

AST(RON



Netherlands Institute for Radio Astronomy

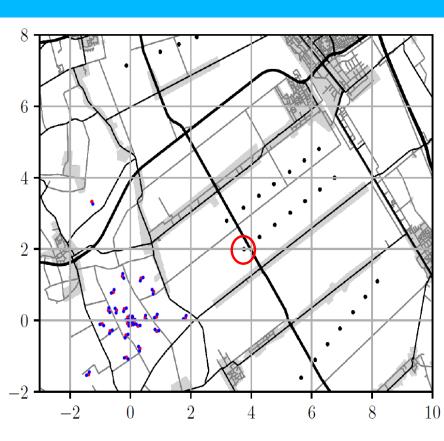


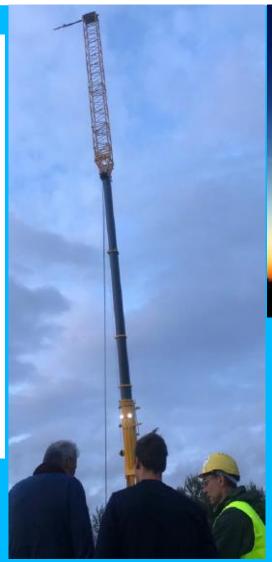
Outline Talk



- Status of Windmill project
- Various solar farms around core and remote stations











- First NORDEX N131/3.9 MW prototype WTG including infrastructure (a few million Euro)
- Build completely from scratch with all possible EMC measures
 - Very good emc engineers/ consultants
 - German university support
- Goal is a level of 35 dB below CISPR-11.

(50dBuV/m @ 10 meter distance at 100 meter height in 120 kHz BW (average) in the 30-240MHz frequency band)

- When pass, the other 44 WTG could be build.
- NORDEX EMC-Dossier under NDA (+190 pages)
 - Assessment of EMI classes
 - A probably no problem (conventional lamp)
 - B might be a problem (control cabinet)
 - C for sure could be a problem (inverter)

For type B and C they have done EMC emission measurements in test labs.

The EMI should be below a certain level including some safety margin.

When a component failed the test, mitigation measures where taken and re-measured.

This systematic approach gives better isolation of the systems and easier to improve.

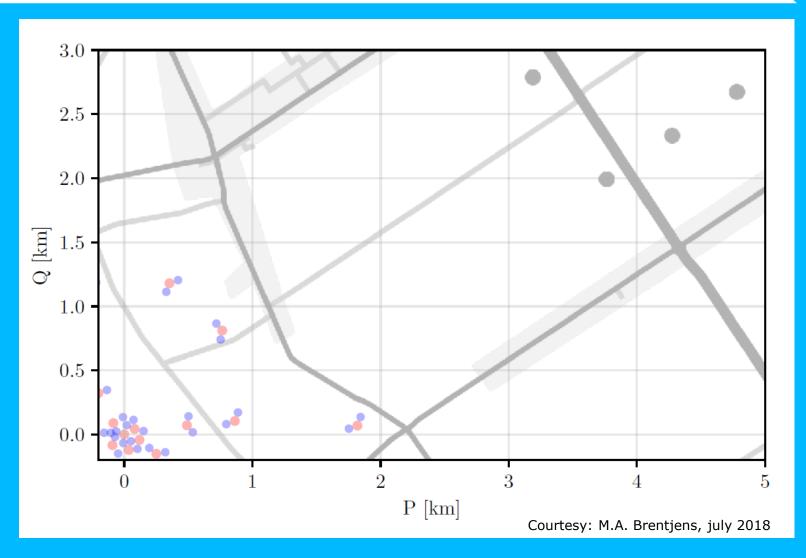


- Several test observations have been executed (Astron, AT, S&T)
 - Check the duration of battery lifetime
 - Remote switching of 0 and -35 dB level
 - Different comb frequencies (100kHz, 1MHz)
 - Test the dedicated LOFAR observing mode
 - Parallel processing of the recorded data (long queue)
 - External company S&T has to post process and image the data
- Reference transmitter places at 100 meter height crane!
 - Both LBA and HBA test observations
- Official measurement campaign first two weeks of September 2019
 - Measurements in a few subbands in LBA and HBA
 - Same subbands are used by reference signal from transmitter
 - WTG and reference transmitter are special isolated by a few hundred meter
 - 0 dB level 1 or 2 seconds is sufficient, for -35 dB you have to integrate minutes
 - Also you need to flag RFI and subtract strong radio sources (A-team,..)
 - The residual power of the strong sources at the location of the windmill should be sufficient low to detect the windmill RFI at -35 dB level with sufficient S/N

