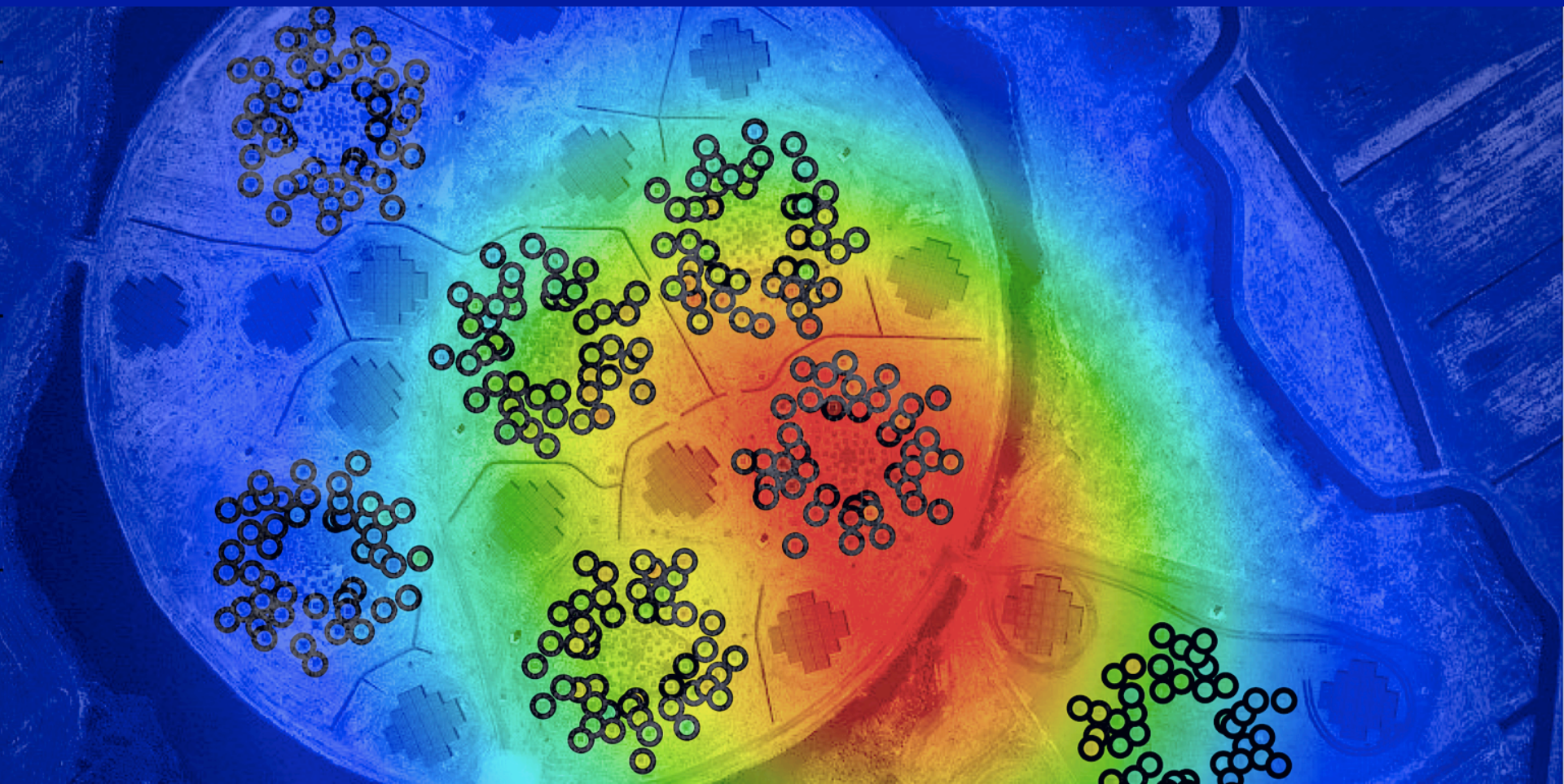


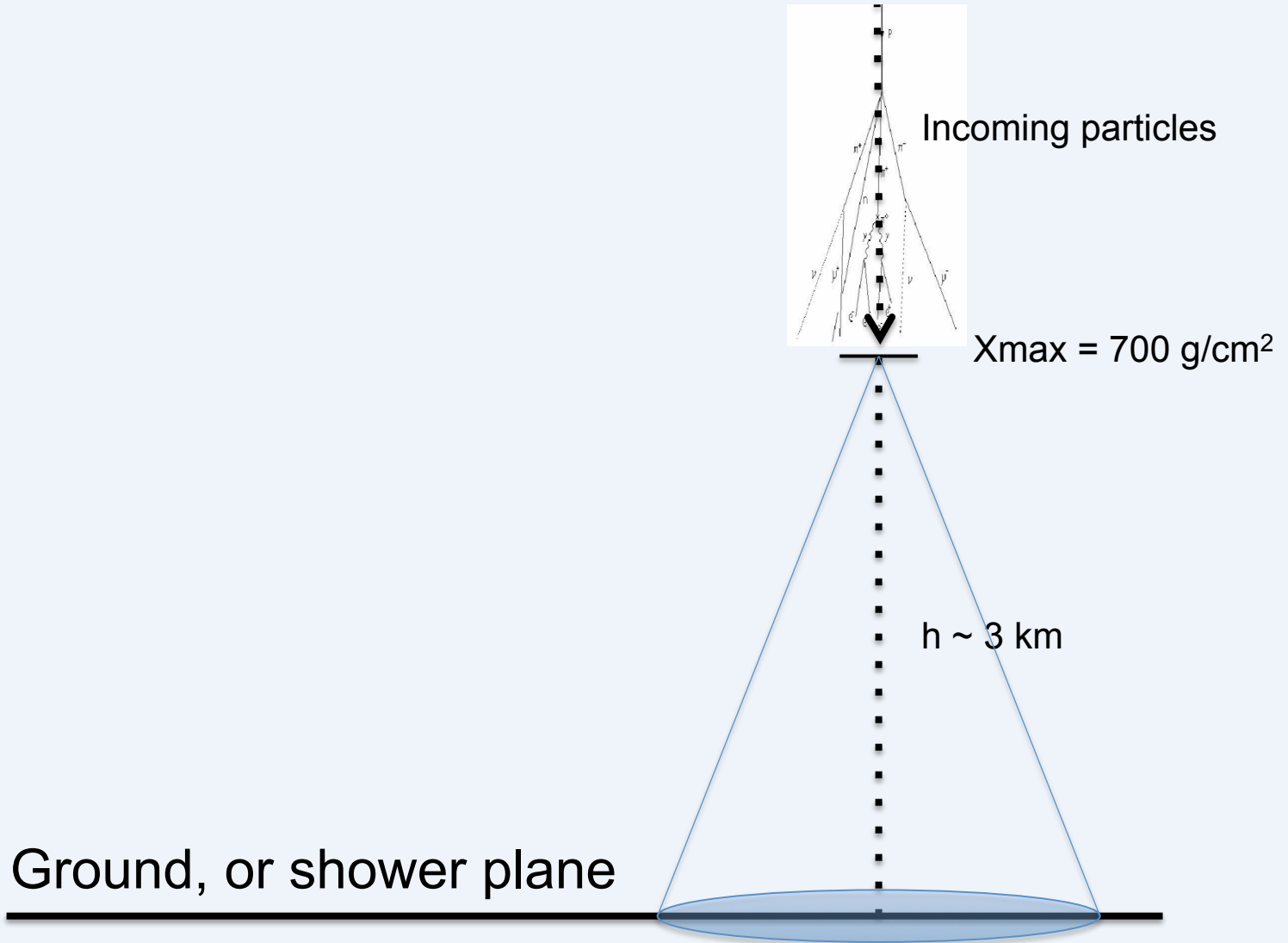
The atmospheric refractive index and radio emission from extensive air showers



