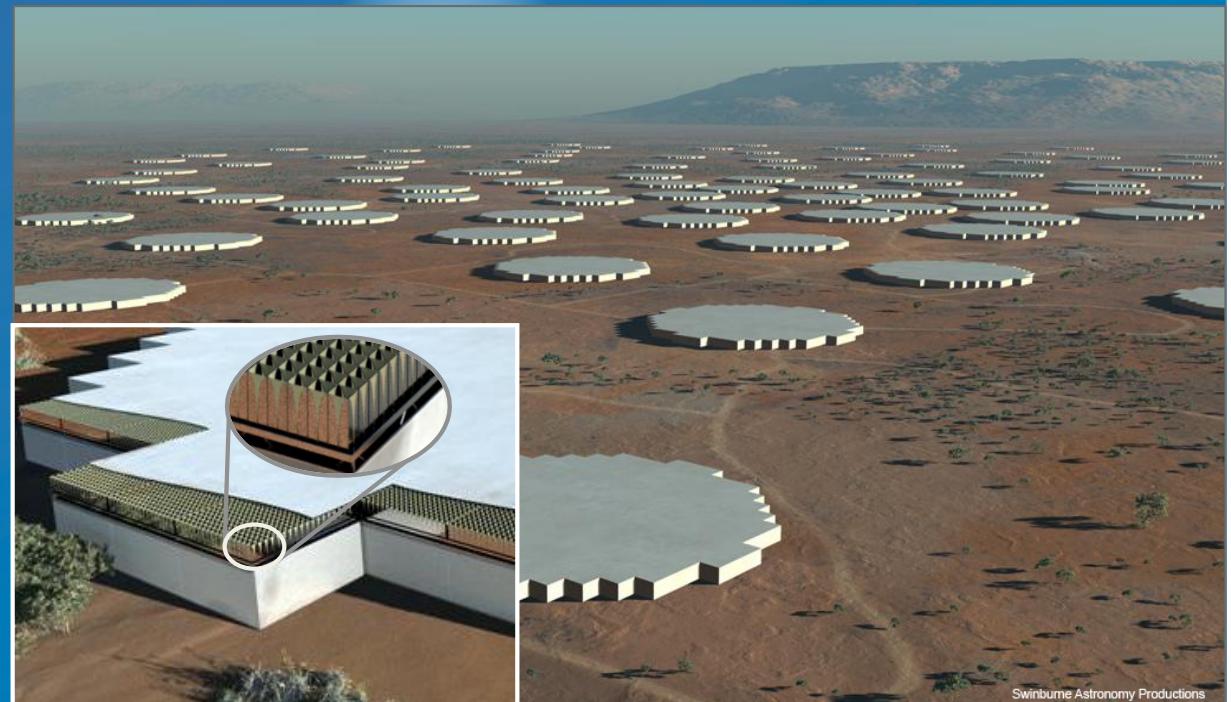


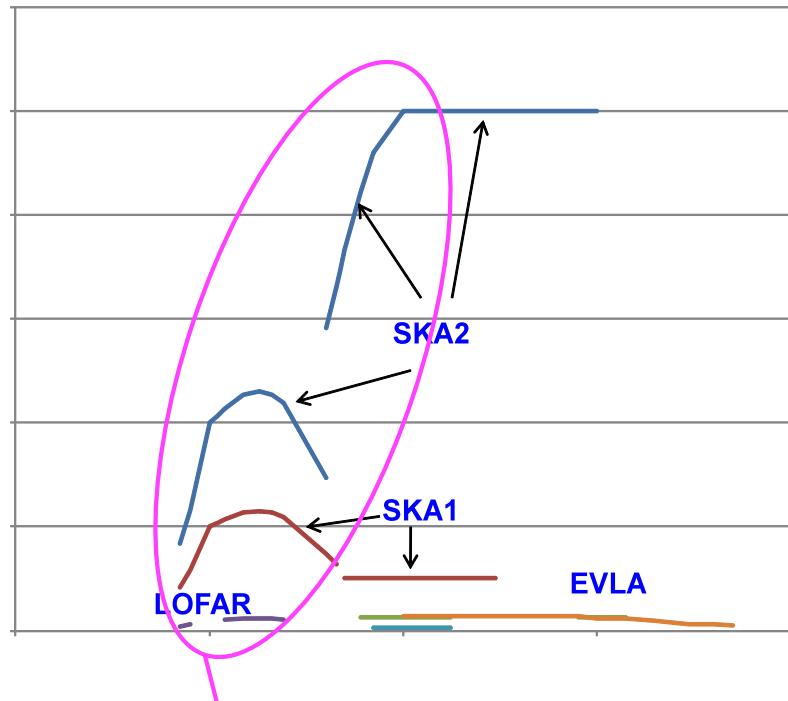


SKA Aperture Arrays

Jan Geralt Bij de Vaate ASTRON
Ilse van Bemmel



Sensitivity Comparison

