

Tracing HI Gas Cycles and Global Star Formation with the ALFALFA H α Survey

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Galaxy Gas Cycles

Inflow



HI reservoir -raw material

Reaccretion?



Feedback



HI Mass



Radial flows and conversion to H₂

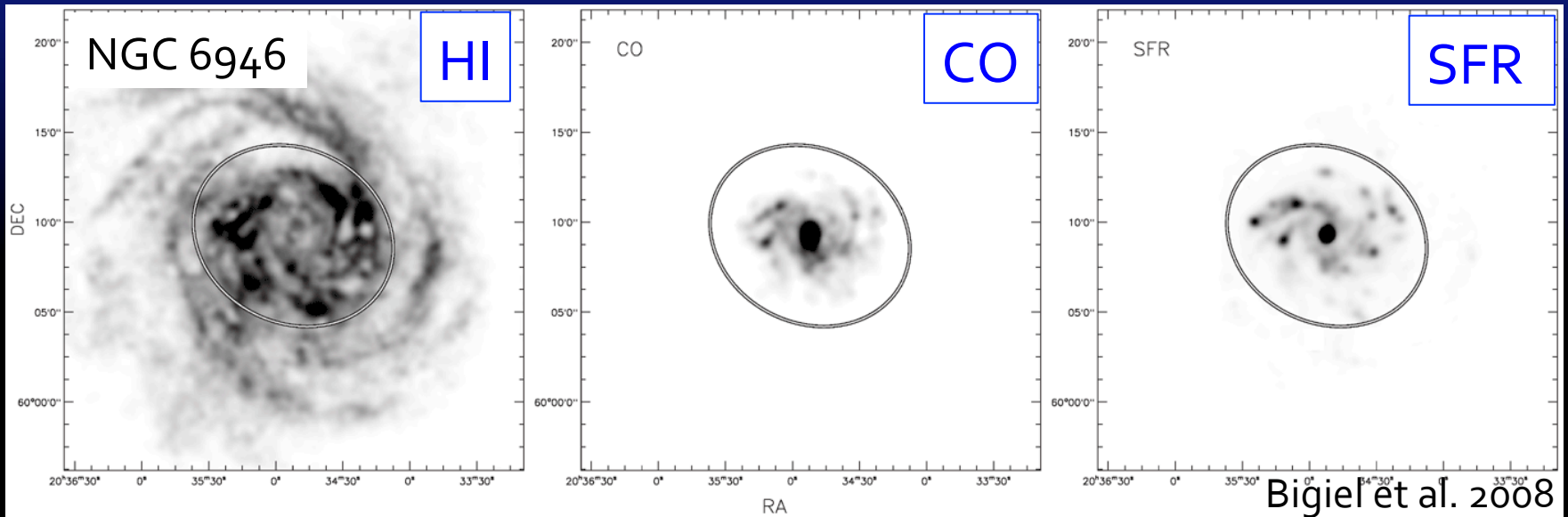


Star Formation



Ionization, Outflows, Photodissociation

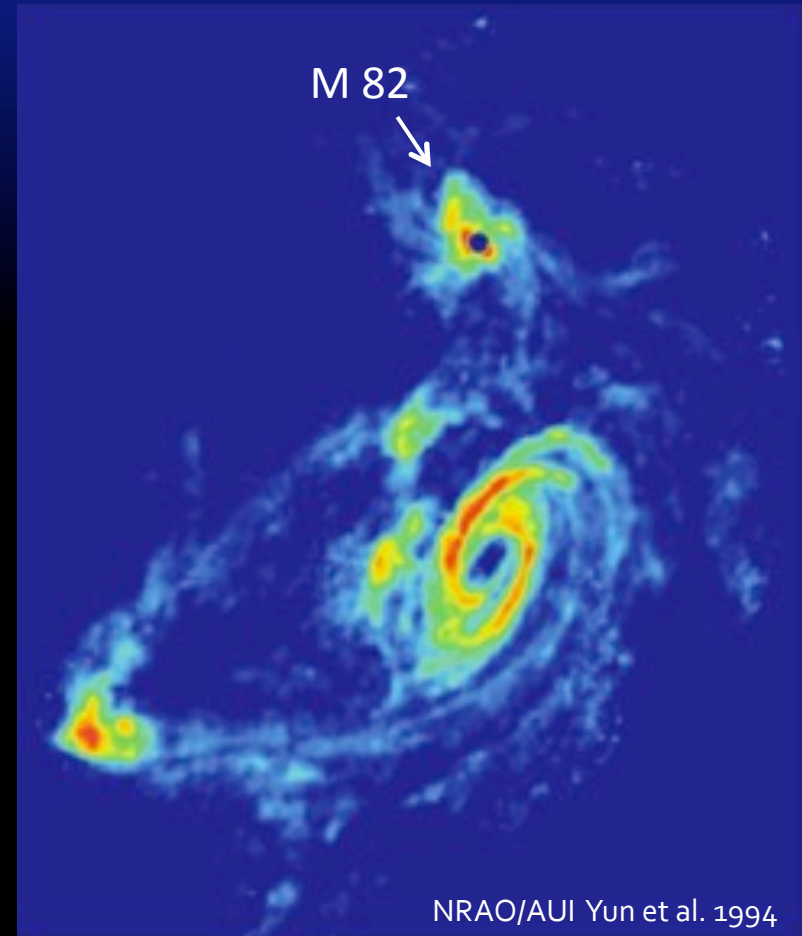