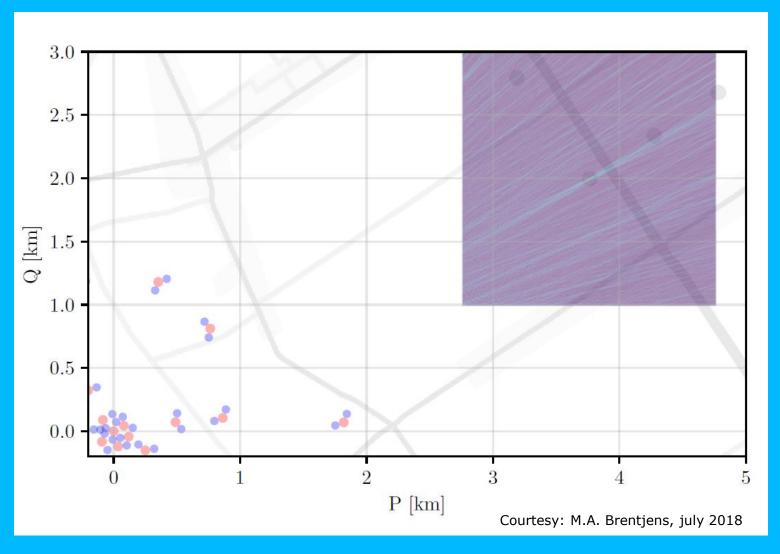


- The WTG is measured in three different modes
 - Operational mode
 - Minimum mode
 - Completely switched off
- We expected final results before end of the year.
- When less than 35 dB reduction, modifications are needed and new measurement campaign.
- When more than 35 dB reduction, they can start building the rest.
- The 50 dB reduction is difficult or impossible to measure with LOFAR.