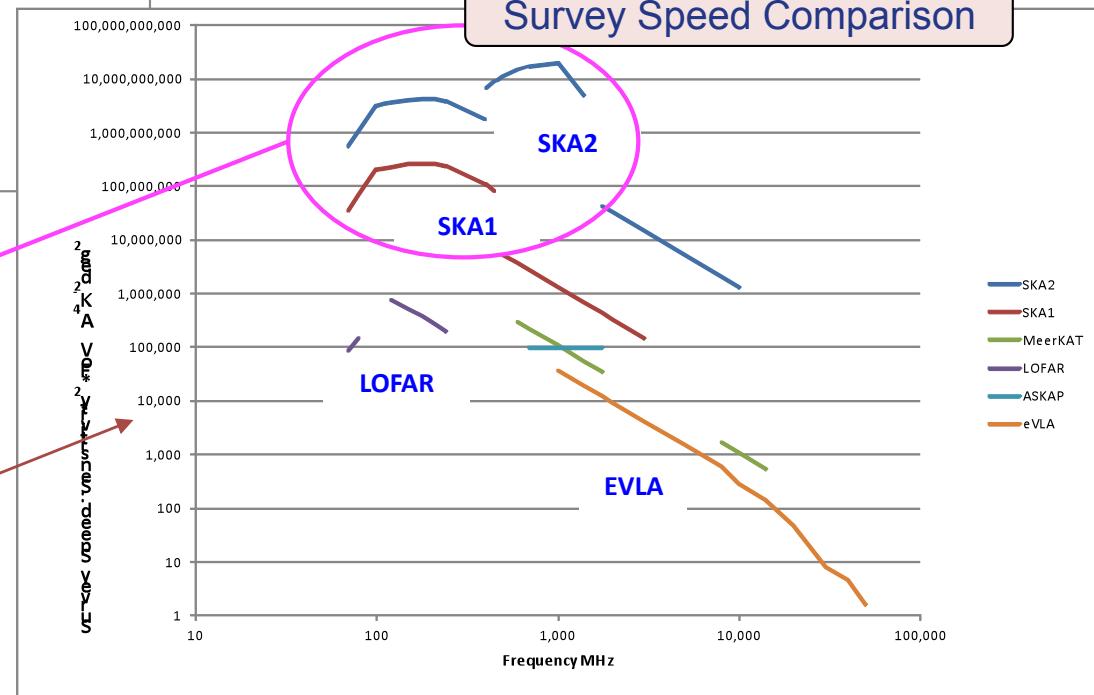


**Aperture
Arrays**

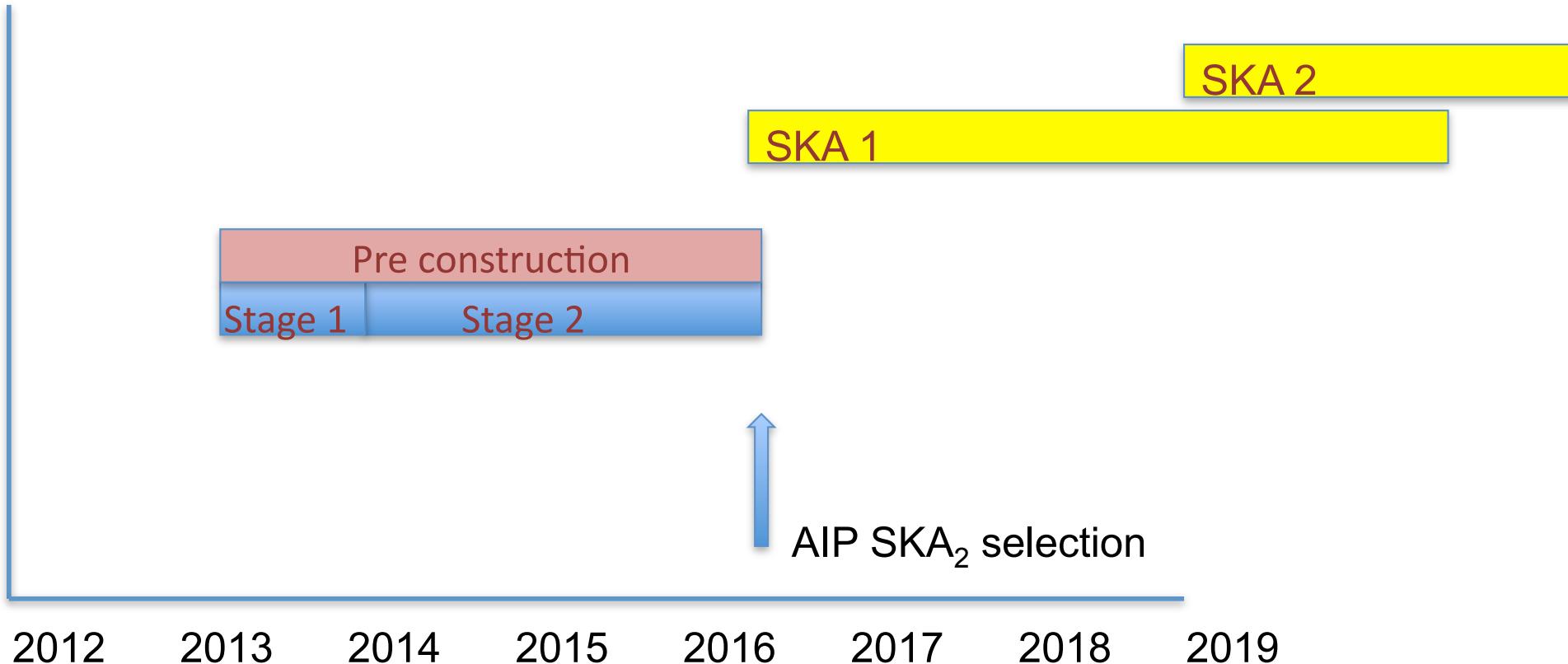
Note: log scale!

SKA₁ & SKA₂ will have much higher sensitivity & survey speed than existing instruments

