

# IMAGING IONOSPHERIC STRUCTURES

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on behalf of the LOFAR-EOR team

# Introduction

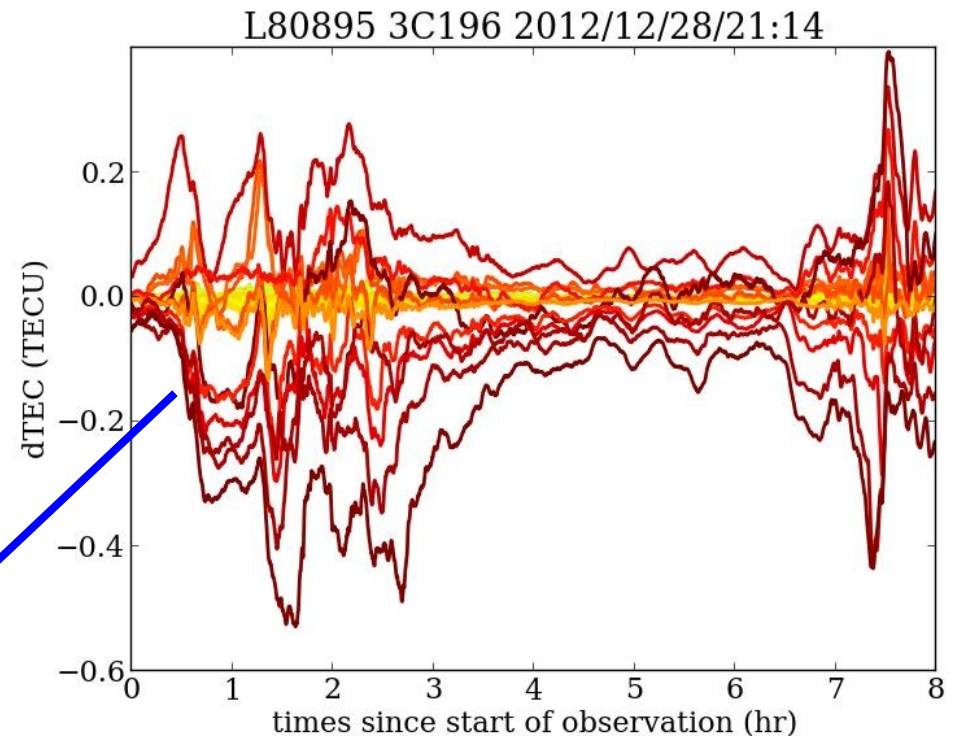
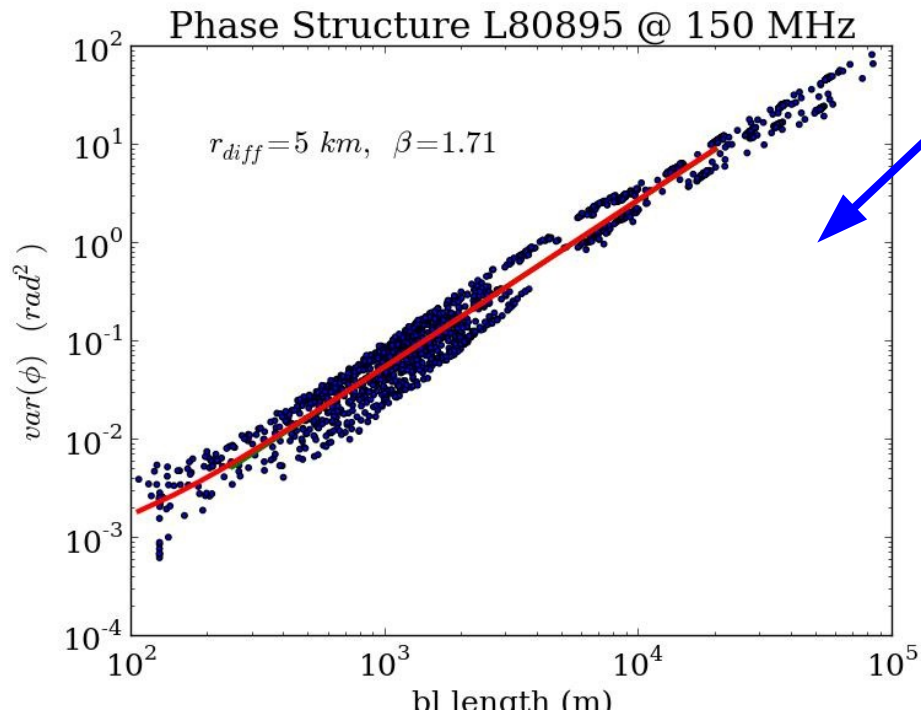
- Gathering ionospheric information from LOFAR data:
  - using calibration phases
  - snapshot images
- Earth magnetic field aligned structures in the ionosphere
  - visible using both methods
- Use snapshot images to visualize moving structures
- all data: LOFAR EOR 3C196 data 2012/2013

# Ionospheric Structure

Calibration phases

large frequency range  $\rightarrow 1/\nu$   
dependence  $\rightarrow$  ionospheric  
phases

time average  $\rightarrow$  phase  
structure function



3C196 observation. dTEC vs. time

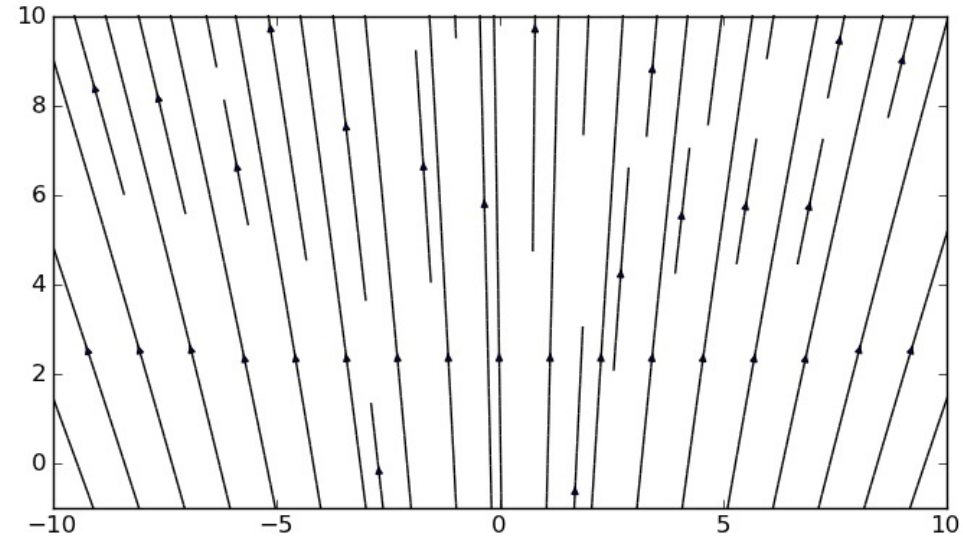
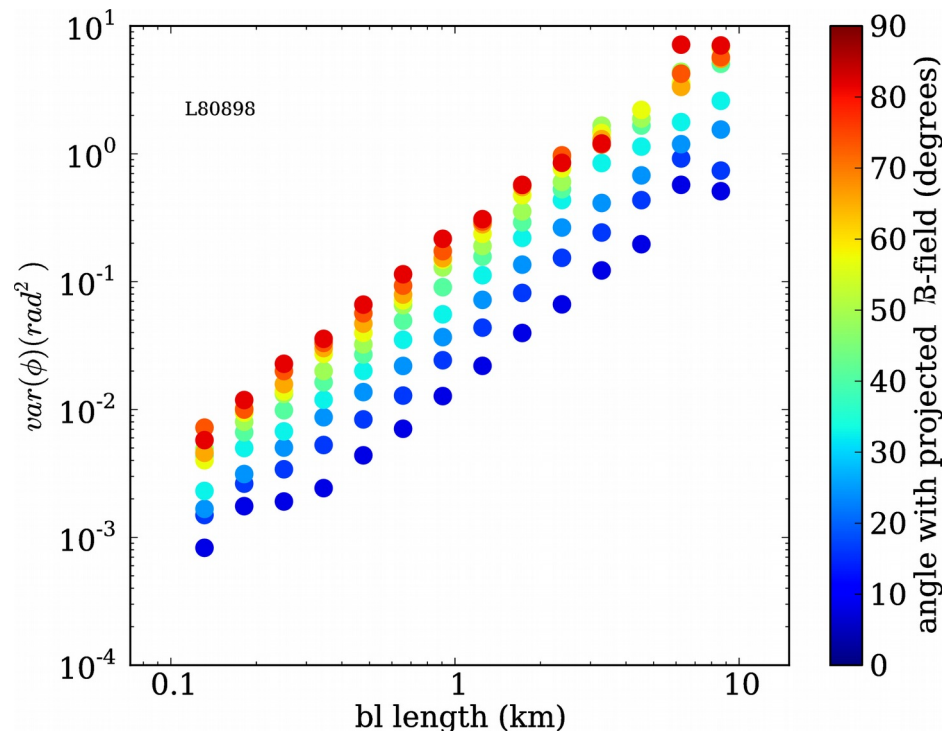
results on 29 EOR 3C196  
observations using this technique  
submitted to Radio Science

# Field Aligned Structure

*bandlike* structure → orientation of the baseline

Earth magnetic field aligned?