Global HI



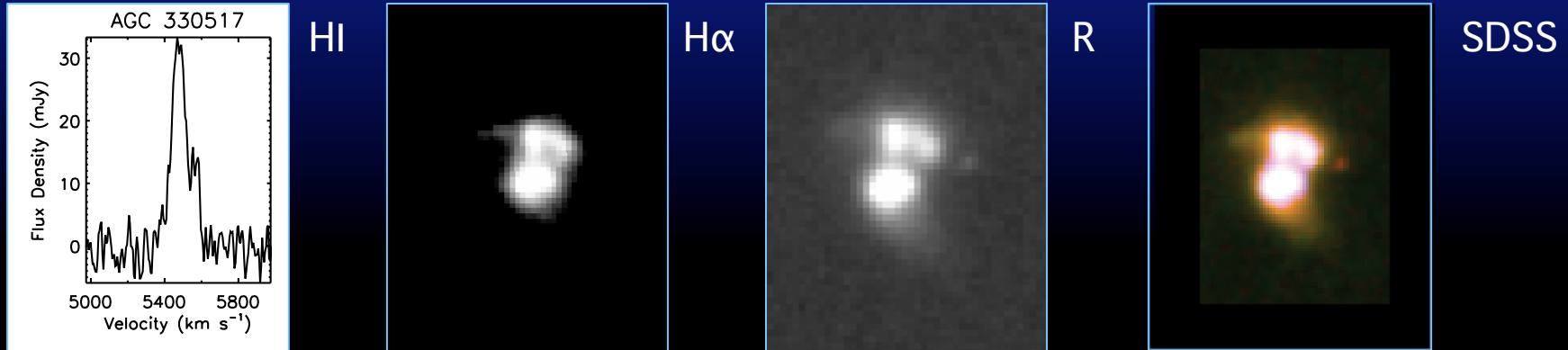
- HI mass traces
 - total gas supply
 - potential for future SF
 - recent accretion (*e.g. Moran et al. 2012*)
- More HI = more total SF
(*e.g. Catinella et al. 2010, Huang et al. 2012*)

HI at high SFR

- Starbursts are HI-rich?
 - Gordon & Gottesman 1981; Catinella et al. 2010
- Starbursts are HI-poor?
 - Oey et al. 2007
- May reveal disturbances



