

Faraday Tomography with LOFAR

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Background: IC342 as seen by LOFAR (Stokes I)



Detecting Magnetic Fields

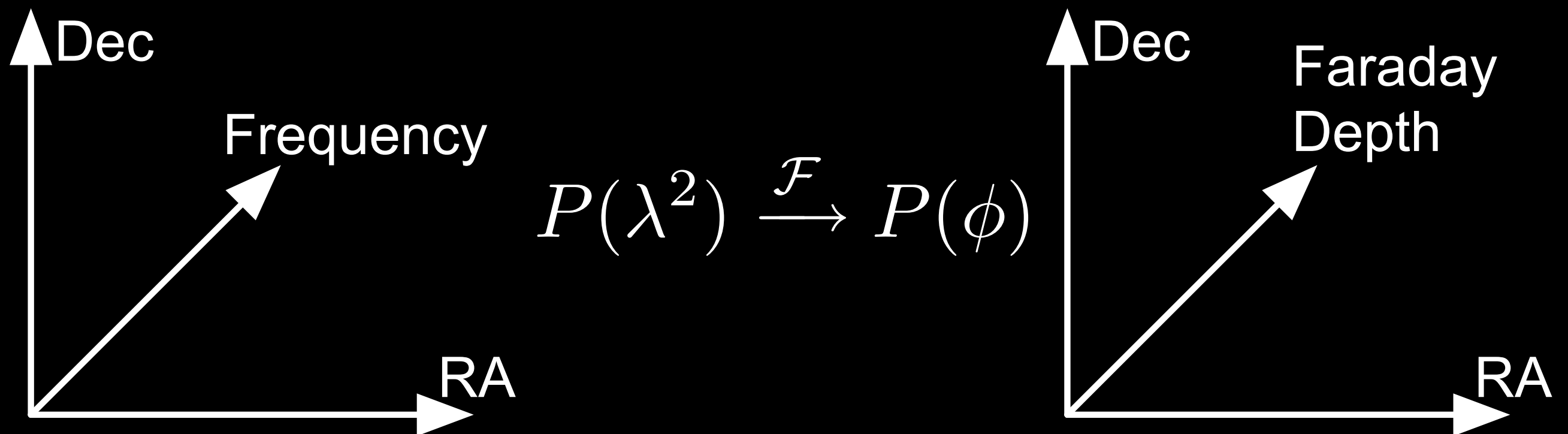
- cosmic rays + magnetic field =
polarized synchrotron emission
- polarization + free electrons + magnetic field =
Faraday rotation

$$\text{Change in polarization} \propto \lambda^2 \int_0^d n_e \vec{B} \cdot d\vec{l}$$

Faraday depth, ϕ

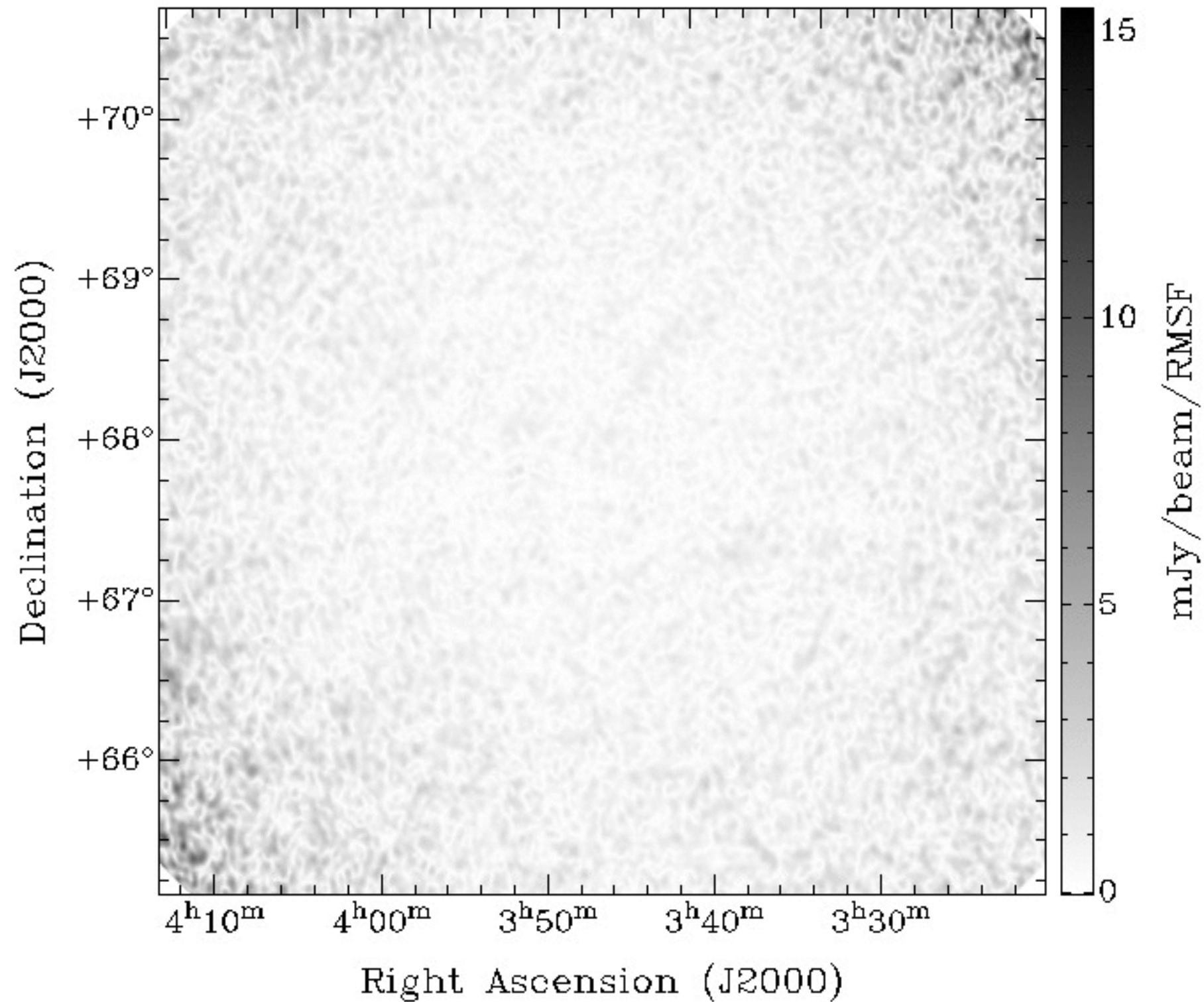
Faraday Tomography

- Broad-band radio polarization cubes can be transformed into **Faraday depth cubes**:



The IC342 field

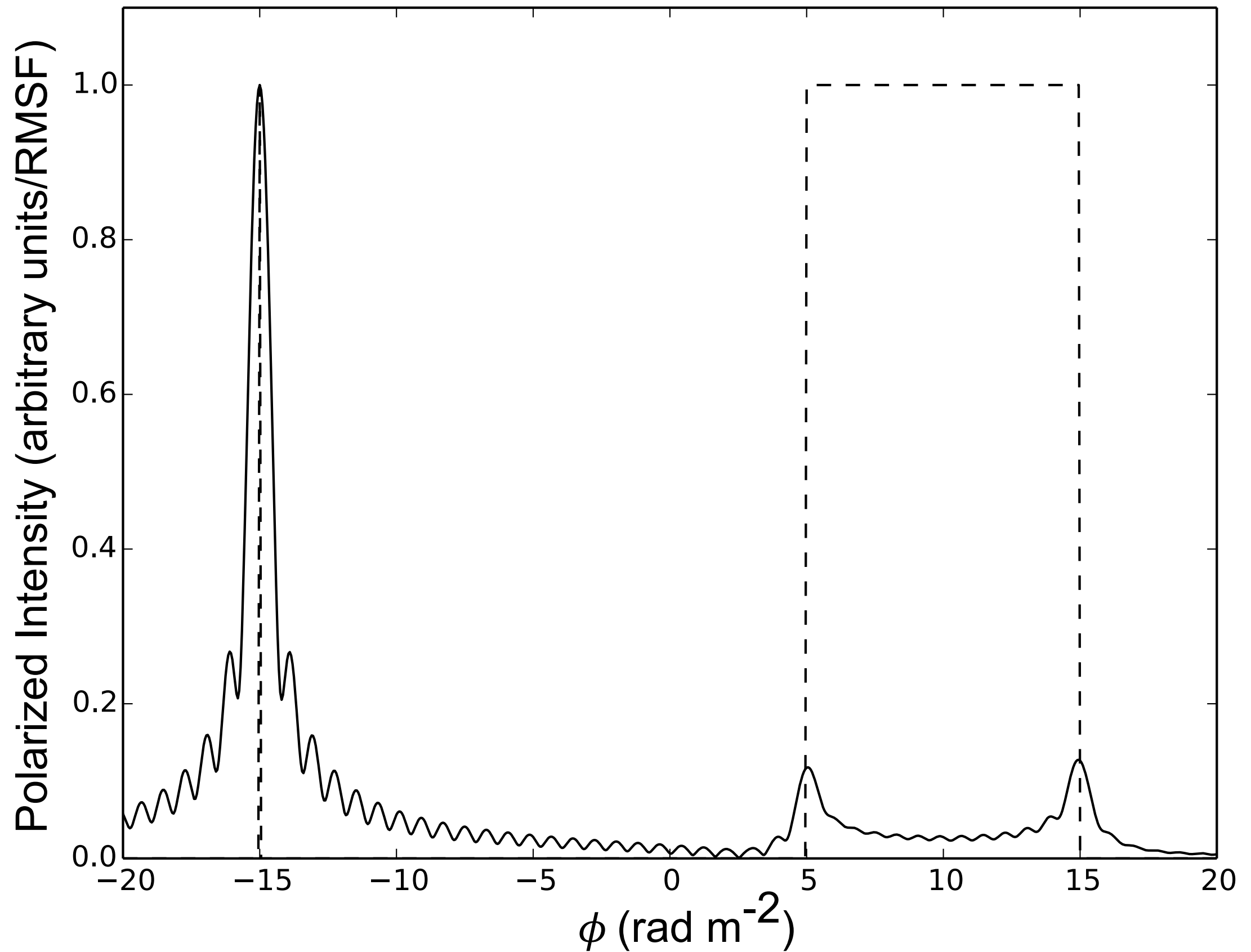
Phi: $-1.000000e+01$



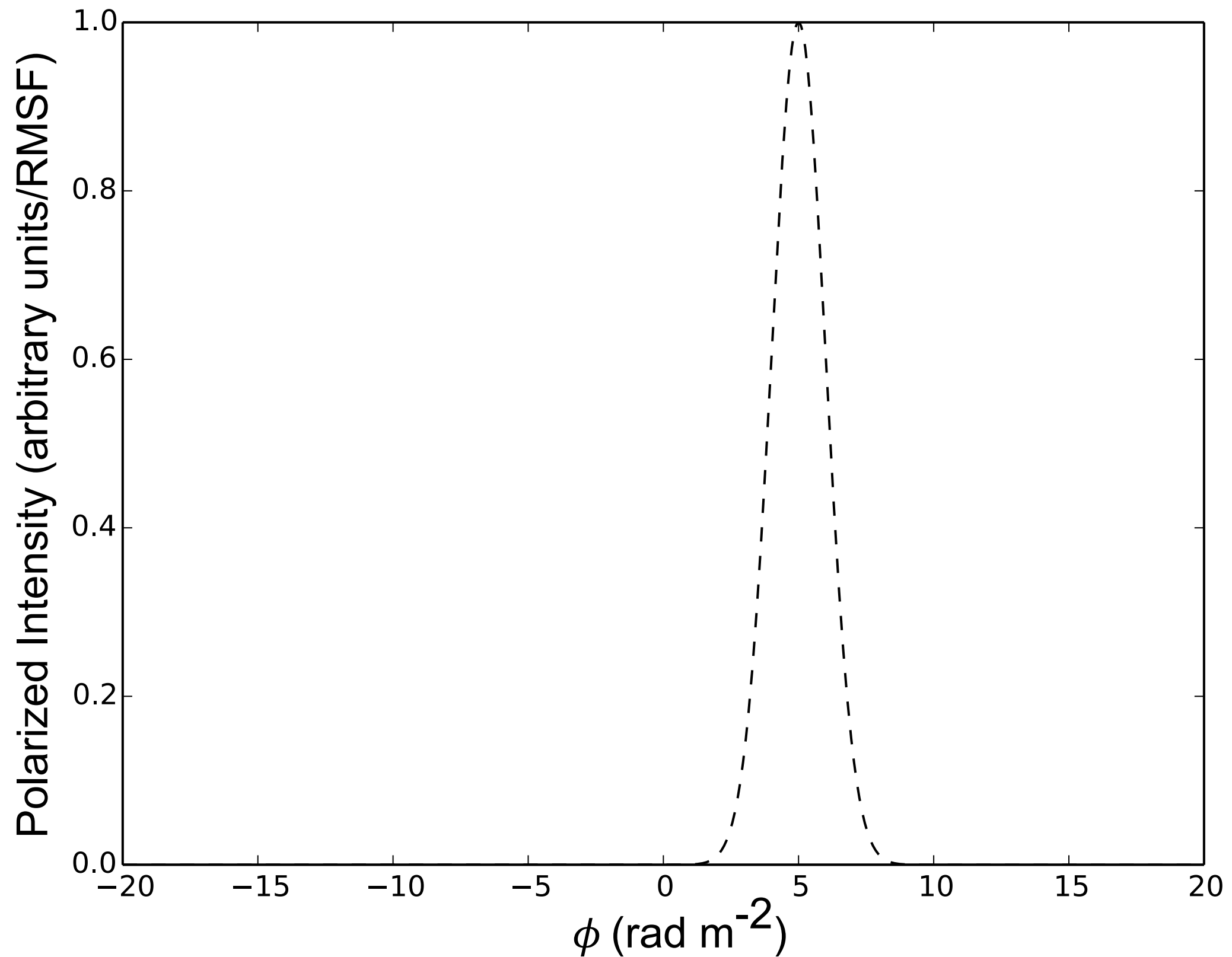
The Problem with (low-frequency) Polarization

- Faraday tomography is a Fourier-based technique: **incomplete sampling** leads to a PSF, sidelobes, missing short spacings/large structures.
- Modelling needs to account for this: how do **physical features** get transformed?

