LOFAR Workshop 2014 Summary



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ILT Board: M. Garret (ASTRON), H. Rottgering (NL), H. Falcke (NL), M. Brüggen (D), J. Conway (S), P. Best (UK), M. Tagger (F) R. Vermeulen (Dir.)

Keynote sucks ...



- Apple is the new Microsoft ...
- People punished for using latest version of Keynote that wants to force me buying a new version:
 - Heald
 - de Gasperin
 - Rowlinson

The Enemy



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The good guys/gals: de Gasperin, Mevius, (van Weeren)



EOR 3C196 – Million:1 DR



EOR NCP



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eepest (low-frequency) 4.00e-04 3.50e-04 radio image! 3.00e-04 2.50e-04 2.00e-04 1.50e-04 20,000 sources, Billion pixels 150 MHz, 2" pixels, 25 μJy noise, 200 hrs data 9.97e-05 Yatawatta

Clusters

Halo E-W ~1.07 degrees 1.8 Mpc!

Coma cluster

Bridge of emission connecting Halo and Relic

> Relic ~800kpc Bridge connecting relic and NGC4789

Radio emission over 2 degrees ~3.3 Mpc NGC 4839

NGC4789

Bonafede





Large Fields



Spectral aging ... Radio Archeology





Magnetism





Long baselines LOFAR LB Snapshot survey





Long Baselines M82





Spectroscopy(!): 1st Extragalactic Carbon Radio Recombination Line in M82





Transients – something is varying ...





"Trapping" Ghosts ...



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Stewart/Rowlinson

Fast Radio Transients



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DM vs RA and DEC vs RA for pulsar(red) and coincidence triggers(blue)

We are in principle ready to look for Fast Radio Bursts regularly ...



Two more pulsars found



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Cooper

100.000 pulsar candidates scanned ...

Millisecond Pulsars well detected (39/55)



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Can scattering be overcome (Archibald, Stinebring)?

Kondratiev

Mode switching





B0943+10: Mode changing but what is changing?



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Polarization properties of pulsars and birefringence (Noutsos)

Solar Imaging Bonanza



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Morosan, Breitling, Mann

Lunar Imaging Bonanza



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Vedantham

Ionosphere: Absorption & Scintillation





Radio Emission from Cosmic Rays





SETI@LOFAR



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Siemion



EOR Limits



LOFAR keeps growing





Science Output



- Refereed commissioning papers
 - 15 since 2011, 201 cites
 - Top Cited:
 - Stappers 2011 (pulsars) 67 cites
 - Van Haarlem 2013 (LOFAR) 53 cites
 - Van Weeren 2012 (Abell 2256) 18 cites
 - High Impact: 1x Science (Hermsen) + 1x Nature??
- Papers mentioning LOFAR (title/abtract):
 - 747 papers with 4869 cites
 - Expectation management needed ...



Top 8 "LOFAR" papers

1	2007MNRAS.377.1043M	164.000	05/2007	<u>A</u> Z	<u>F</u> <u>G</u>	X	<u>R</u> <u>C</u>		<u>U</u>	
	McQuinn, Matthew; Lidz, Adam; Zahn, Oliver; Dutta, Suvendra; Hernquist, Lars; Zaldarriaga, Matias	The morphology of HII regions during reionization								
2	□ <u>2004ApJ6157M</u>	124.000	11/2004	AZI	EE	x	<u>R</u> <u>C</u>		U	
	Morales, Miguel F.; Hewitt, Jacqueline	Toward Epoch of Reionization Measurements with Wide-Field Radio Observations								
3	2003MNRAS.3468710	113.000	12/2003	ΑZ	<u>F</u> <u>G</u>	x	<u>R</u> <u>C</u>		<u>U</u> H	
	Oh, S. Peng; Mack, Katherine J.	Foregrounds for 21-cm observations of neutral gas at high redshift								
4	2002ApJ572L.123I	112.000	06/2002	AZI	EF	x	<u>R</u> <u>C</u>		U	
	Iliev, Ilian T.; Shapiro, Paul R.; Ferrara, Andrea; Martel, Hugo	On the Direct Detectability of the Cosmic Dark Ages: 21 Centimeter Emission from Minihalos								
5	□ 2006ApJ63820B	109.000	02/2006	AZI	EE	x	<u>R</u> <u>C</u>		U	
	Bowman, Judd D.; Morales, Miguel F.; Hewitt, Jacqueline N.	The Sensitivity of First-Generation Epoch of Reionization Observatories and Their Potential for Differentiating Theoretical Power Spectra								
6	2003APh19477F	97.000	07/2003	AZI	E	x	<u>R</u> <u>C</u>		U	
	Falcke, Heino; Gorham, Peter	Detecting radio emission from cosmic ray air showers and neutrinos with a digital radio telescope								
7	□ 2006MNRAS.369.1577C	96.000	07/2006	ΑZ	<u>F</u> <u>G</u>	X	<u>R</u> <u>C</u>	<u>S</u> <u>N</u>	<u>U</u>	
	Cassano, R.; Brunetti, G.; Setti, G.	Statistics of giant radio haloes from electron reacceleration models								
8	□ 2008MNRAS.385.1211P	94.000	04/2008	ΑZ	<u>F</u> <u>G</u>	x	<u>R</u> <u>C</u>	<u>s</u> o	U	
	Pfrommer, Christoph; Enßlin, Torsten A.;	Simulating cosmic rays in clusters of galaxies - II. A unified scheme for radio haloes and relics with predictions of the y-ray emission								

Progress







Progress (Imaging)

• There is progress!

- It seems slow, but from year to year we make huge progress:
 - MSSS detecting 10 src/sqdeg (~200.000 sources) fully automatic (Heald)
 - LOFAR users are able to get to 1-2 mJy images "routinely" (with quite some effort still)" (Coppejans, Heesen, Mahony, Bonafede)
 - Several experimental methods available that allow to go to science quality thermals noise imaging (de Gasperin)
 - "Extreme peeling" (van Weeren) actually exercised on HBA data
 - LBA data still needs some work but also progress

Breakthroughs



- Historical "technical" breakthroughs
 - Thermal noise-limited science quality imaging
 - Highest resolution low-frequency images (VLBI)
- Unique LOFAR science breakthroughs:
 - Deepest EOR limit
 - 1st extragalactic Carbon RRL
 - Cosmic ray radio emission finally understood, composition measurement with highest precision and duty cycle
- A wealth of impressive pulsar results
 - New pulsars, intermittency, emission process and location

Lots of science



- Many more (potential) science results
 - MSSS lots of great science opportunities (!)
 - First LOFAR transient: What is this? Where is the rest?
 - Beautiful images (workshop banquet!)
 - Magnetic fields turbulence, CR diffusion
 - Resolved low-freq. solar imaging, tracing type III bursts
 - Spectral aging, spectral curvature (needs more work: many sources don't look much different from VLA ... find rare birds!)
- Overall science output still relatively low ...
 - So much data, so little to publish? Pick low-hanging fruits!
 - Better is the enemy of good (and also enemy of excellent)

Final thoughts



- We are not done yet
 - Transient capabilities
 - Implement new calibration schemes
 - long baselines (user software, reliability)
 - Polarization calibration
 - Publish papers, watch out for new things ...
- Fantastic progress by a very dedicated, highly-qualified, young generation of radio astronomers
 - We are in the final half-time of the match that is the tough one ...
 - We need perseverance and more smart ideas
 - But thanks to PhD students, postdocs, science support (!) and staff ...