Arthur Corstanje, RU Nijmegen, LOFAR CR KSP

LOFAR Science Workshop, Apr 5, 2016

Simplified air shower footprint on the ground



Simplified air shower footprint on the ground

Cherenkov angle: $\cos \theta = \frac{1}{\beta n}$

$\beta > 1/n$ for electrons > 20 MeV

$n \sim 1.0003$

-> $\theta \sim 1.2^\circ$

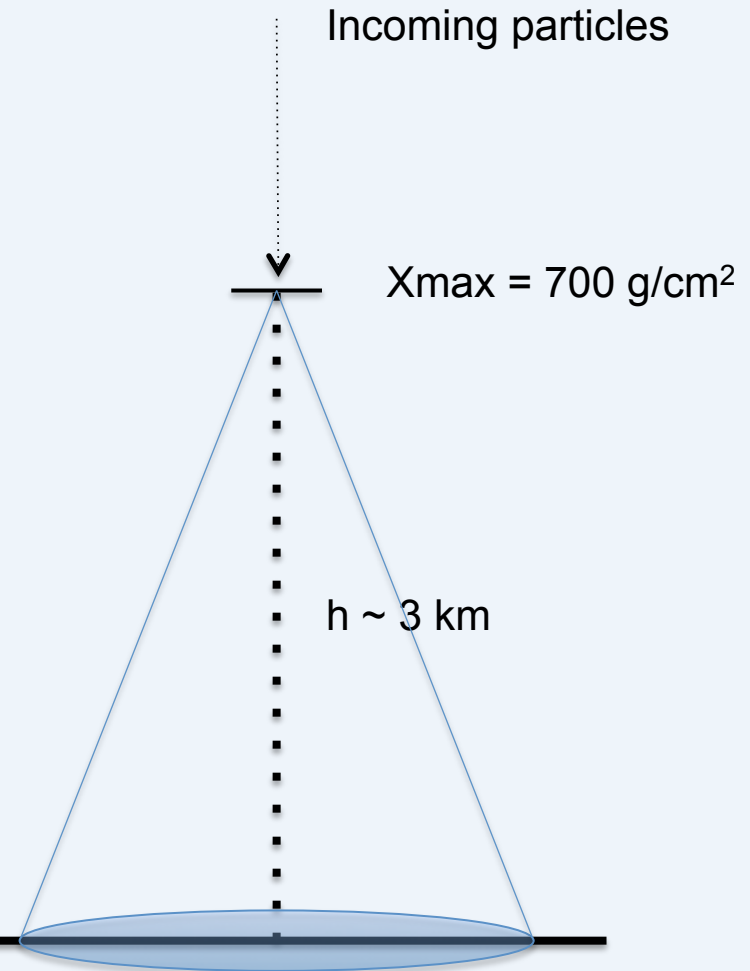
Variations at ground:

$N = 288$ (dry air 0°C)

