

An aerial photograph of a vast desert landscape filled with numerous circular radio telescope dishes. The dishes are arranged in a grid-like pattern across the reddish-brown terrain. In the foreground, a white car is parked next to one of the dishes, providing a sense of scale. The background shows distant mountains under a clear blue sky.

# Signal processing *(in RA)*

A Radio Observatory is a *signal processing and computing facility* with an unusual source of data...

**... Astronomers are always wanting  
more (of everything) !**

## Why?

- To see what we are looking for!

## What?

- Add up signals
- Take bits away from signals
- Select the bits of signals we want
- Process signals from one form to another

## When?

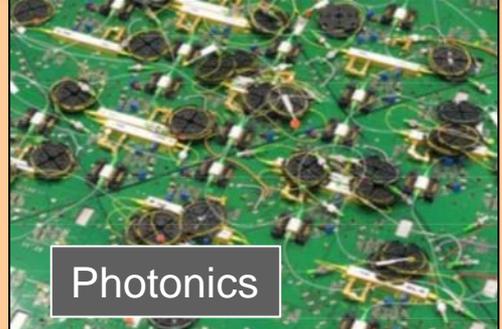
- At the time
- After an observation

# Some signal processing

**Metal mirrors**



**Cables & Switches**



Photonics

**Some Others**



Lipson Beaver strips

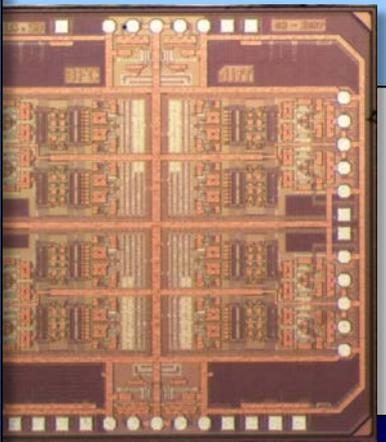
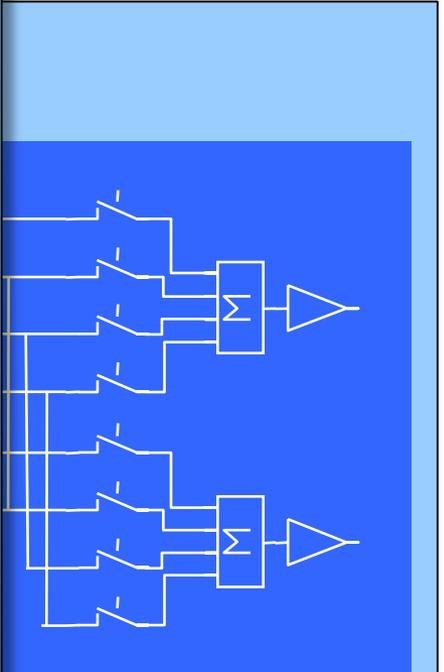
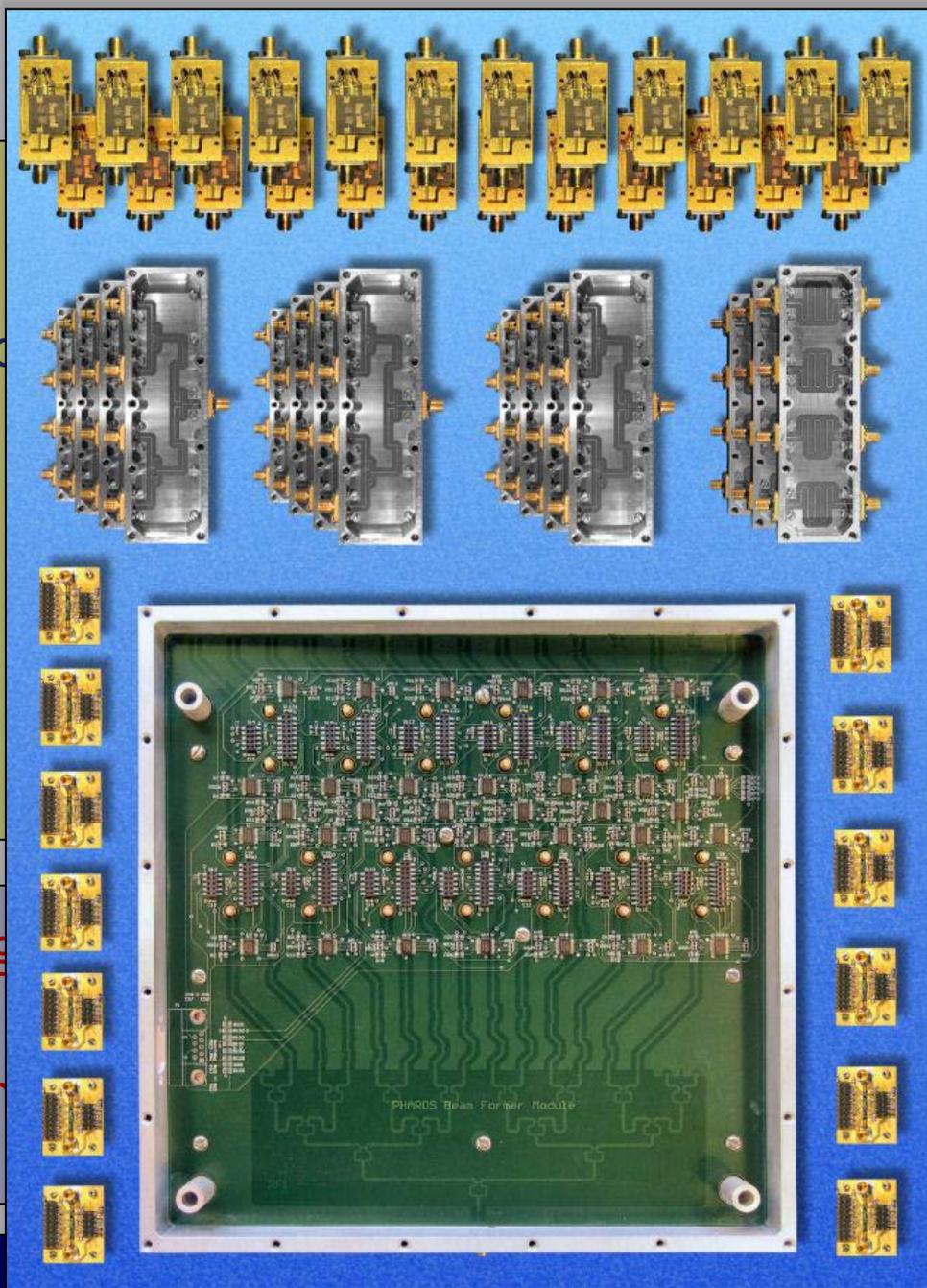


LC Delay Line

## Analogue

- Phase rotation
- Time delay
- Level adjust
- Summing channels
- etc.

- Special design
- Channels=el
- Stability, con



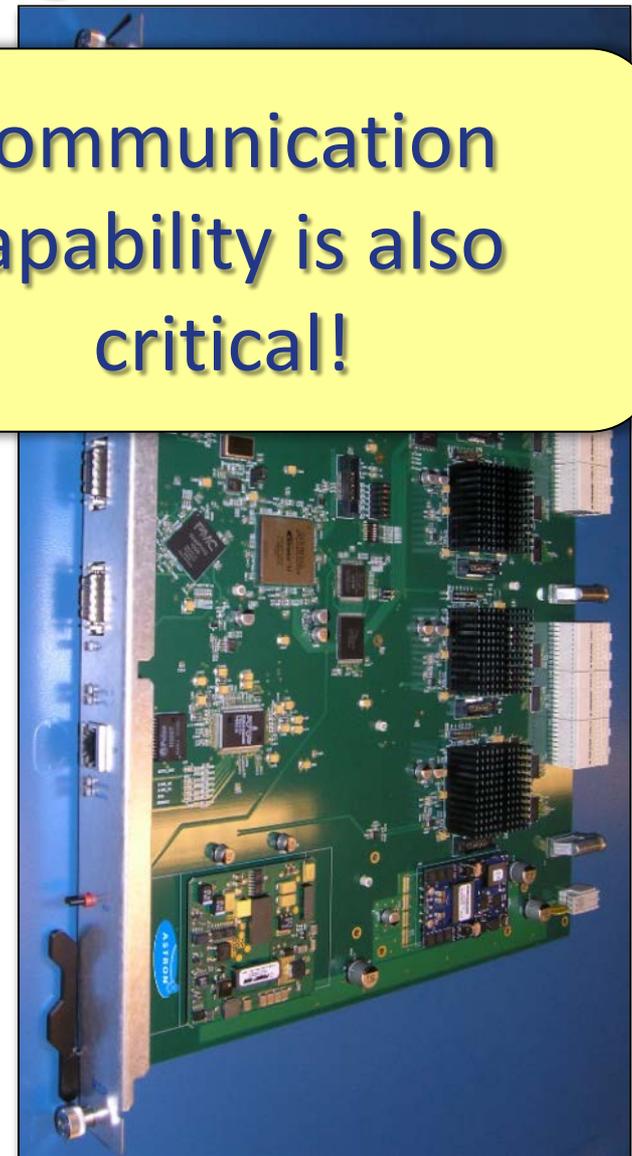
## Digital is:

- Flexible
- Compact
- Programmable
- Driven by the **ICT industry**
- Highly integrated ...