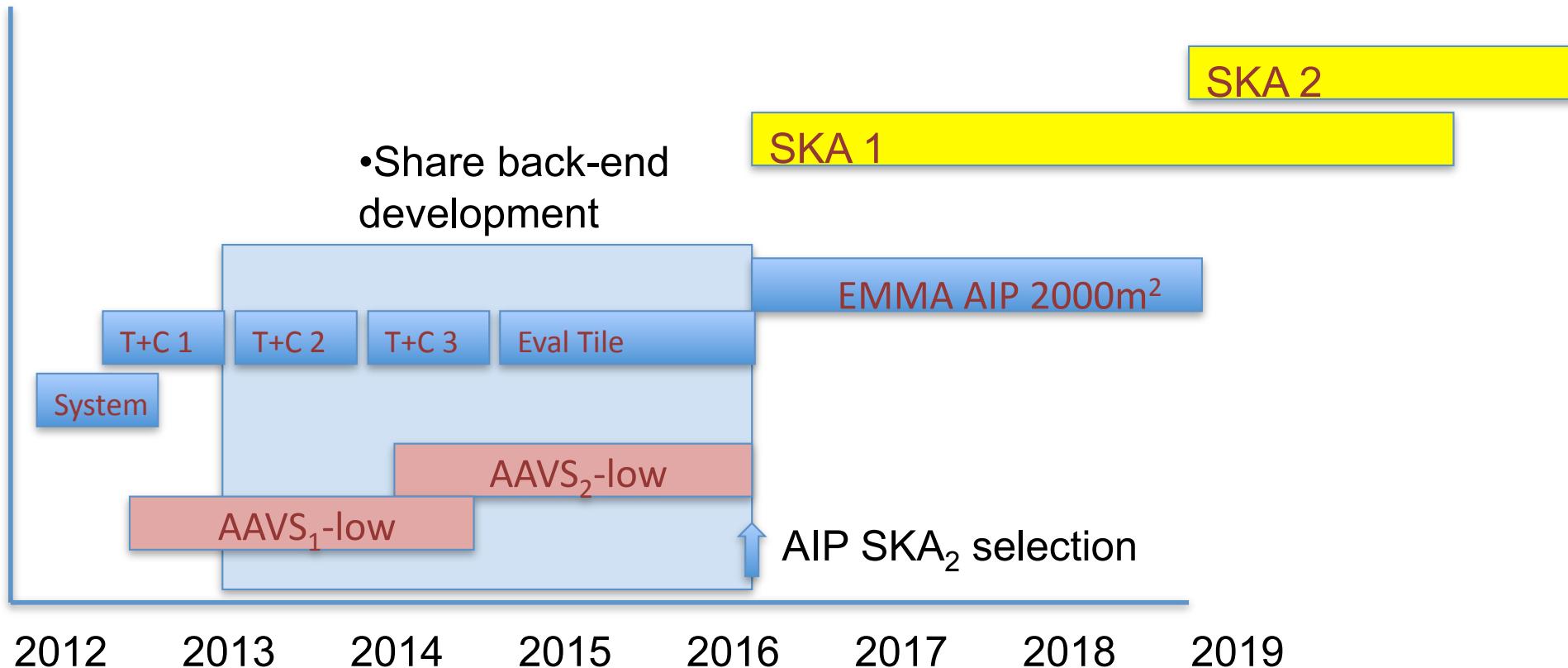
Survey Speed Comparison



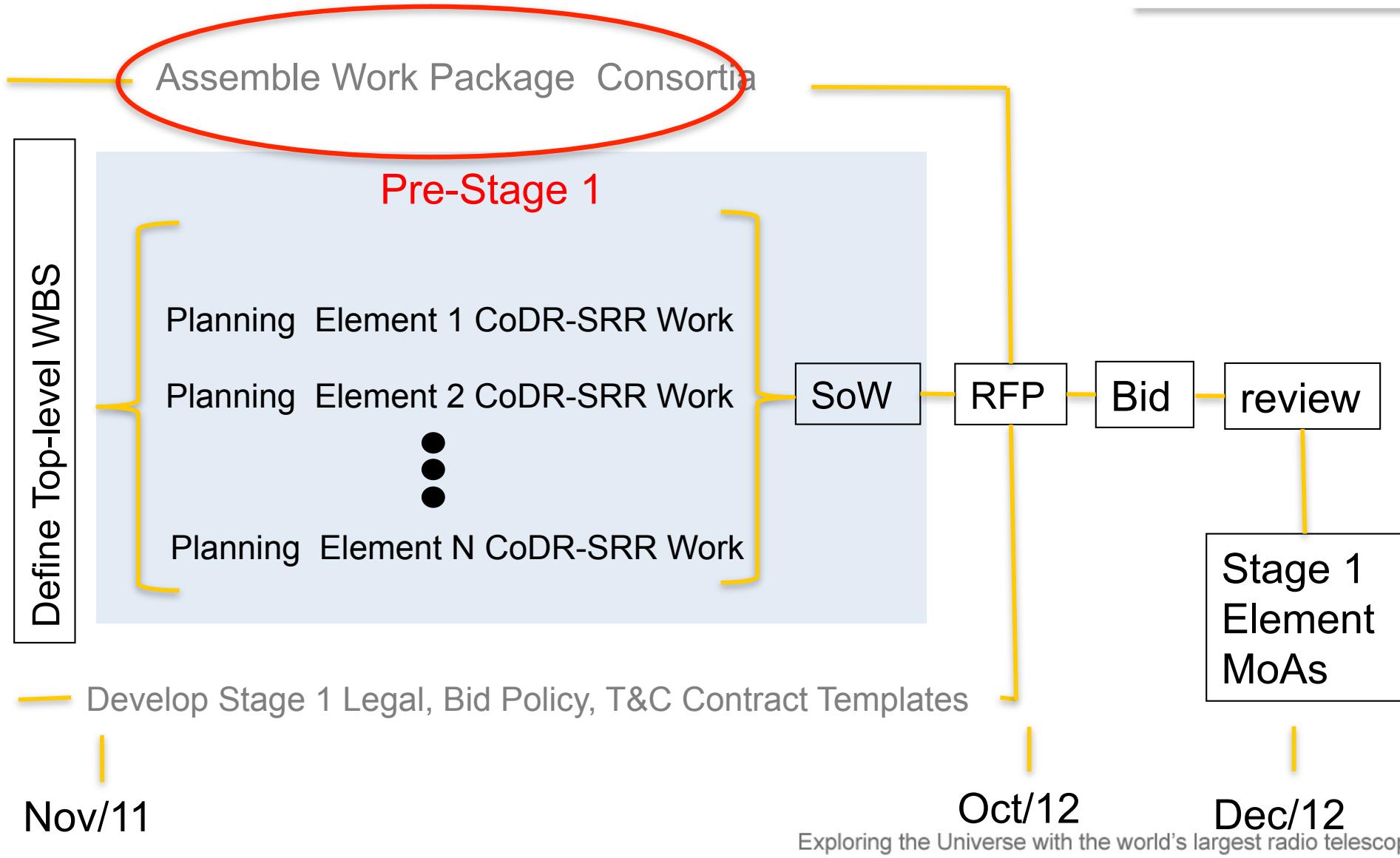
SKA Schedule



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Assembling Element Level RFPs for Stage 1 Work



Expressions of Interest



- EoI's semi public
- 133 respondents
- No real surprises

Low Frequency Aperture Array (LFAA)

80 Respondents

(* Lead = Expression of interest to lead Element; Full = full coverage of the work within the Element; Cons = Consortium; Y = Yes; N = No; P = Participate)

- Adaptive Array Systems Limited
- Agilent Technologies UK
- Alberta Centre for Advanced Micro&Nanotech Products
- Altera
- ARiYA Project Managers
- Arup
- ASTRON (Lead: Y; Full: Y; Cons: Y)*
- Aurecon (Lead: P; Full: N; Cons:N)*
- BCF Solutions
- Bigen Africa Services
- Callisto Limited/Callisto France
- Cambridge Consultants Ltd
- Canadian SKA Industry Consortium (NRC led)
- CCD Design & Ergonomics
- Cisco International Limited
- Clearspeed Technology Limited
- Cobham Technical Services
- CommAgility Limited
- COMMUNICATIONS AUDIT UK
- DA-Integrated
- Daniels Electronics Ltd.
- Denel Aerostructures a Division of Denel Group
- e2v SAS
- EnSilica
- Filtronic Broadband Limited
- Fluor Ltd
- FormaShape
- Fujitsu Semiconductor Europe GmbH
- GE Intelligent Platforms
- GHD Pty Ltd
- Hatch Associates
- IBM
- IBM United Kingdom Limited
- INAF - Istituto Nazionale di Astrofisica
- Institute for Radio Astronomy & Space Research
- Instituto de Física de Cantabria (IFCA, CSIC-UC)
- Jet Propulsion Laboratory
- Lockheed Martin Australia
- MARAND PRECISION ENGINEERING
- METHODE ELECTRONICS
- Micreo Limited
- Minex Engineering Corp.
- Mott MacDonald Ltd
- National Institute of Aerospace Technique
- National Instruments
- National Physical Laboratory
- National Research Council of Canada
- NCRA-TIFR
- Nexeya Systems
- Norsat International Inc
- NXP Semiconductors
- Observatory Sciences Ltd
- Omnisys Instruments AB
- Parsons Brinckerhoff
- Prudent Energy Corporation
- Reutech Radar Systems
- RFEL Ltd
- Roke Manor Research
- Sanyati Holdings (Lead: Y; Full: N; Cons:N)*
- SCISYS
- SELEX Galileo
- SELEX Sistemi Integrati S.p.A.
- Siemens Industry
- Siemens Nederland
- SKA South Africa (Telescope manager)
- SSI Engineers and Environmental consultants
- STFC – Technology Department
- STFC RAL Space Department
- Systems Engineering & Assessment
- Tata Consultancy Services
- TEK Microsystems
- Teledyne Defence Limited
- Telespazio VEGA UK Ltd
- The Boeing Company
- Tronicon Industries Inc.
- UNIVERSIDAD CARLOS III DE MADRID (UC3M)
- Universidad de Cantabria. Dept of Communications Engineering
- University of Cambridge
- University of Malta
- WorleyParsons Europe Limited (Lead: Y; Full: Cons:N)*

Aperture Arrays



- LFAA + MFAA
 - ASTRON lead
 - ICRAR
 - INAF-IRA
 - JIVE
 - Observatoire de Paris
 - Raman Research Institute
 - Université de Bordeaux 1
 - University of Cambridge
 - University of Malta
 - University of Manchester
 - University of Oxford

