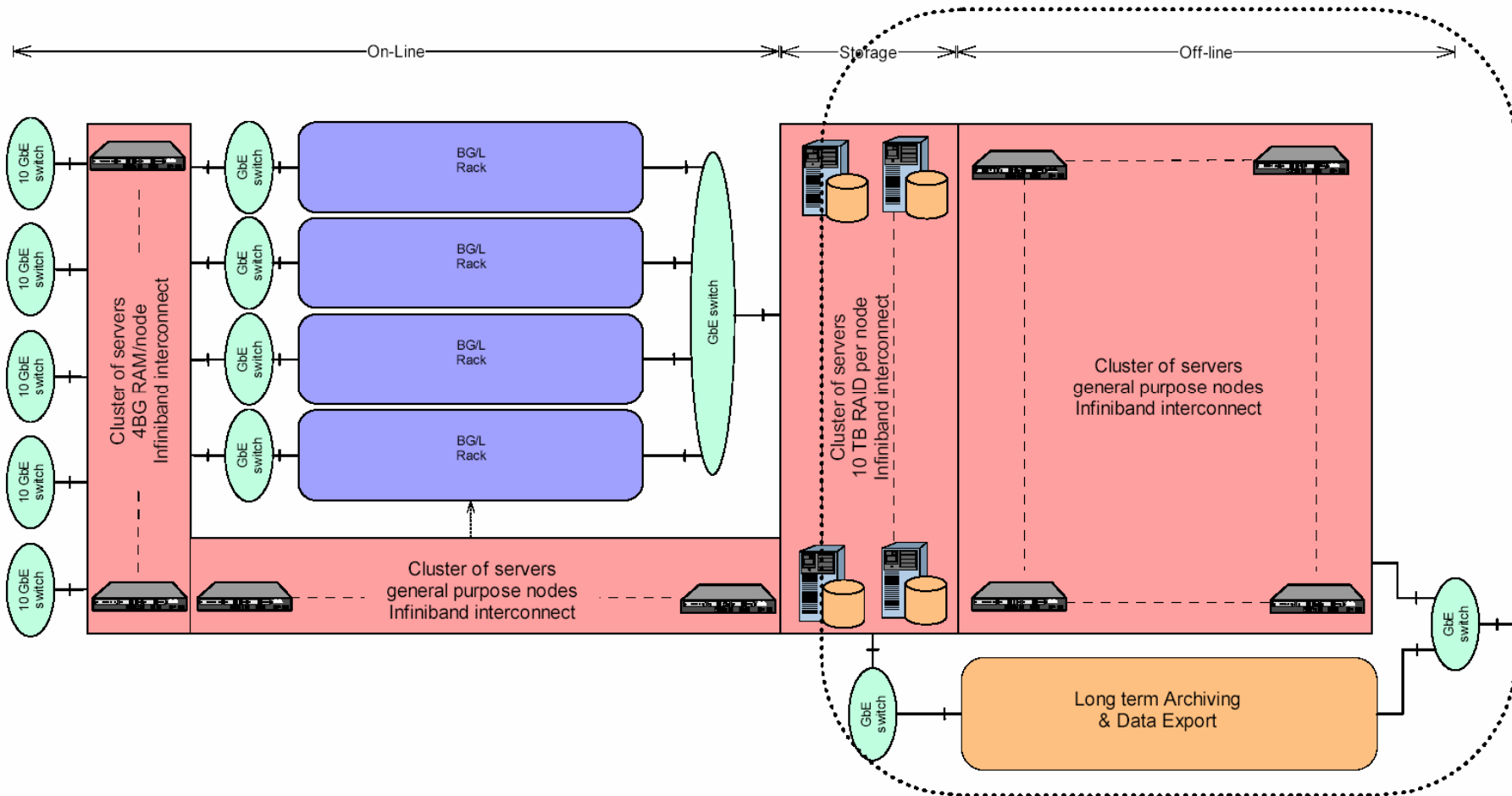
- Input data rate: ~ 460 Gbps
- Output data rate: ~ 2 Gbps
- Processing capacity: ~ 1.5 Tmul/s
- Storage capacity: 96 Gbyte