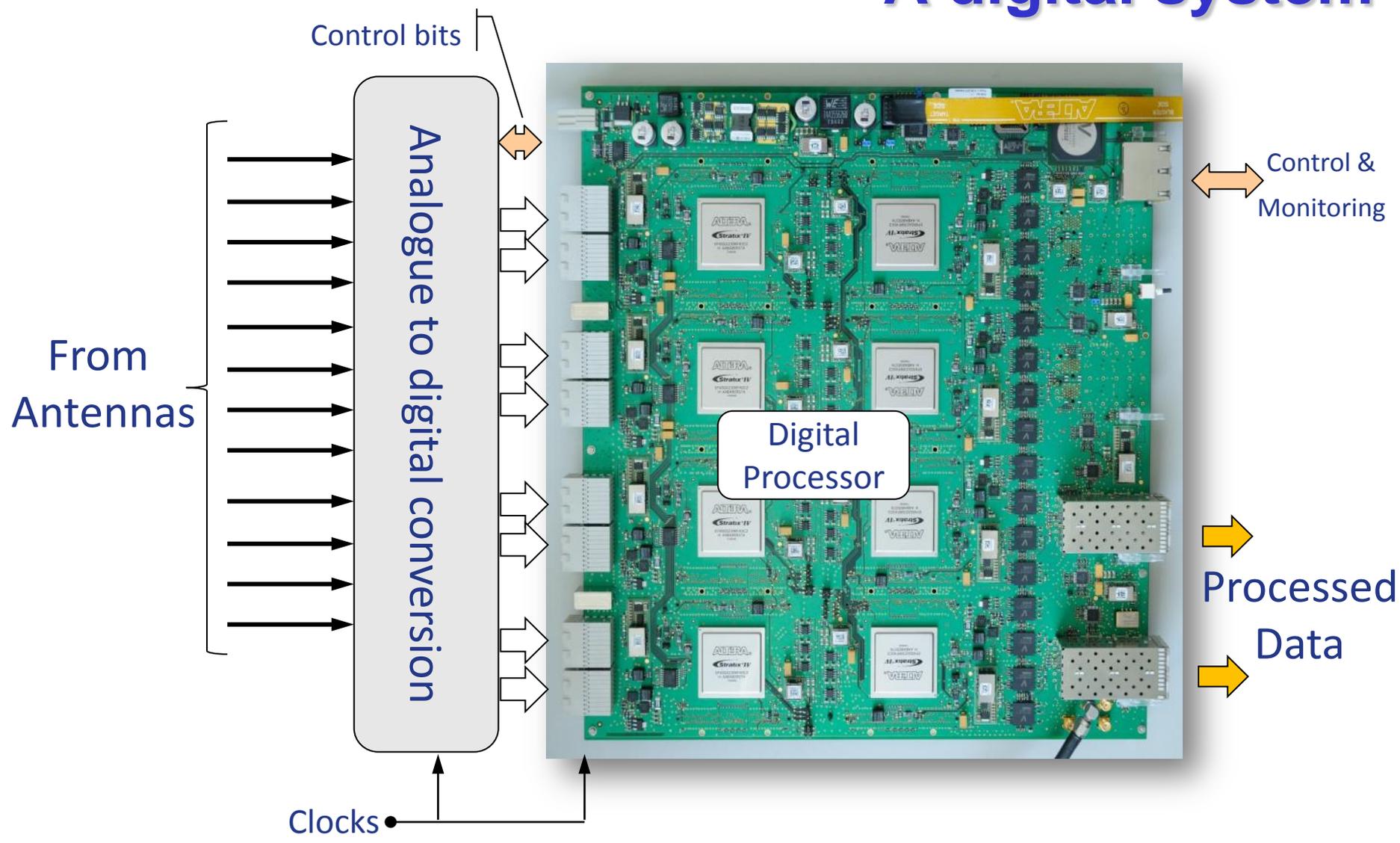
## Digital is also:

- Tricky
- Power hungry
- Abstracted
- Interference causing

Communication  
capability is also  
critical!



# A digital system



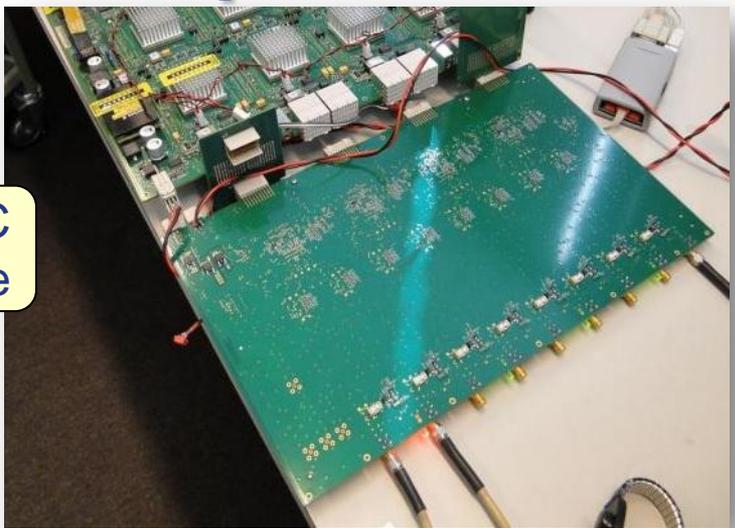
# Digital Processing types

Type	Pros	Cons
Custom/Asics	<ul style="list-style-type: none"> <li>Low power</li> <li>High performance</li> <li>Low cost in high volume</li> </ul>	<ul style="list-style-type: none"> <li>Long Development</li> <li>Expensive NRE</li> <li>Inflexible</li> </ul>
FPGA	<ul style="list-style-type: none"> <li>Low-Medium power</li> <li>Programmable</li> <li>High performance</li> </ul>	<ul style="list-style-type: none"> <li>Expensive parts</li> <li>Hard to program</li> <li>Specialist boards</li> </ul>
CPU	<ul style="list-style-type: none"> <li>Very easy to program</li> <li>Very flexible</li> <li>COTS systems available</li> </ul>	<ul style="list-style-type: none"> <li>Low performance for SP</li> <li>High power</li> </ul>
Many core processors	<ul style="list-style-type: none"> <li>Medium Power/performance</li> <li>Very flexible</li> <li>Easy to program</li> <li>COTS systems available</li> </ul>	<ul style="list-style-type: none"> <li>Limited I/O</li> </ul>

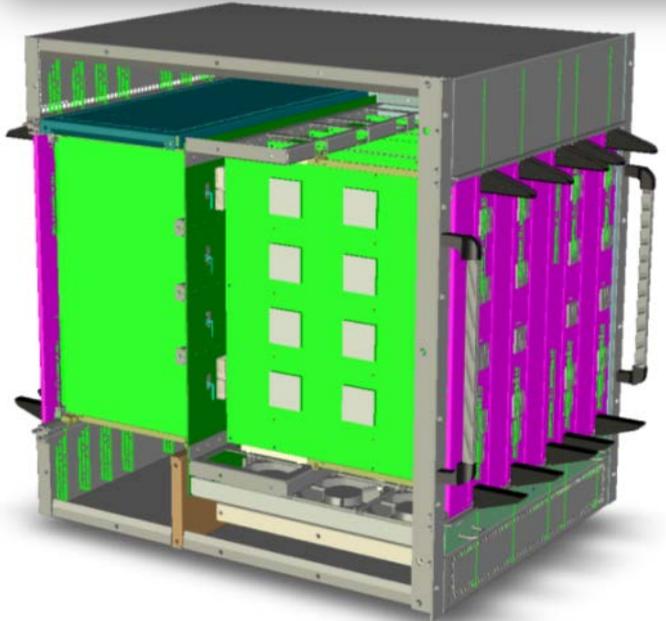
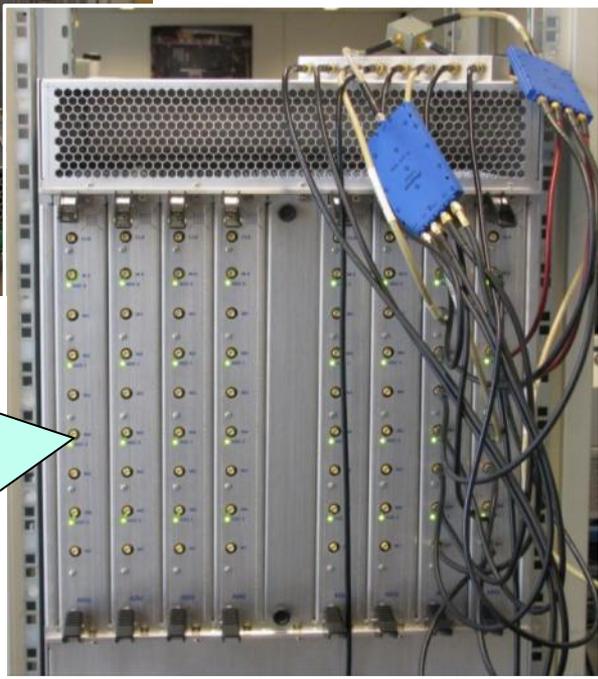
# Uniboard 1 Implementation



ADC Interface



- Shelf:
- 4 Processors
  - 8 ADC interfaces
  - 64 inputs (32 elements)



- Field of view
- Bandwidth
- Resolution
- Processing
- Dynamic range
- Flexibility
- Interference excision