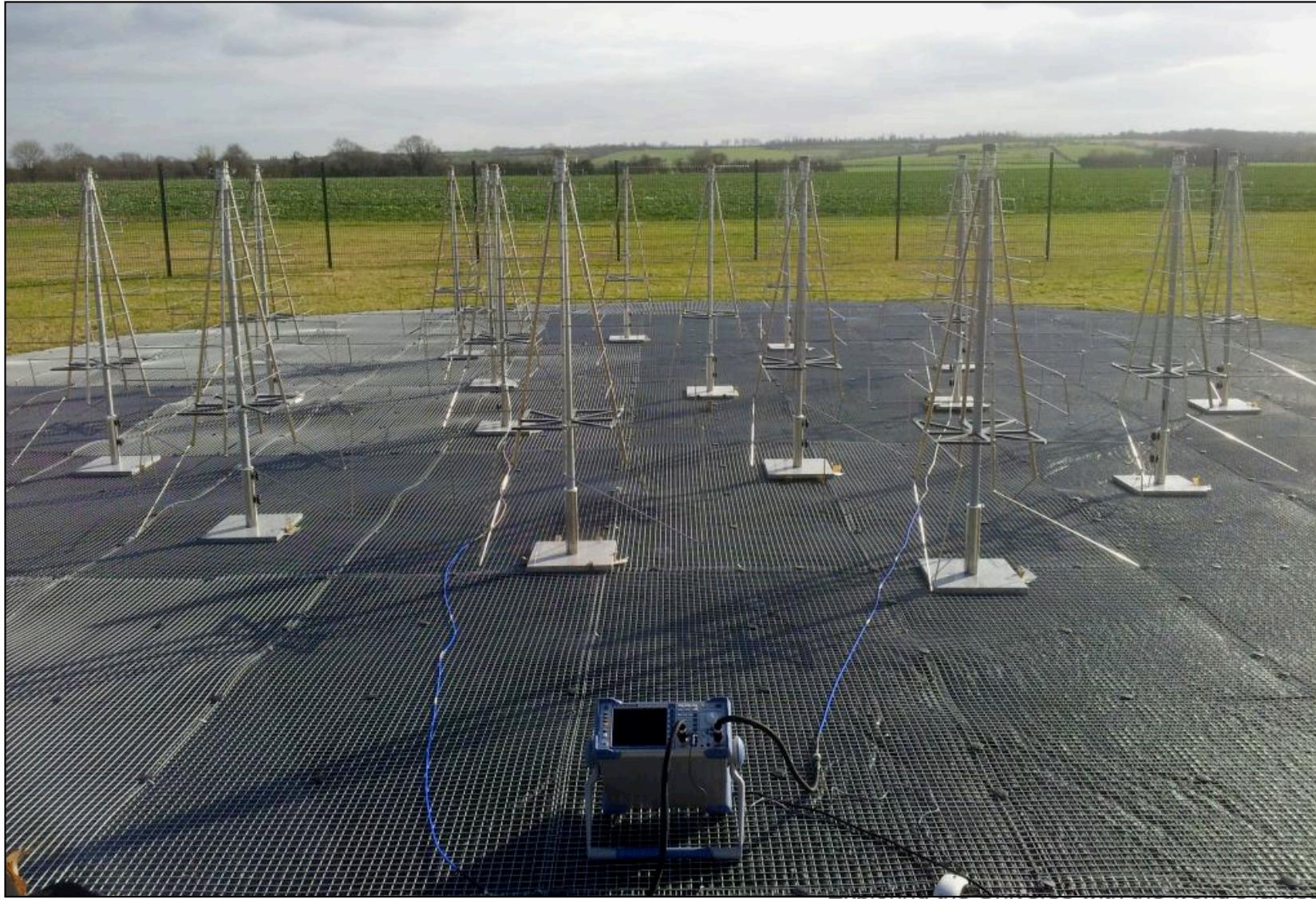
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LFAA Planning



- AAVS0: several elements
- AAVS1: 256 elements
 - 4 x 64 or 16 x 16
- AAVS2: 11500 elements (2% SKA1)
 - 23 x 500

AAVS0: Log Periodic Dipole



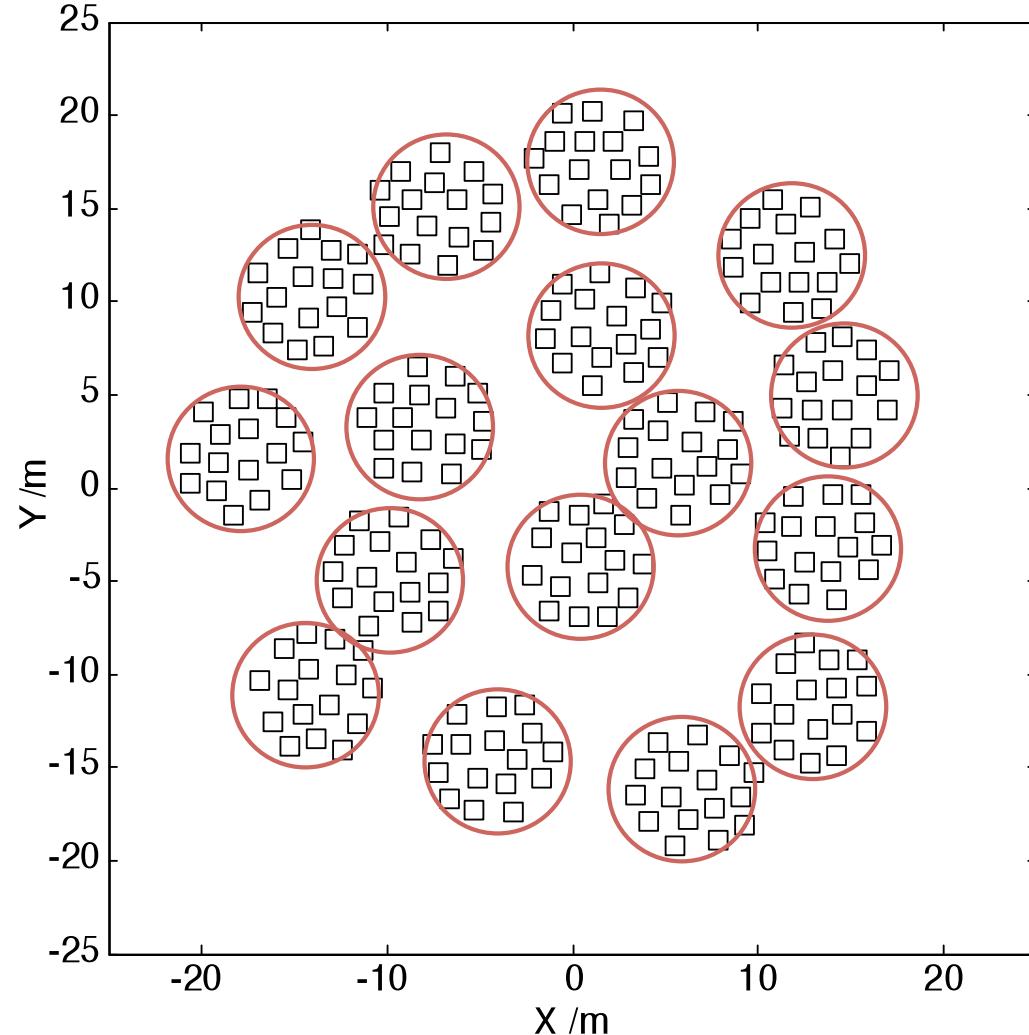
AAVS0: Vivaldi



LFAA AAVS1



Possible AAVS1 configuration suitable for beamforming and correlating



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LFAA AAVS1: system tests



- Electronics
- Beamformer
- Stability (gain)
- Reliability (time stability)
- A/T
- Calibration