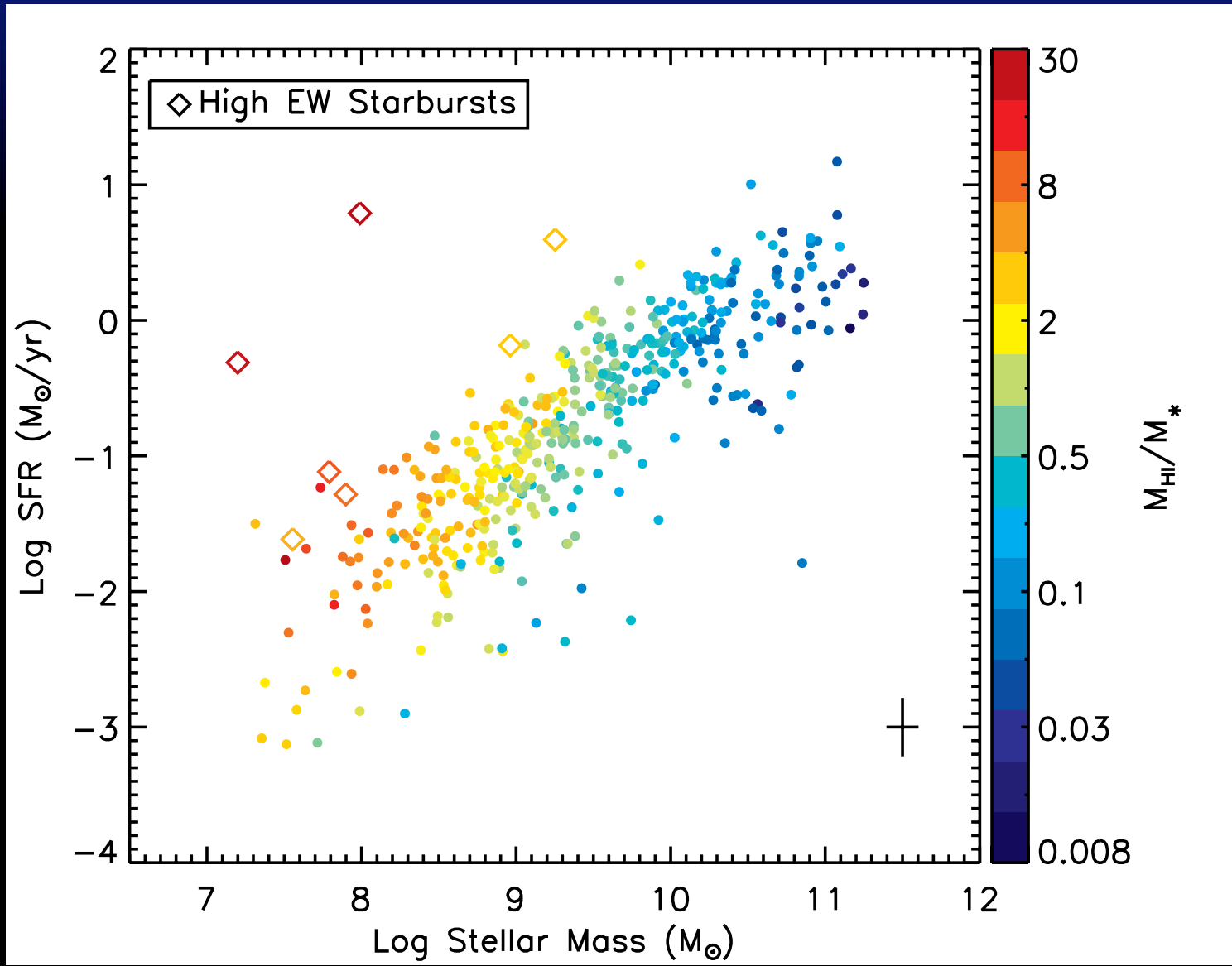
ALFALFA H α



- ALFALFA: Arecibo Legacy Fast *ALFA* Survey
 - Blind, volume-limited 21 cm HI survey; R. Giovanelli, P.I.
 - > 30,000 detections, 7000 square degrees
- ALFALFA H α
 - Volume-limited subset of ALFALFA
 - 20-100 Mpc
- KPNO: H α and *R*-band imaging
- Fall sample (565 galaxies) complete
- Starbursts: H α equivalent width (EW) > 80 Å