More, smaller receptors

gives

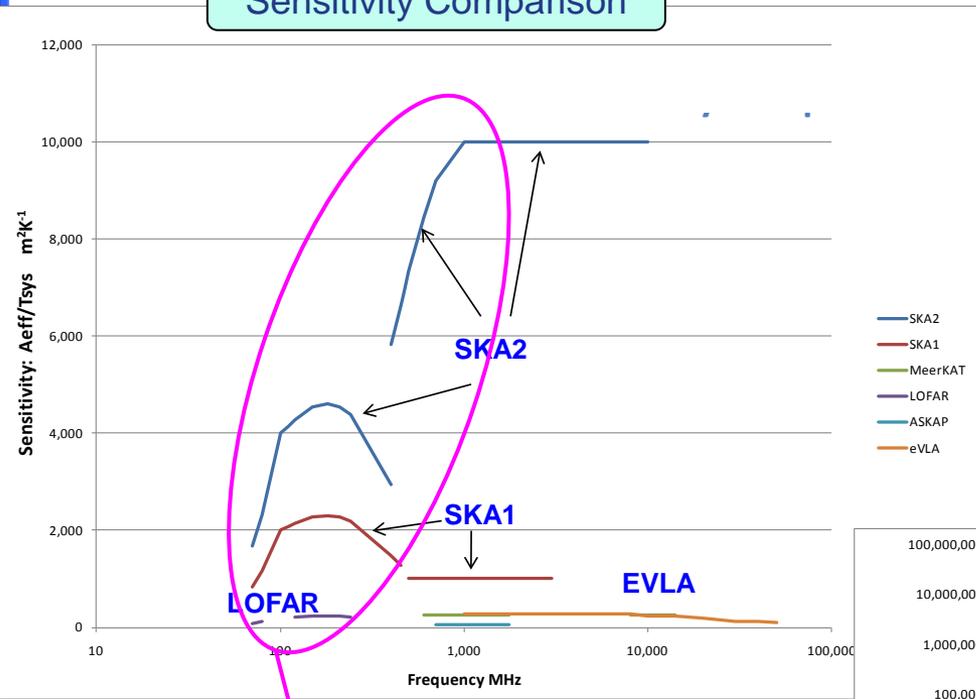
Bigger Field of View

using

Much more Signal Processing

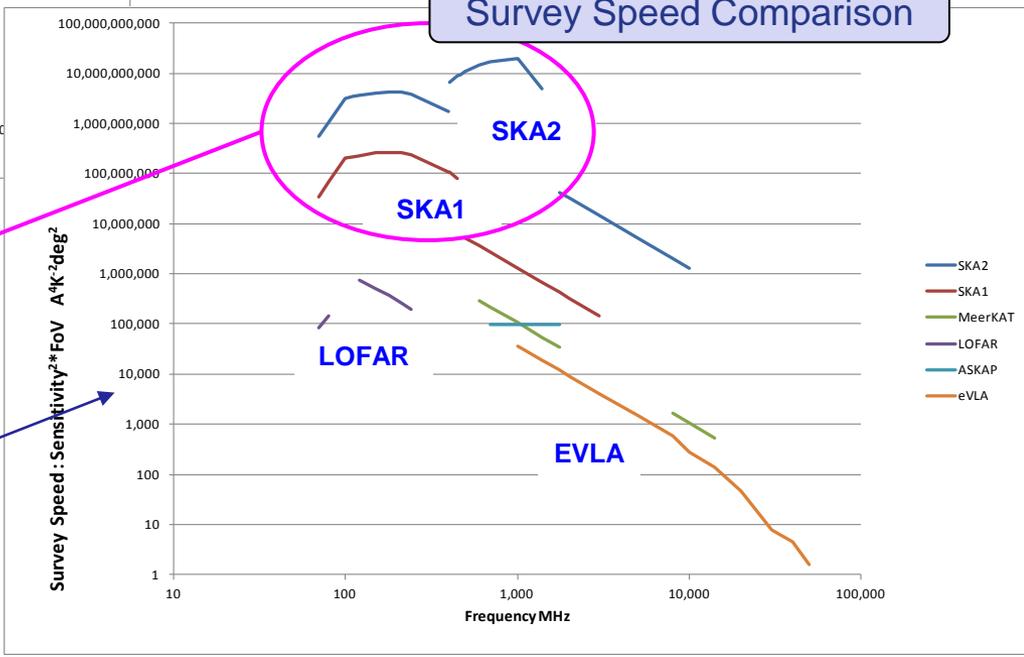
# Small dishes & Phased arrays

Sensitivity Comparison



SKA<sub>1</sub> & SKA<sub>2</sub> will have very high sensitivity & survey speeds

Survey Speed Comparison

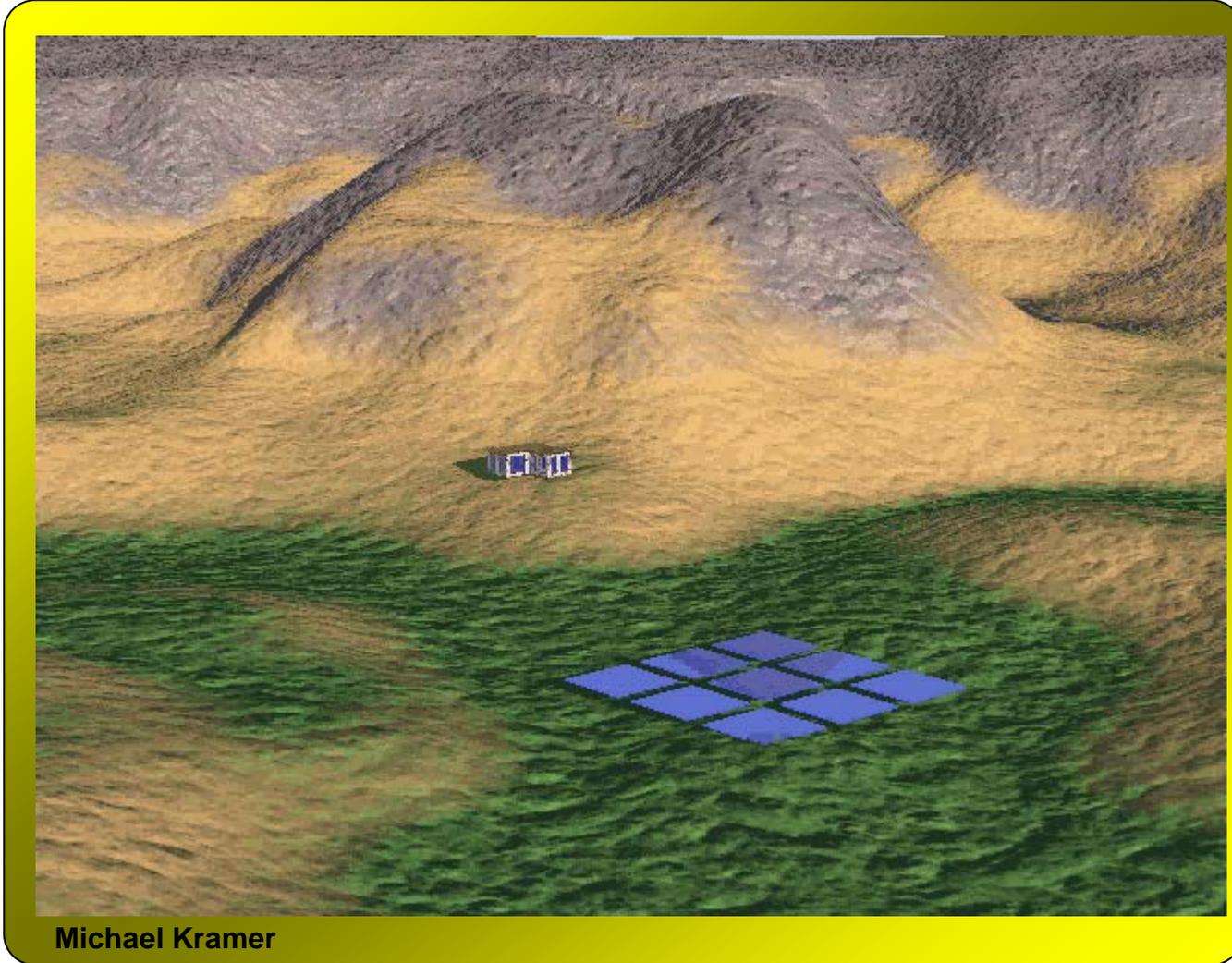


Aperture Arrays

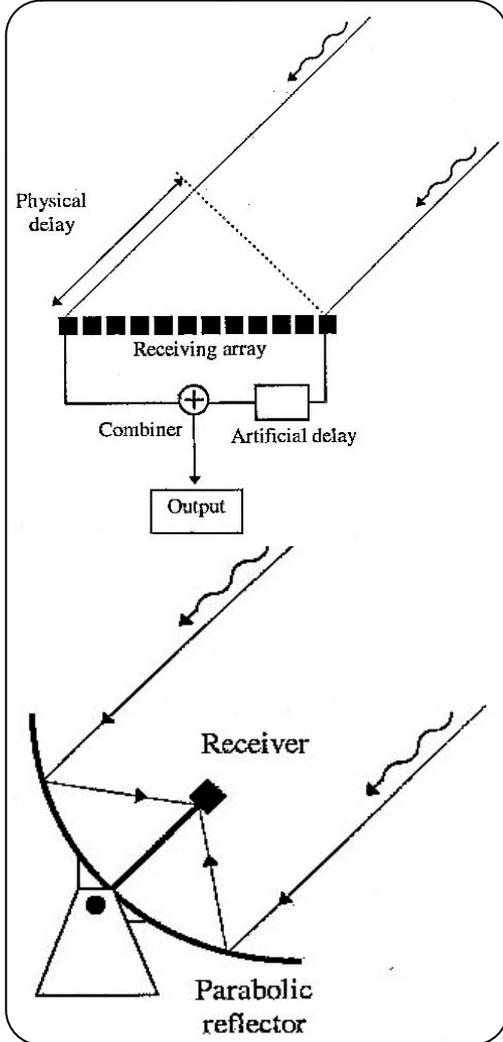
Note: log scale!