projected field lines along LOS  
single ionospheric height



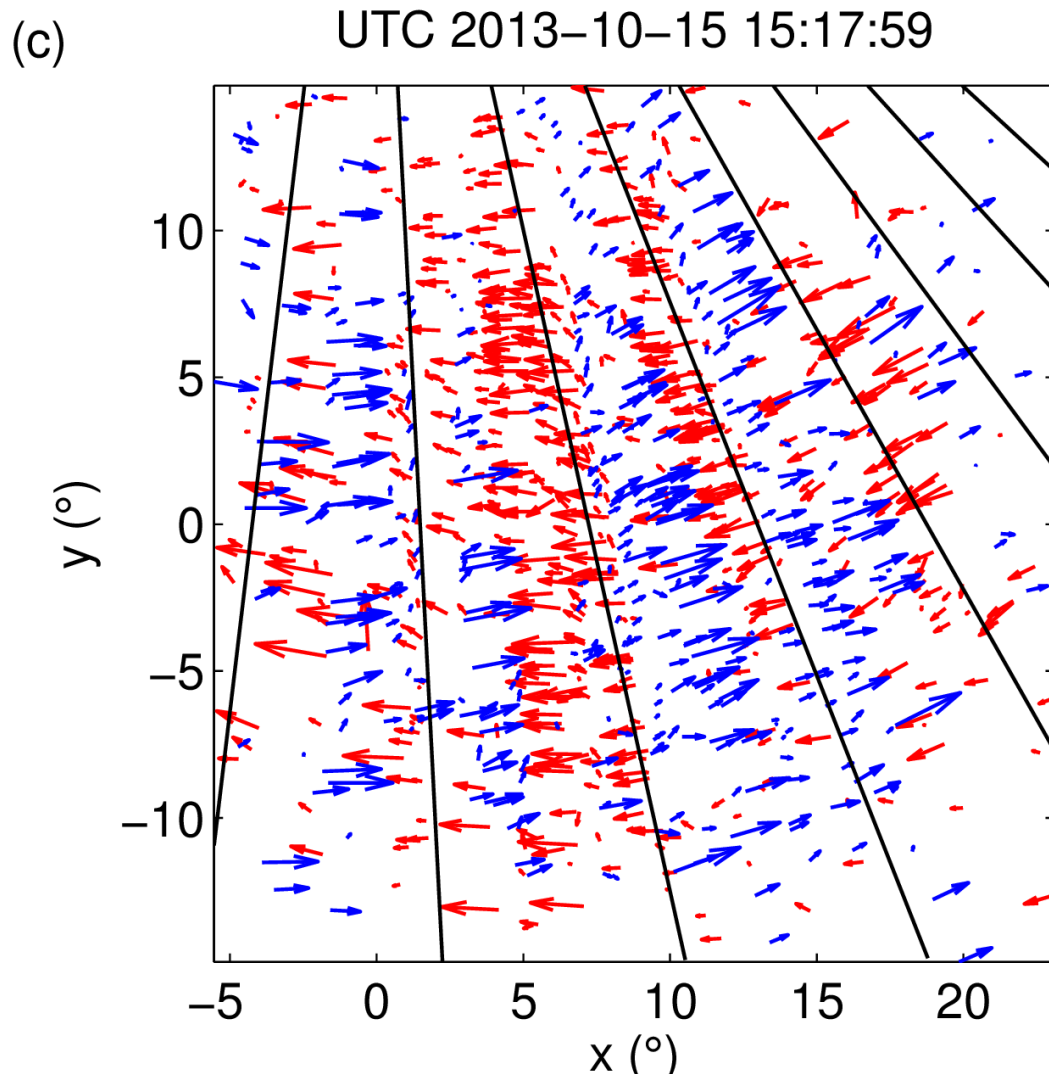
Earth magnetic field lines along viewing angle (NS/EW beam angle, zenith = 0,0)

perspective view → time dependent orientation  
bin data in according to angle  
wrst projected field lines  
field aligned structure observed in ~ 50 % of the observations

Earth magnetic field : WMM

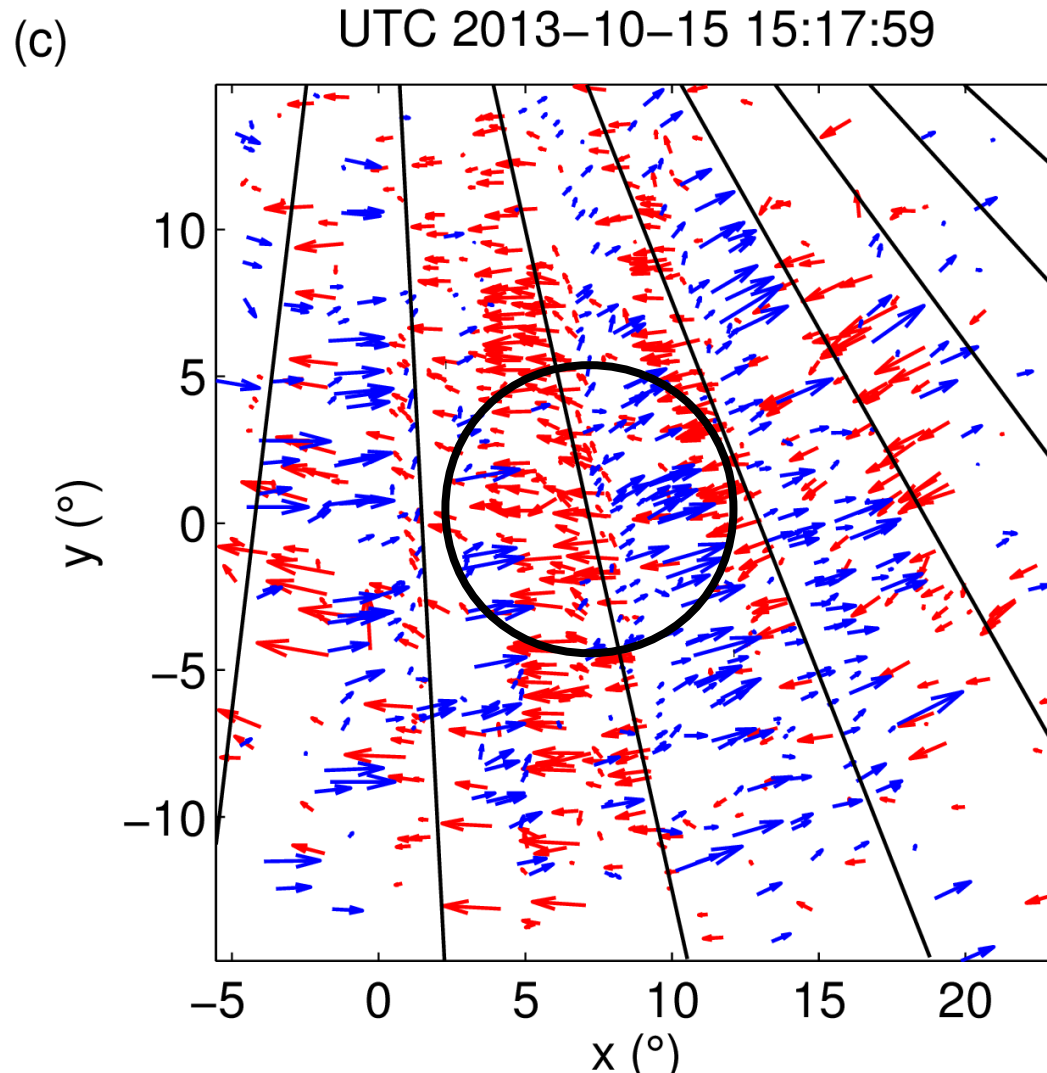
# Loi et al. GRL 2015

## *Real-time imaging of density ducts between the plasmasphere and ionosphere*



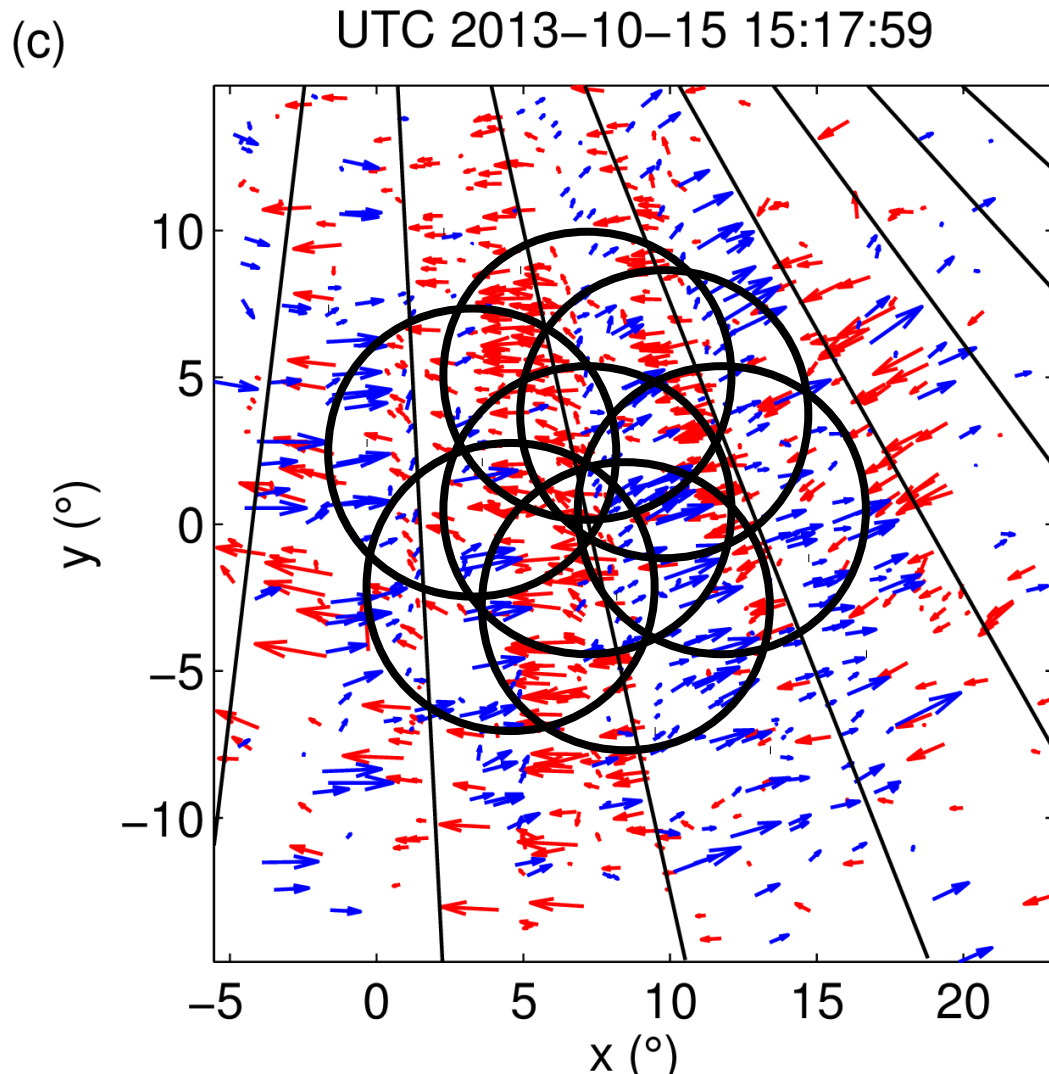
MWA data  
snapshot images of  
source shifts  
ionospheric gradient  
→ position shift  
elongated slowly  
moving field aligned  
structures