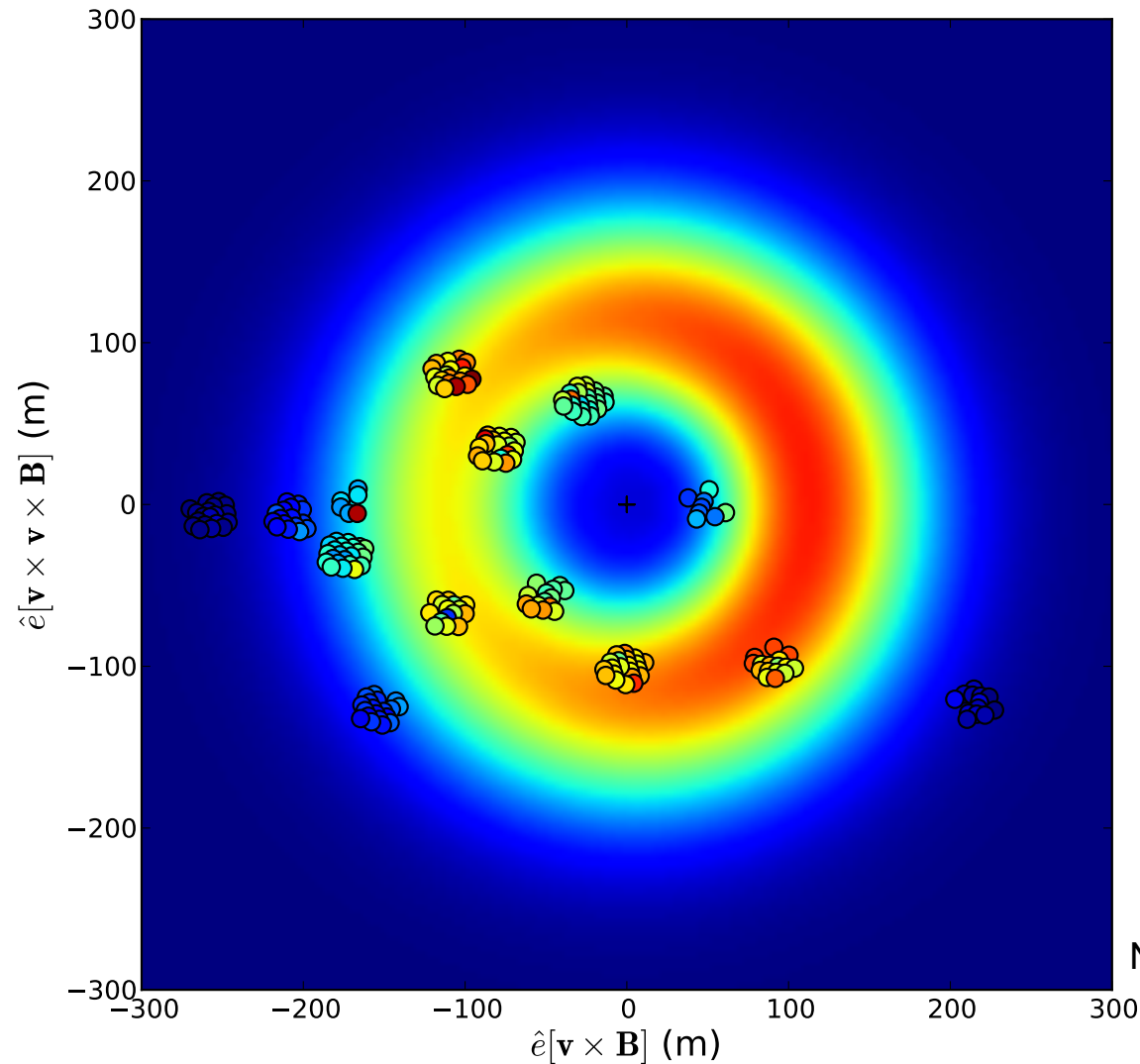
to $N \sim 330$ ppm

where $N = 10^6 (n-1)$

$p = 1013$ hPa = 1034 g/cm²



Cherenkov ring at frequencies > 120 MHz



Simulation (CoREAS)
with LOFAR HBA
detection (circles)

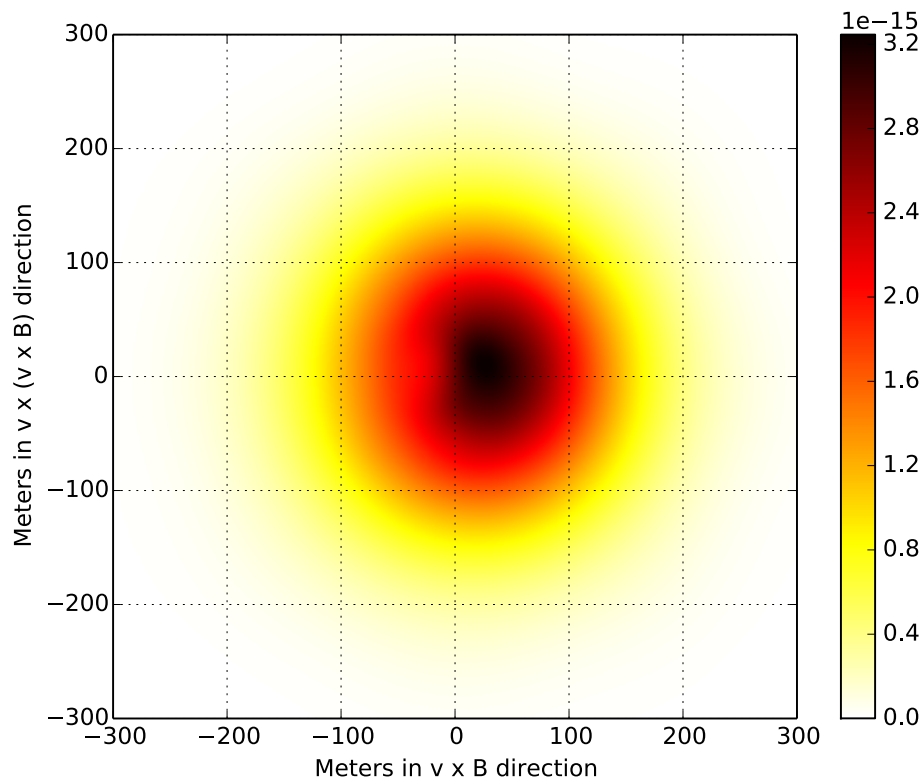
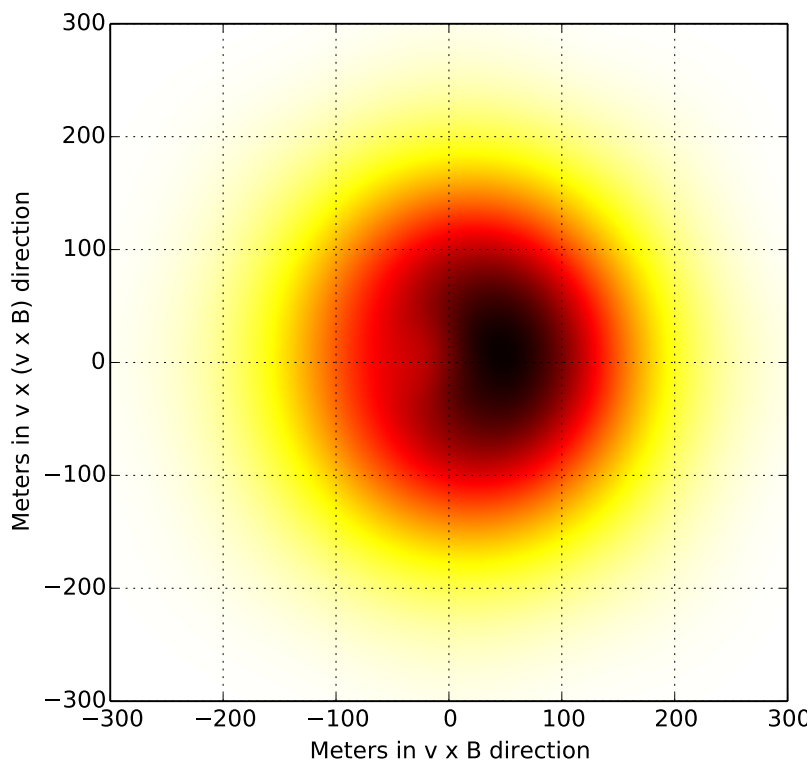
Ring becomes sharper
at higher frequencies

Nelles et al., *Astropart. Phys* 65 (2015)

Footprint at low frequencies, 30 – 80 MHz

$X_{\max} = 630 \text{ g/cm}^2$

$X_{\max} = 700 \text{ g/cm}^2$



CoREAS simulated footprints of radio intensity

The effect of variations in refractive index (simplified)

- Assuming all radiation coming from X_{\max} level
- Assuming footprint size scales with Cherenkov angle

