HI Absorption 2018 Discussion Synergies and Follow-up Chairs: Nina Hatch & Elizabeth Mahony Date: 30 August 2018 - What do we want/need to get out of HI absorption (after detection) - Images - Metallicity - It is not just about redshift, but also about line ratios - optical imaging of the galaxy, is the galaxy distburbed ... - high-res radio (continuum) data - DLAs are more relevant for the intervening systems - here you want opt+NIR imaging - with adaptive optics also for NIR => JWST, but MUSE also possible - separate targets from HI detections - non-detections are hard and it will be difficult to get follow-up time - but it is probably still interesting, e.g., to see why it is interesting - CO, OH - OH is about a factor 10 weaker than HI - relationship between CO and HI is seems to be complicated - CO and HI may not trace the same region - bias against CO - What do we want? - Stacking - will stacking work in the light of large velocity offsets - answer will depend on associated versus intervening systems - parametrisation of stacking need to be explored - anything else than spectroscopic redshifts stacking not possible with photometric redshifts - is it too optimistic to get spectroscopic redshifts for all radio sources - FLASH predicts about 150000 sight lines - photometric redshift needs multi-band photometry - it may depend on the science question - a two-step approach may be better - photometric redshift for the majority of targets - and spectroscopic redshifts for specific tartets - stacking in absorption is harder - contamination within in beam - need a small-enough beam - Apertif should be okay - if stacking is done on peak of continuum, it should be fine - Filippo tested with ATLAS3D, but with the given sample it was okay - but if sample size would increase by a factor of 10, it can become important - for larger redshifts it can still become a problem - Current optical facilities - situation in the north better because of more smaller facilities WEAVE can go to south - J-PLUS/J-PAS (can go down to equator) - in the south - S-PLUS in Brazil - identical to J-PLUS - first release perhaps this year - WAVE is similar to WEAVE, but probably no additional option - VISTA will put a public call for proposal => proposal as a community, we should start thinking about it now - in the future, this will be solved with EUCLID and LSST starting 2024 - what to do in the years till then - photometric redshifts can be provided in the near future - also available for the south, Kenneth Duncan knows about them

- Where do we go from here?

- Additional coordination for follow-up?
- success rate with ALMA for survey follow-up perhaps lower than specific
- target
 - Coordination with simulations/theory
 - from the list of Lilian's project idea what do we want from the simulations

 - Small-scales are still a problem
 - simulations have a relatively large amount of freedom
 - simulations probably not to use as a prediction
 - but still look at statistics of population