# HBA beam



Do we observe the same structures with LOFAR?  
single beam: too small FOV

# flanking fields



Do we observe the same structures with LOFAR?  
single beam: too small FOV  
multiple beams  
standard observing mode for EOR: 1 central beam + 6 flanking fields  
18 SB each

# Imaging Structures

ionosphere: linear gradient → position shift

higher order term → distorted source

use only short baselines: **CS only**

$$\Delta\theta = c/v^2 \nabla_{\perp} \text{TEC}$$

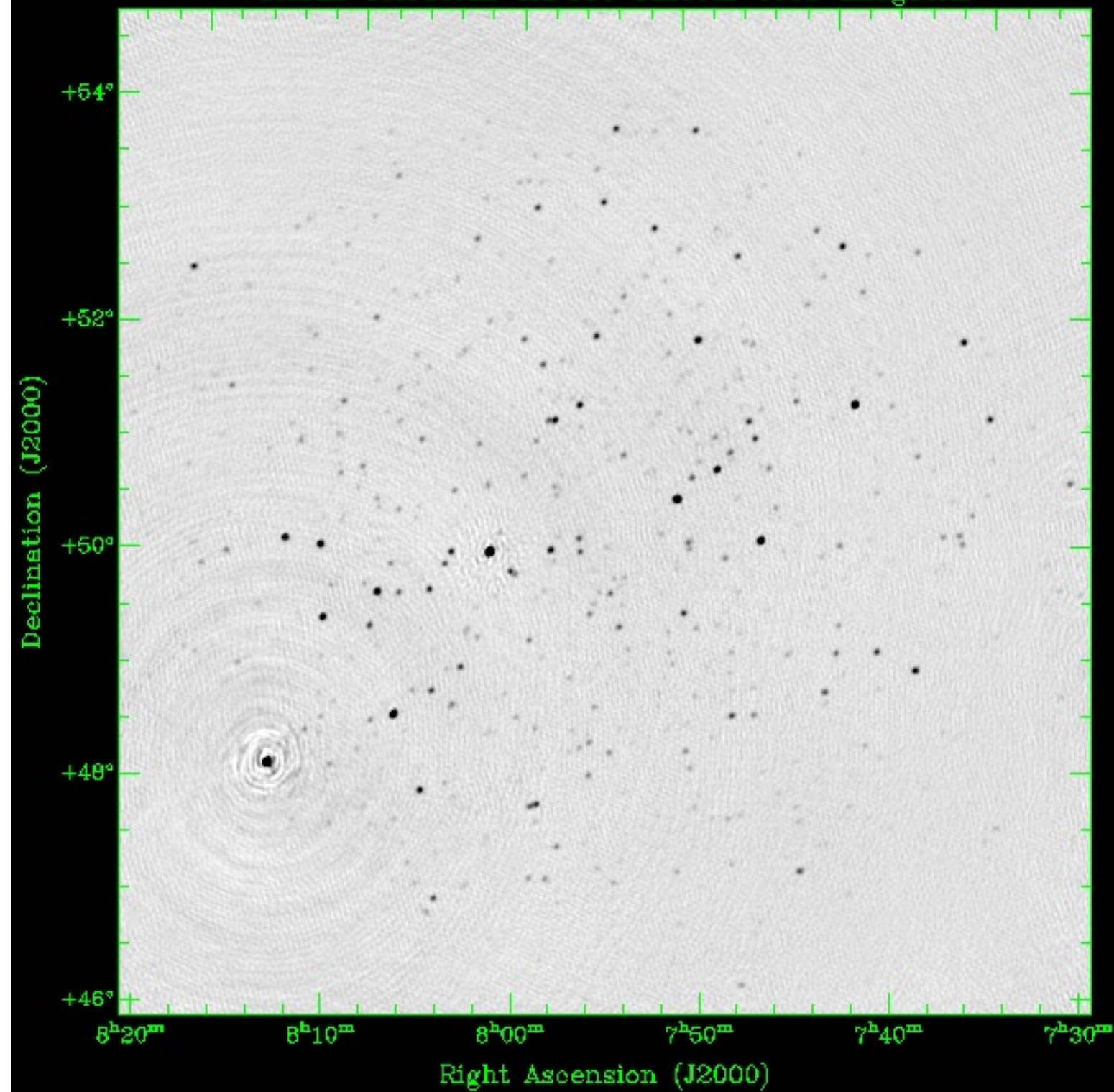
I. correct all with calibration gains of central field

- **subtract 3C196 from central field**

II. image corrected data (**wsclean**) → extract sources for reference (**pybdsf**)



wsclean→correctall→SAP004→subset3-0000-image.fits



wsclean:  
combine SBs  
to create 3  
images with  
different  
frequency

~ 4 SB each,  
due to  
missing files

# Imaging Structures

ionosphere: linear gradient → position shift

higher order term → distorted source

use only short baselines: **CS only**

$$\Delta\theta = c/v^2 \nabla_{\perp} \text{TEC}$$

I. correct all with calibration gains of central field

- **subtract 3C196 from central field**

II. image corrected data (**wsclean**) → extract sources for reference (**pybdsf**)

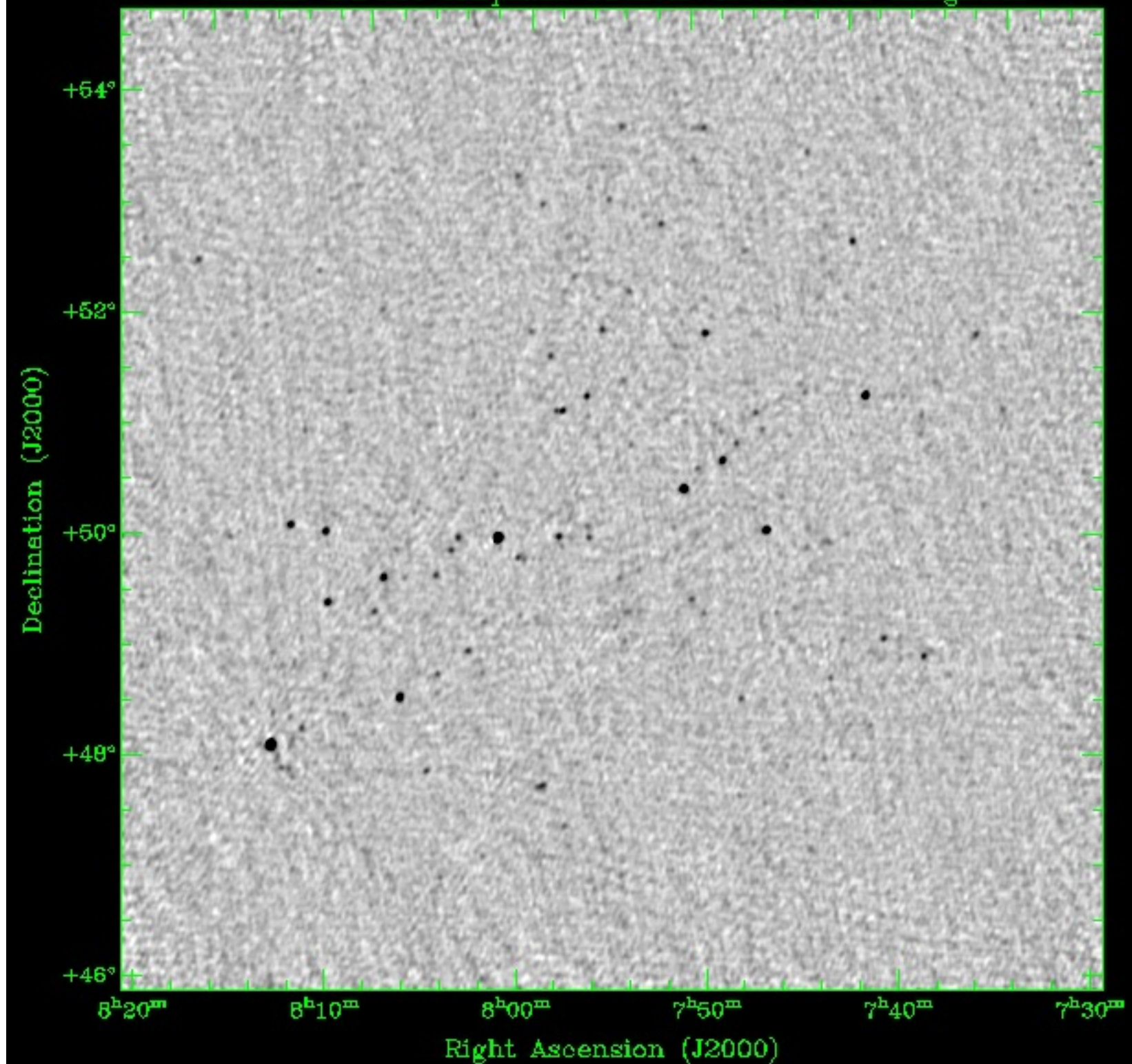
III. remove TEC-phases from calibration gains, correct all