Fitted $X_{\max} = 700 \text{ g/cm}^2$

Actual $X_{\max} = 715 \text{ g/cm}^2$

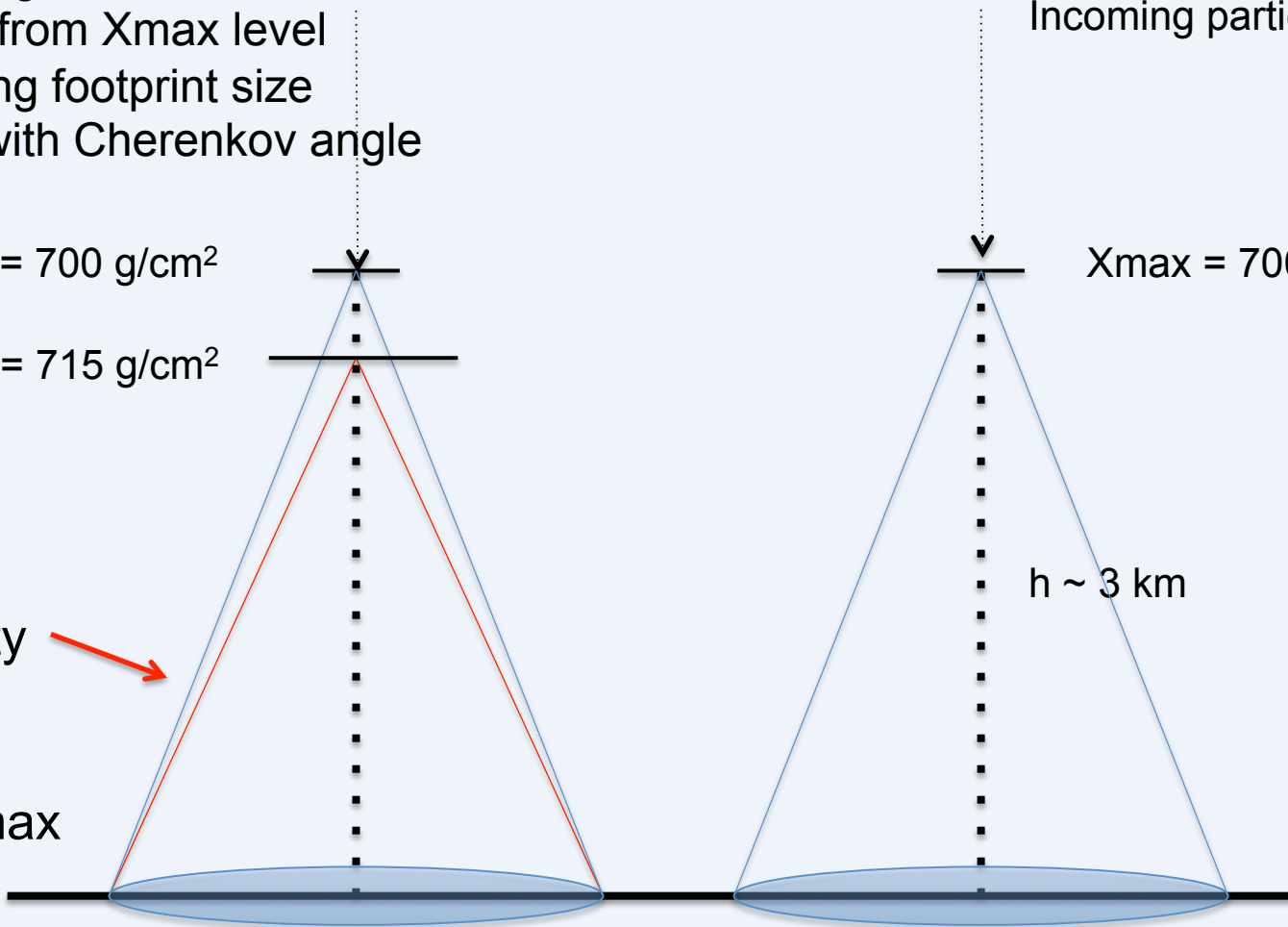
Higher refractivity

Mimics lower X_{\max}

Incoming particles

$X_{\max} = 700 \text{ g/cm}^2$

$h \sim 3 \text{ km}$



Standard atmosphere (US/Intl.)

Constant temperature lapse rate: $L = 6.5 \text{ K / km}$

$$p_0 = 1013.25 \text{ hPa} \quad (\text{variable})$$

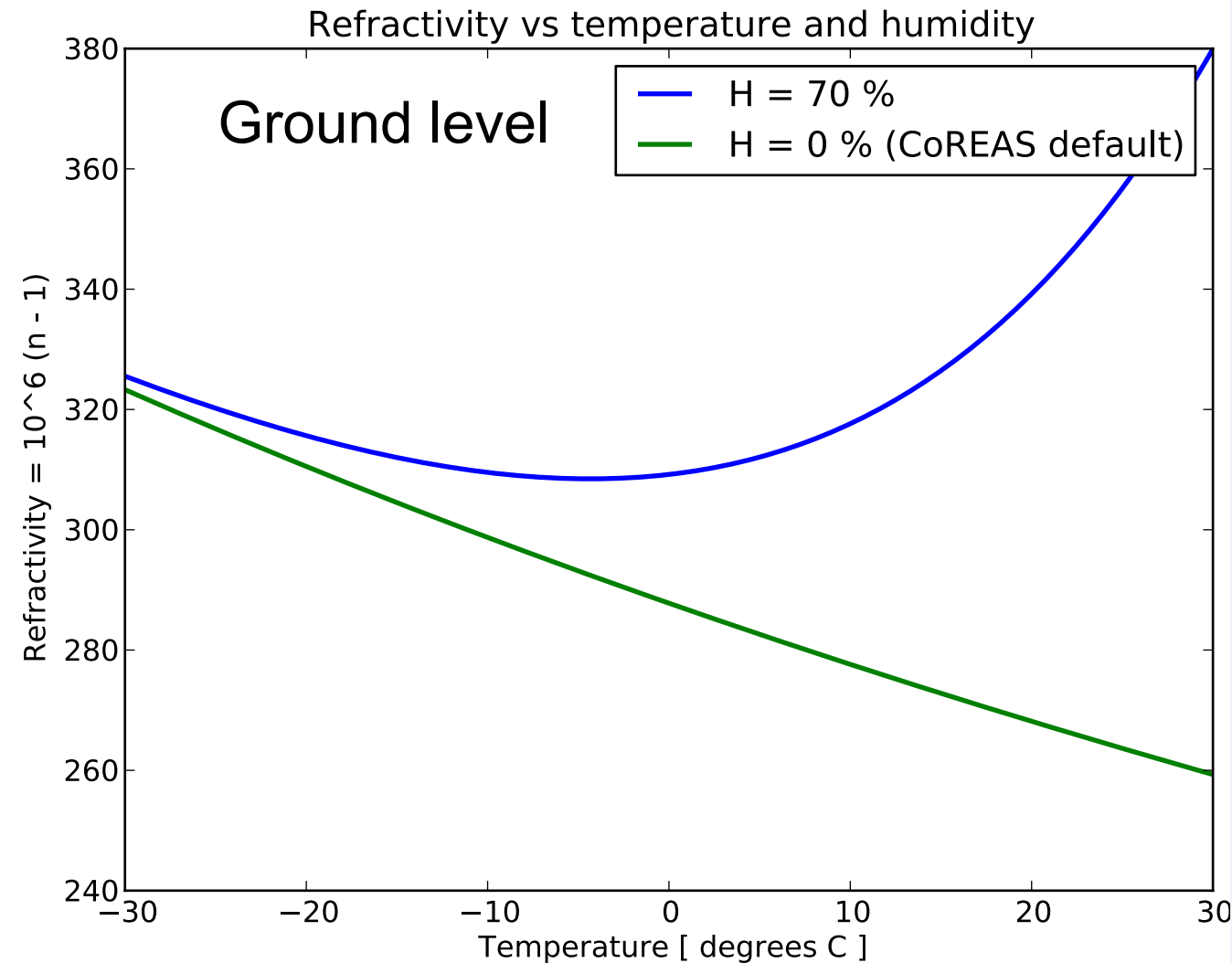
$$T_0 = 273.15 + 15 \text{ K} \quad (\text{“ ”})$$