# Aperture Arrays

# Aperture Arrays



Michael Kramer



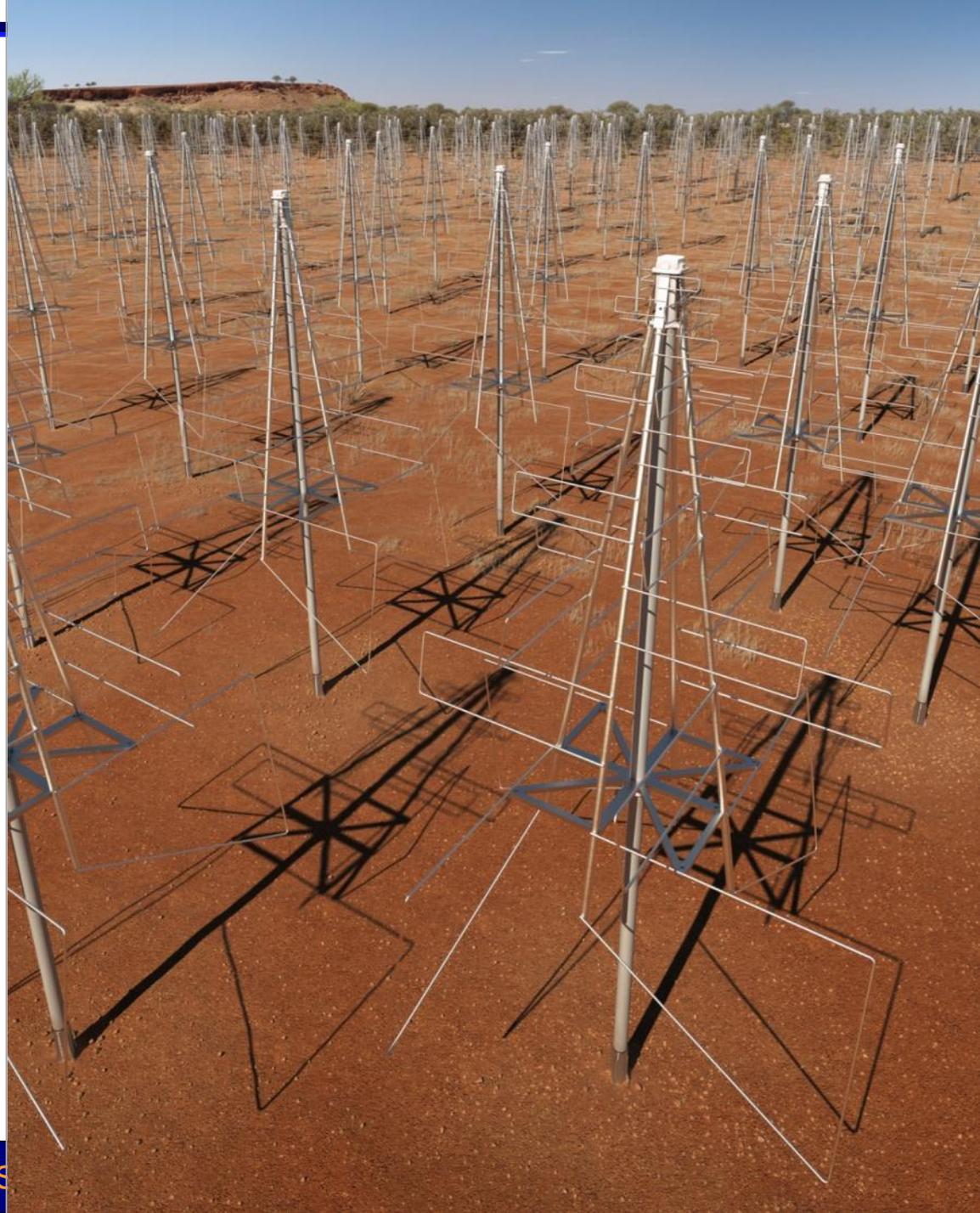


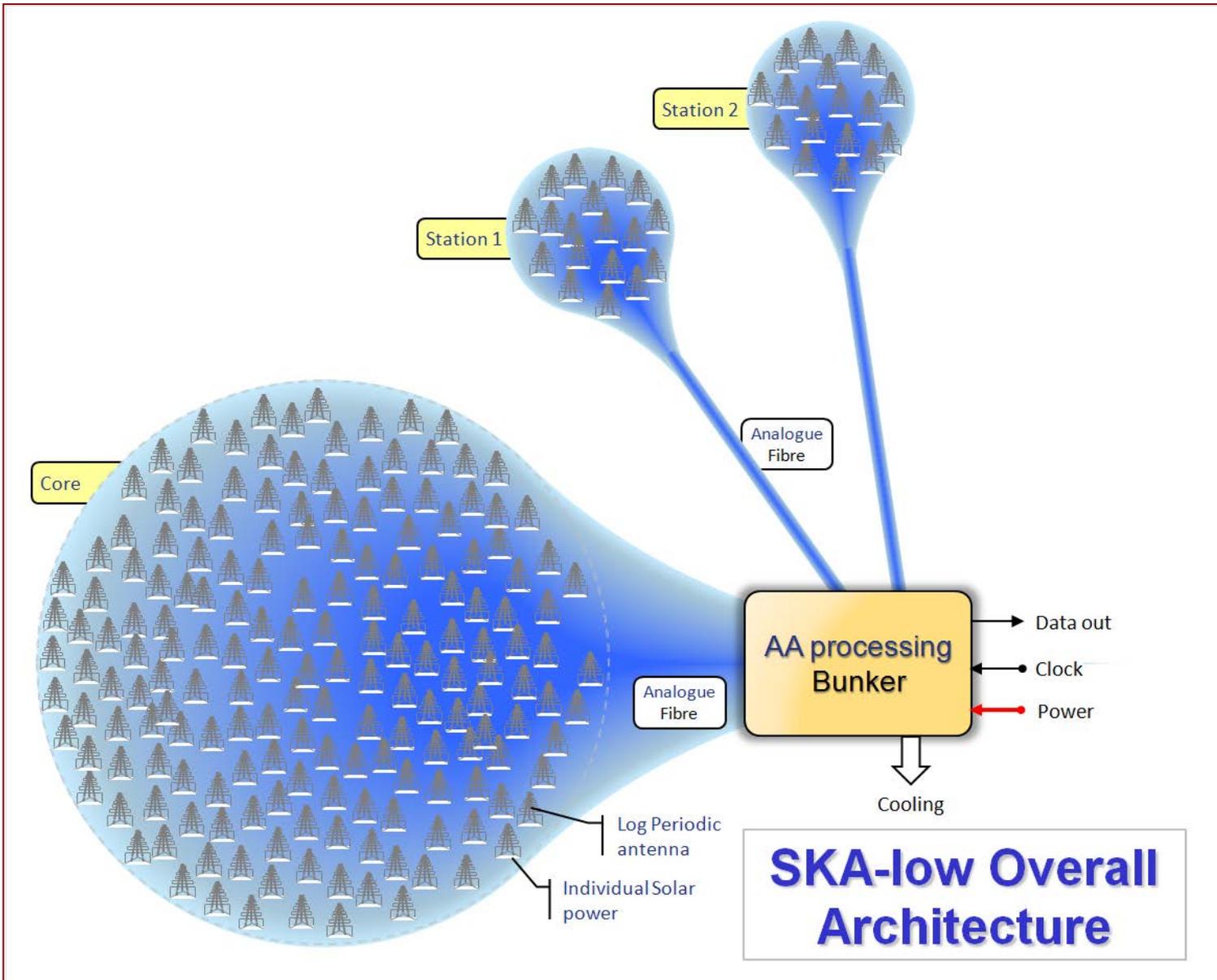
LOFAR station



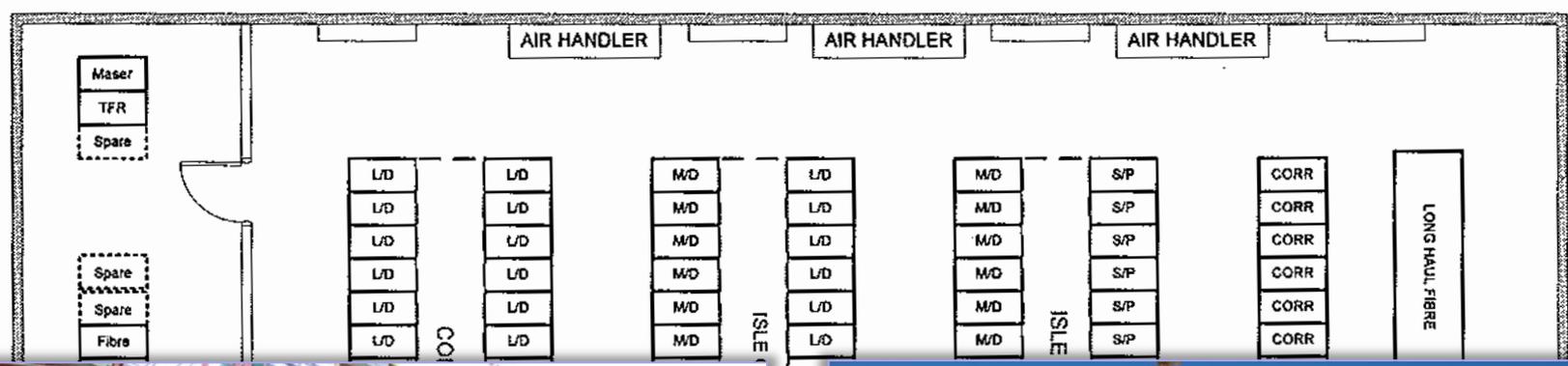
# SKA-low Design Overview (proposed)

Target	Parameters
50 – 650	MHz
256,000	Antennas
300	MHz B/W
~250 – ≥911	Stations
10 – 20	Tb/s
Multiple	Beams