# Imaging Structures (2)

IV. 1 minute snapshot images ([wsclean](#))

- 3 frequencies

wsclean→SAP004→snapshot→subset3-t0101-0000-image.fits



1 minute  
snapshot  
image

SAP004

L86767

# Imaging Structures (2)

IV. 1 minute snapshot images ([wsclean](#))

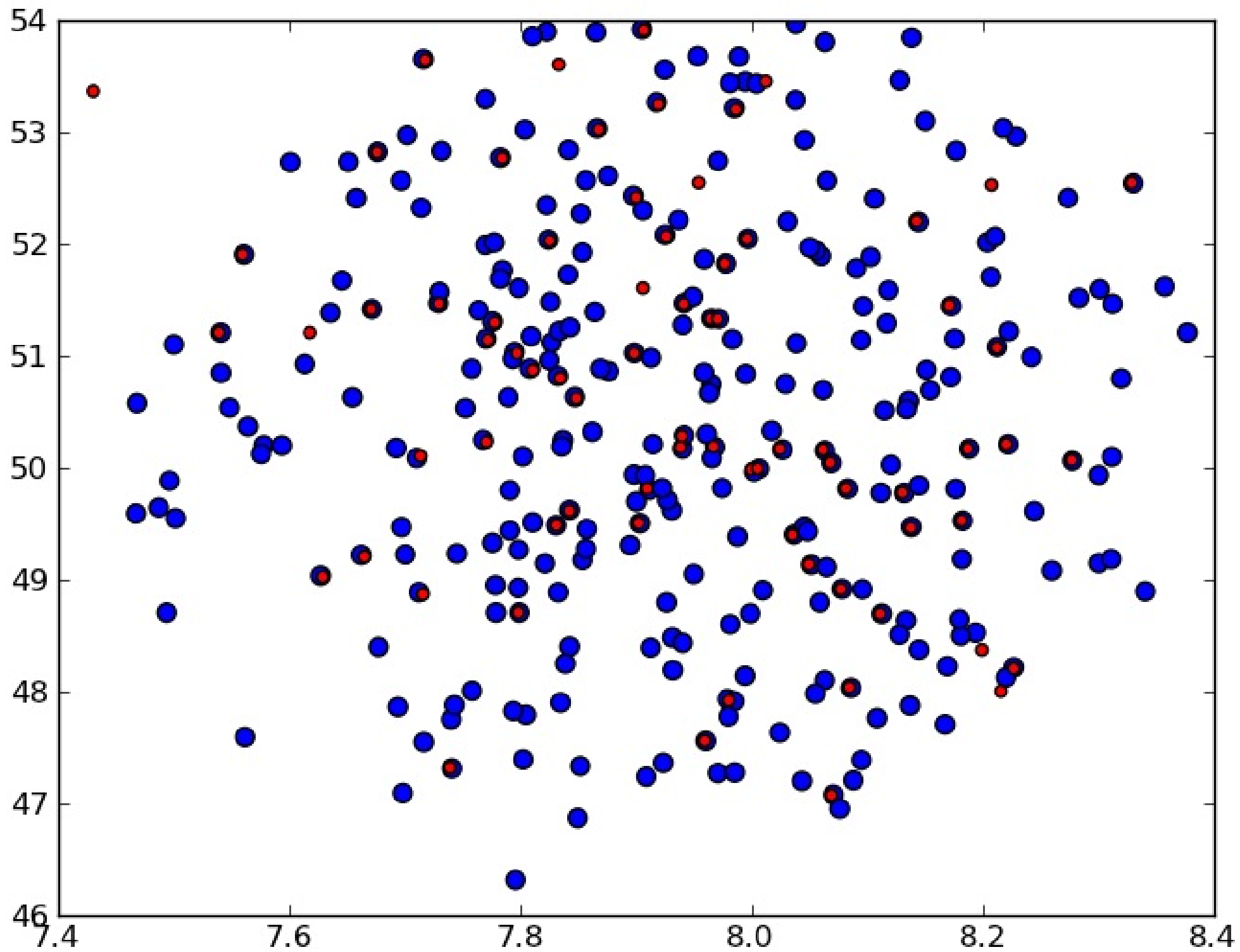
- 3 frequencies

V. source extraction ([pybdsf](#))

- default values
- about 400 sources (including double matches)

VI. match with reference sources





# Imaging Structures (2)

IV. 1 minute snapshot images ([wsclean](#))

- 3 frequencies

V. source extraction ([pybdsf](#))

- default values
- about 400 sources (including double matches)

VI. match with reference sources

VII. image source position shifts

[Selected 5 observations:](#)

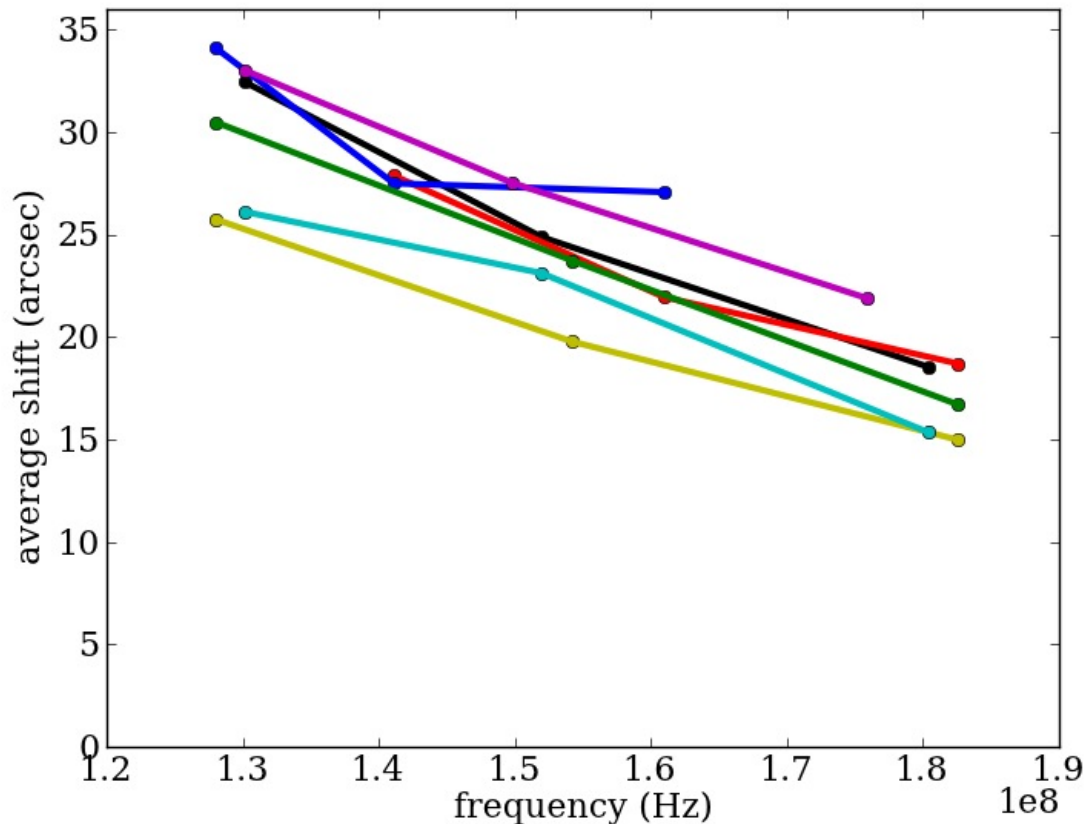
[4 with field aligned structure function](#)

[1 other](#)

[processed on new eor cluster dawn](#)

# frequency dependence

average position shift bright sources



images were dividing  
all SBs in 3 groups,  
using the

*-joinchannels*

*-channelsout*

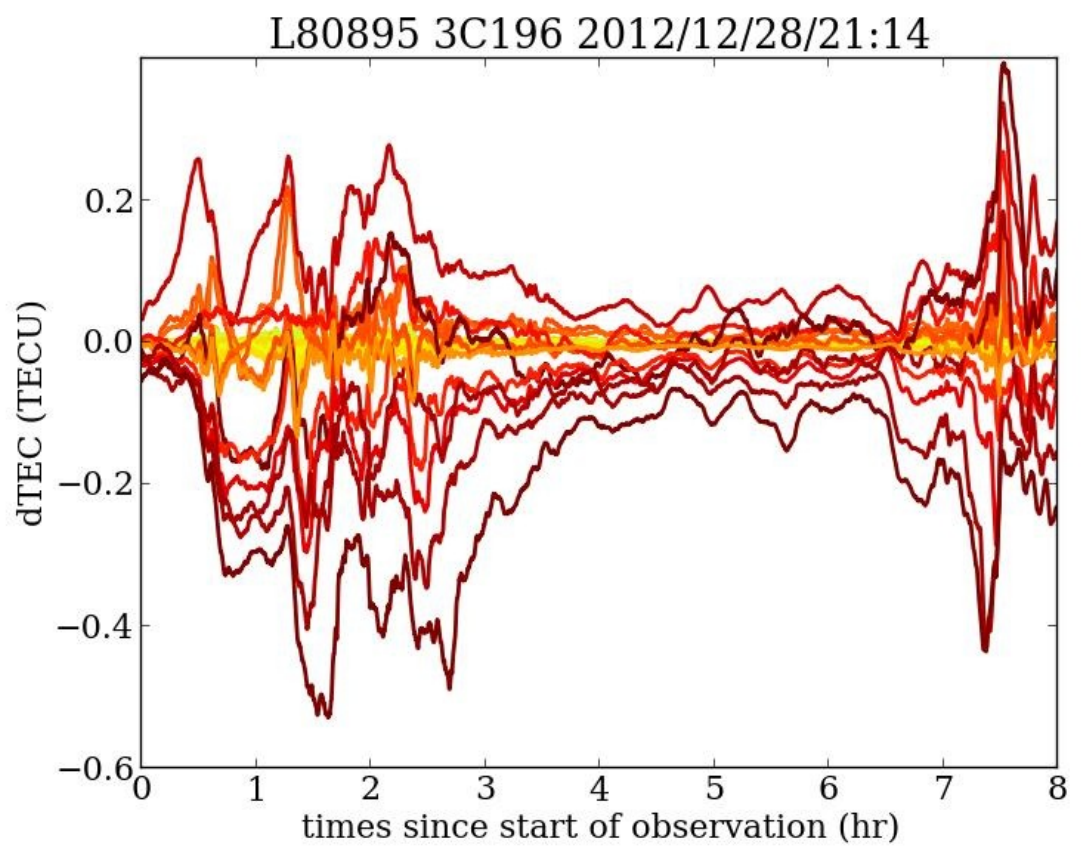
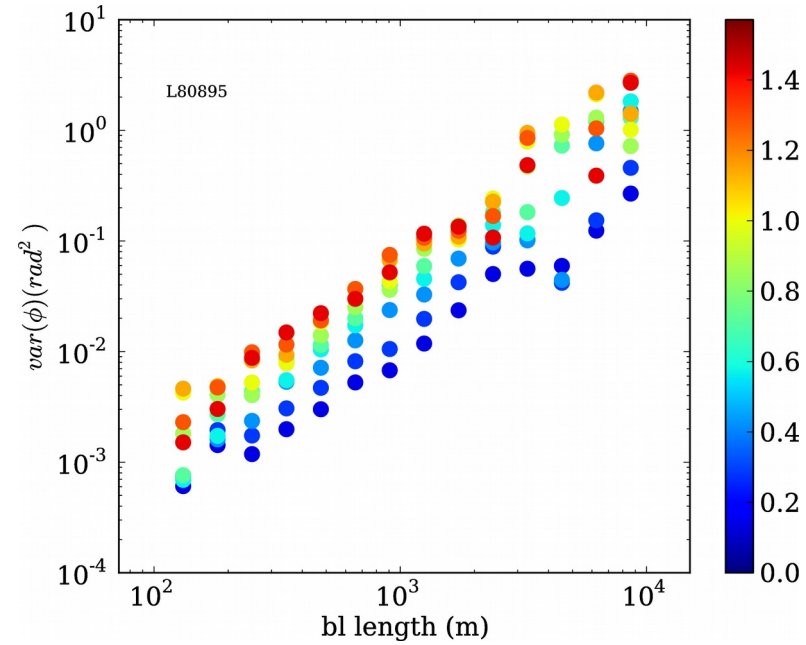
options in wsclean