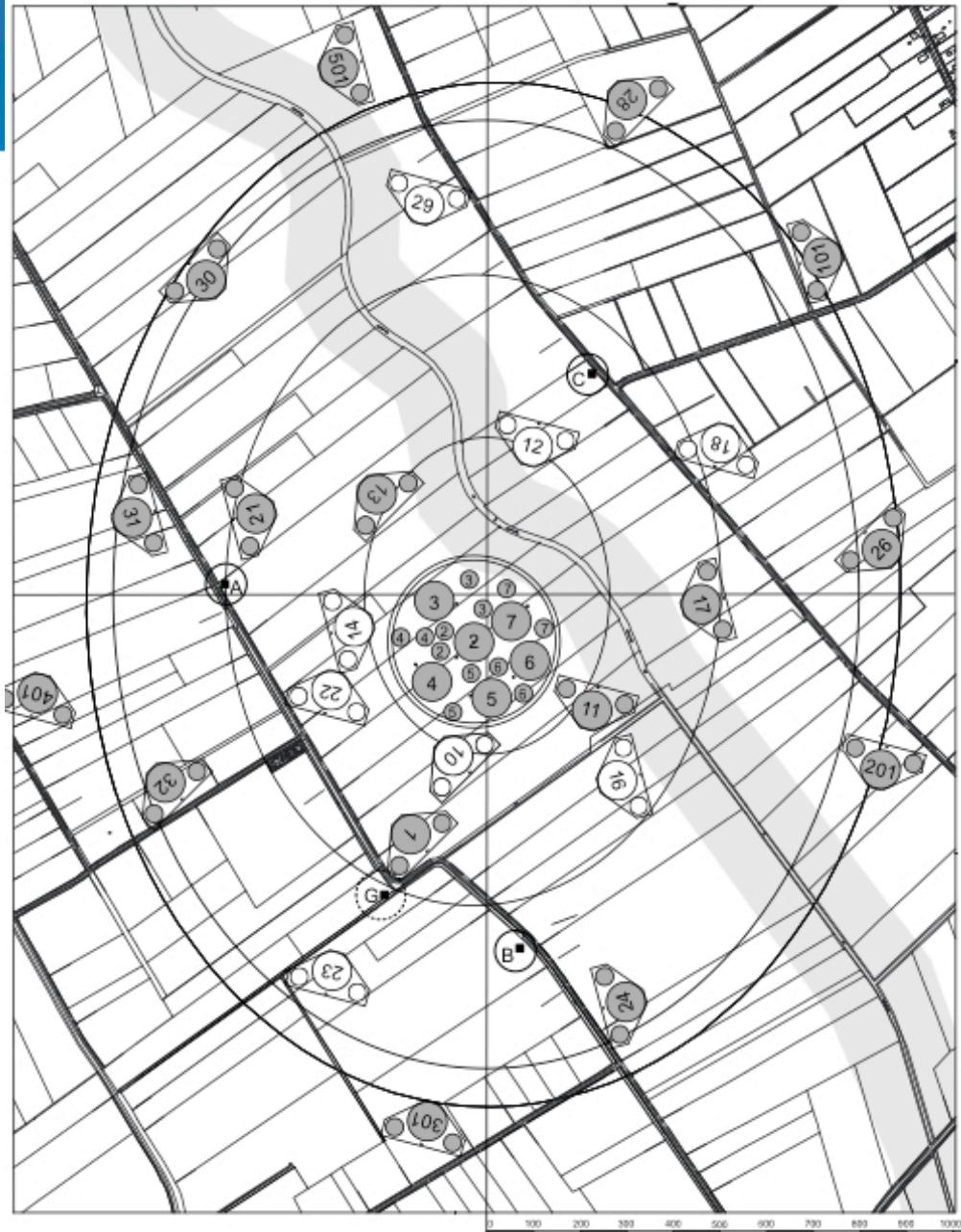
(input from Nima Razavi)

LFAA AAVS1: astronomical verification



- Pulsar detection
- Drift scans of bright sources
- Tracking bright sources
- Imaging capability
- Beam switching experiments?
- Multi-beam experiments?

AAVS2



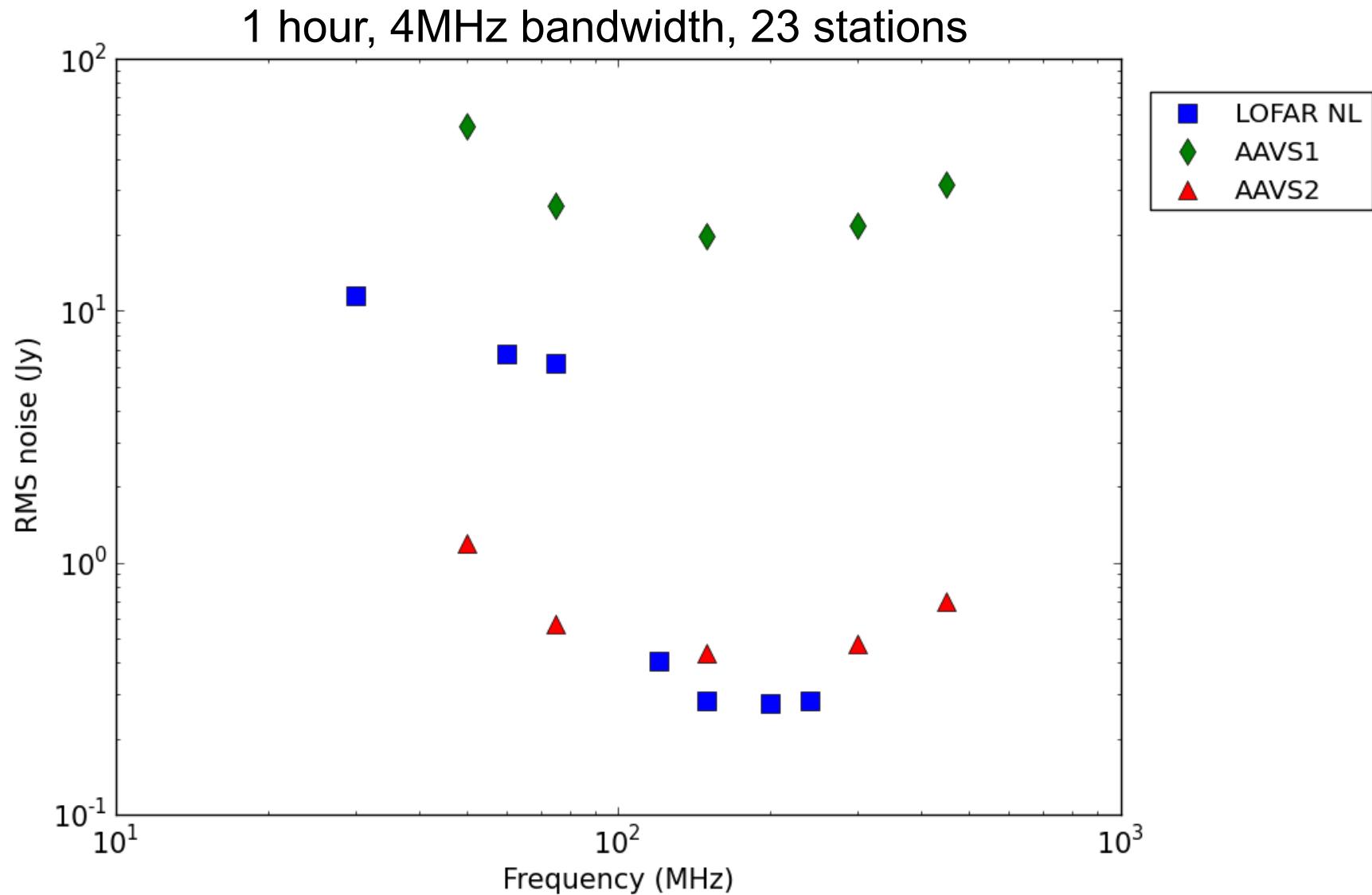
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LFAA AAVS2 technical tests