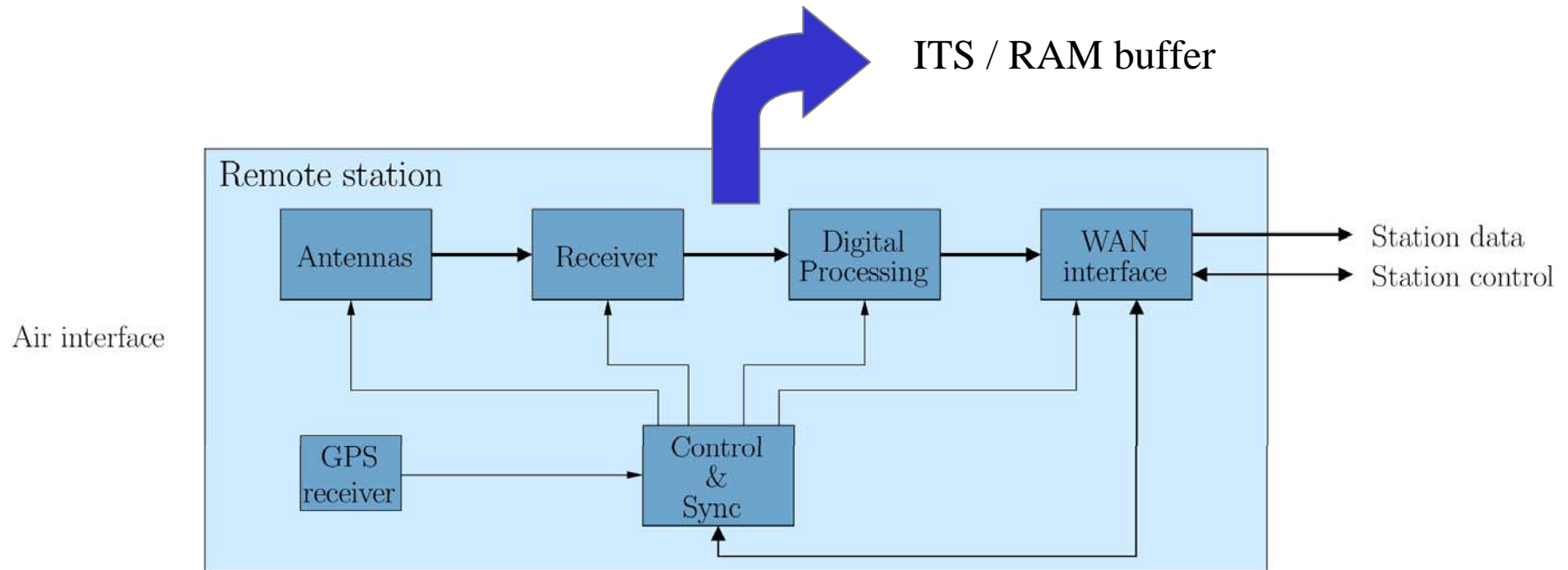


LOFAR Remote Station FTS-1 Hardware



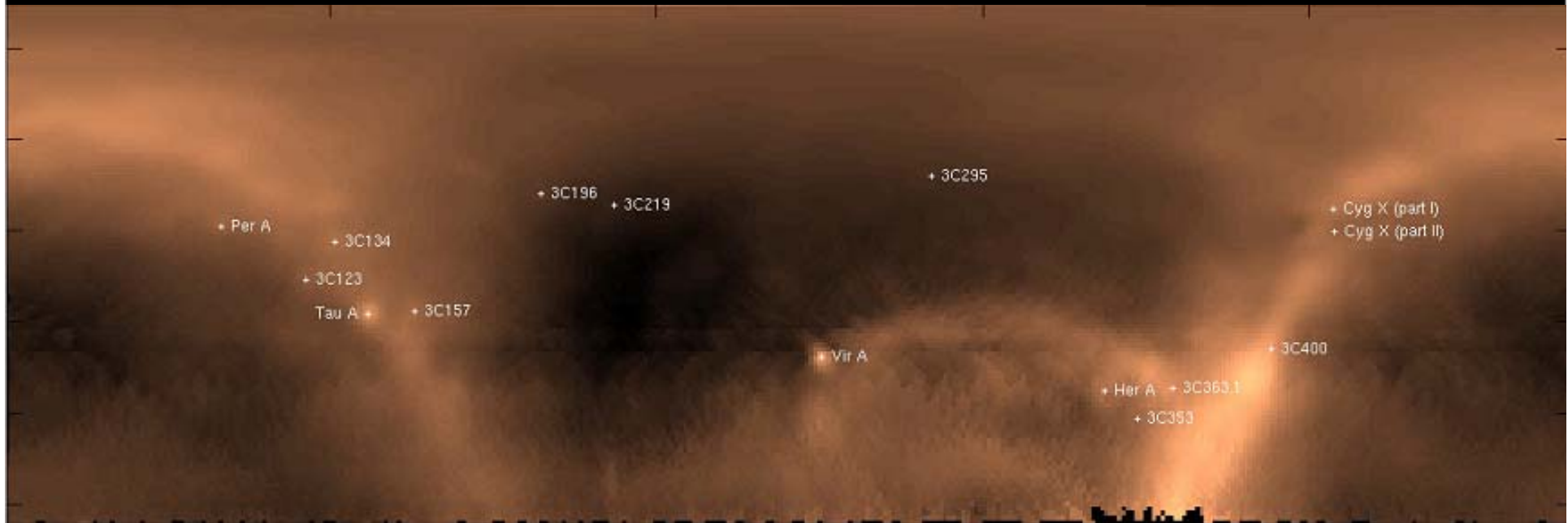
LOFAR Central Processor





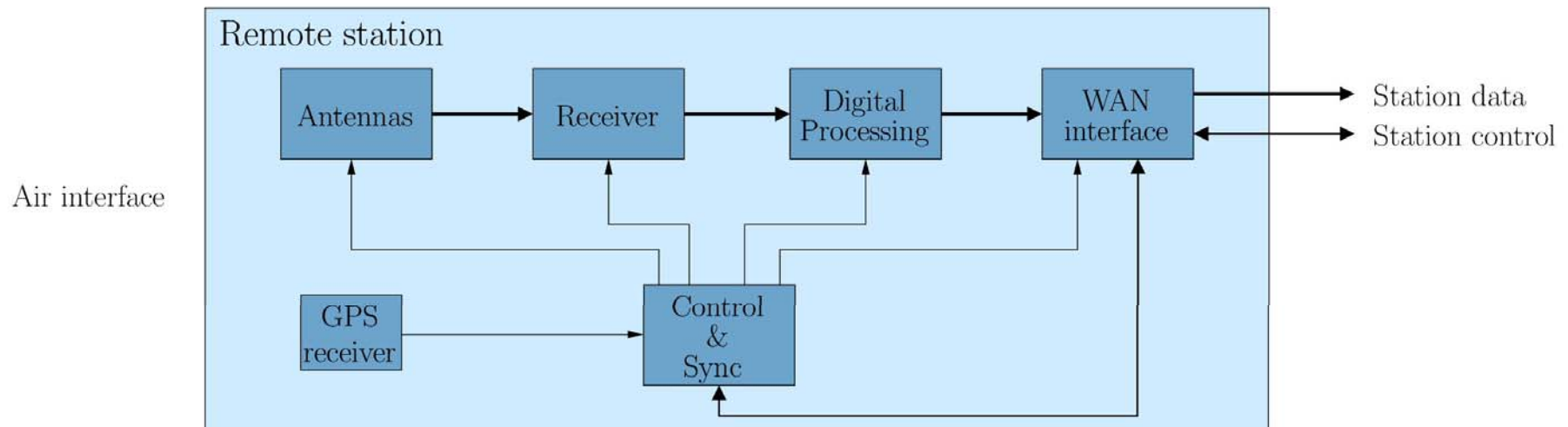


Survey of the full sky made with a LOFAR Test Station



source	flux @178MHz (Jy)	source	flux @178MHz (Jy)
Perseus A	58	Hercules A	325
3C123	175	3C363.1	90
3C134	66	Cygnus X (part I)	230
Taurus A	1420	Cygnus X (part II)	410
3C157	210	3C353	203
3C196	59	3C295	73
3C219	44	3C400	400
Virgo A	970		