Then:
$$p = p_0 \left(1 - \frac{L}{T_0} h \right)^{\frac{gM}{RL}}$$

Refractivity N can be expressed as a function of pressure, temperature and (relative) humidity

- Different for radio vs IR/visible !
- Especially the water vapor contribution

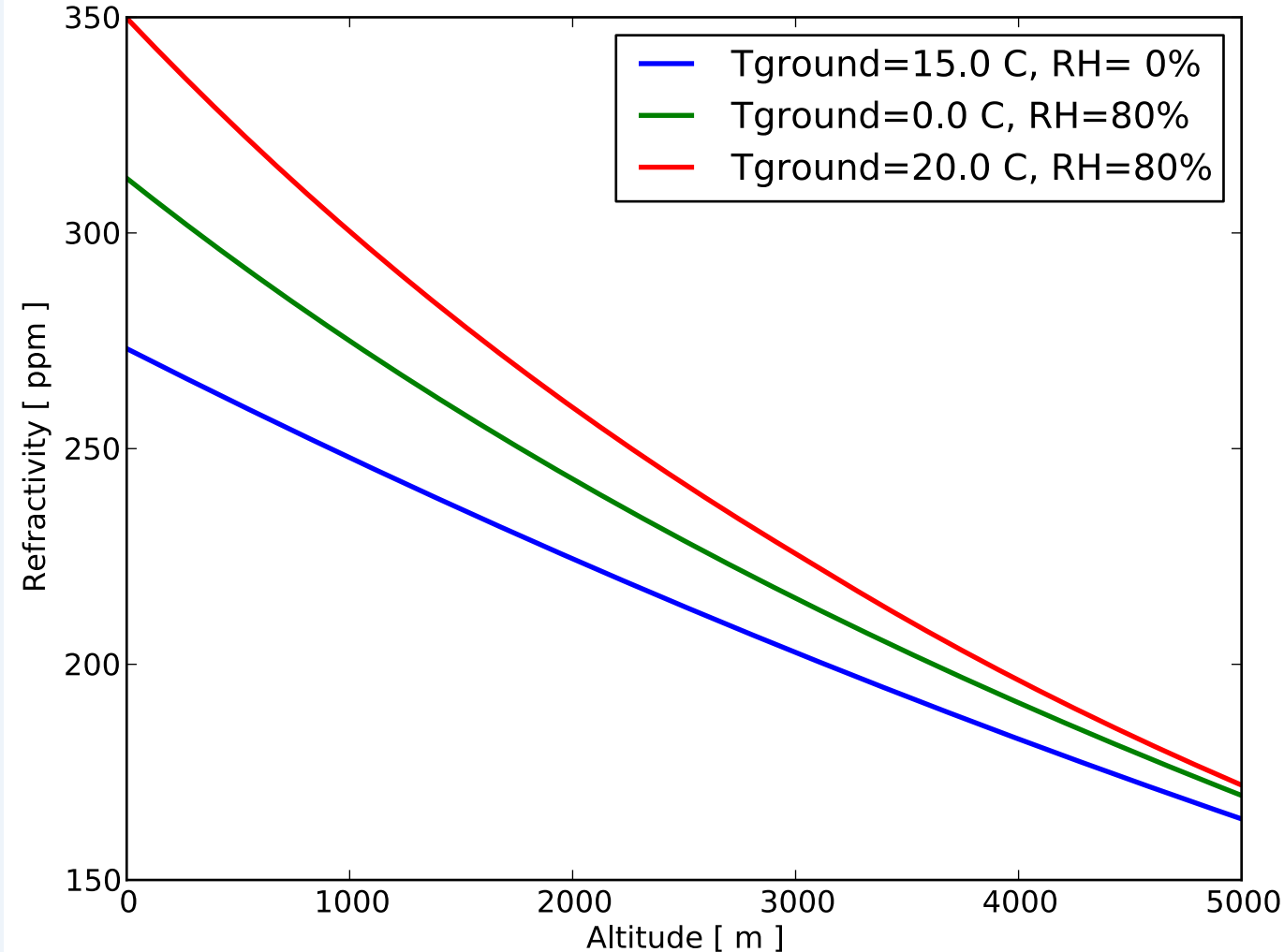
Variations in refractive index



Refractivity
proportional to
density (dry)
= green line