- Signal transport
- Calibration
- System stability
- Correlation modes
- Post processing
- Beam

LFAA AAVS2



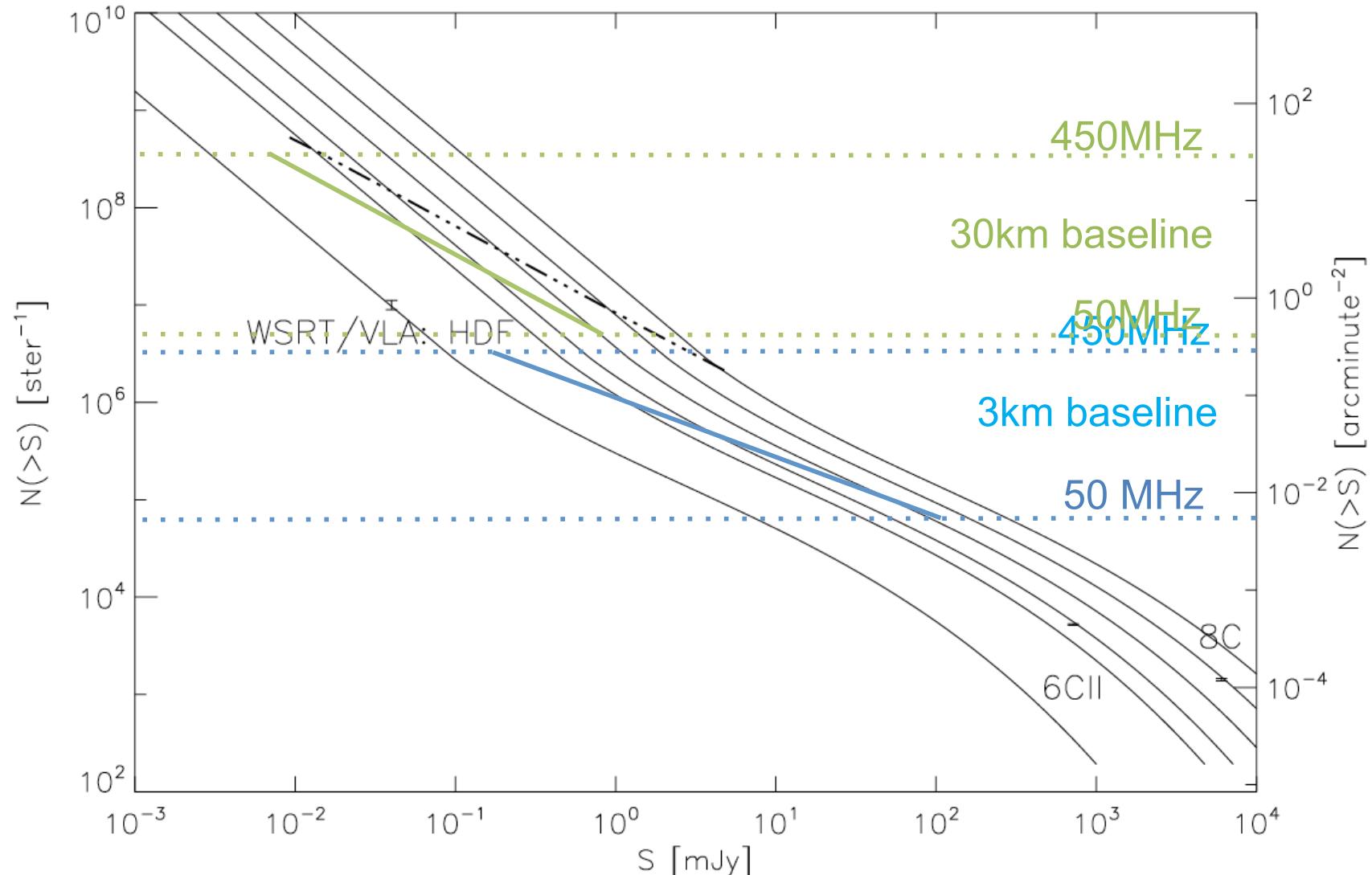
LFAA AAVS2 astronomical verification



- Imaging
- Polarization (RM synthesis)
- Pulsars
- Solar system (dynamic spectra)
- Ionosphere and space weather

LFAA AAVS2 is science capable!

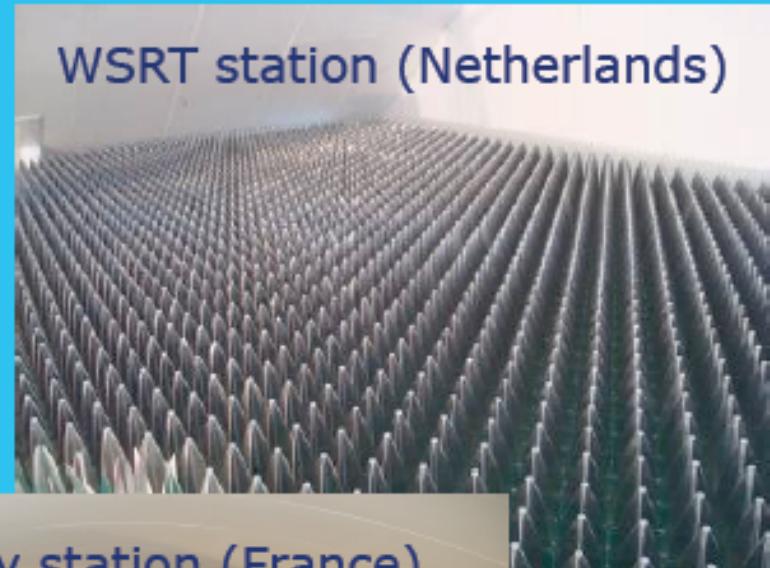
LFAA AAVS2 confusion limit



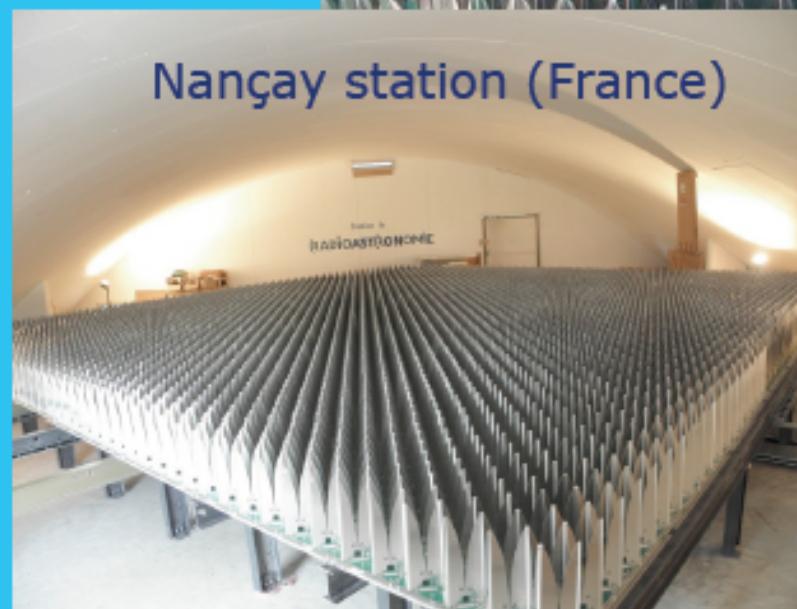
EMBRACE: AA-mid technical demonstrator

ASTRON

- $A/T \sim 1 \text{ m}^2/\text{K}$
- 400-1500MHz
- TWO analogue beams
- ~200 digital beams (195.3kHz wide)
- ~40MHz bandwidth
- single pol