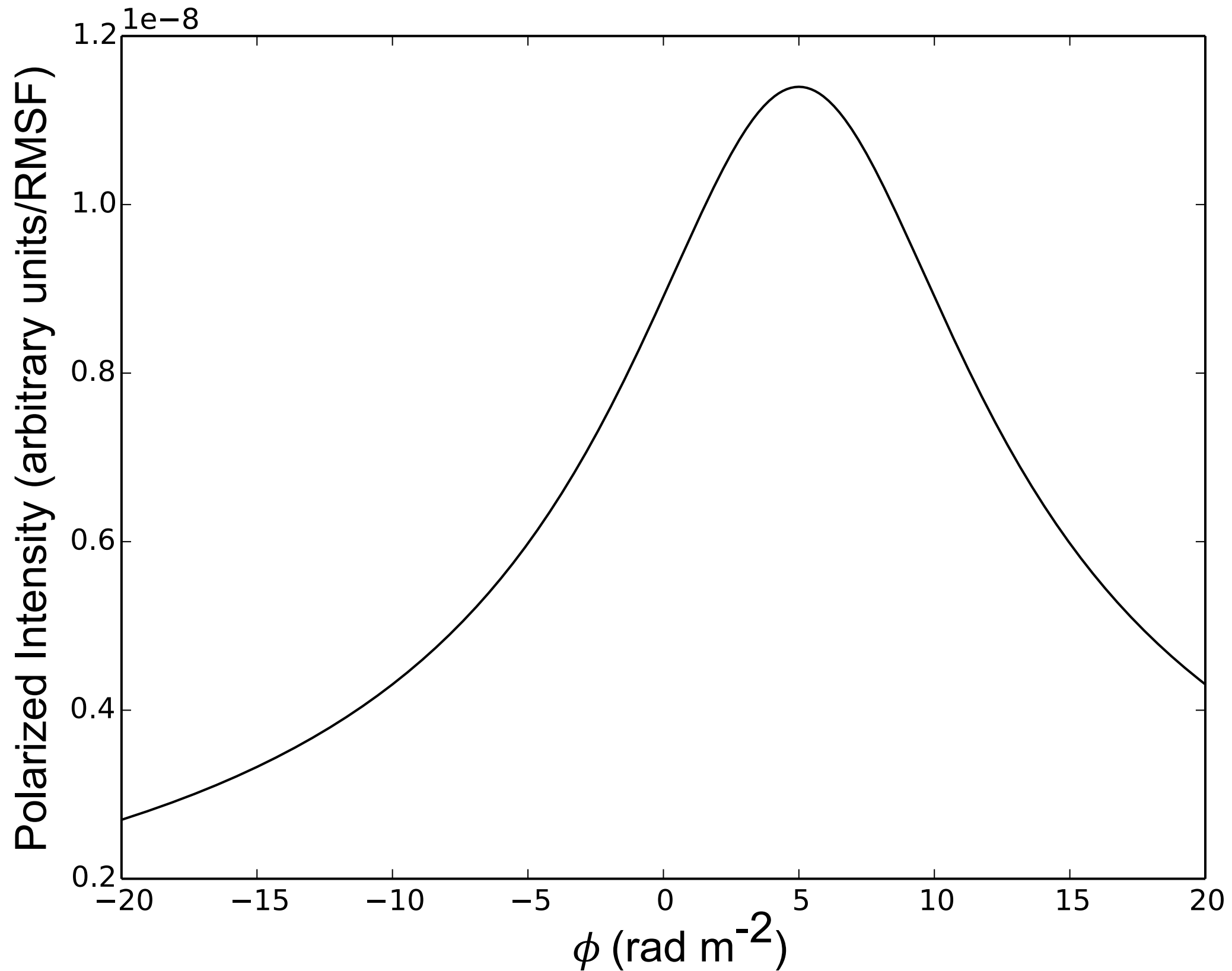
Simulated Spectrum - Tophat slab



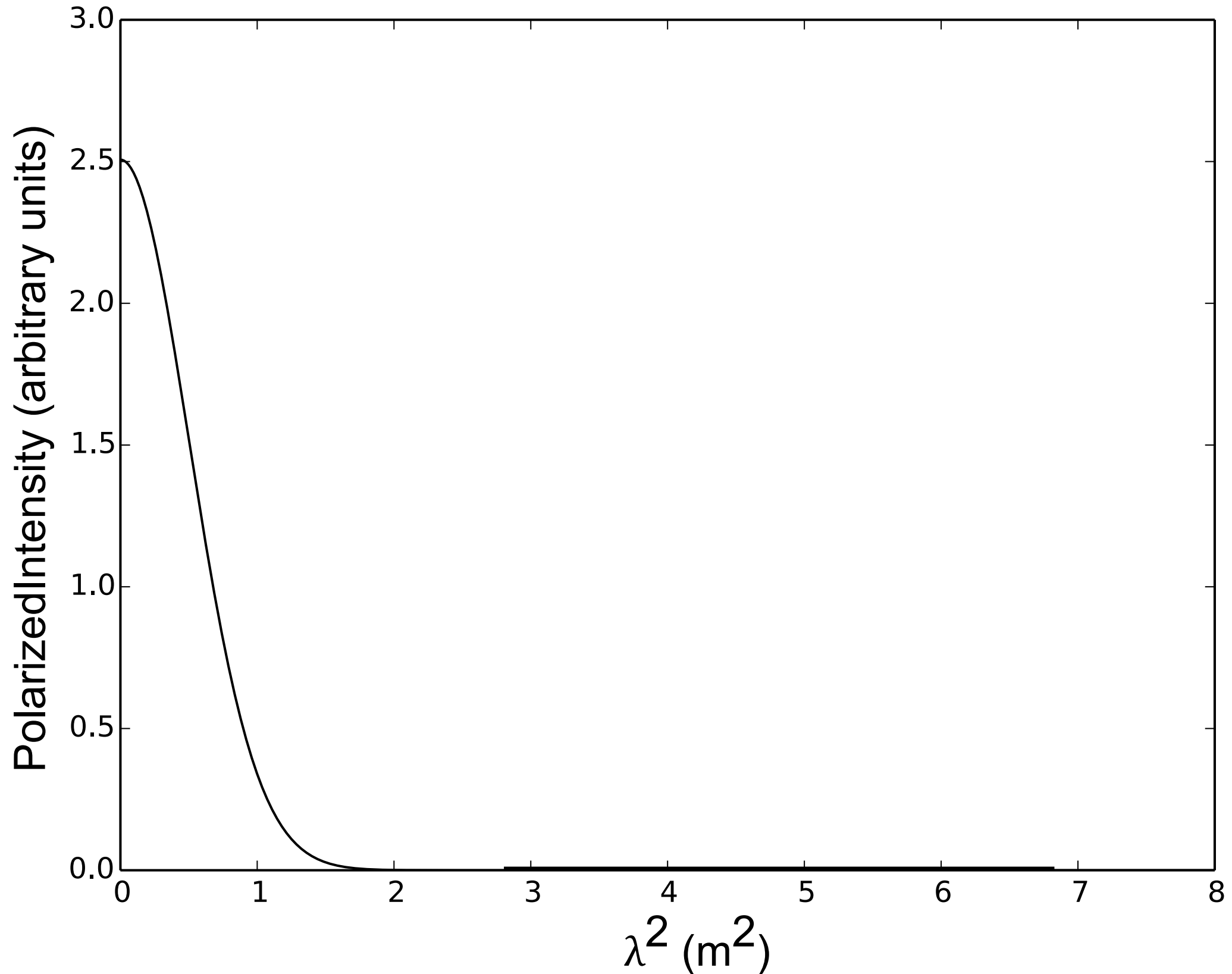
Simulated Spectrum - Gaussian