From HI to Stars

See talk by Saintonge

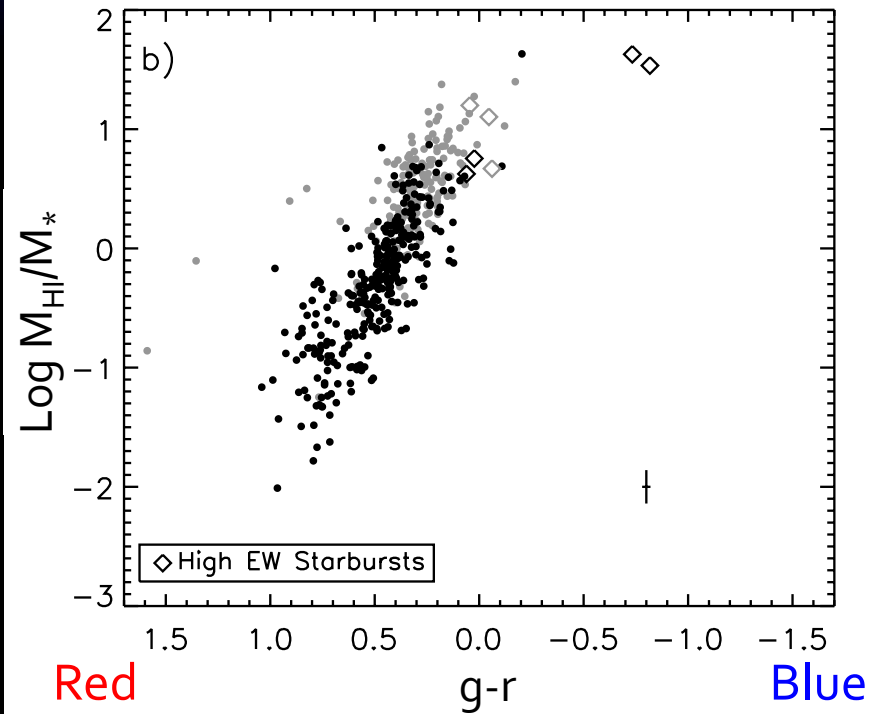
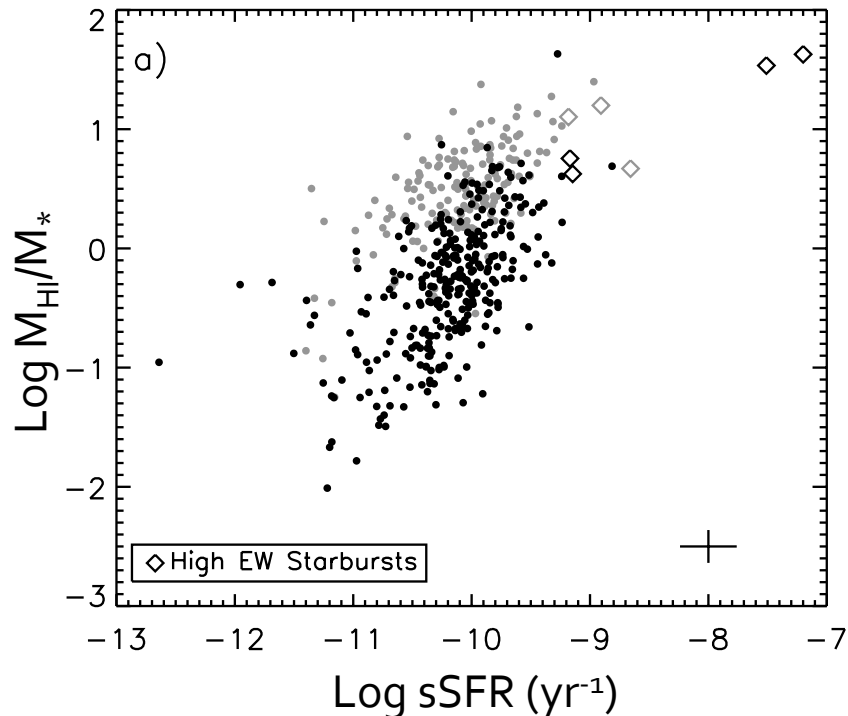


The HI and SFR Connection

e.g., Catinella et al. 2010;
Huang et al. 2012

Weak...

or strong?

