Astronomy with the WSRT

Deep observations of the radio sky

The WSRT is often used to produce these deep images. The deepest so far is a region of about one square degree centred on the First-Look "verification" strip of the Spitzer IR Space Telescope (see figure).

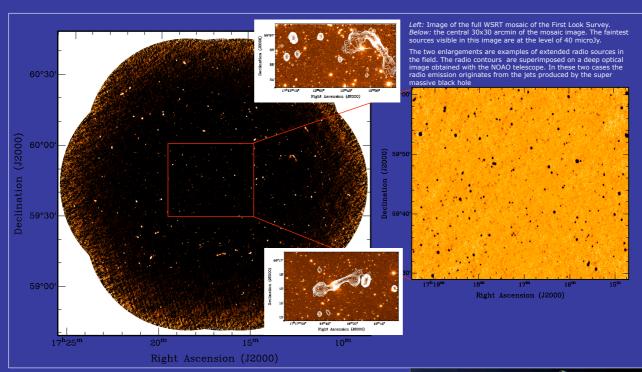
The observations made use of the 160 MHz broad-band IF system (8 bands of 20 MHz with 128 channels in each band) covering the frequency range 1311-1450 MHz. The final image (a mosaic of seven fields) reaches a noise level of 8.5 μ Jy - the deepest WSRT image made to date. More than

1000 sources have been detected in the field.

What do we want to learn?

- What kind of galaxies inhabit the early Universe? How different was the Universe then?
- When did the first galaxies and massive stars form ?
- At what rate did these galaxies form stars in the early Universe?
- Do they have different characteristics than nearby galaxies?

A key factor to be able to fully exploit deep radio surveys is the need for multi-wavelength source identifications at other wavelenghts e.g. optical/near IR. The distance of the radio sources can also be obtained via the redshift information derived from optical spectra or from radio and sub-mm spectral indices. indices.



Under the magnifiying glass:

a trick to observe the most distant radio sources.

The magnification from gravitational lenses makes it possible to do a detailed study of extremely **distant** systems that otherwise would be too faint to observe.