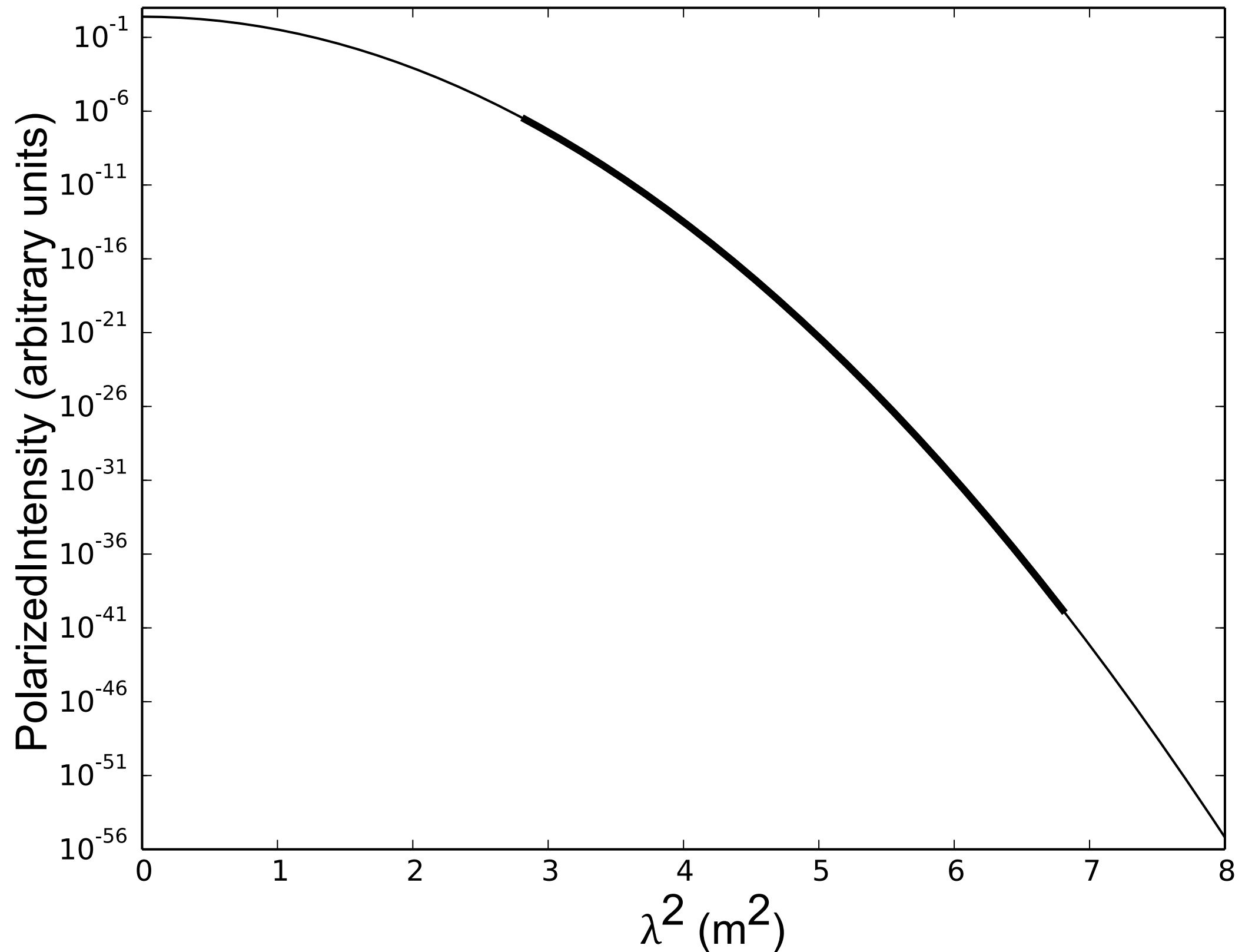
Simulated Spectrum - Gaussian



Simulated Spectrum - Gaussian



Simulated Spectrum - Gaussian



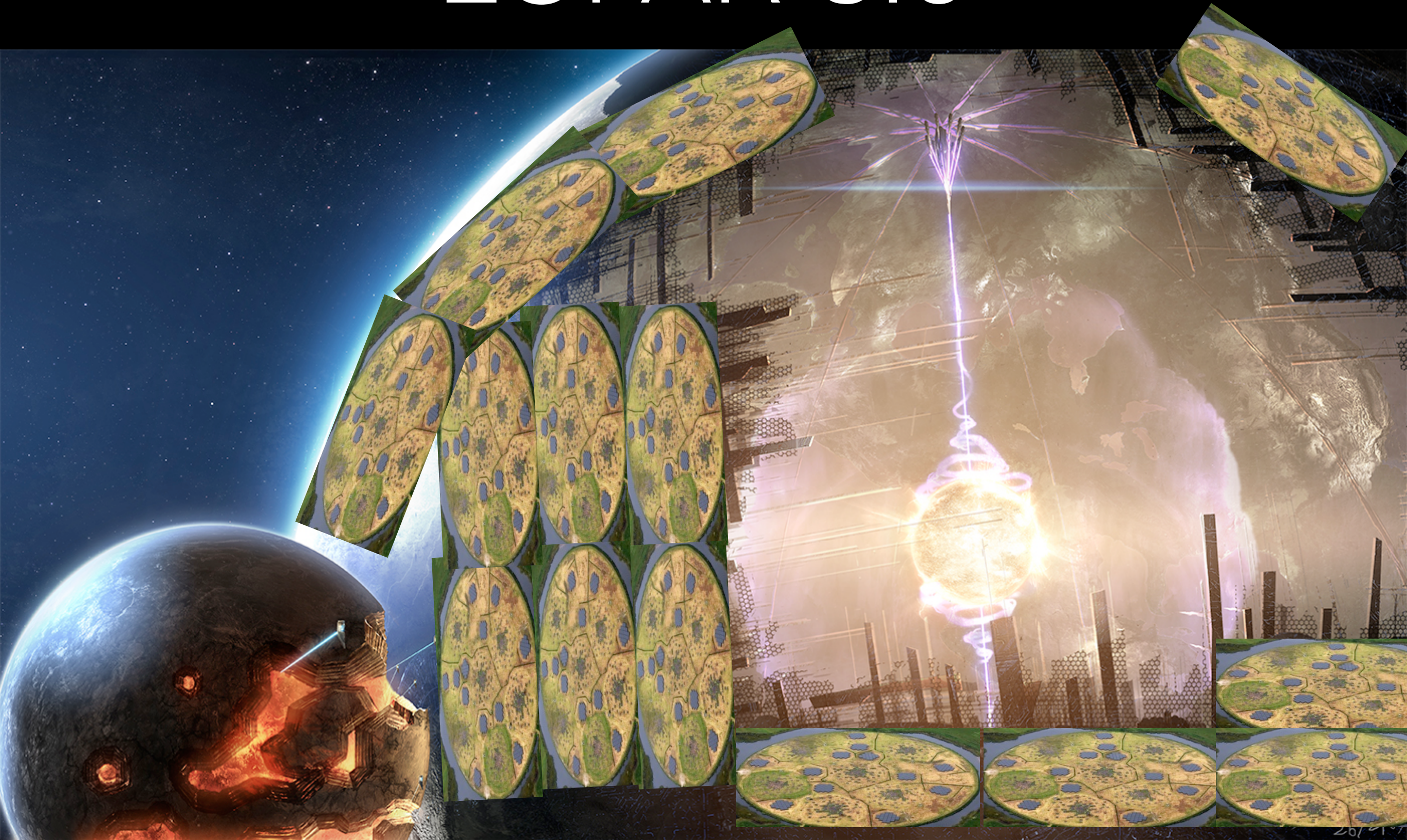