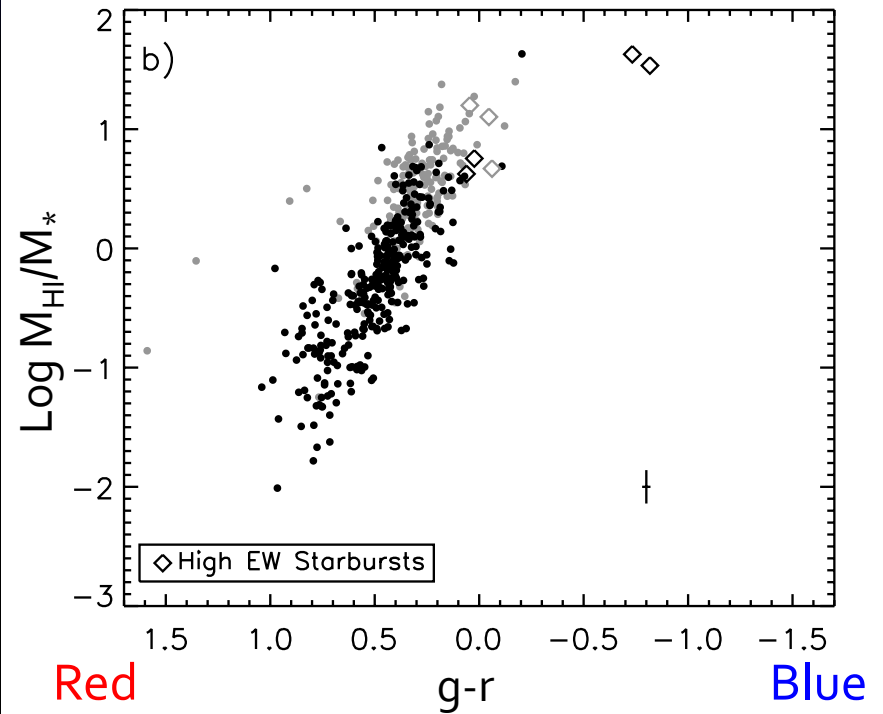
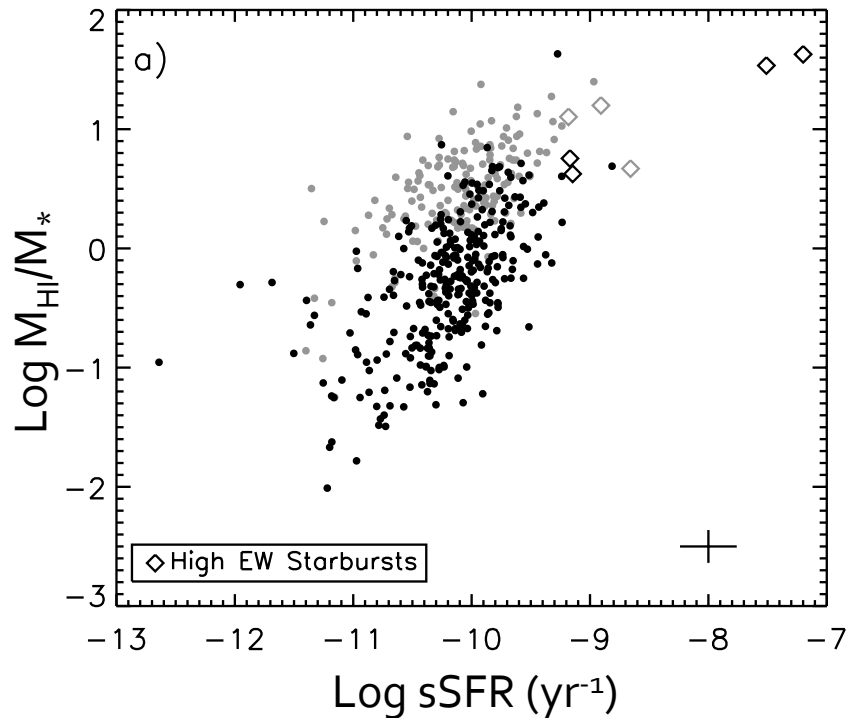