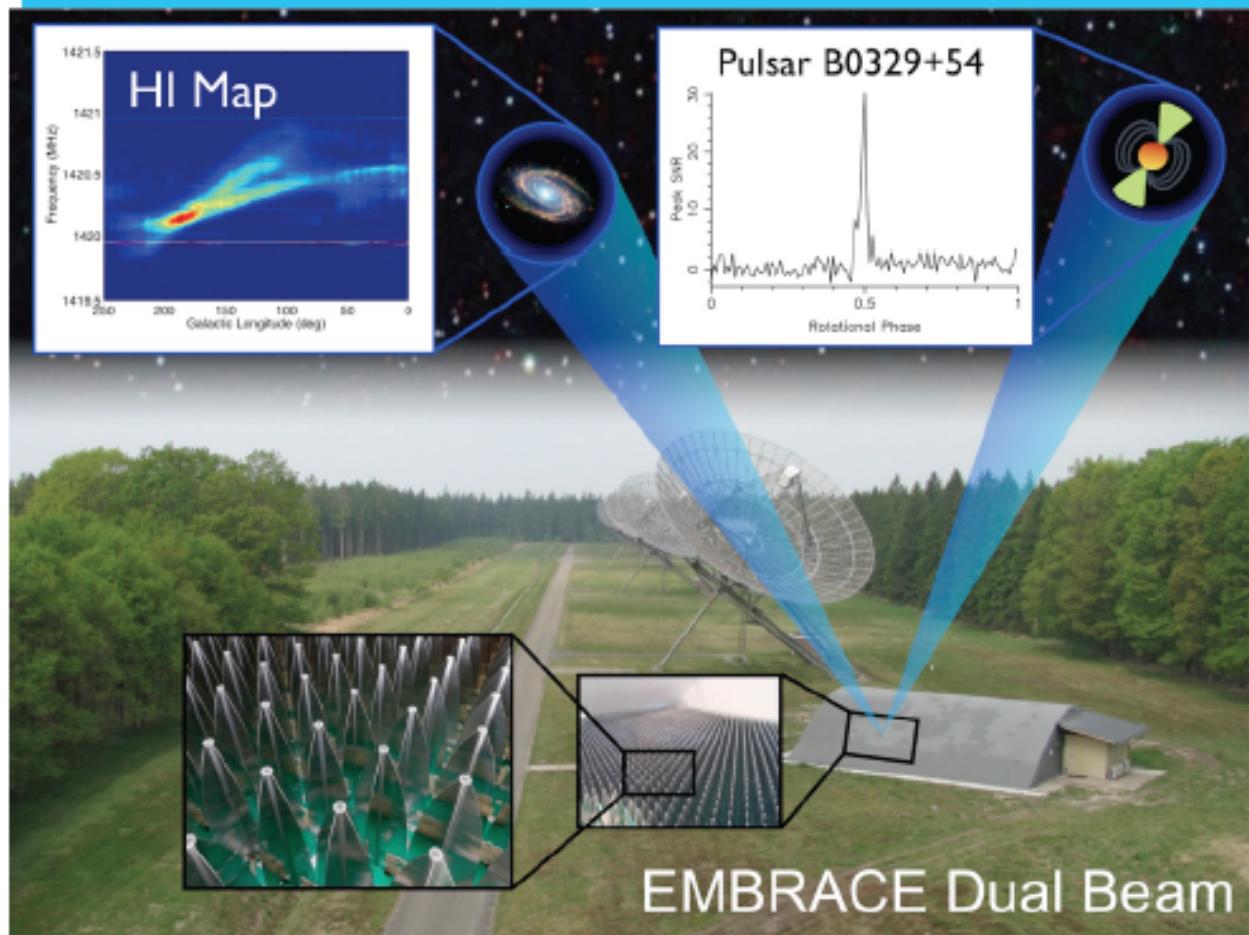
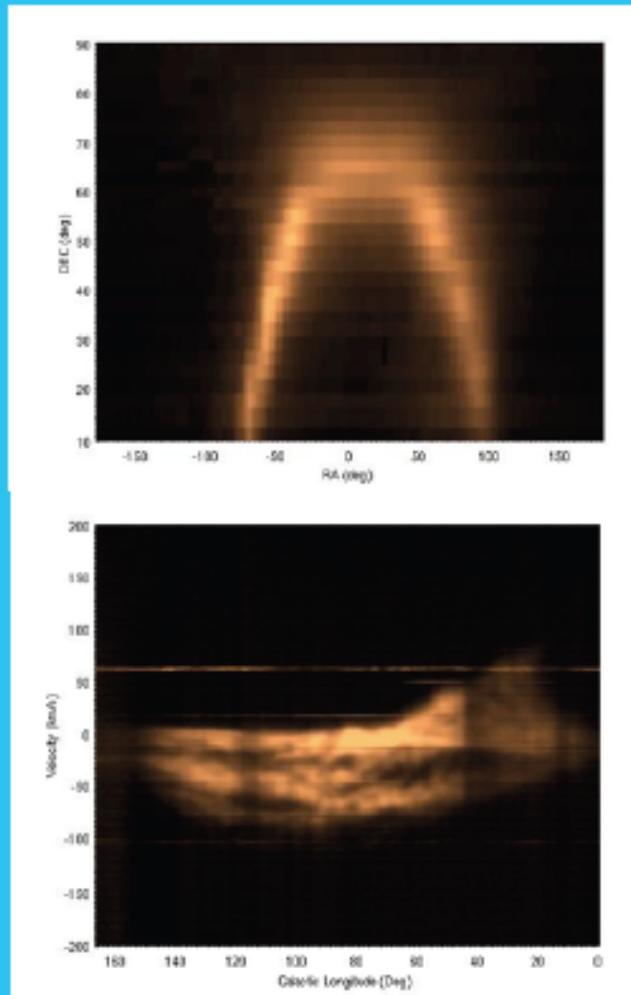
WSRT station (Netherlands)



Nançay station (France)

EMBRACE results: WSRT

ASTRON



EMMA specs

ASTRON

- 2000m^2 ($A/T \sim 40 \text{ m}^2/\text{K}$)
- multiple stations
- full Stokes
- 2 FoV, 64 digital beams
- $\sim 80 \text{ deg}^2$ per FoV
- 450-1450 MHz
- $T_{\text{sys}} \sim 50\text{K}$
- bandwidth 500MHz