missing data:

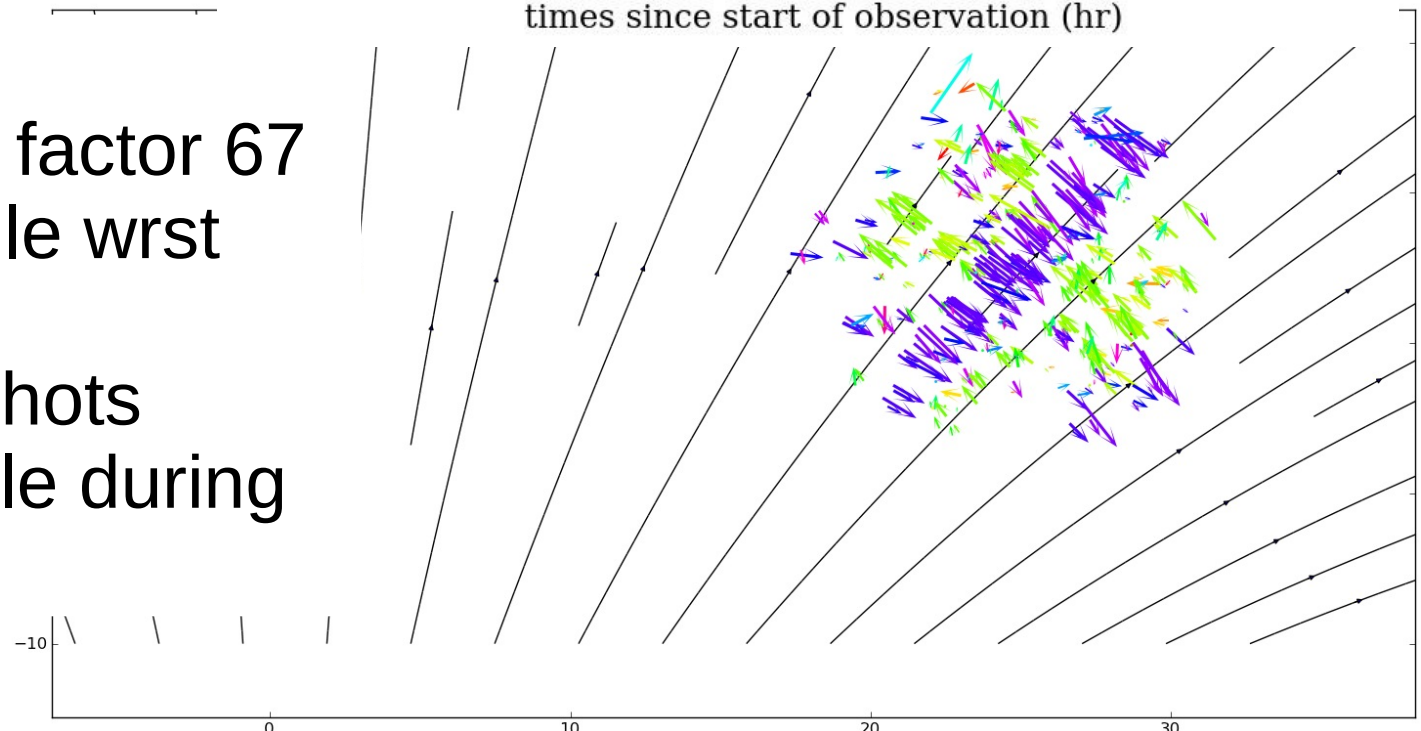
~4SB per image

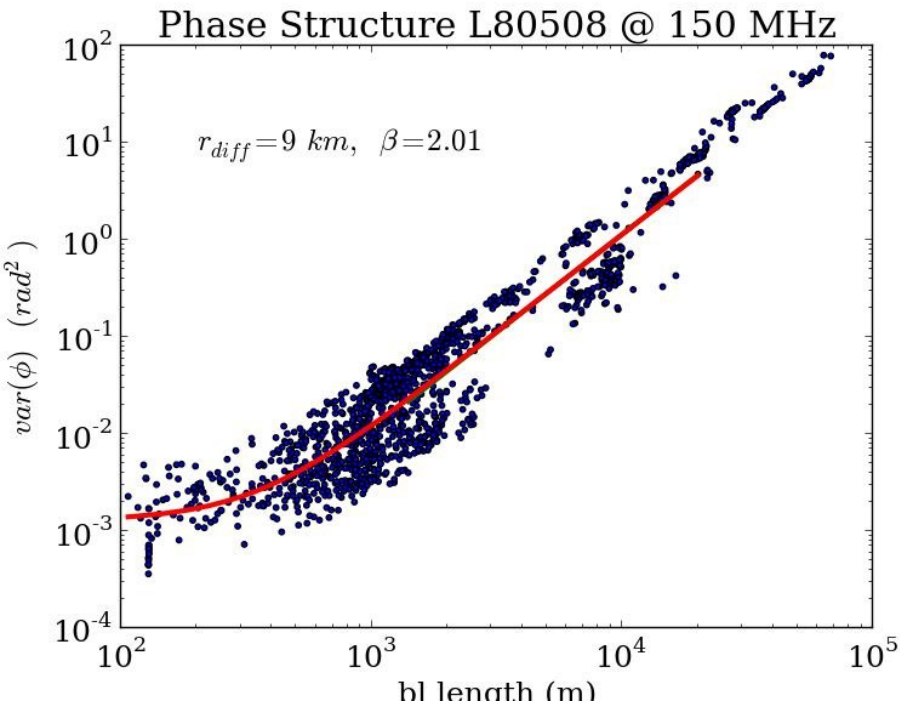
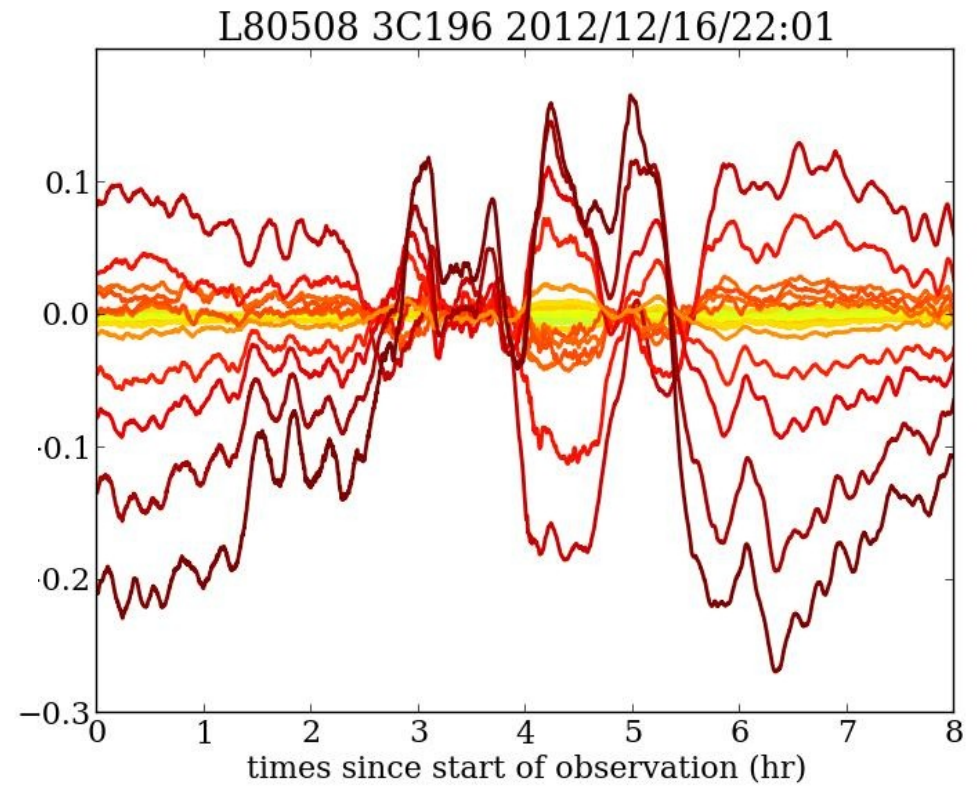
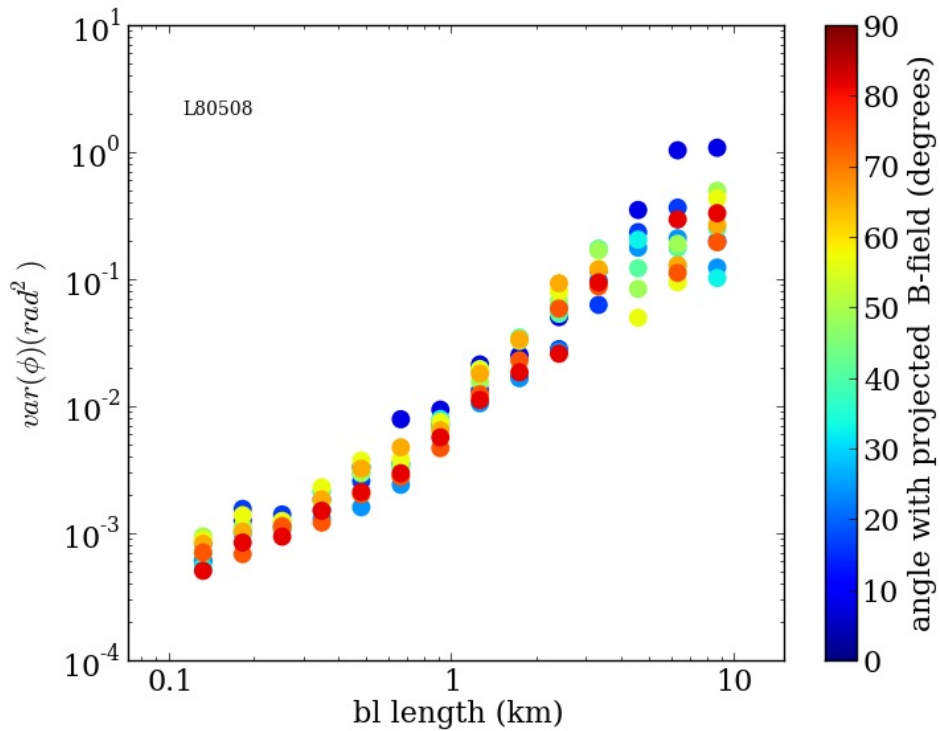
ionosphere: shift  $\sim 1/\nu^2$