LOFAR 3.0

LOFAR 3.0

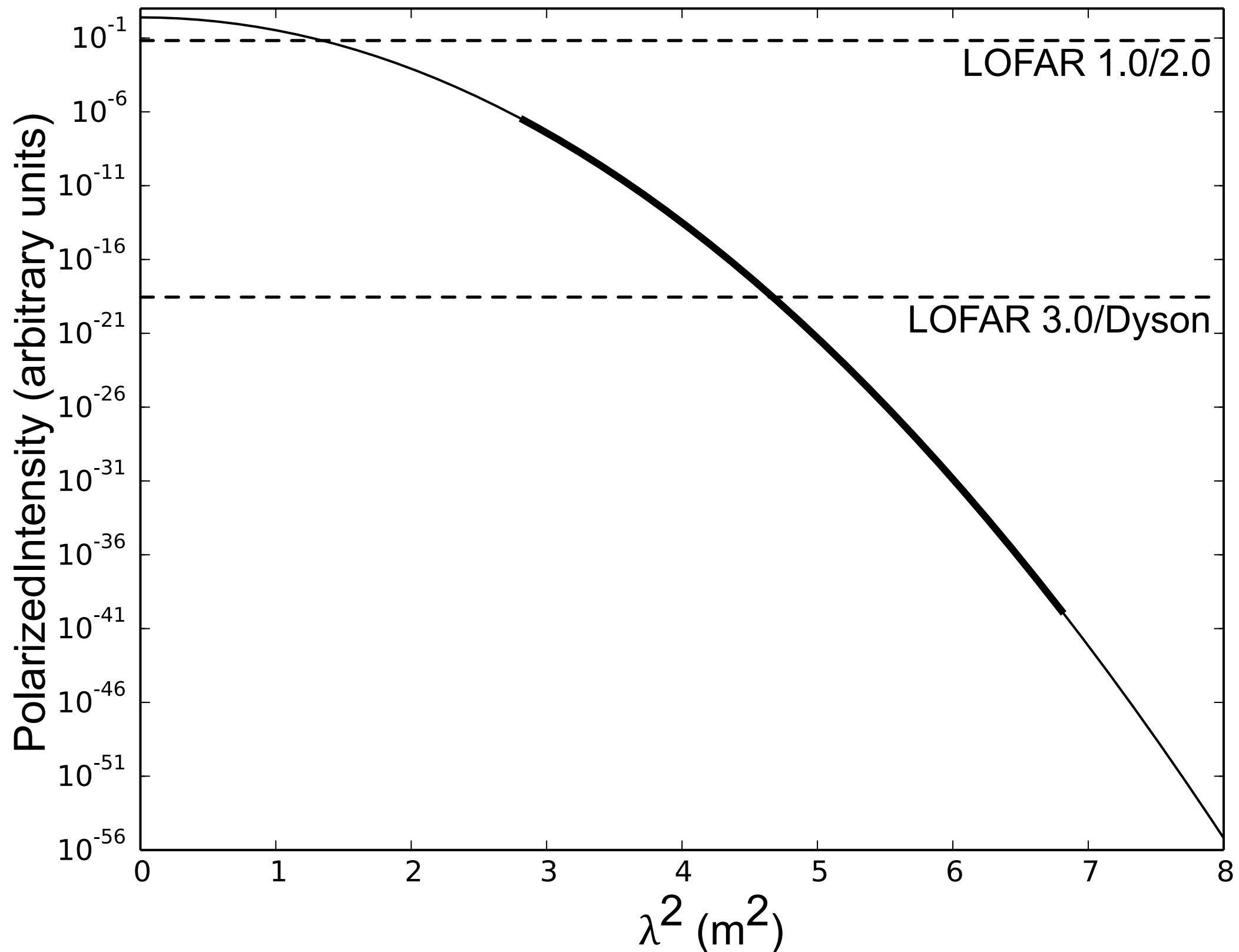


Image credit: "Shield World Construction" by Adam Burn

LOFAR 3.0



Detecting a Gaussian



LOFAR 4.0?