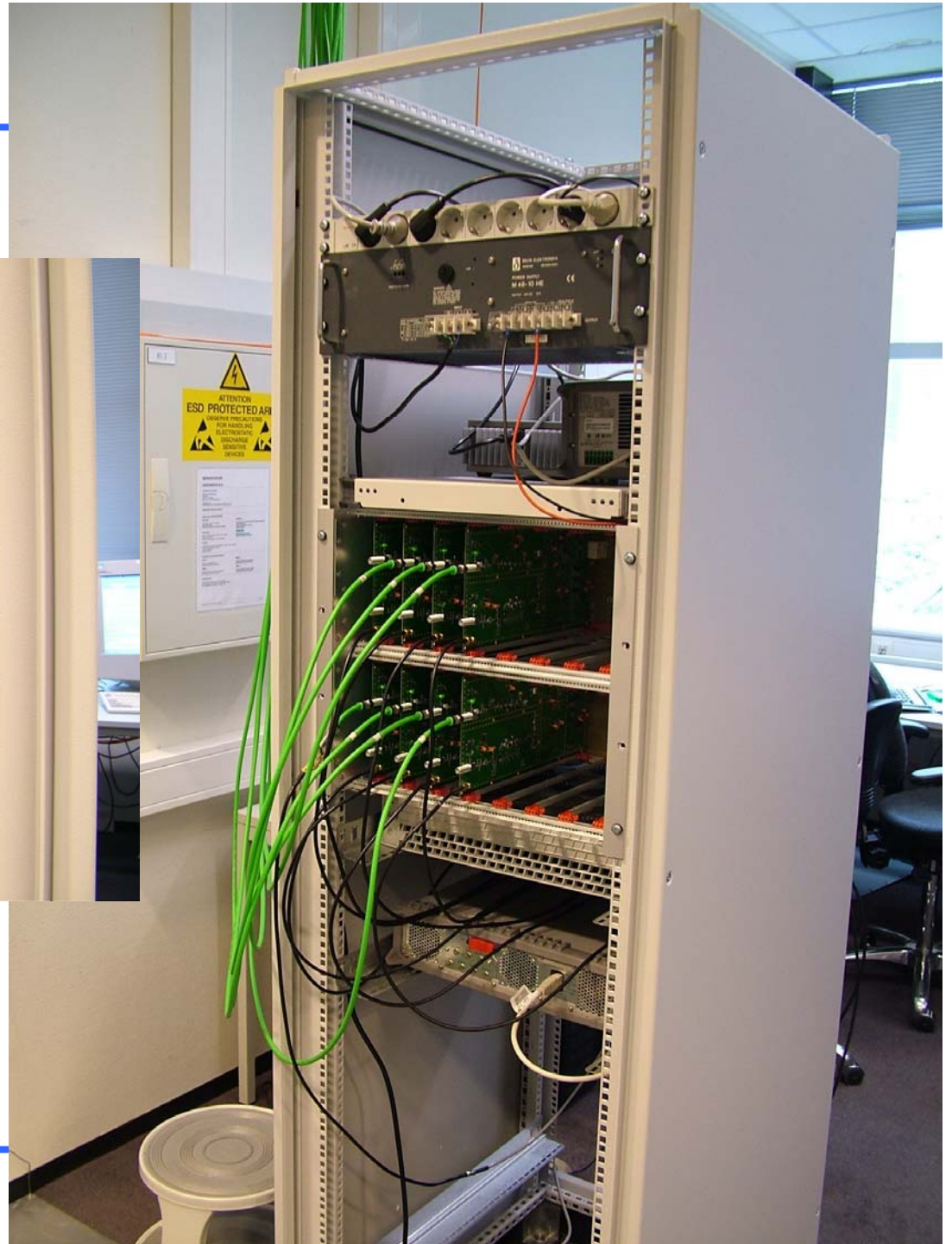
This map was produced using 86 snapshots with 6.7s of integration and 9.7kHz channels. First RFI free channels between 29.5 and 30.5 MHz were selected using a median filter. The selected channels were calibrated on the four strongest 3C sources (Cas A, Cyg A, Tau A, Vir A). After the calibration a first map was made by a flux conserving projection from the (l, m) grid of the individual snapshots to the (α, δ) grid of the all sky image. This map was dominated by the averaged out sidelobes of Cas A and Cyg A. Therefore these two sources were cleaned from the image resulting in this map.





