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HI ABSORPTION OBSERVATION USING FAST

HI absorption meeting @ ASTRON Aug 29, 2018

An incomplete list of previous single dish studies

- Arecibo: 1 associated absorption in 8983 sources stronger than 7.7mJy in 517deg^2 (Darling+2011)
- HIPASS: 4 associated HI absorptions in 204 sources stronger than 250mJy (Allison+2014)
- GBT: 4 intervening absorbers out of 17 systems (Zwaan+2015)

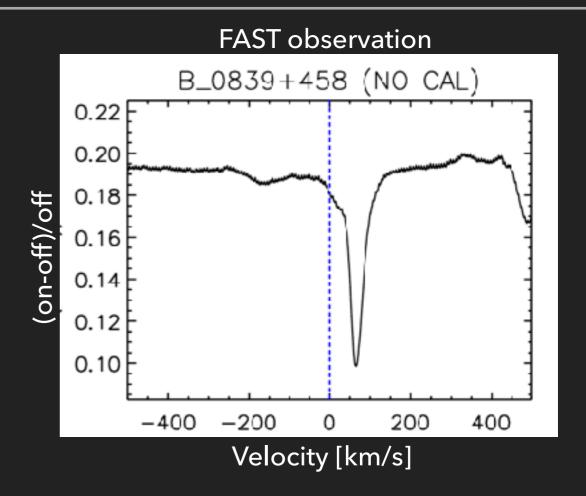
Brief summary of FAST status

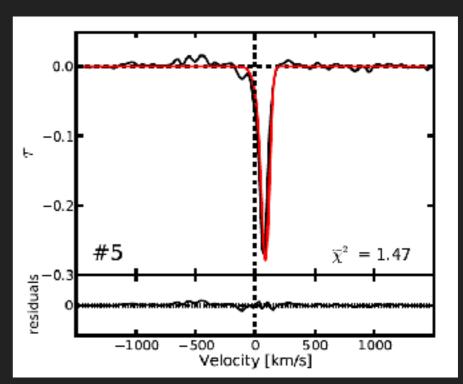
- ▶ Can track ~> 2hr
- Dec range:
 - 0 to 52 deg (full sensitivity)
 - -14 to 66 deg (partial sensitivity)
- ▶ Beam size: 2.9′
- ▶ 19-beam receiver is on
 - ▶ Tsys ~ 18 K
 - ▶ 1.05-1.45 GHz
- Ultra wide-band receiver
 - 0.27 1.62 GHz



FAST test observation

- ▶ B 0839+458
 - 331 mJy @ 1.4 GHz
 - z = 0.192
 - FAST observed for 18min (on-source)
 - WSRT observed for 4 hrs