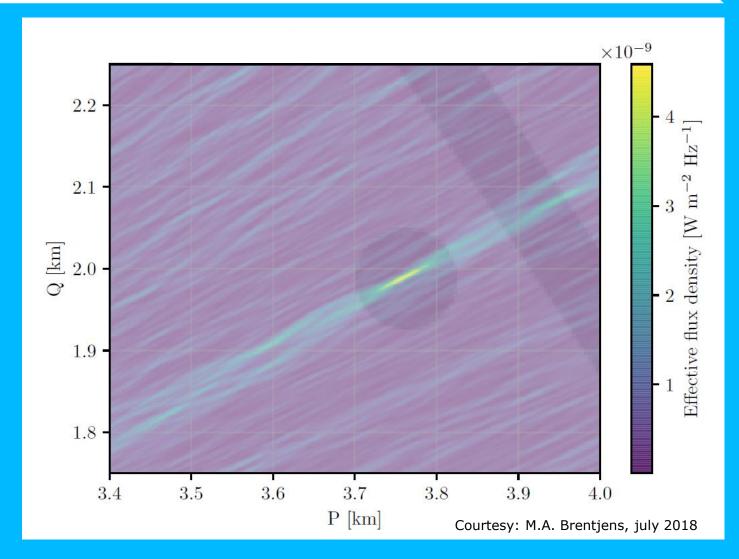




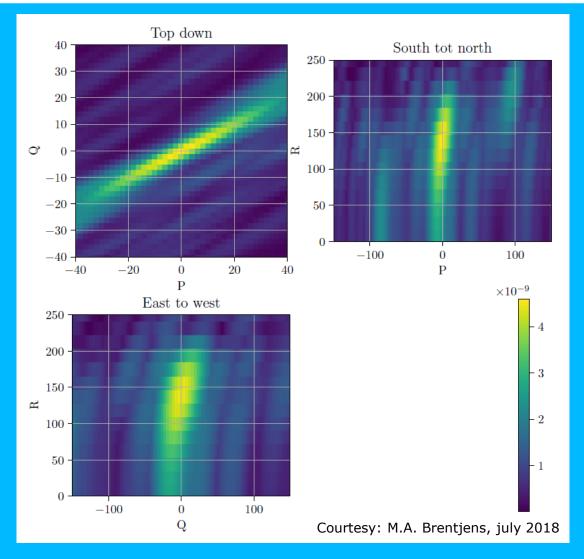












Conditions for LOFAR radio telescope and wind farm co-existence **ASTRON**

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LOFAR



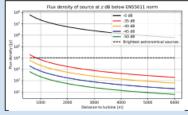


Core area: 3,456 phased array antennas on 4 km² area

Total: 6,336 antennas Antennas grouped in 52 'stations'

Imaging by using radio interferometry Transient research by using tied-array beams

Windfarm challenge



Plans for 45, ~240 m tall wind turbines near LOFAR core (~4-10 km from Superterp), max.

Covenant requiring radio-quiet wind turbines

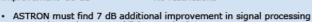
Covenant

EM interference reduction < 35 dB:

35 dB < improvement < 40 dB 40 dB < improvement < 50 dB improvement 50 dB

Consequence

No permission to operate 56-62 12 h idle Reduced idle time to be negotiated No restrictions





In case of conflict: binding arbitration

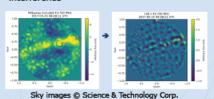


Measurement approach

- Use imaging radio interferometer plus cal source to measure below ambient noise levels
- · Use near-field visibility models:

$$V_{ij} = \sum_{k=1}^{K} I_k e^{2\pi i \nu (\vec{q}_{ij} \cdot \vec{l}_k)/c} \implies V_{ij} = \sum_{s=1}^{S} \sqrt{I_{si} I_{sj}} e^{2\pi i \nu (||\vec{r}_{sj}|| - ||\vec{r}_{si}||)/c}$$

 Sensitivity determined by ability to subtract astronomical sources and unrelated interference



Field measurements

Test set-up verification done using drones, to check system parameters such as

- integration time and number of antennas
- source and interference subtraction capability

Measurements on-going

- biconic transmit antenna mounted on 100 m tower near turbine
- LOFAR stations in near-field imaging mode