SKA single station

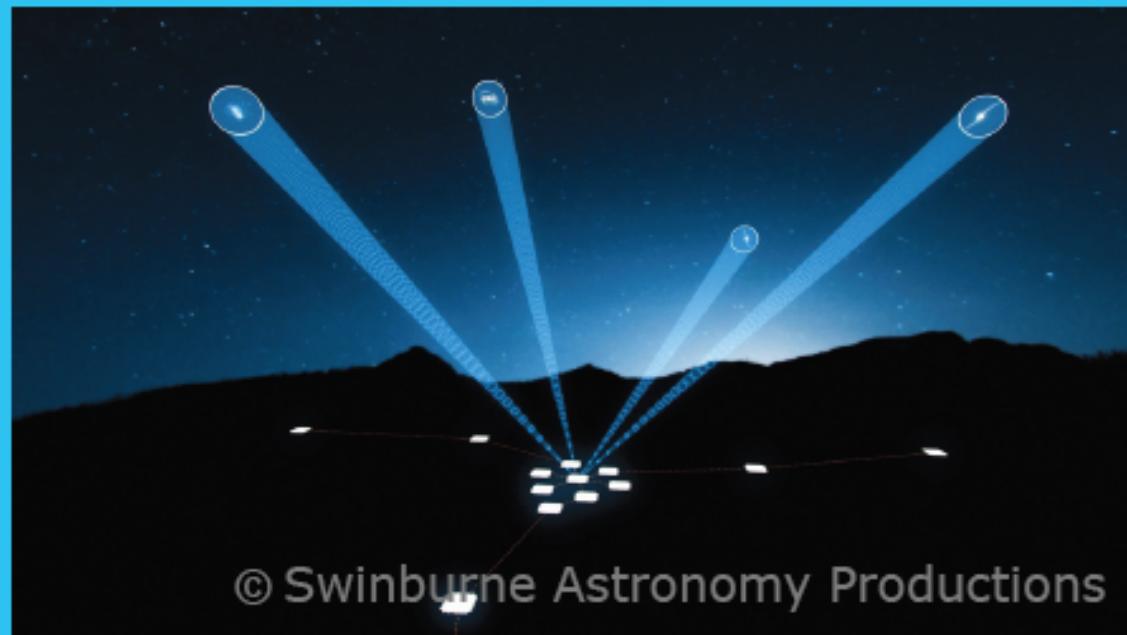


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EMMA configuration



- 14 stations
- 13.5m diameter
- central core: >50%
- longest baseline between 300-1000m



Comparison



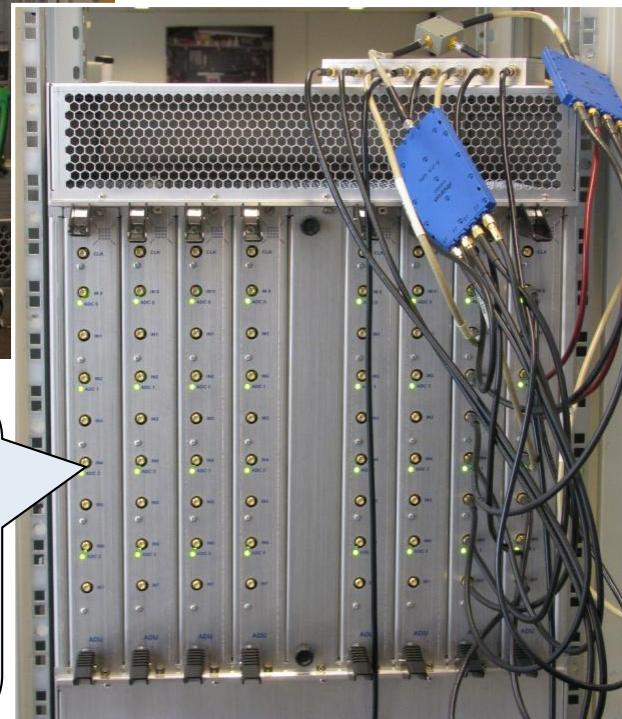
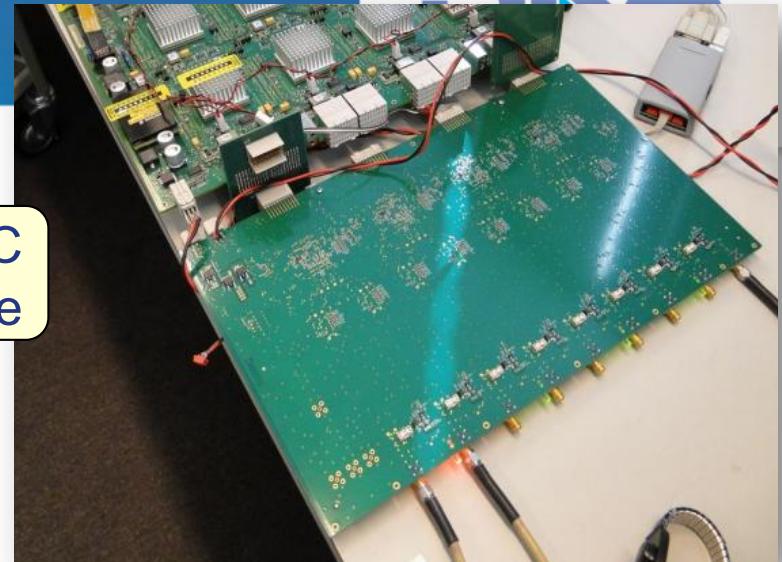
	EMMA	APERTIF	ASKAP	EVLA	MeerKAT-1
Frequency (GHz)	0.4-1.45	1.0-1.7	0.7-1.8	1.0-50	0.9-1.75
Bandwidth (GHz)	0.5	0.3	0.3	0.5 (8.0)	0.35
FoV (deg ² , 1.4GHz)	78	8	30	0.3	0.6
z_{\max} for HI absorption	2.55	0.42	1.03	0.42	0.58
S_{rms} (μJy , 1h, full BW)	37	30	35	7.6	14.6
S_{rms} (μJy , 1h, 100MHz)	84	49	61	17	27
S_{rms} (mJy, 1h, 5 km/s)	5.5	3.7	4.0	1.1	1.8
A/T (m ² /K)	40	105	58	246	150
SSFOM $\times 10^4$ (m ⁴ /K ² / deg ²)	12.5	8.9	13.8	1.8	1.4
SSL($\tau < \tau_0$)/N _t	1	0.92	0.73	5.3	5.6

based on online information of late 2011

Uniboard Implementation

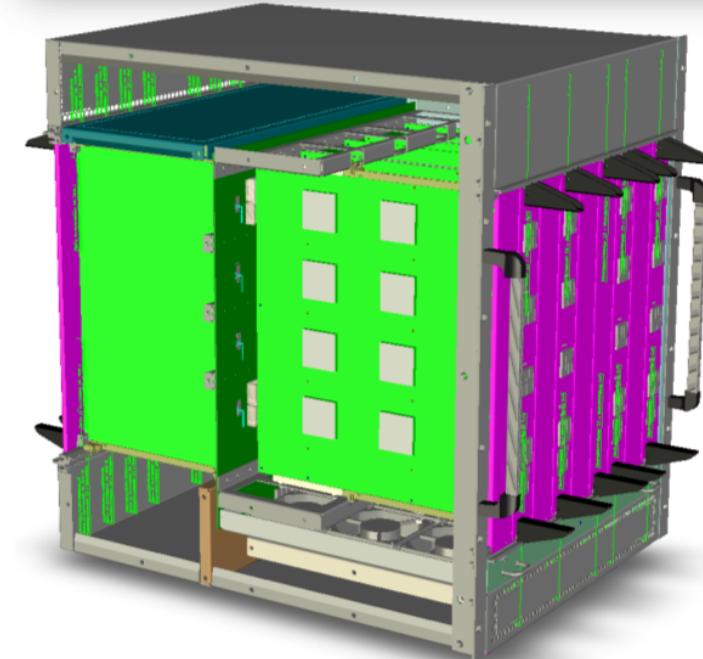


ADC
Interface



Shelf:

- 4 Processors
- 8 ADC interfaces
- 64 inputs
(32 elements)



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Today



- To use the outcome of the meeting to shape-up our plans to get to the specification, design/development and verification of AA's
- To identify risks in the current approach and work out alternatives if needed

Leading up to:

Aperture Array Design and Construction Consortium
AADC