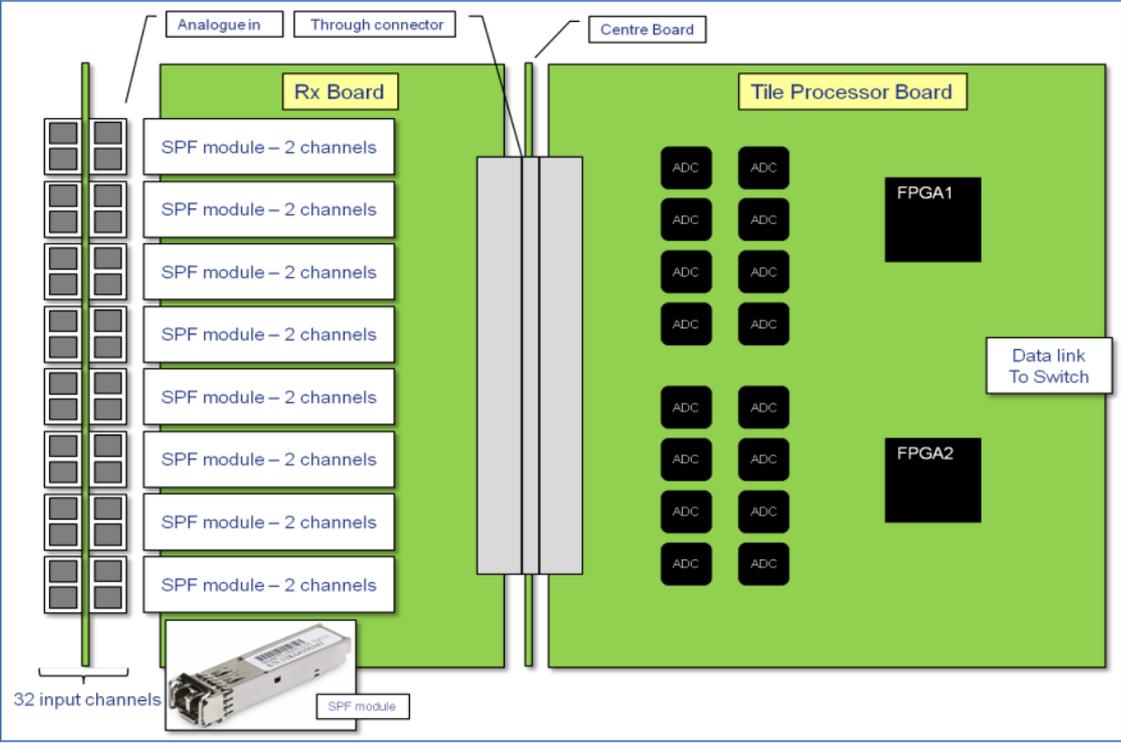


# Example bunker...

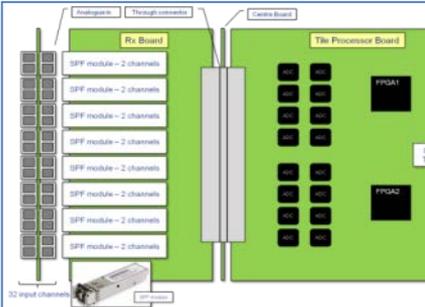
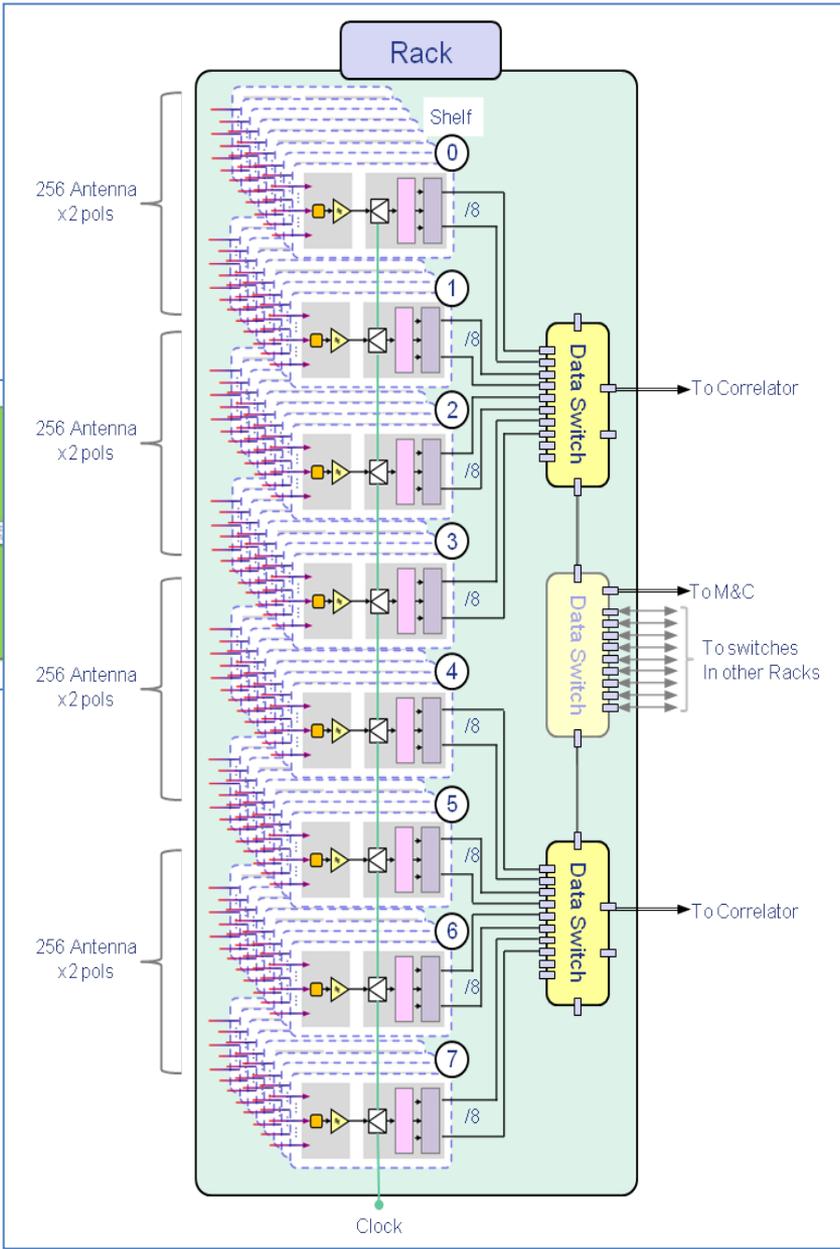


# SKA-low processing trail:

# SKA-low processing trail:



Analogue inputs  
16 dual pol. antennas  
Digitise and beamform



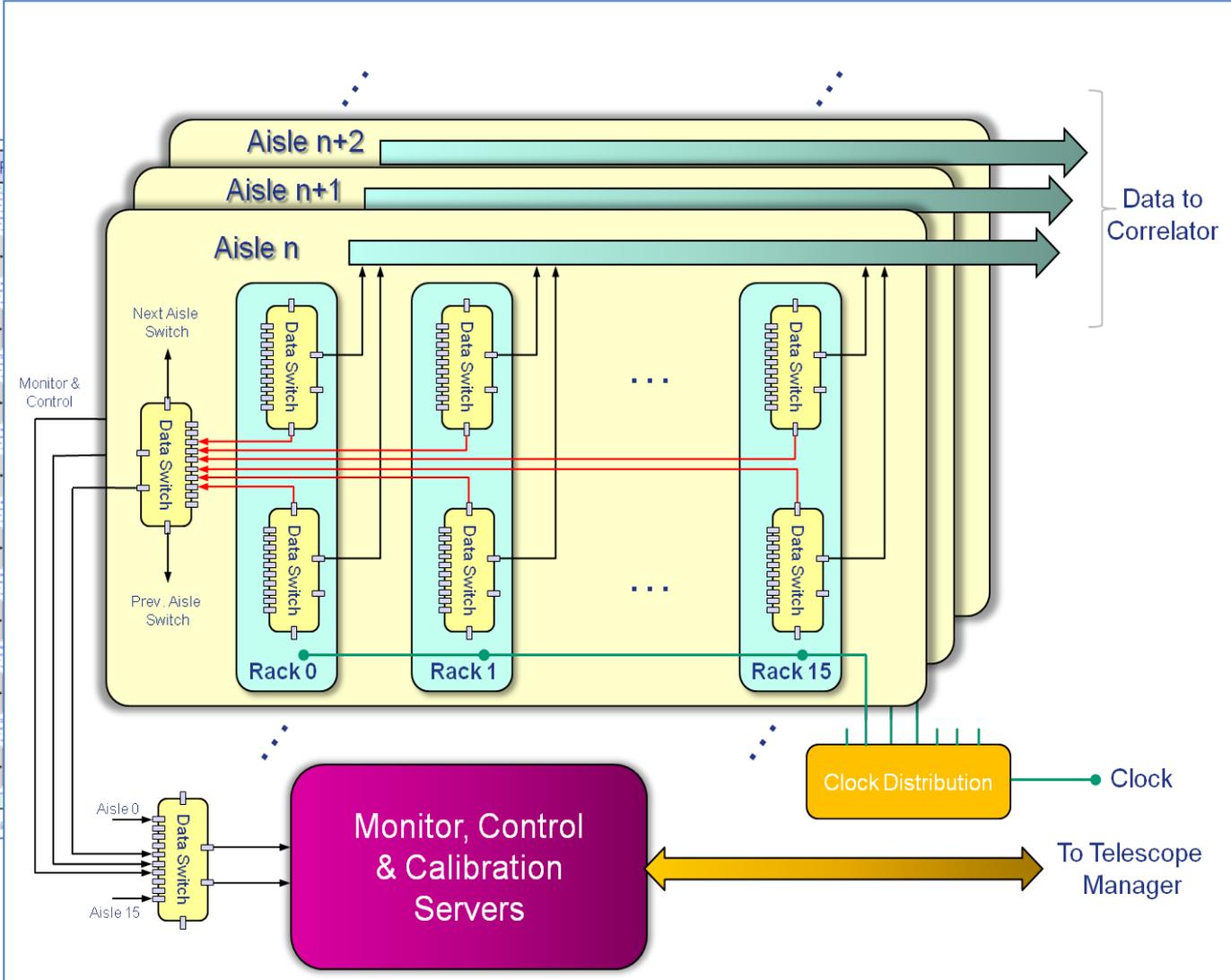
64 Tile processors in a rack

2048 fibres into each rack

Switch network to handle the data

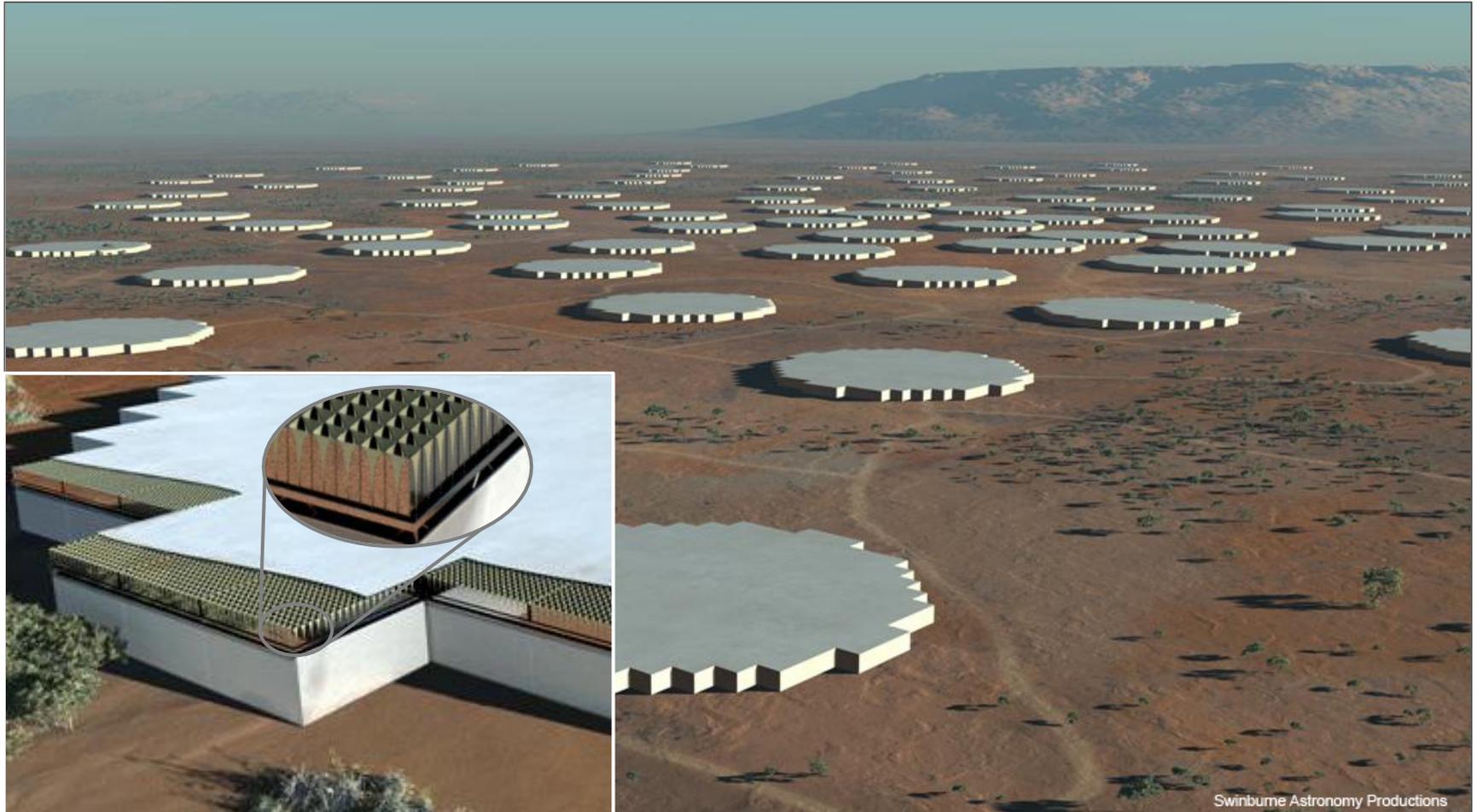
Power supplies





256 racks in 16 aisles  
 Flexible beamforming  
 High speed switched network  
 Overlaid monit. & control