Absolute
humidity
independent
of pressure

Variations in refractive index

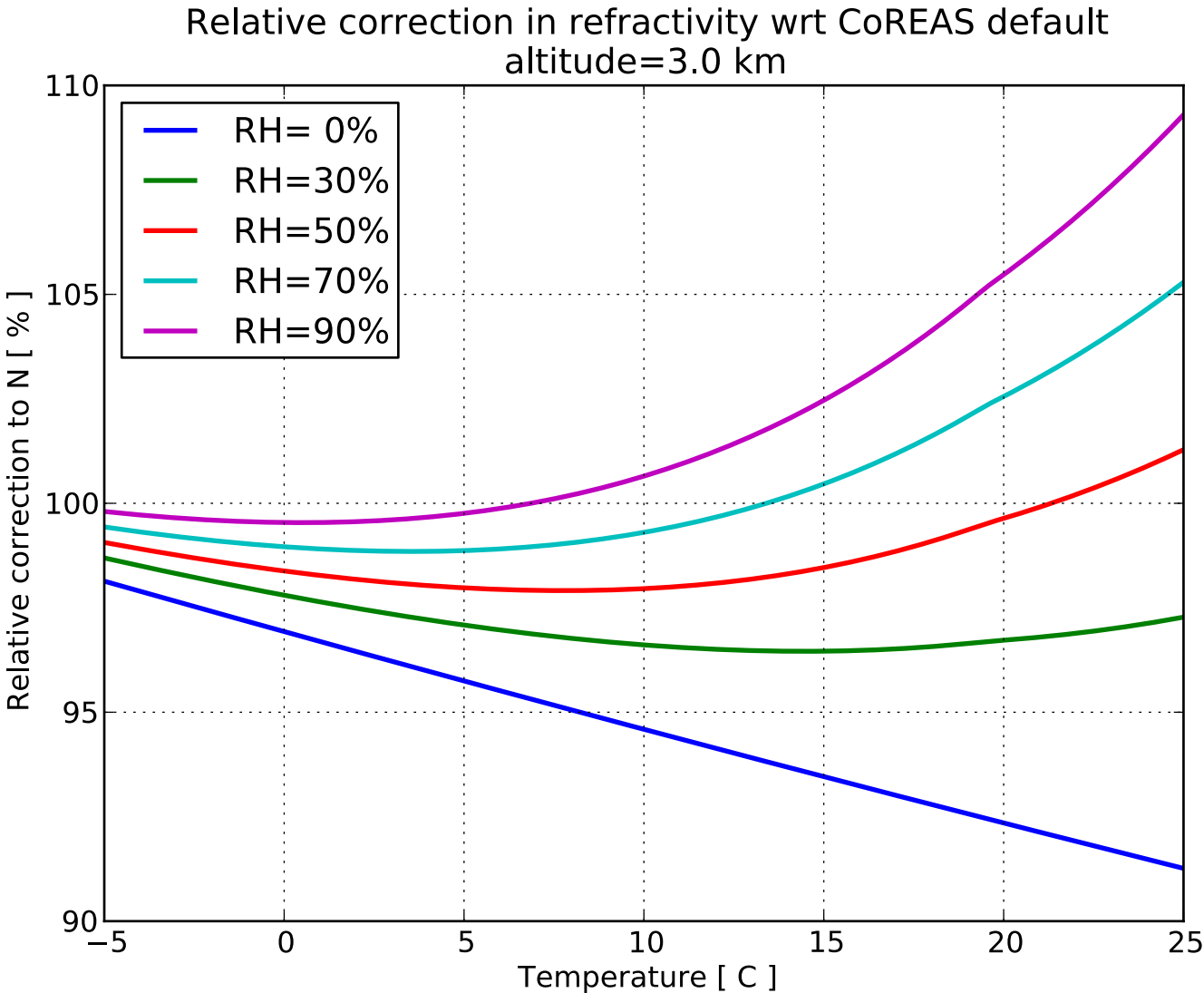


Humidity raises refractivity at all altitudes (radio)

Absolute humidity drops with altitude (colder)

Relative humidity (RH) assumed const

Variations in refractive index



Percentage
scaling of N

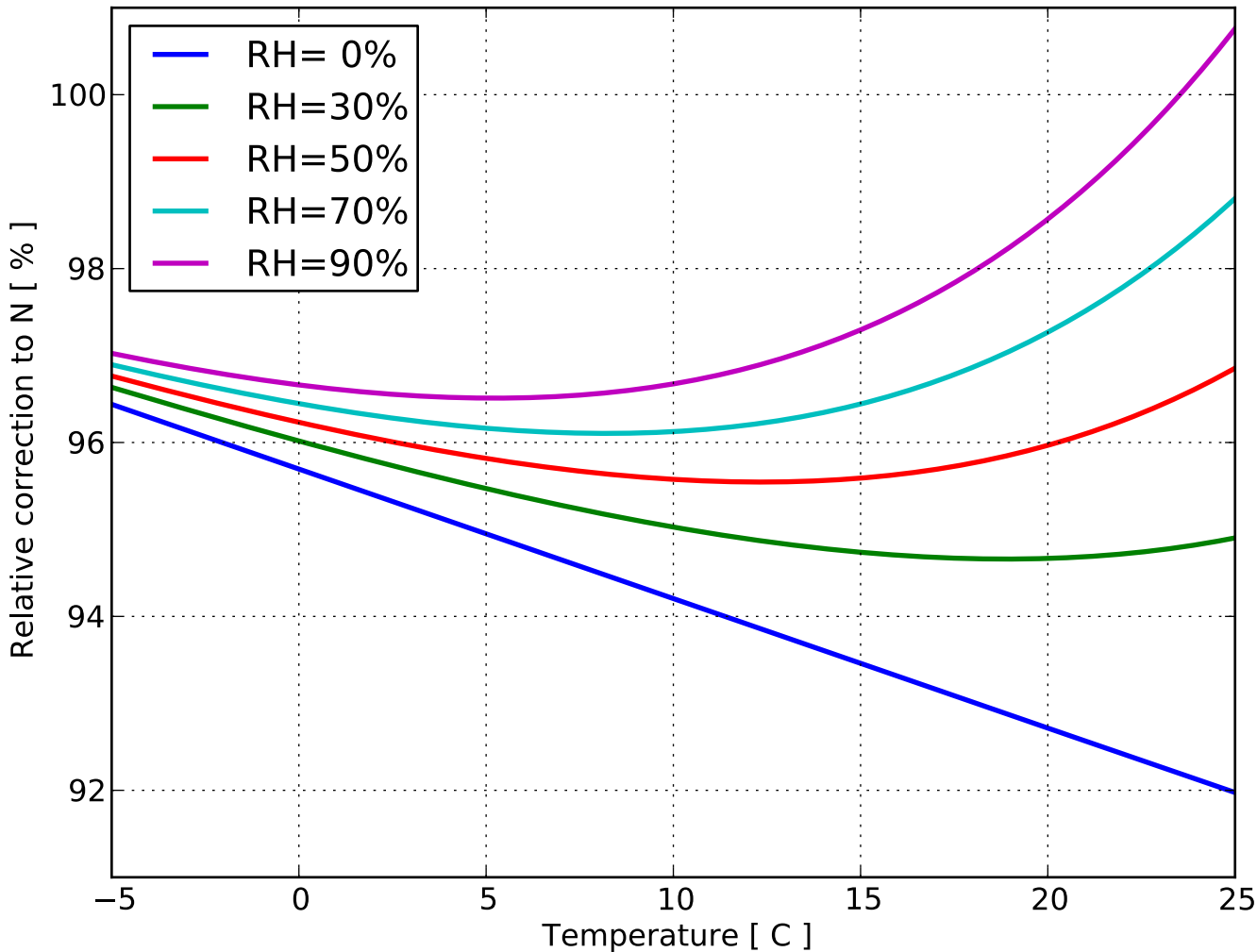
3 km up (zenith)