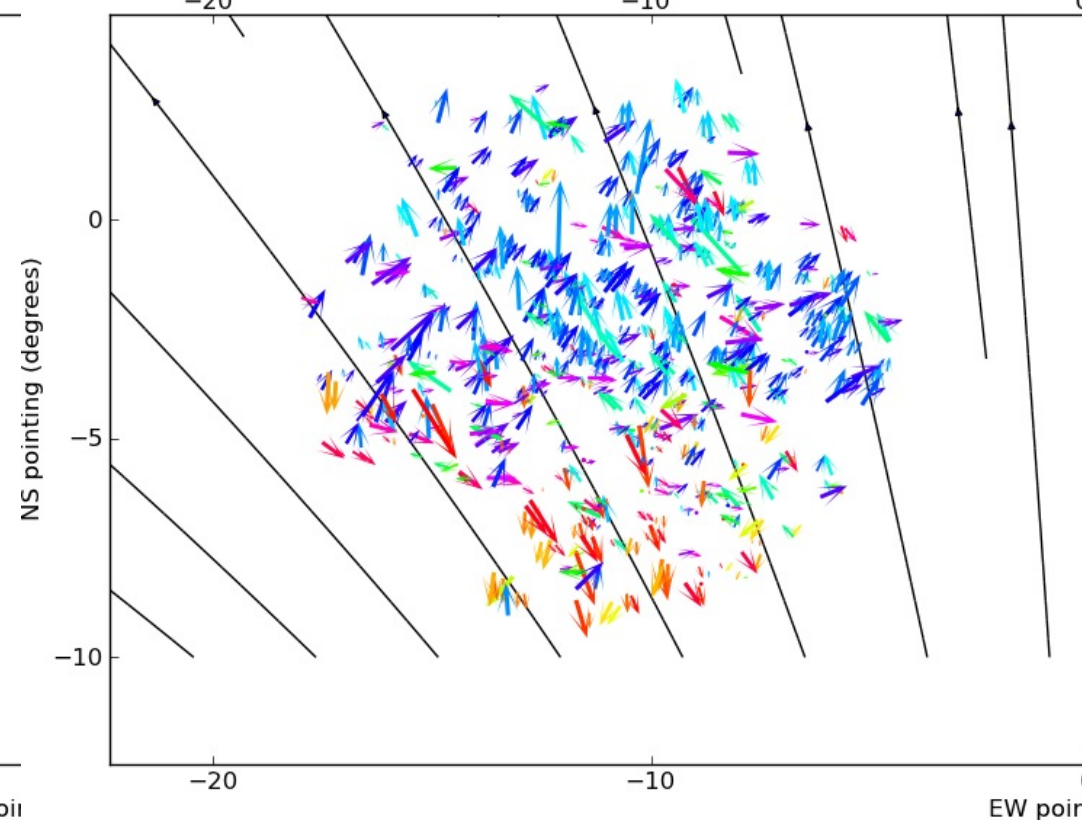
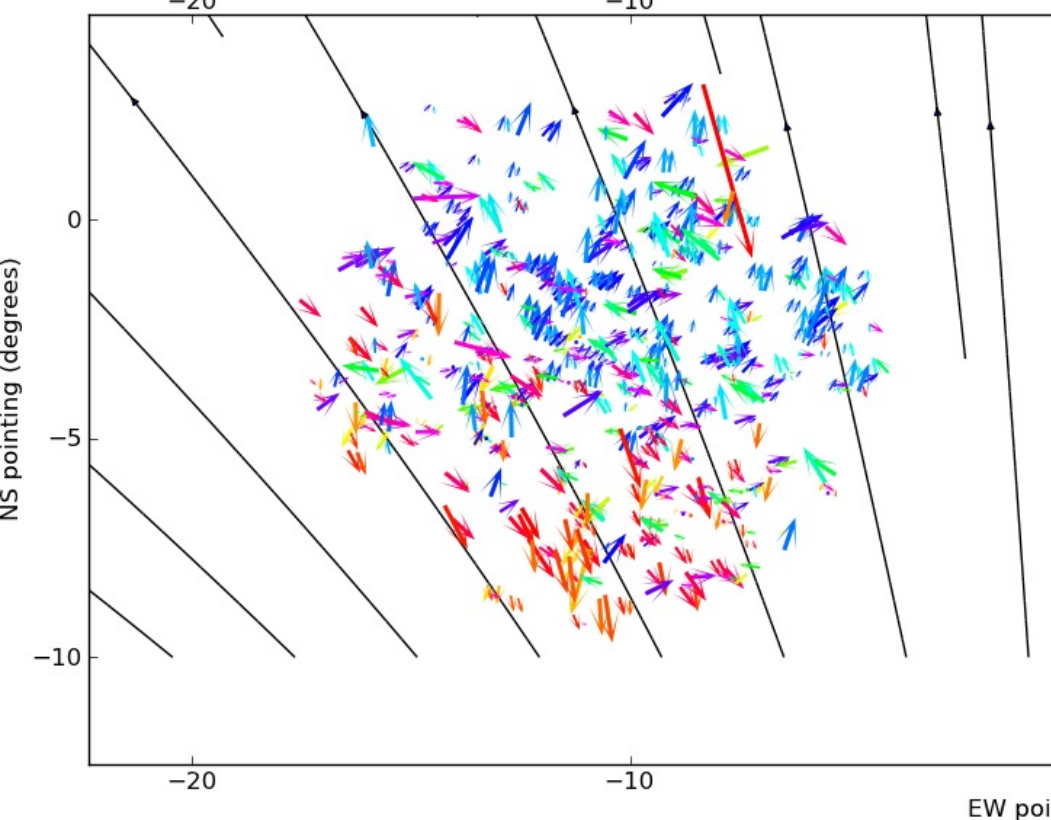
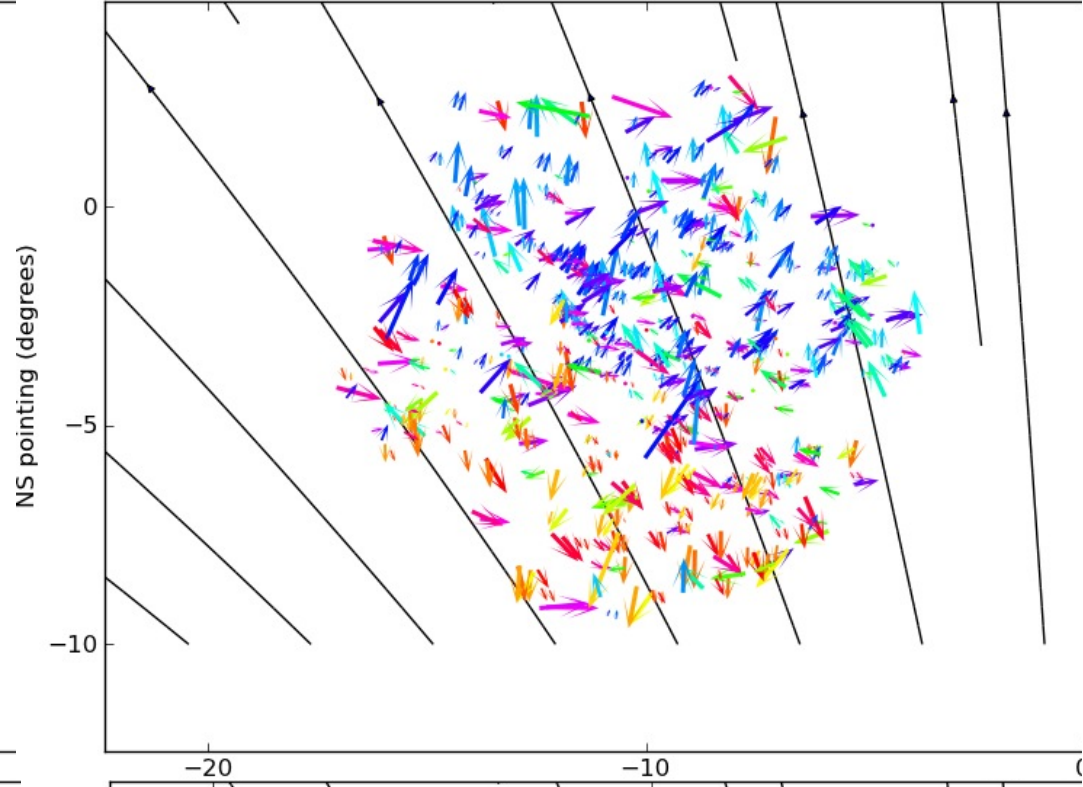
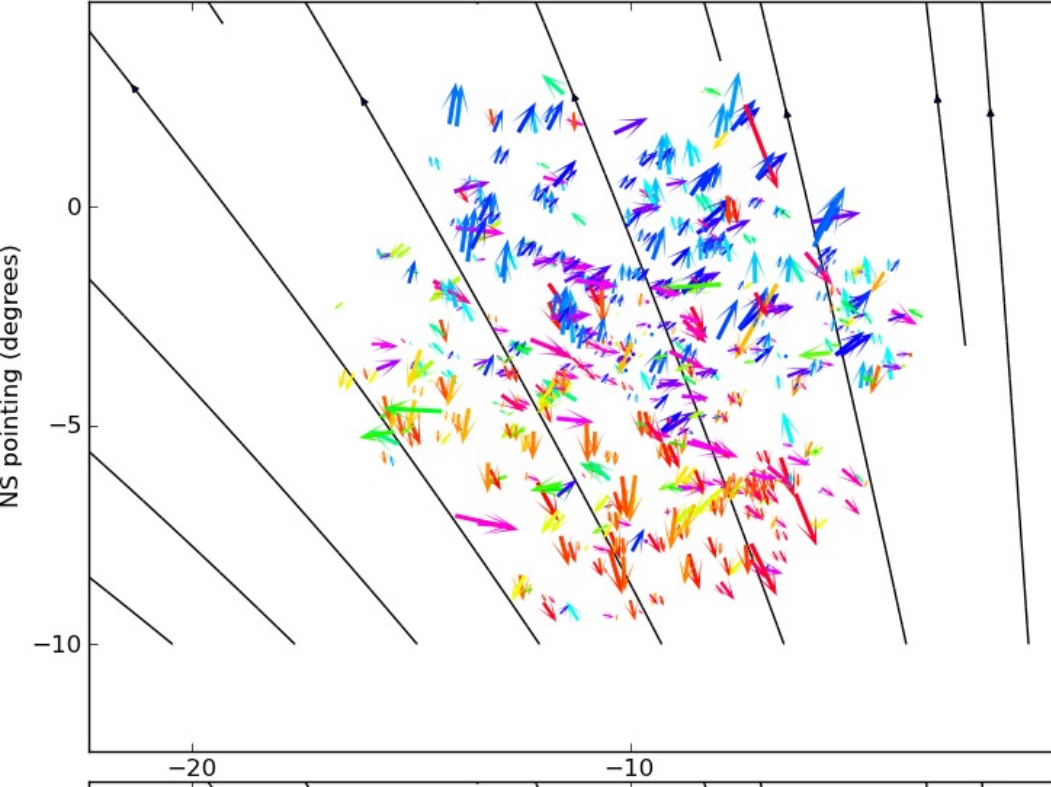
L80895  
 arrows scaled with factor 67  
 color indicates angle wrst  
 local field lines  
 2 (1 minute) snapshots  
 structure only visible during  
 first hour