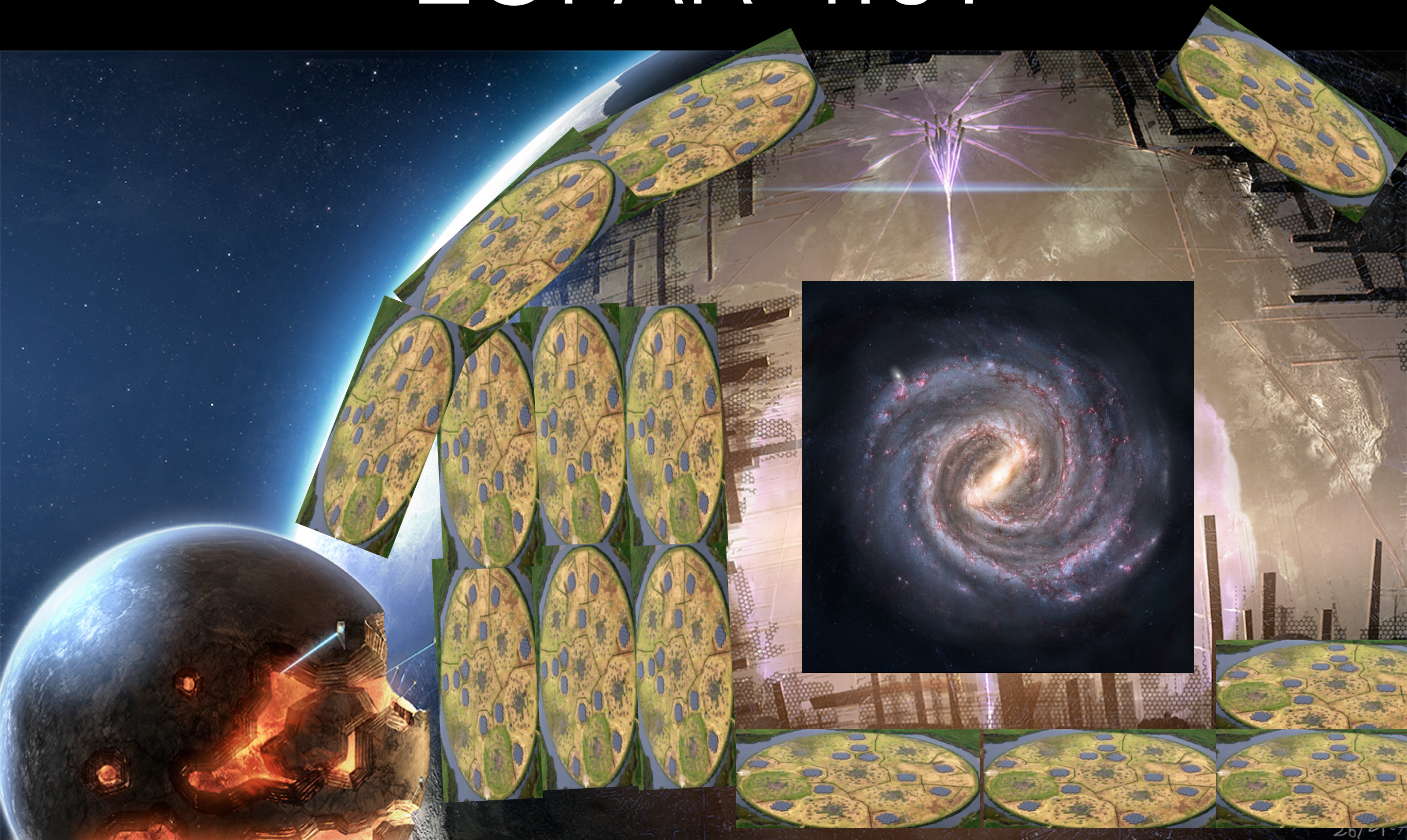
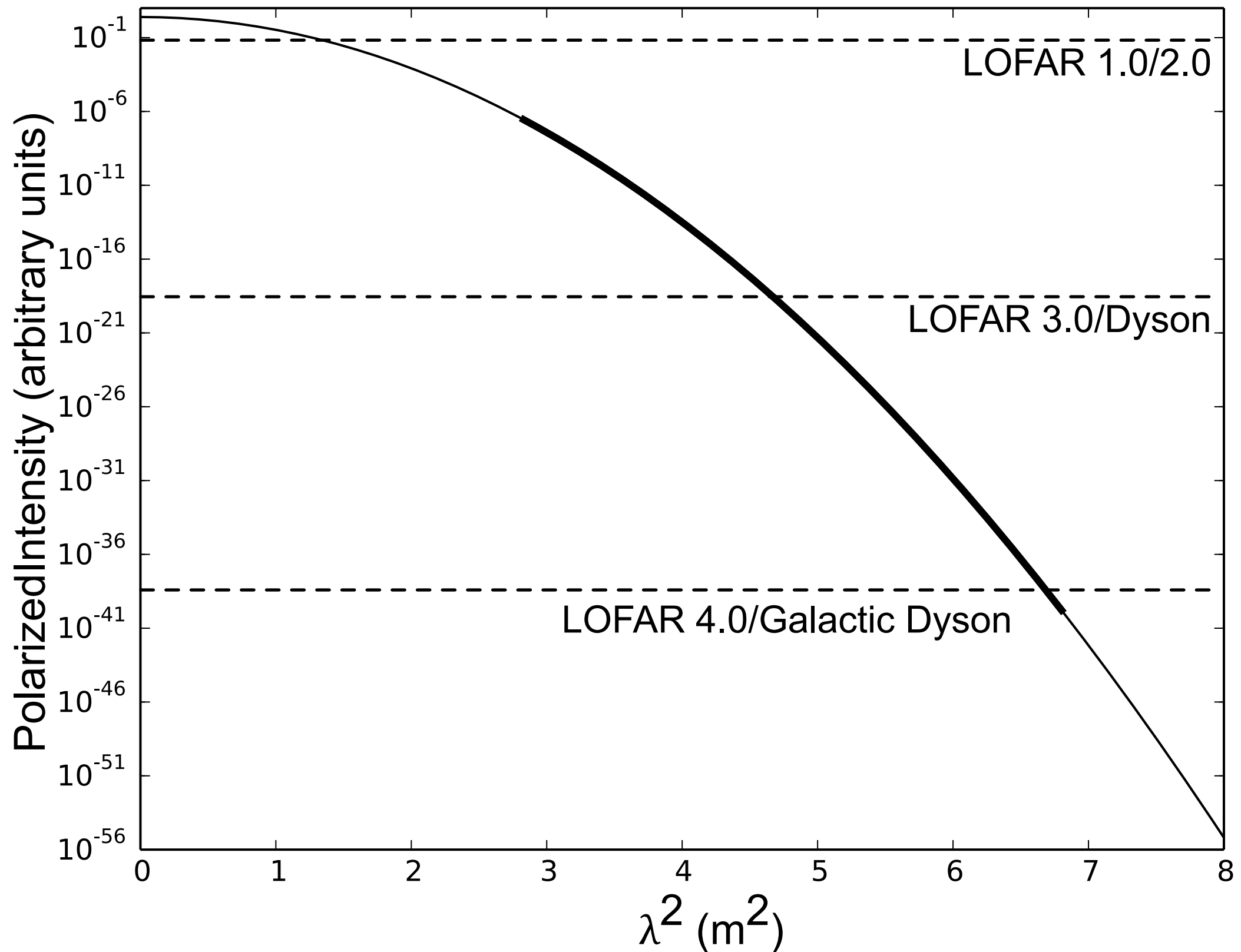