Versus ground
temperature and
relative humidity

RH assumed
constant

Variations in refractive index

Relative correction in refractivity wrt CoREAS default
altitude=5.0 km



Percentage
scaling of N

5 km up
for $\sim 45^\circ$ shower

Systematically
too low, $\sim 4\%$

- Cannot use
one number for
N at ground

Effect of varying refractivity on Xmax measurements

- Use a fitting method as in composition analysis:

Normal N

50 simulated showers

Fit 49 showers to the test shower

(lateral power distributions)

- Make plot of fit quality versus Xmax
- Minimum indicates best-fitting Xmax

10% higher N

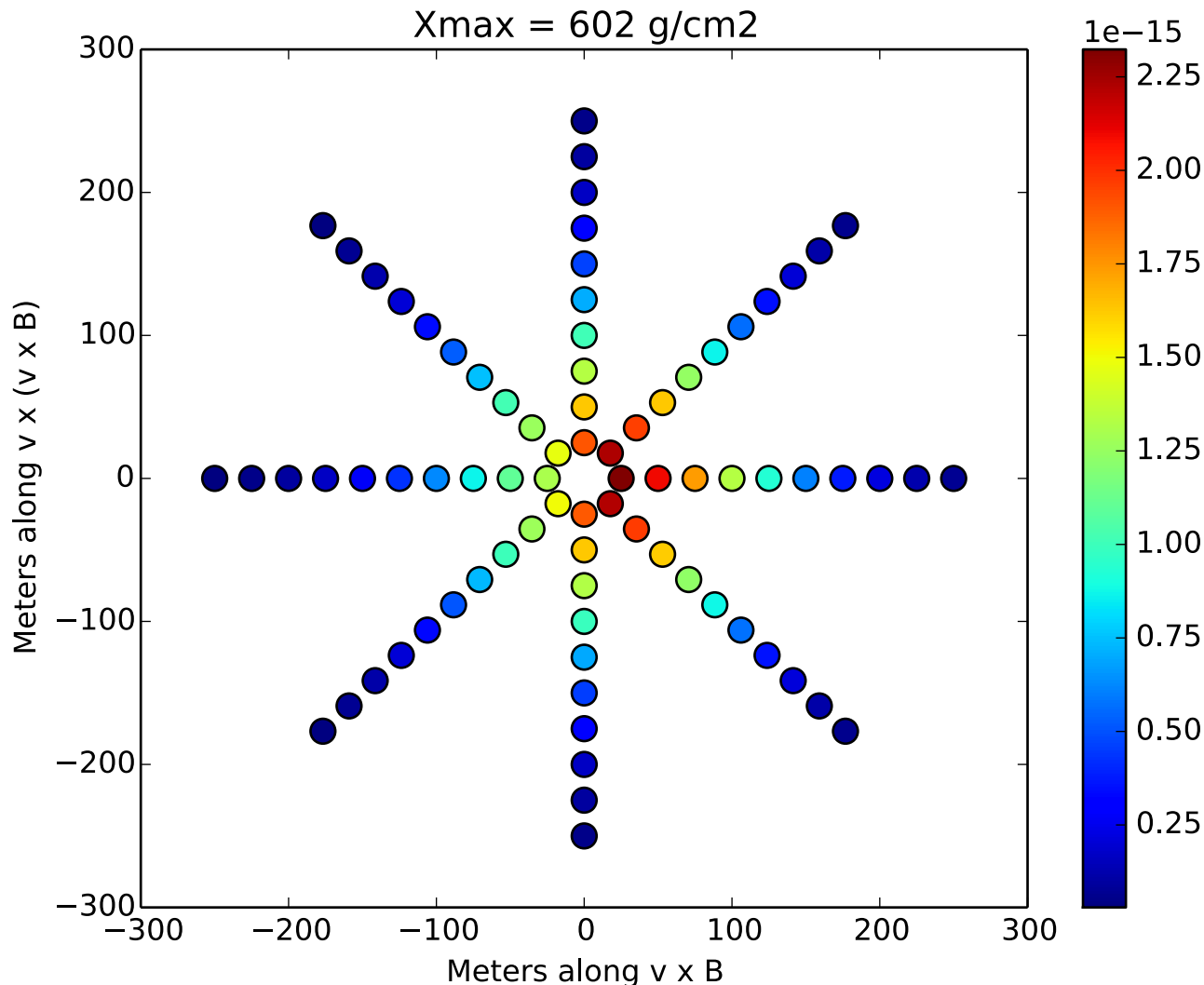
50 simulated showers

take one as 'test shower'

('measured data')