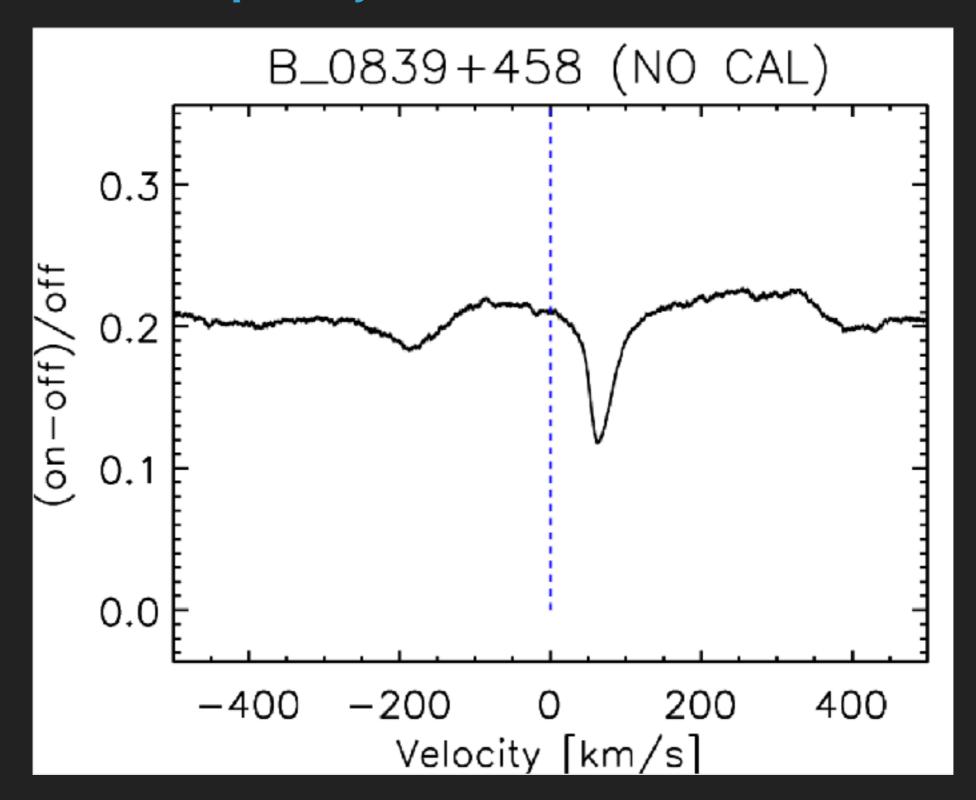




WSRT observation (Gereb+2014)

FAST drift scan capability

10s integration ~ <drift scan integration time



Commensal Radio Astronomy FAST survey (CRAFTS)

- Drift scan using the 19-beam receiver
- 220 full days to cover the FAST sky between
 -14° and +66° of DEC

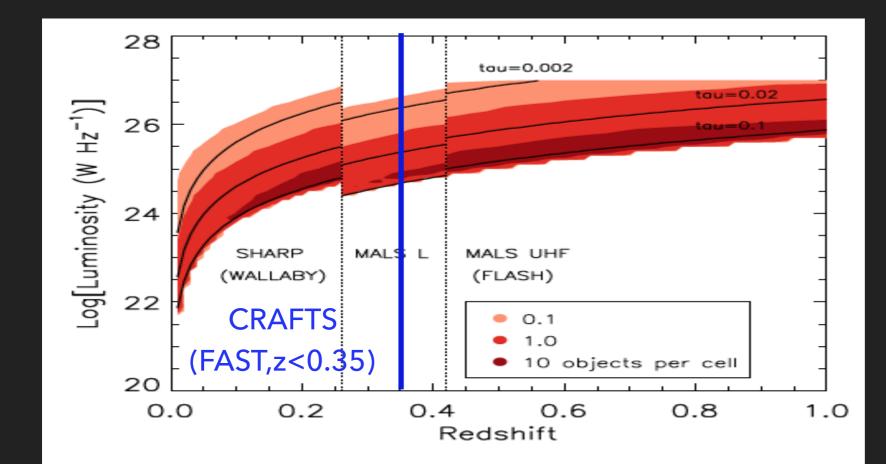


Li+2018

http://crafts.bao.ac.cn/pulsar/

UPCOMING HI ABSORPTION SURVEYS & FAST-CRAFTS

Table 3: Summary of various upcoming H+21 cm absorption line surveys						
Survey	Redshift	Time per pointing	Spectral r.m.s.	Sky coverage	Total time	Number
-	[H ₁ 21 cm]	[hrs]	[mJy]	[deg ²]	[hours]	of lines of sight
Apertif - SHARP	0-0.26	12	1.3	4000	6000	25000
-						(> 30 m J y)
ASKAP – FLASH	0.4 - 1.0	2	3.8	25000	1600	65000
						(> 90 mly)
ASKAP – Wallaby	0-0.26	8	1.6	30000	8000	132000
						(> 40 mly)
MeerKAT – MALS	0-0.57	1.4	0.5	1300	1333	16000
(Lband)						(> 15 mly)
MeerKAT – MALS	0.40 - 1.44	1.7-2.8	0.5-0.7	2000	2125	33000
(UIII4-band)						(> 15 mJy)



- FAST-CRAFTS
- > z: 0-0.35
- integration time: ~15s
- rms: 5mJy?
- sky coverage: wholesky from -14 to 66 Deg

Maccagni+2017

FAST observation plan I: drift scan

- Yu+2017: very preliminary CRAFTS drift scan estimation
 - rms ~ 2.3 mJy
 - for 10sigma detection, only F > 0.11 Jy background sources can be detected
 - ~ 10,000 luminous (F>0.11Jy) radio sources/month
 - detect ~ 200 sources/month

FAST observation plan II: tracking

- search for HI in promising sources such as MgII absorbers and DLAs
- search for OH absorption lines in detected 'high redshift'
 HI absorption systems

FAST observation problems

- ▶ RFI
- Baseline ripples, instabilities
- large overhead for switching RA&Dec

SUMMARY

- HI absorption survey is one of the FAST key projects
- FAST will be mostly in drift scan survey mode (CRAFTS) in the next couple of years but now we are mostly doing tests in tracking mode.
- We have done a few HI absorption observation tests and found that FAST is promising in detecting extragalactic HI absorption systems given the absorption redshift.
- RFI and baseline could be big problems.