Poster @ RFI 2019

Toulouse, France September 23-26







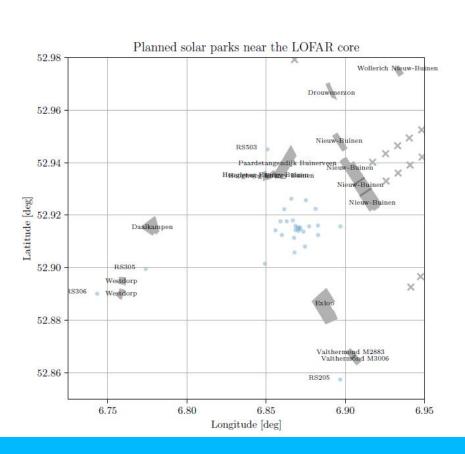






Solar plant emission



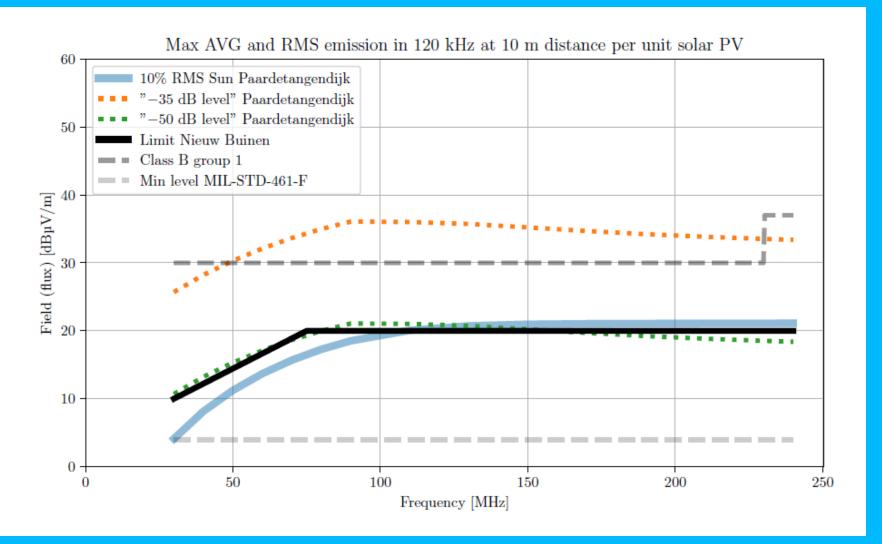




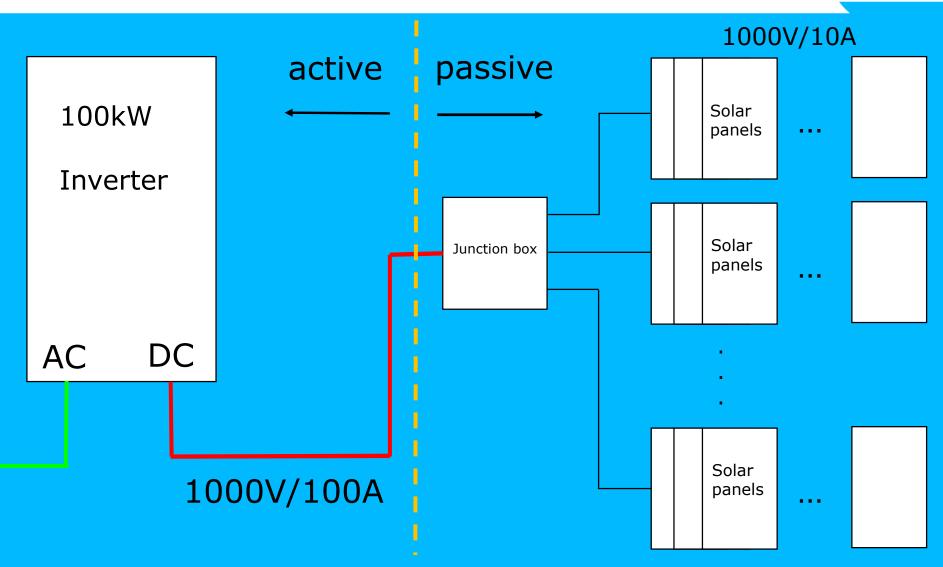
2 - 100 MW installations (total 250 MW around core)