Average offset over all 'test' showers

Fitting shower LDF profiles

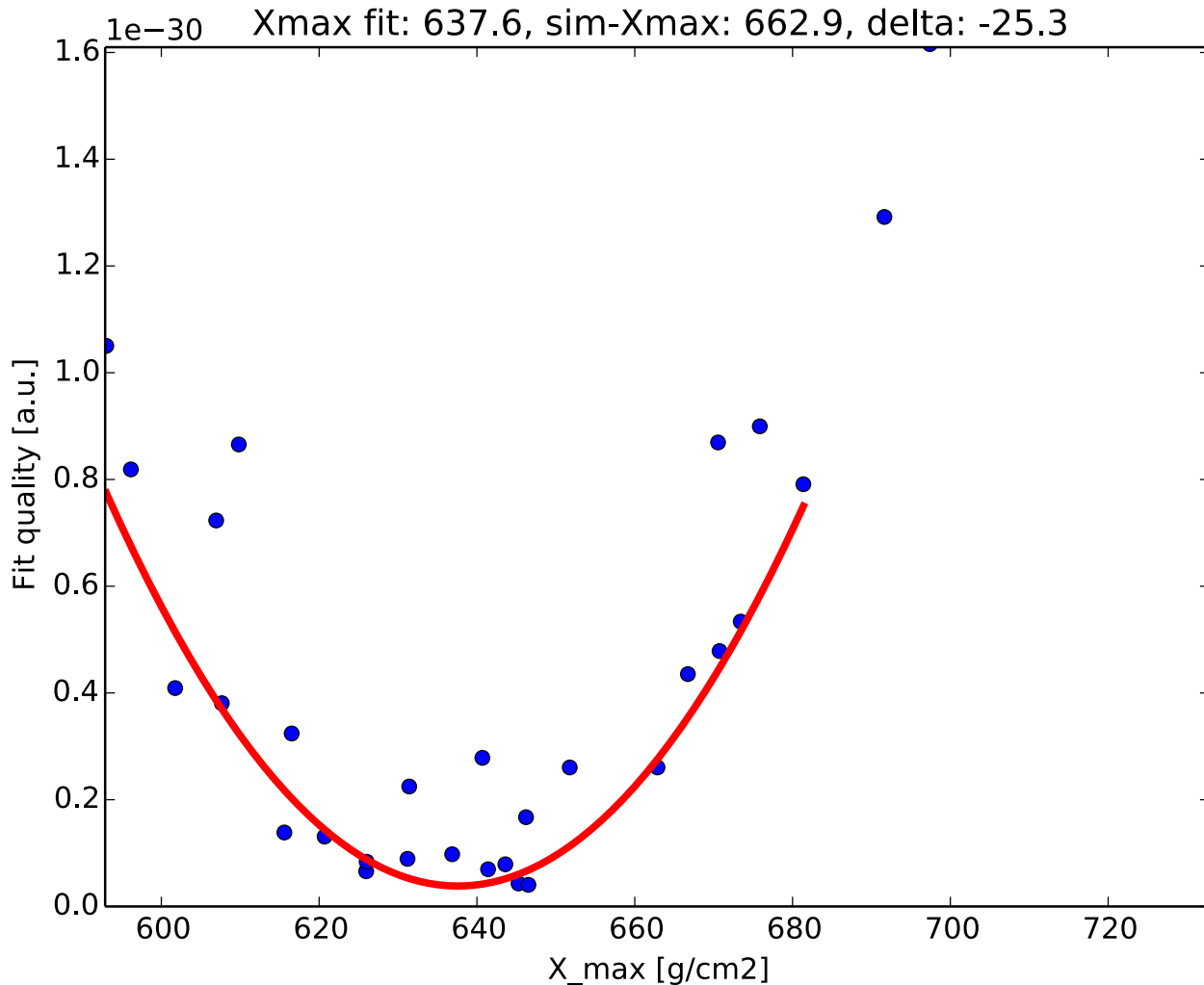


Compared
antenna by
antenna

Least-squares,
MSE = fit quality

x 50

Fitting Xmax



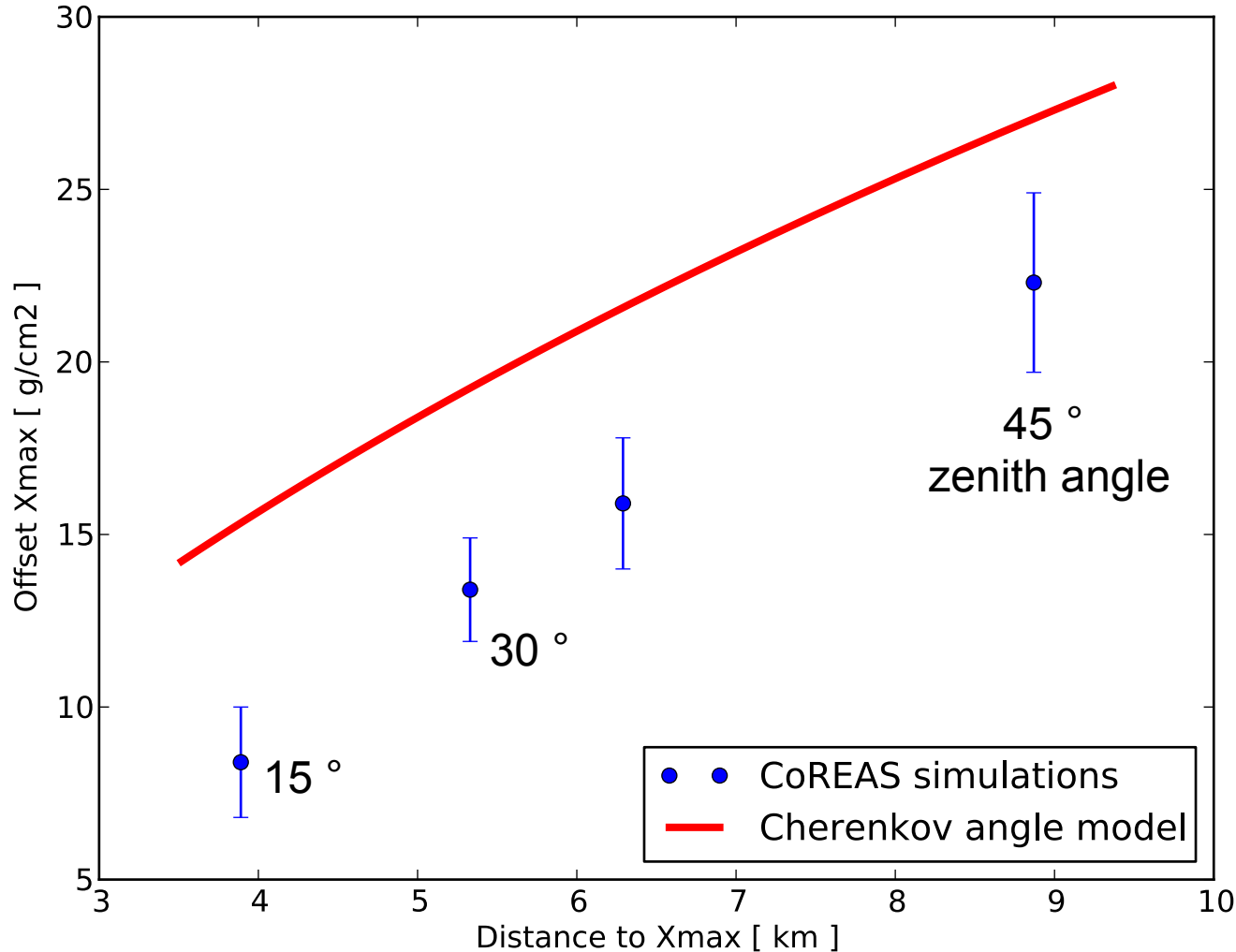
Parabolic fit,
weights $1 / y^2$
for emphasis on
lower envelope

True Xmax is
known (simulation)

Fitted Xmax
measures the
offset from varying
refractivity

Results

Xmax offset for 10 % increase in N



Rough model fits
qualitatively
Up to ~ 25 % scaling

Cherenkov angle
dependence only

Standard
atmosphere

No free parameters!

Significant effect
cf. 16 g/cm²
precision in
composition analysis

Conclusion

- Atmospheric refractivity is one of the major systematic uncertainties in determining X_{\max}
- Bias of 0.9 to 2.2 g/cm² per 1 % change in N
 - About 4 % variability realistic at X_{\max} level
- A simple model describes offsets qualitatively
 - Cherenkov angle scaling, standard atmosphere
- Not fully fixable through only ground level refractivity / one standard atmosphere
 - Would need CoREAS update, and use local weather data