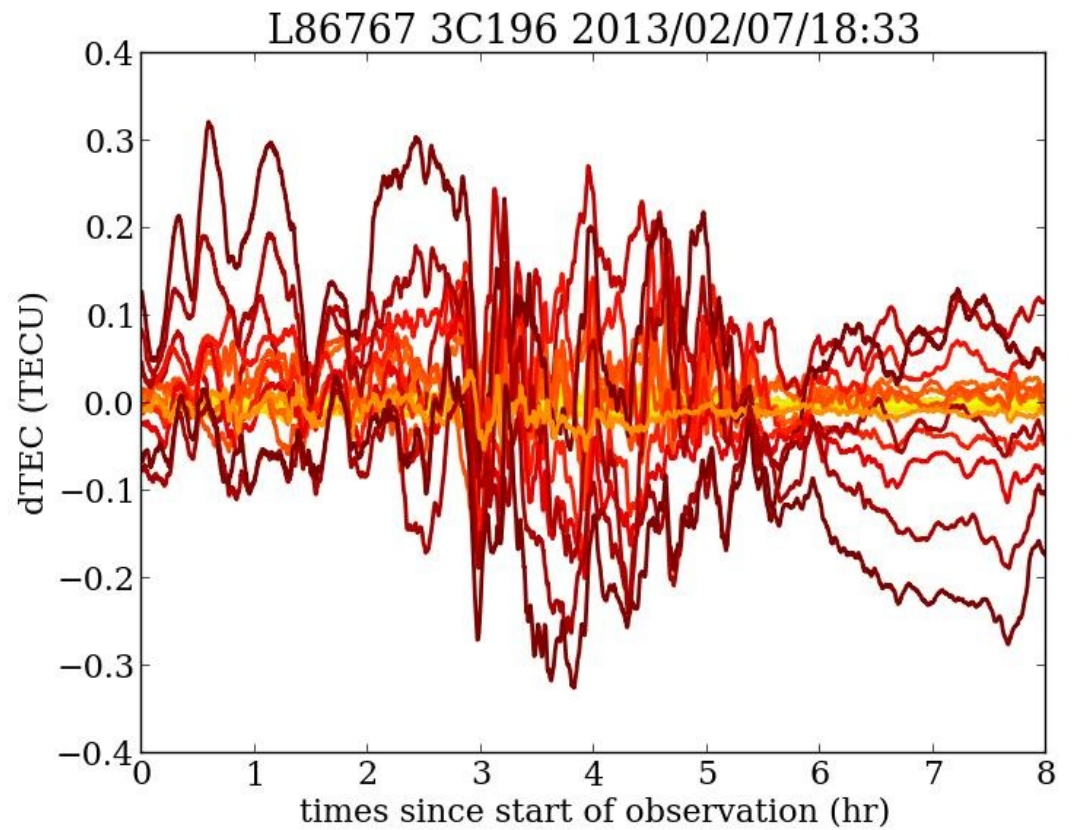
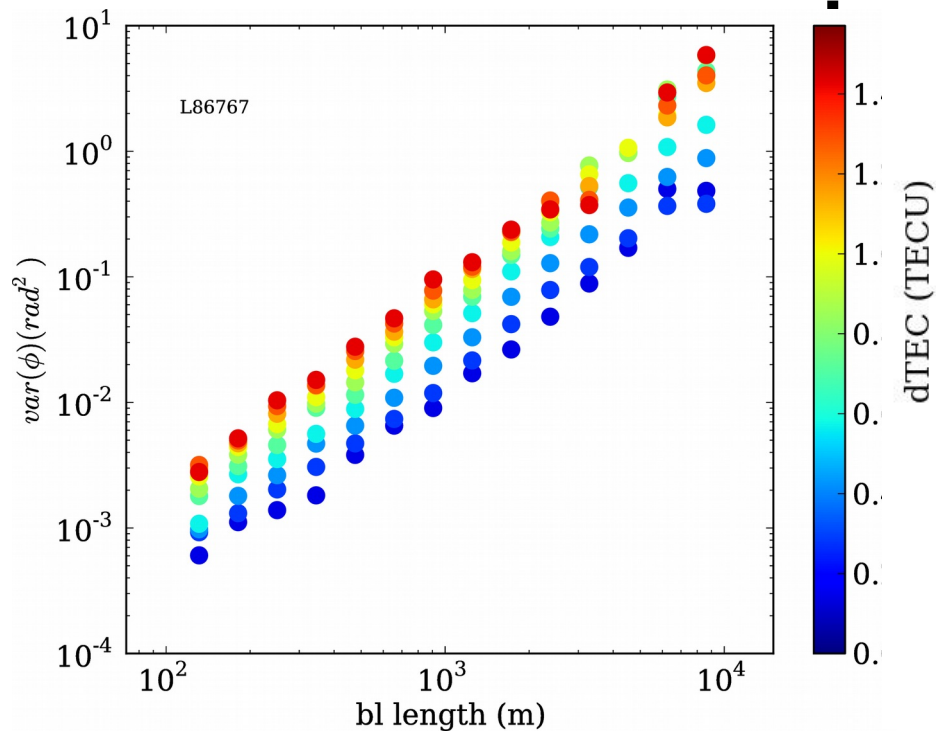