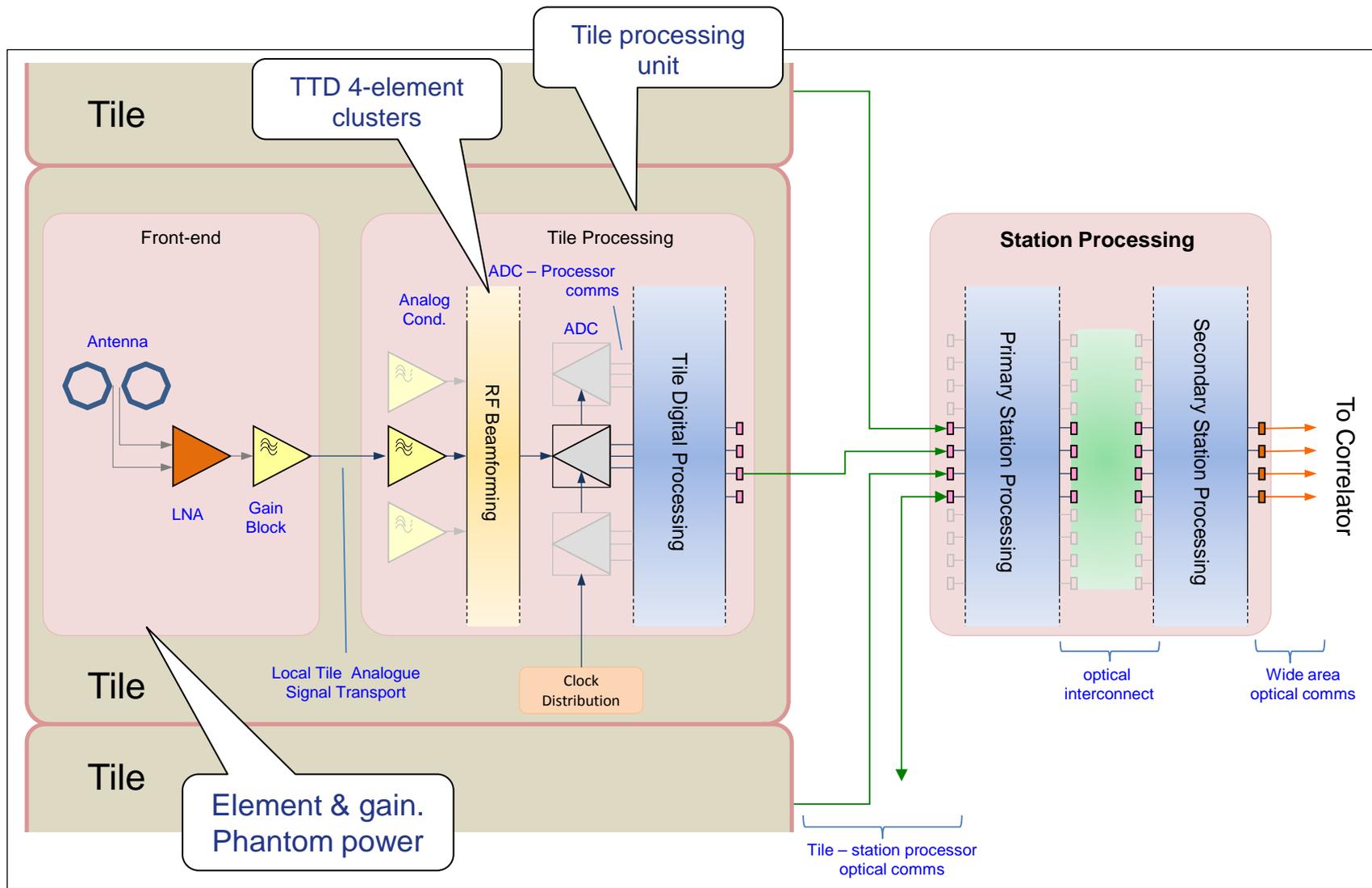
# AA-mid Array – SKA<sub>2</sub>



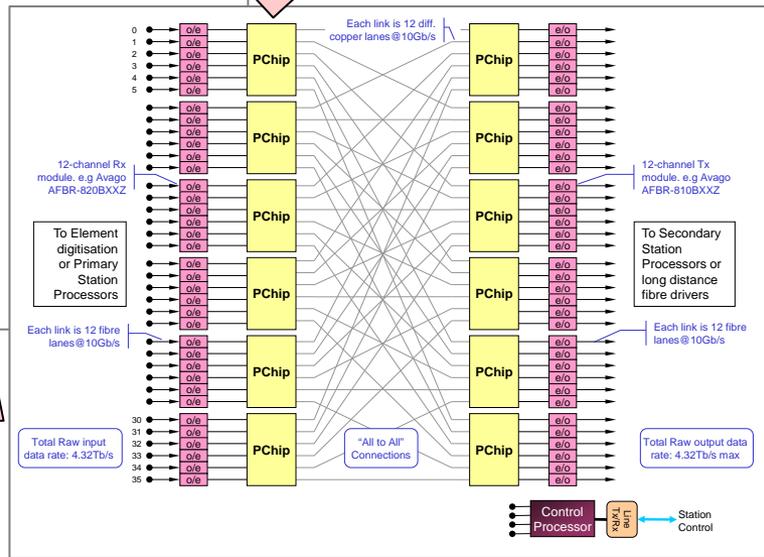
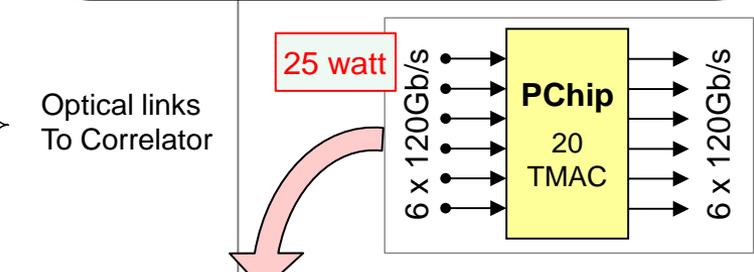
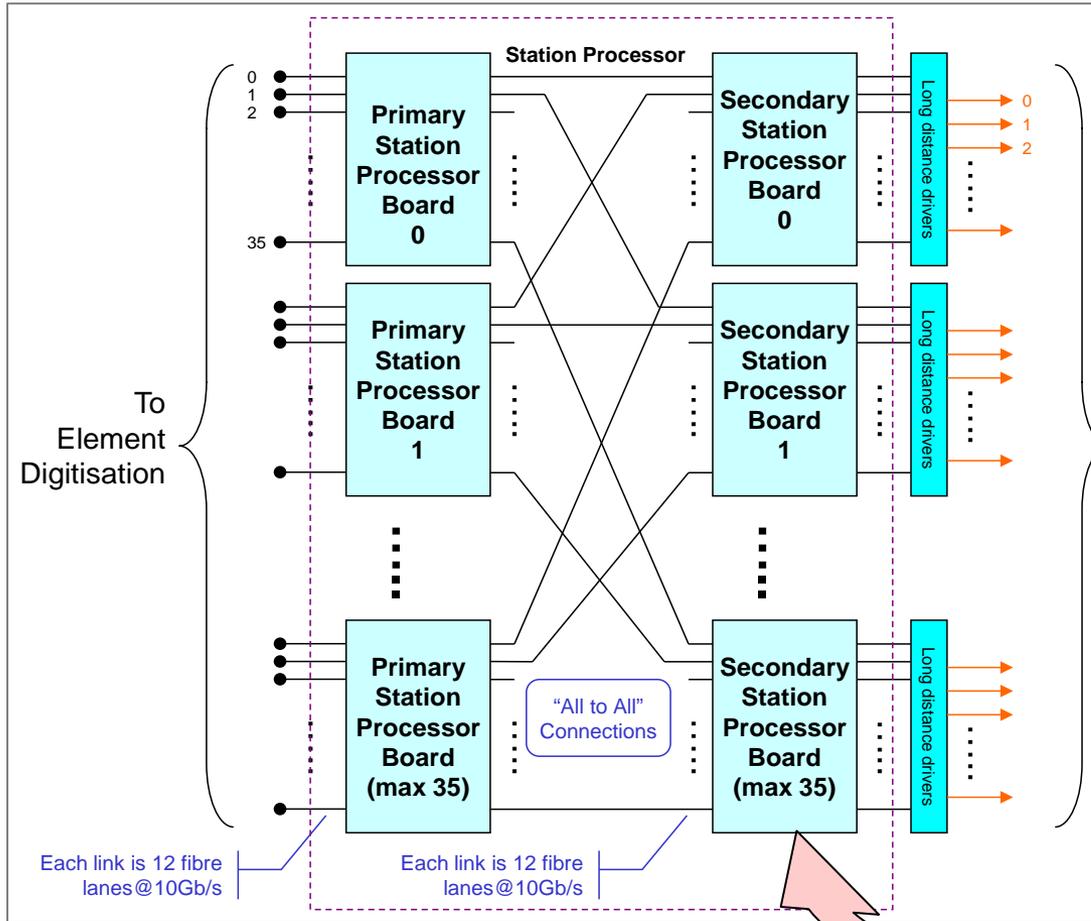
# AA-mid station design

Parameter	Value	Comments
Type of array	Single element	Dense array using Vivaldi or ORA.
Number of elements	<b>110,000</b>	<b>30 million total</b>
Pitch of elements	15 cm	$\lambda/2$ at 1000MHz
No. of polarisations	2	Each element has two receiver chains
Diameter of station	56m	
Cluster size	4 elements	Uses true time delay beamforming
Tile size	16 x 16 elements	Built out of 4 x 4 clusters
No. of Tiles	430	Each tile is ~2.4m square
Number of stations SKA <sub>2</sub>	250	Anticipated number of Phase 2 SKA Stations
Element communication	Copper	Includes Phantom power
Layout	Dense rectangular	Regularly spaced
Frequency range	400-1450 MHz	
Digitisation rate	3GSamples/s	There is no frequency conversion,
Digitisation depth	6/8-bit	Required for RFI environment at these frequencies
Beamforming technology	Digital	Using cluster outputs
Max inst. bandwidth	1000 MHz	Covers operating band of array
Max output data rate	<b>16Tb/s</b>	Organised as 4+4bit complex data