Solar plant emission

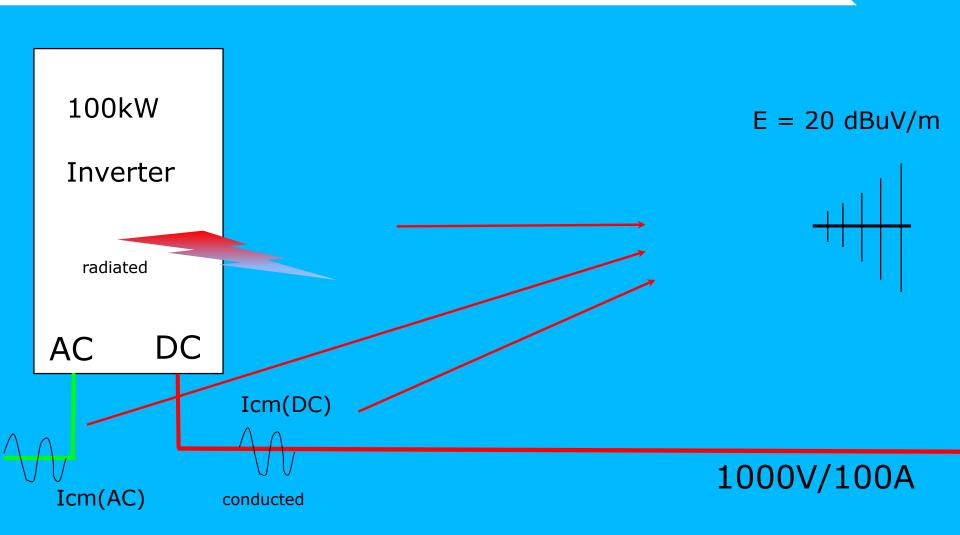




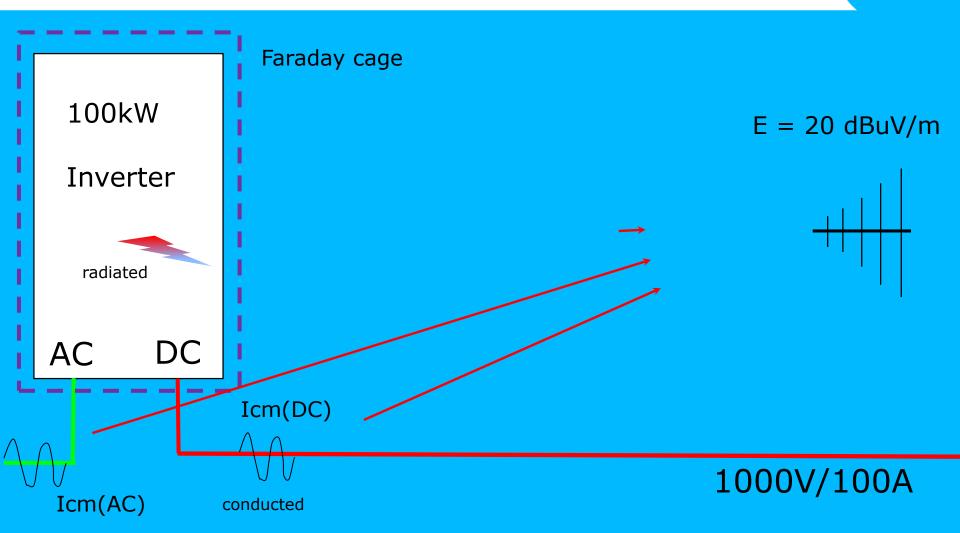




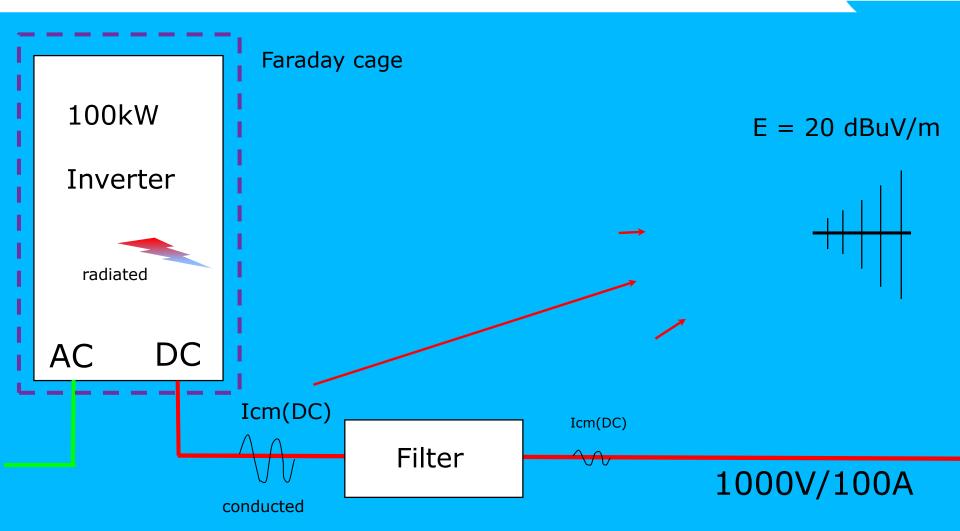




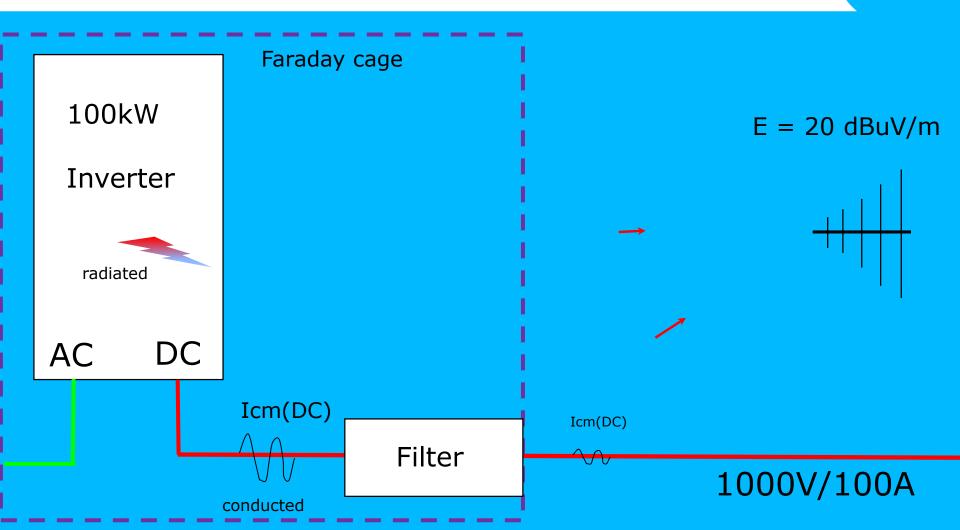












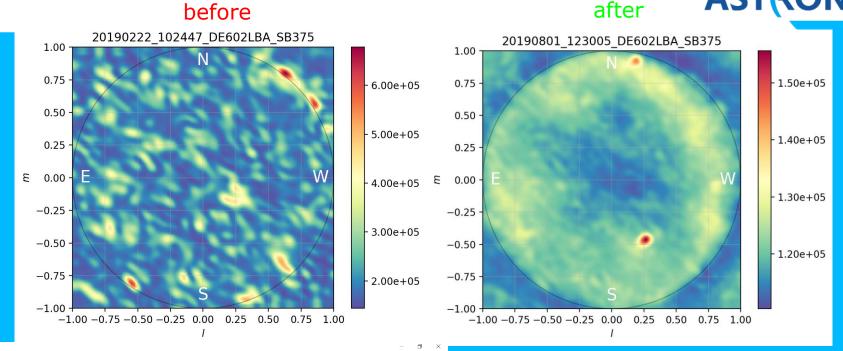
Faraday cage examples

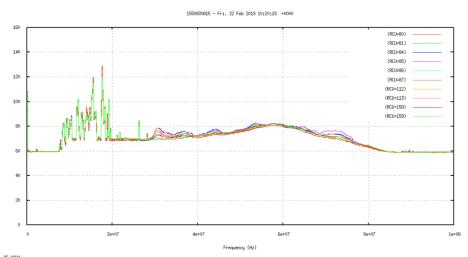




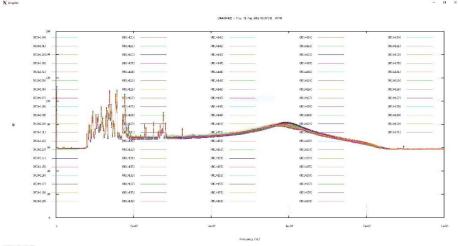
Faraday cage result

after AST(RON





M Gnuplot



Implementation example

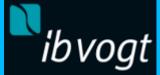
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Measurement 30&31 Sept 2019







Conclusions/Remarks



- The total radiated PV unit emission should stay below the covenant limit.
- Both radiated and conducted emission need to be considered.
- Also Cameras, LED equipment, controller EMI should low.
- The common mode current should be as low as possible on the AC/DC cables.
- The loop area of the cables should be as small as possible. Shielding or use the construction to reduce emission.
- A faraday cage can reduce the radiated emission (look out for slits!). No longer than 0.2m (preferably smaller).
- Create a technical construction file with all technical details and measurements (before/after).