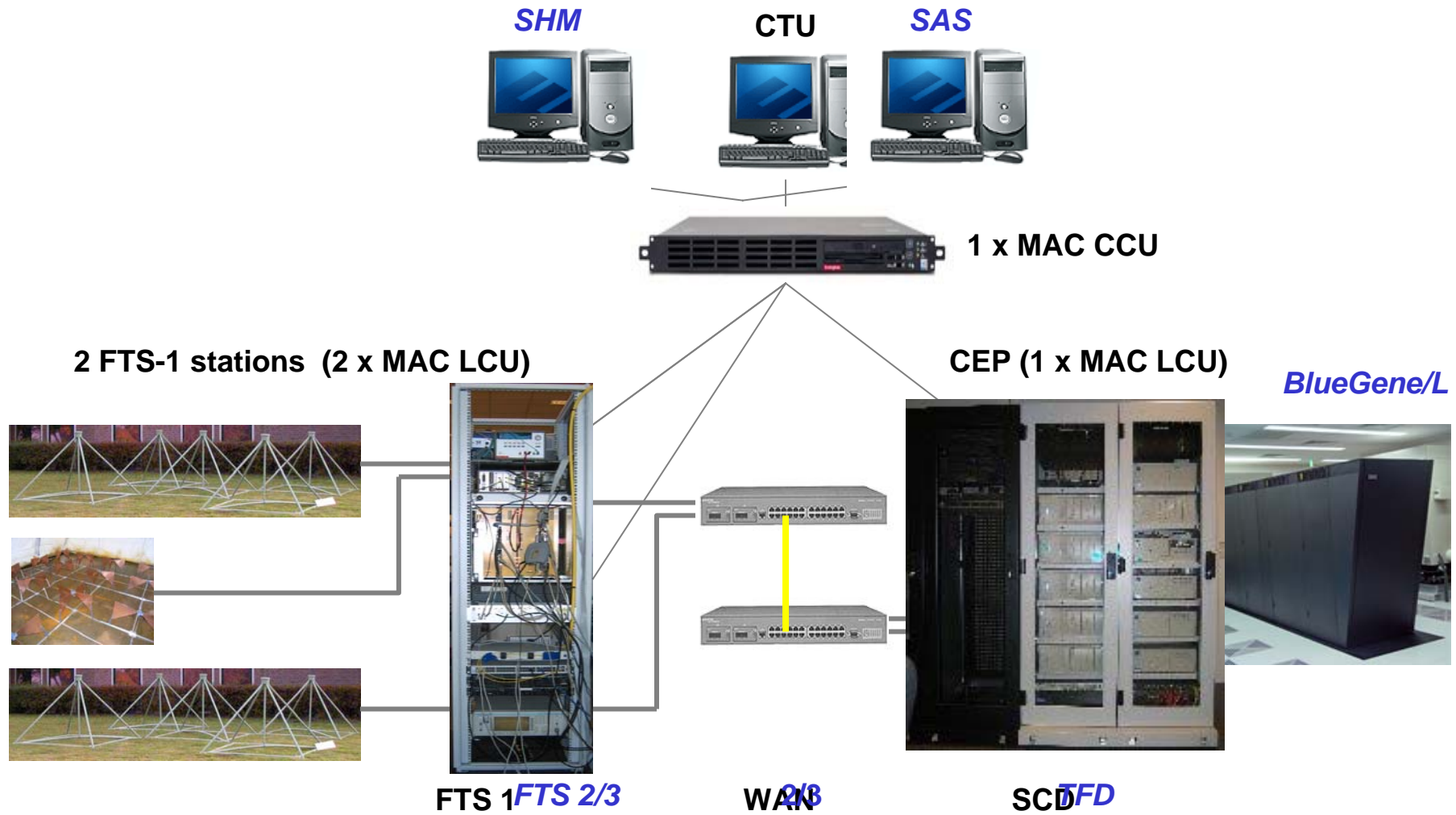








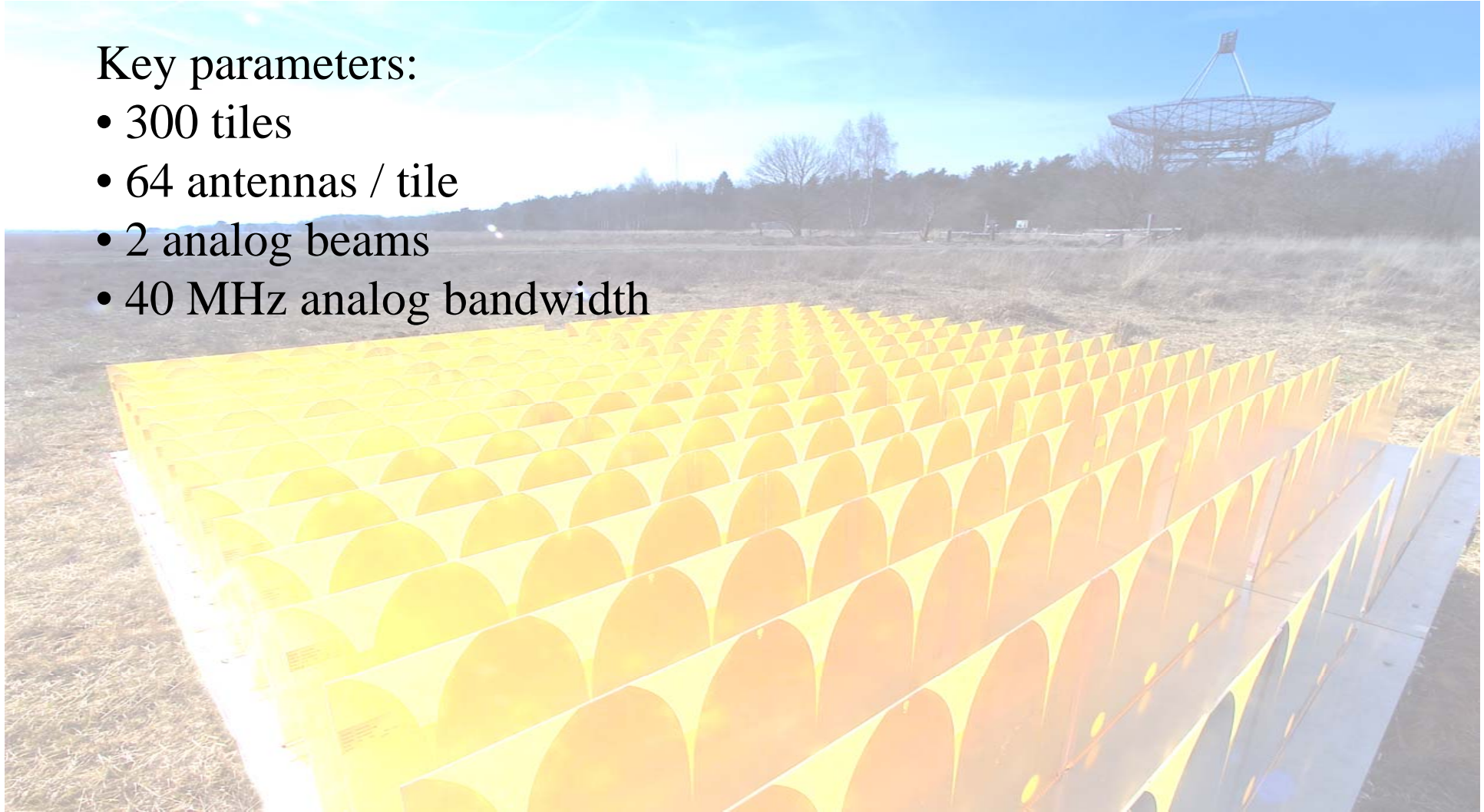
LOFAR Integration

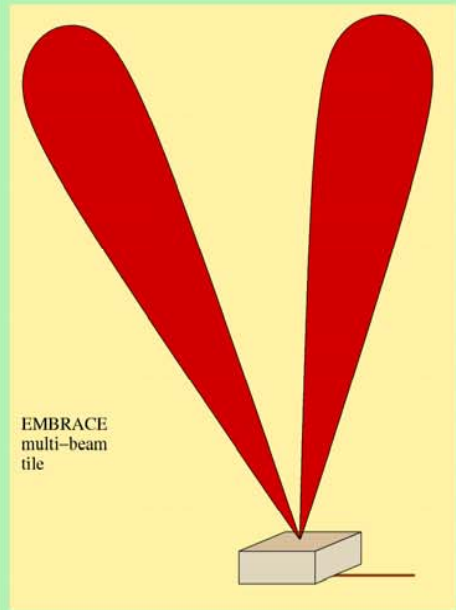


Embrace in a Nutshell

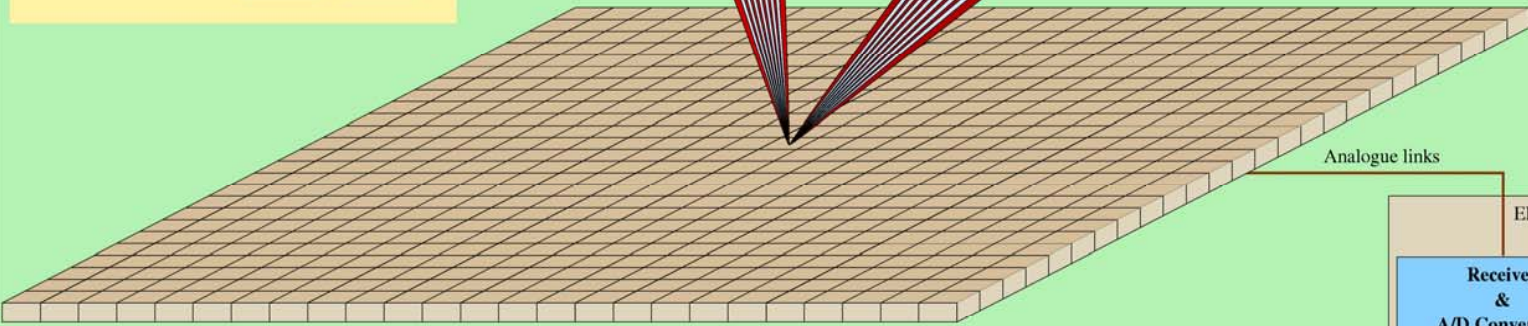
Key parameters:

- 300 tiles
- 64 antennas / tile
- 2 analog beams
- 40 MHz analog bandwidth





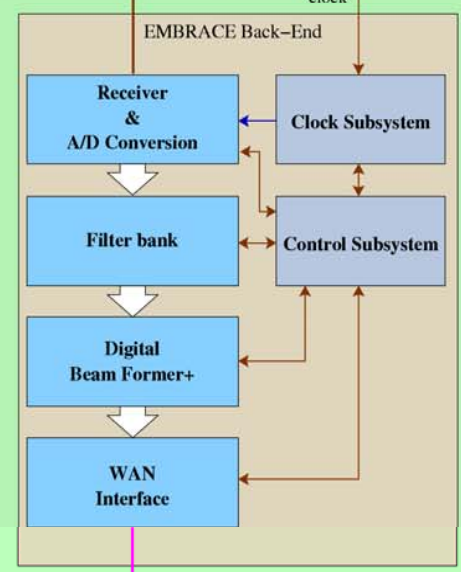
EMBRACE > 600 m²
Aperture Array Station



GWK/2004

Analogue links

GPRS disciplined
clock



towards WSRT correlator



Fibre Optic Data Link

- Hardware of a LOFAR station is re-used
 - stripped version
 - different receivers
 - 3 stations necessary

- WSRT correlator is used

- Digital circuits are stable (if you design it properly ...)
 - ☞ temperature (within a given range)
 - ☞ components
 - ☞ time
- Digital signals can be copied without loss
- Digital signals can be controlled in a flexible way
- Doing it digital becomes cheaper

- Massive amount of data
 - ☞ number of signals (sensors) increase
 - ☞ number of bit to represent a signal increase
 - ☞ bandwidth of signals increase (ADC moves towards the antenna)
- Relative “simple” operations required (implementation is not)
 - ☞ filtering
 - ☞ beamforming
 - ☞ correlation
 - ☞ routing signals
 - ☞ ...
- More flexibility required
 - ☞ exchange beams for bandwidth
 - ☞ select bands, signals, ...

But, price per operation decreases ...

Digital Processing in Focal Plane Arrays