# AA-mid possible signal path

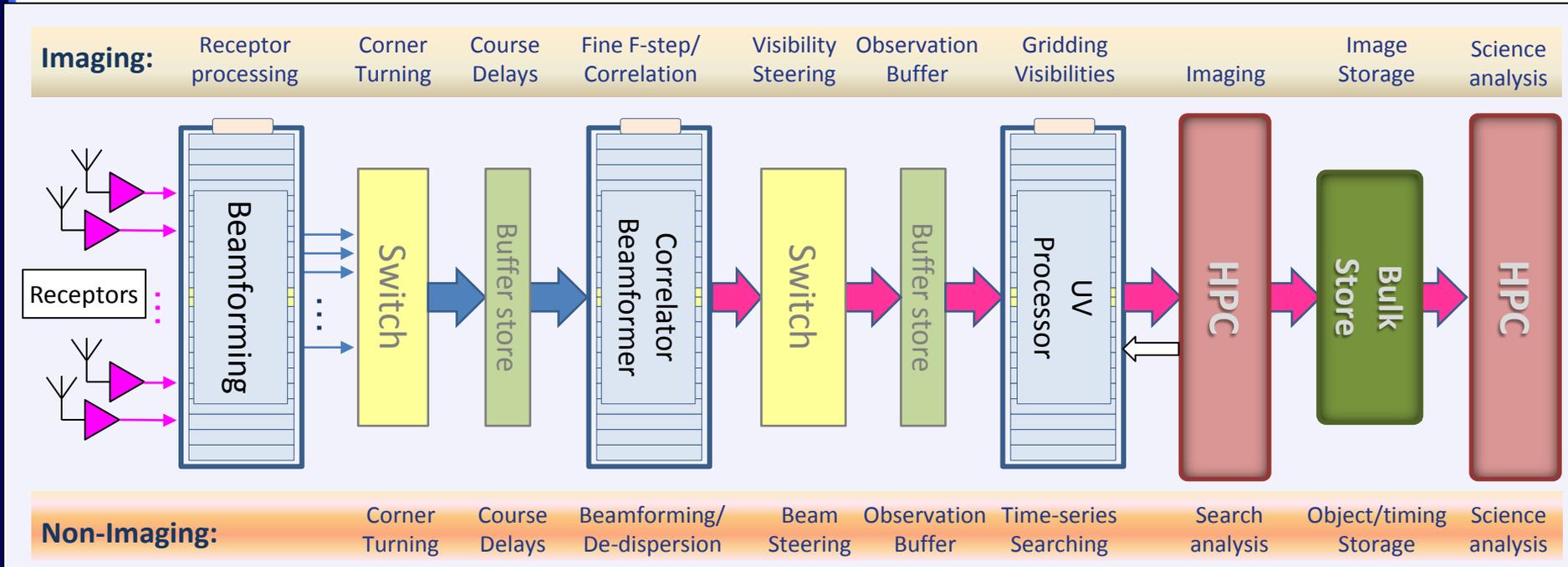


- Requirements:**
- High bandwidth in
  - High bandwidth out
  - Largely cross connected
  - Scalable at various levels
  - Programmable beamforming