Image credit: "Shield World Construction" by Adam Burn

Image credit: Nick Risinger/Wikimedia

Image credit: astron.nl

Detecting a Gaussian



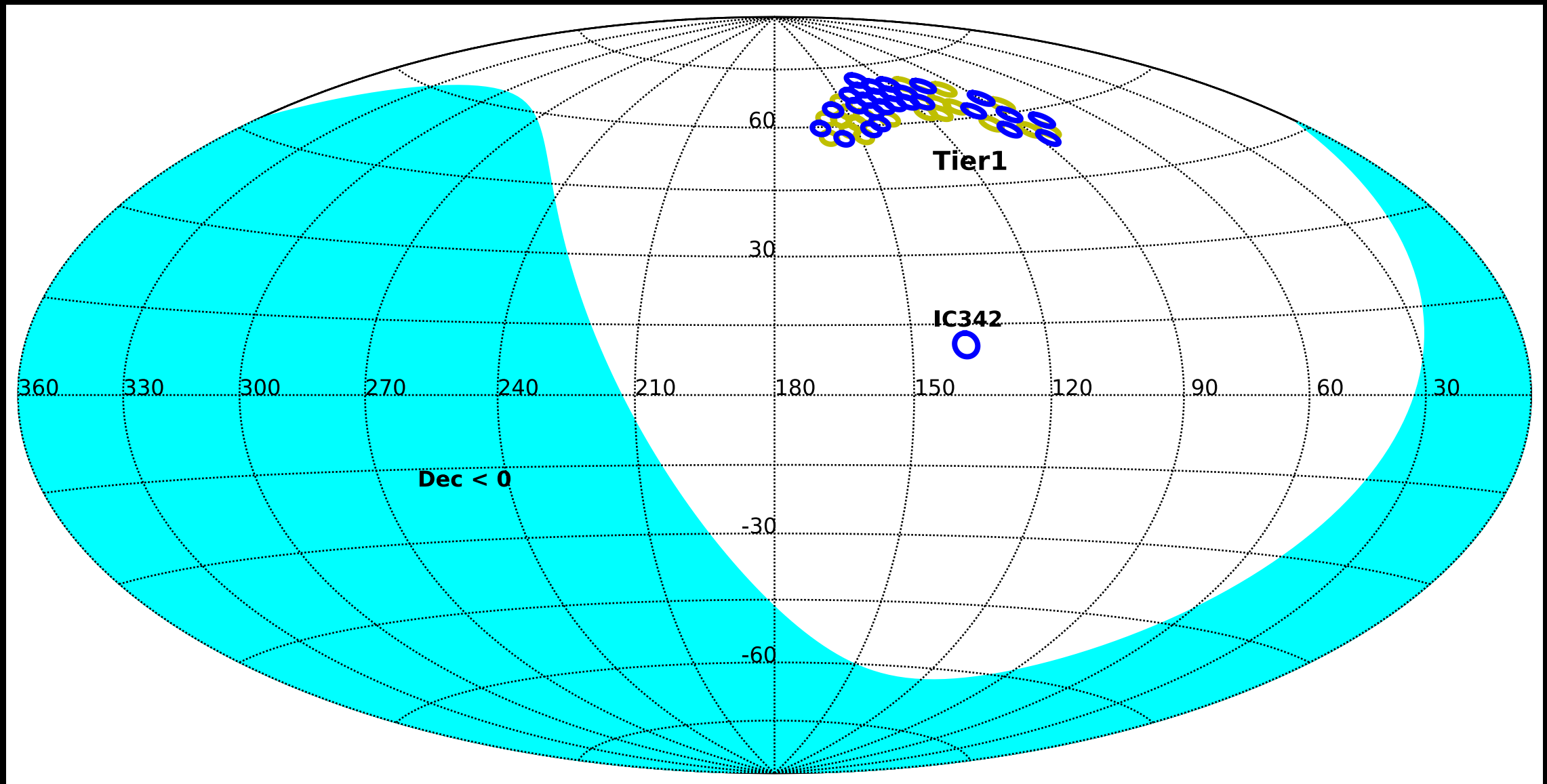
Depolarization

- Delta function: λ^0
- Tophat: λ^{-2}
- Triangle/Quadratic/etc: λ^{-4}
- Gaussian: $e^{-\sigma_\phi^2 \lambda^4}$

Requires volume with **no Faraday rotation**:

$$\phi \propto \int_0^d n_e \vec{B} \cdot d\vec{l}$$

Tier 1 Faraday Tomography



Faraday cubes made

Data available

Summary

- We're finding a wealth of **diffuse polarized structure** with LOFAR
- Faraday tomography has the same limitations as interferometry: **missing large scales**
- LOFAR may be most sensitive to emission from **neutral clouds**
- **Big thanks** to Tim Shimwell and Surveys KSP for sharing their Tier 1 data