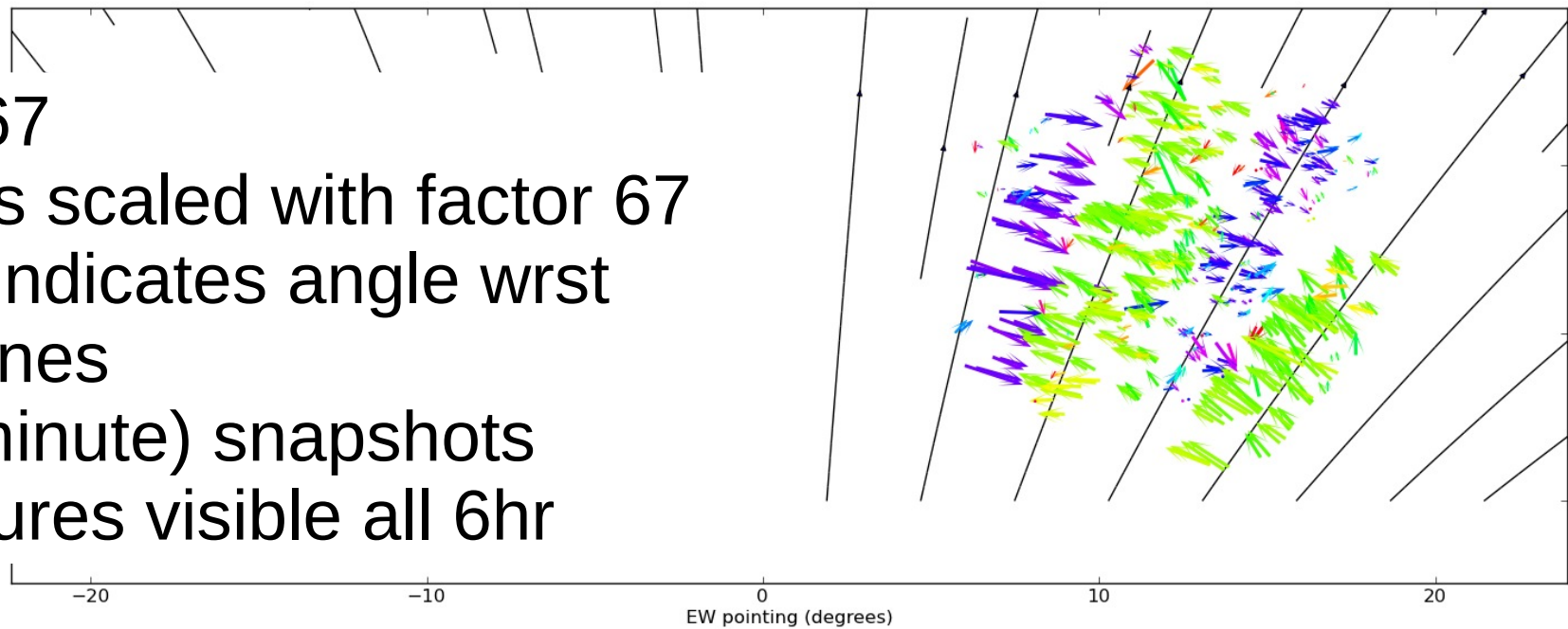
NO field aligned structures  
 larger scale

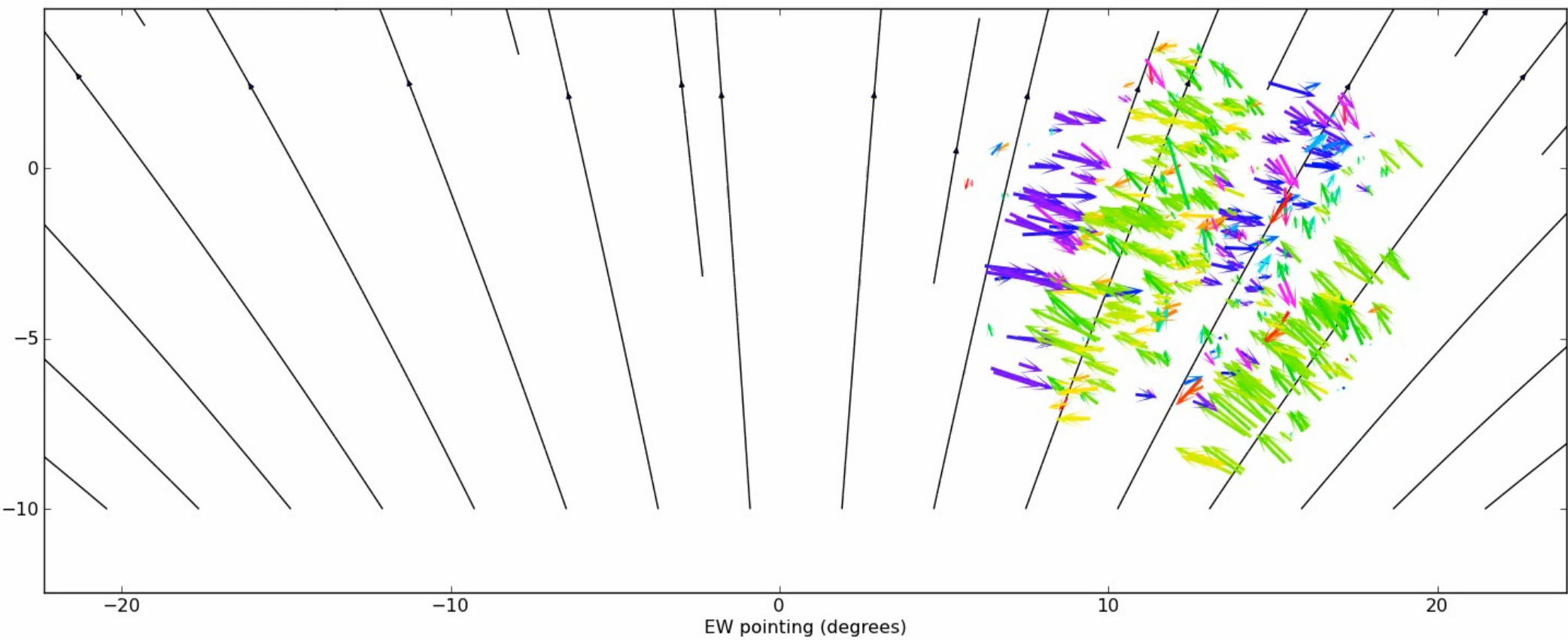






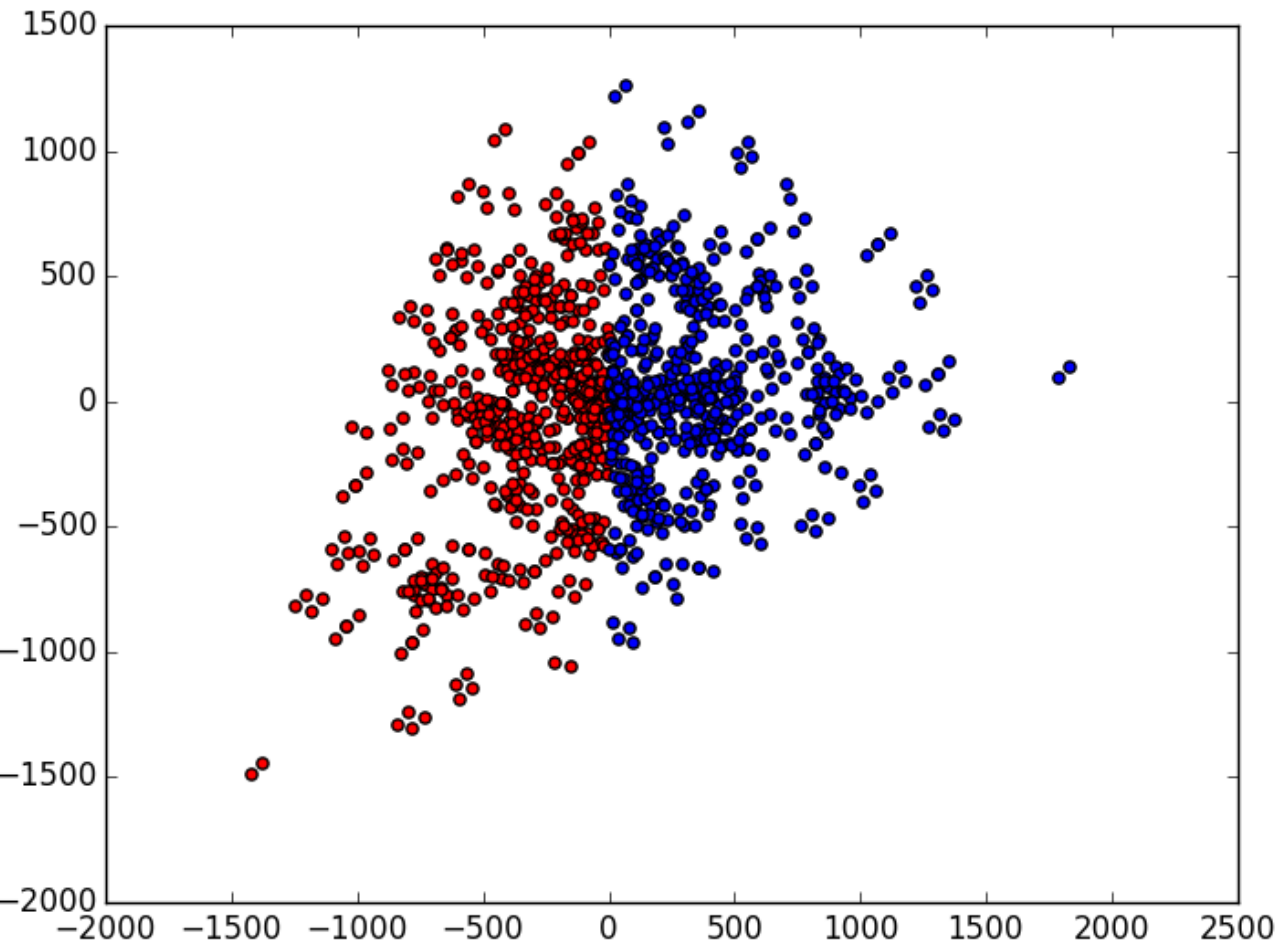
L86767  
 arrows scaled with factor 67  
 color indicates angle wrst  
 field lines  
 2 (1 minute) snapshots  
 structures visible all 6hr







# measure height via parallax



divide baselines in E  
and W

E-W=730 m

create  
snapshotimages with  
both

cross correlate to get  
parallax shift

ongoing

# Conclusion and Outlook

- Two ways of visualizing ionosphere:
  - TEC phases from calibration solutions
    - all baselines
  - source position shifts in snapshot images
    - short baselines
- Earth magnetic field aligned structures observed in structure function of  $\sim 50\%$  of the observations
- density ducts?
  - clear structures in snapshot movies of 1 out of 4 (2 out of 5)
  - hints in short intervals of 2 more observations
- height via parallax measurement ongoing:
  - use nearby remote stations for better stereo angle