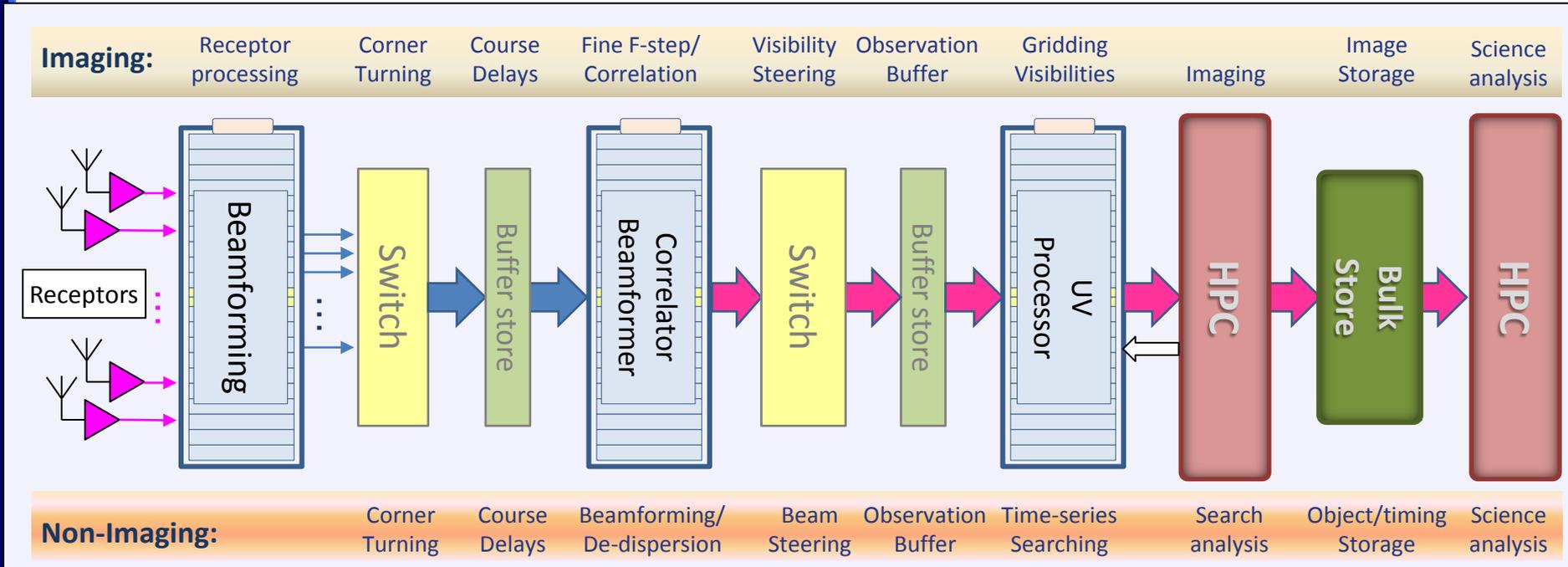


# SKA<sub>2</sub> AA Station processor

# Processing: signal to science



# Processing: signal to science



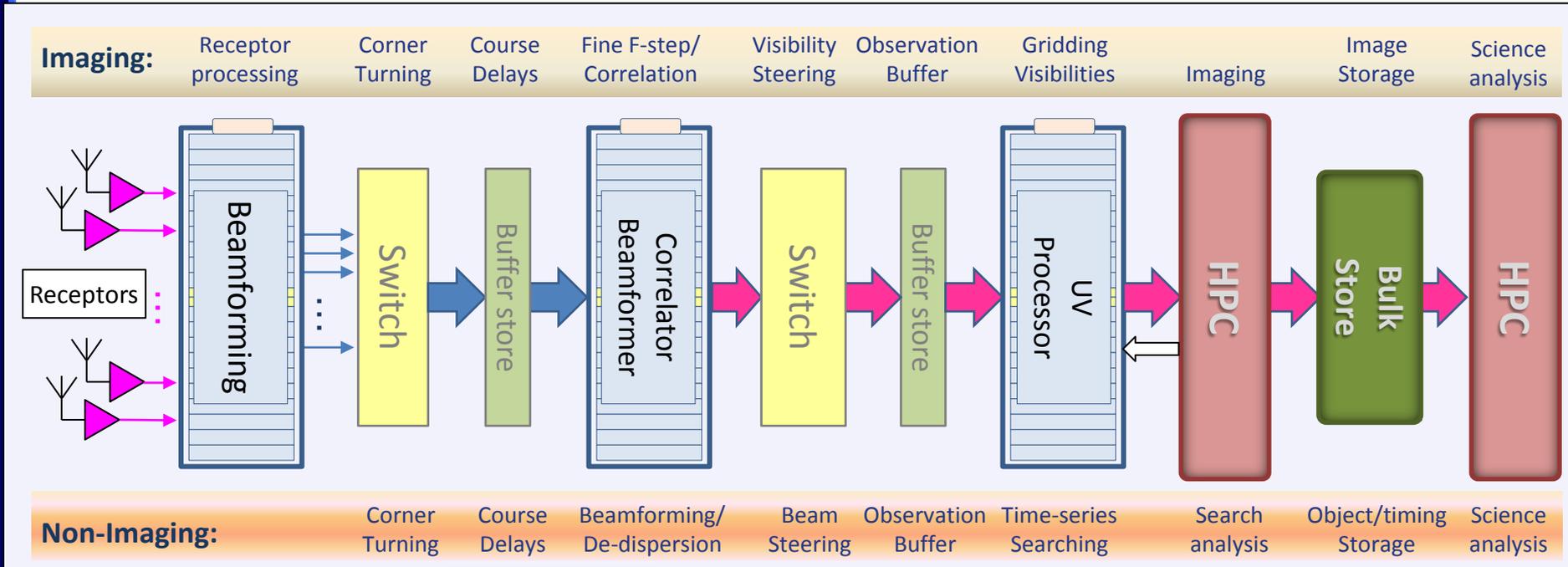
Metal Cable

Analog

Software

'60's – '70's

# Processing: signal to science



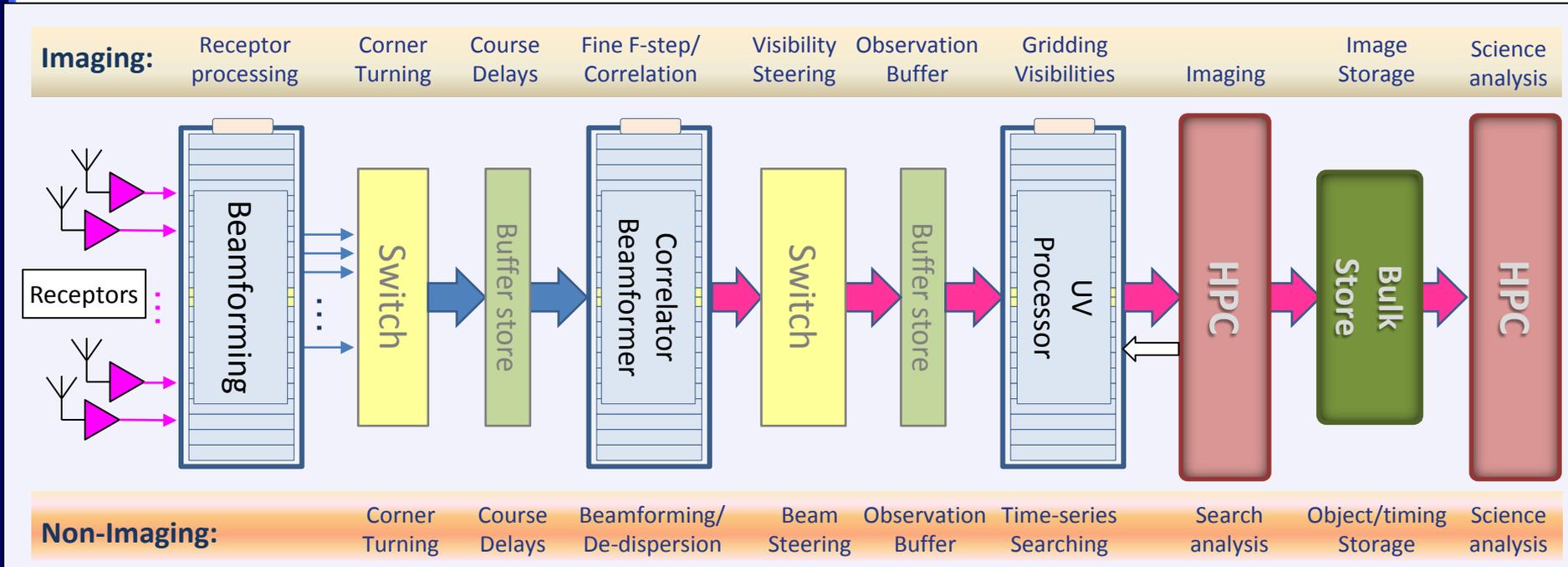
Metal  
Analog

Software!  
Custom ASIC

Software →

'70's – '80's

# Processing: signal to science



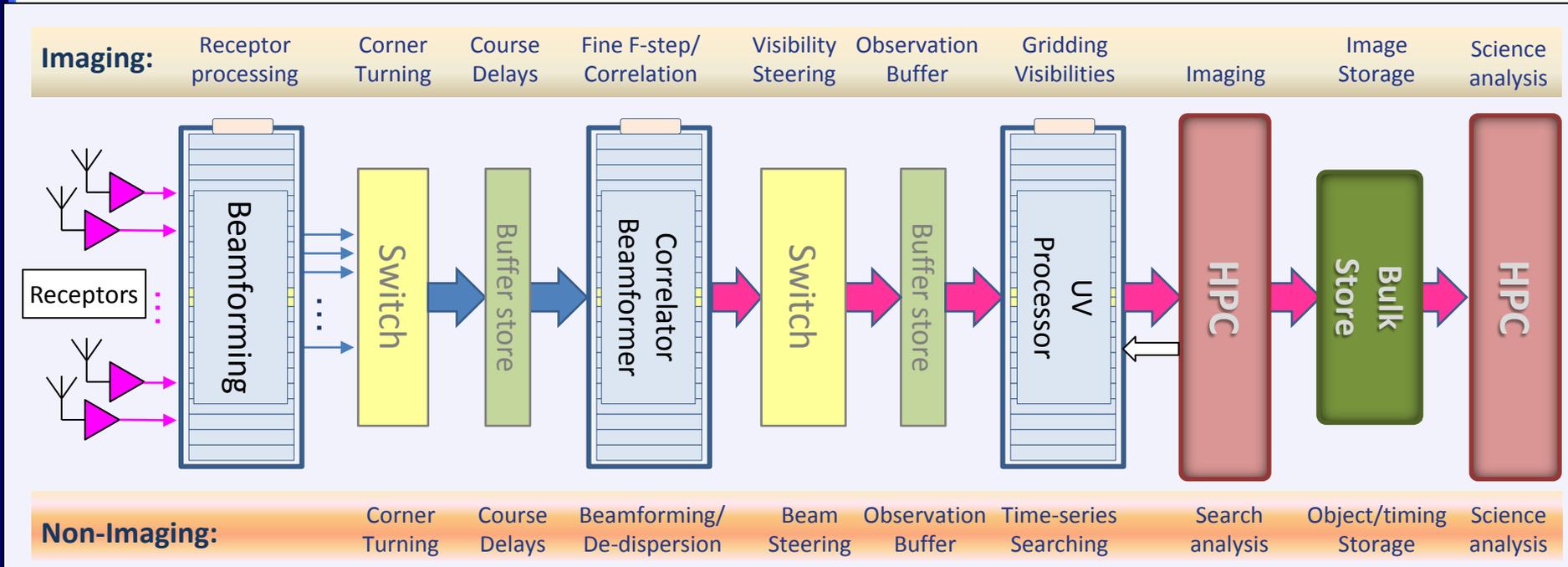
Metal  
Analog

ASIC  
FPGA

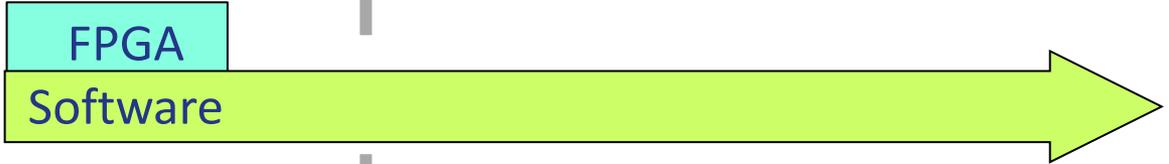
Software

'90's

# Processing: signal to science

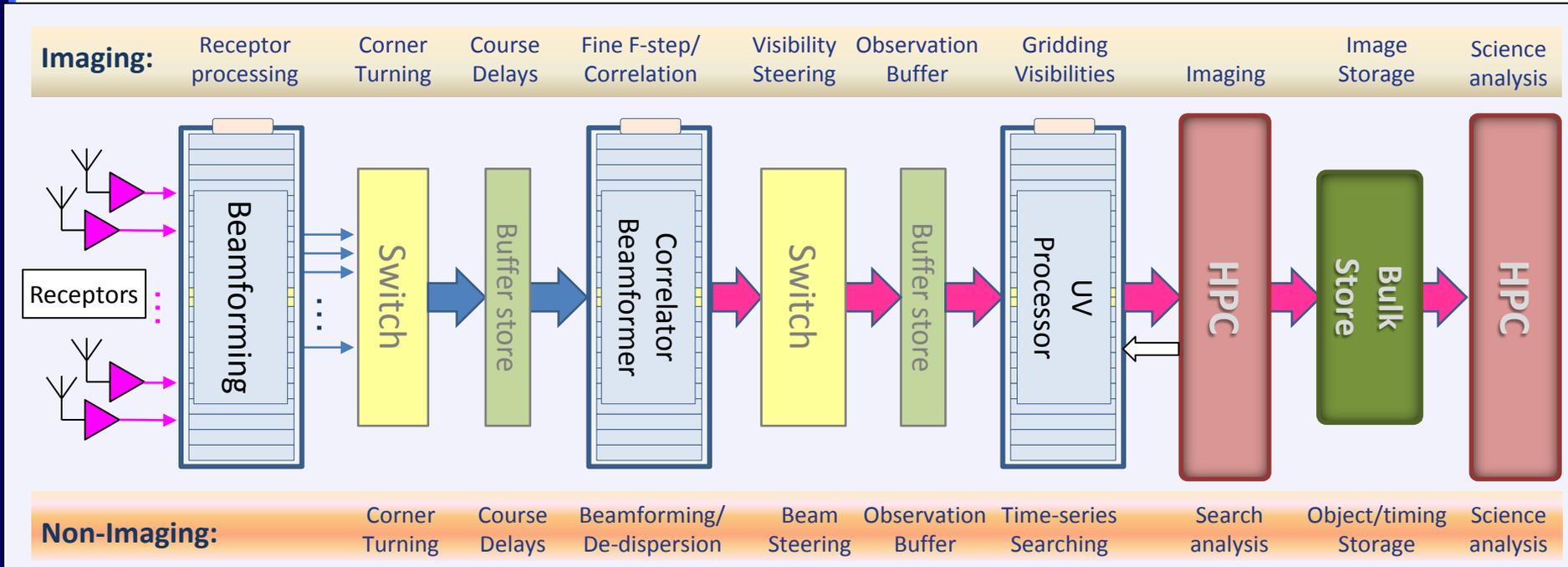


- Metal
- Analog
- FPGA



2000's

# Processing: signal to science

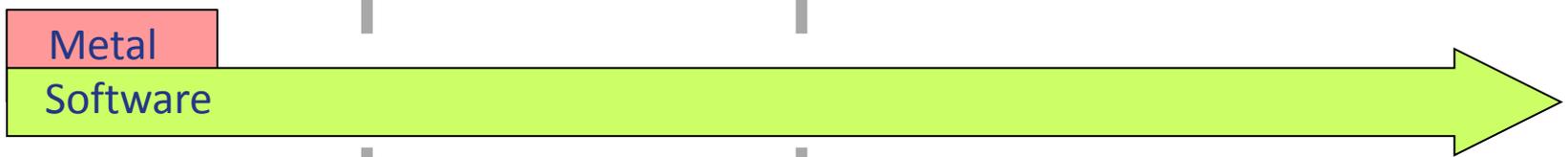
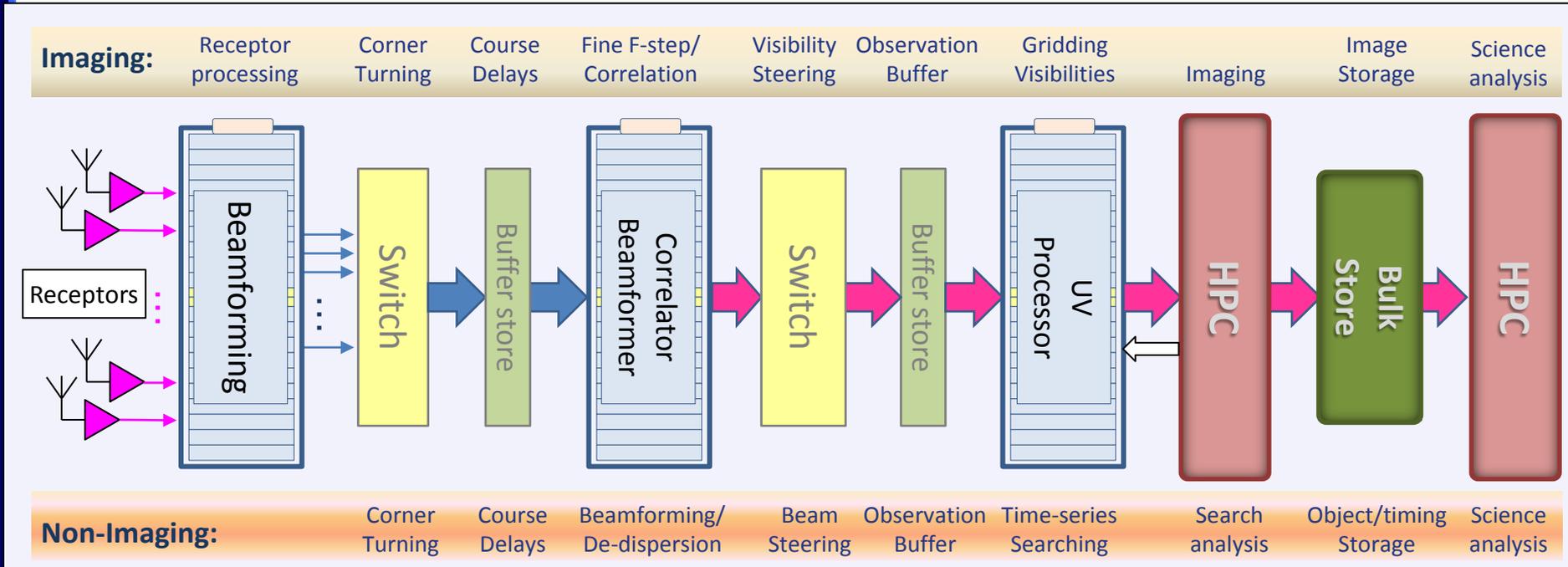


Metal  
FPGA

FPGA  
Software

2010 - 2020

# Processing: signal to science



Future

**And finally...**



**It's always fun travelling with Arnold!!**