1. Link between HI and long-term averaged SF?
2. Link between HI and dust extinction?

The HI and SFR Connection

Following Wen et al. 2014

Before dust correction

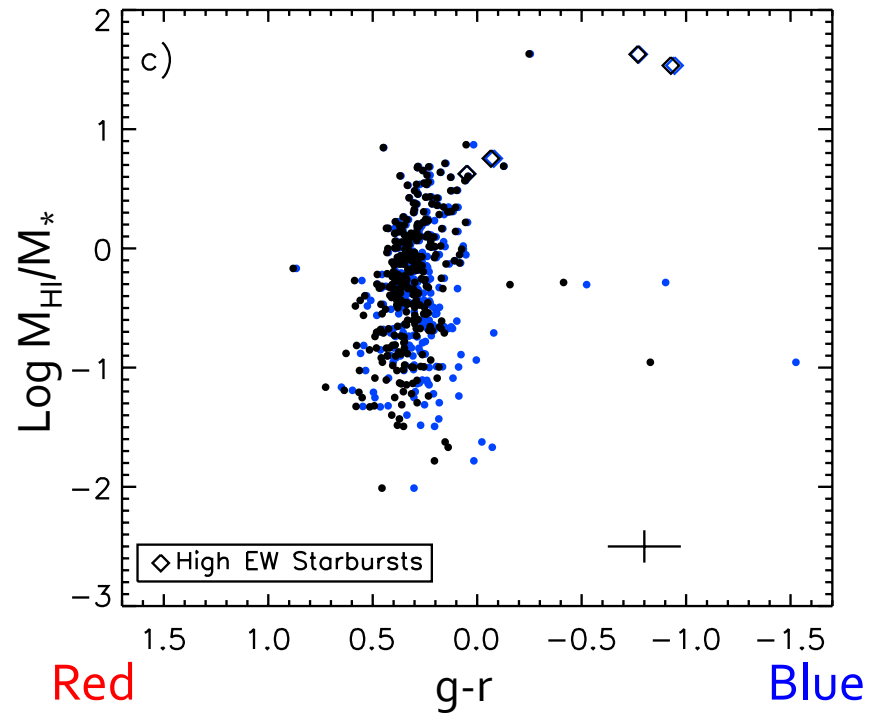
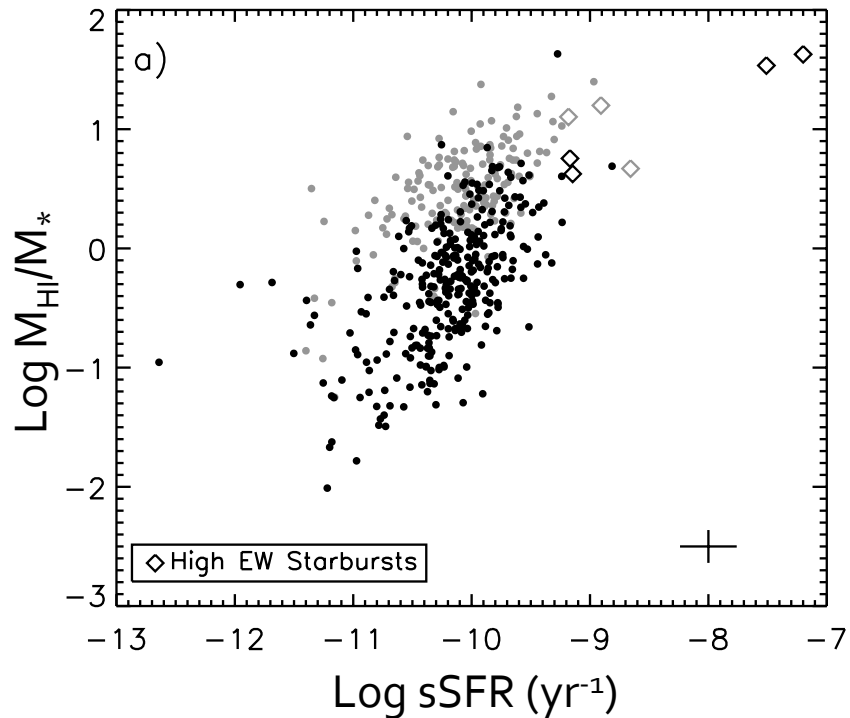


Mass-Metallicity-HI Relation (Bothwell et al. 2013)

The HI and SFR Connection

Following Wen et al. 2014

After dust correction



Mass-Metallicity-HI Relation (Bothwell et al. 2013)

HI and Metallicity

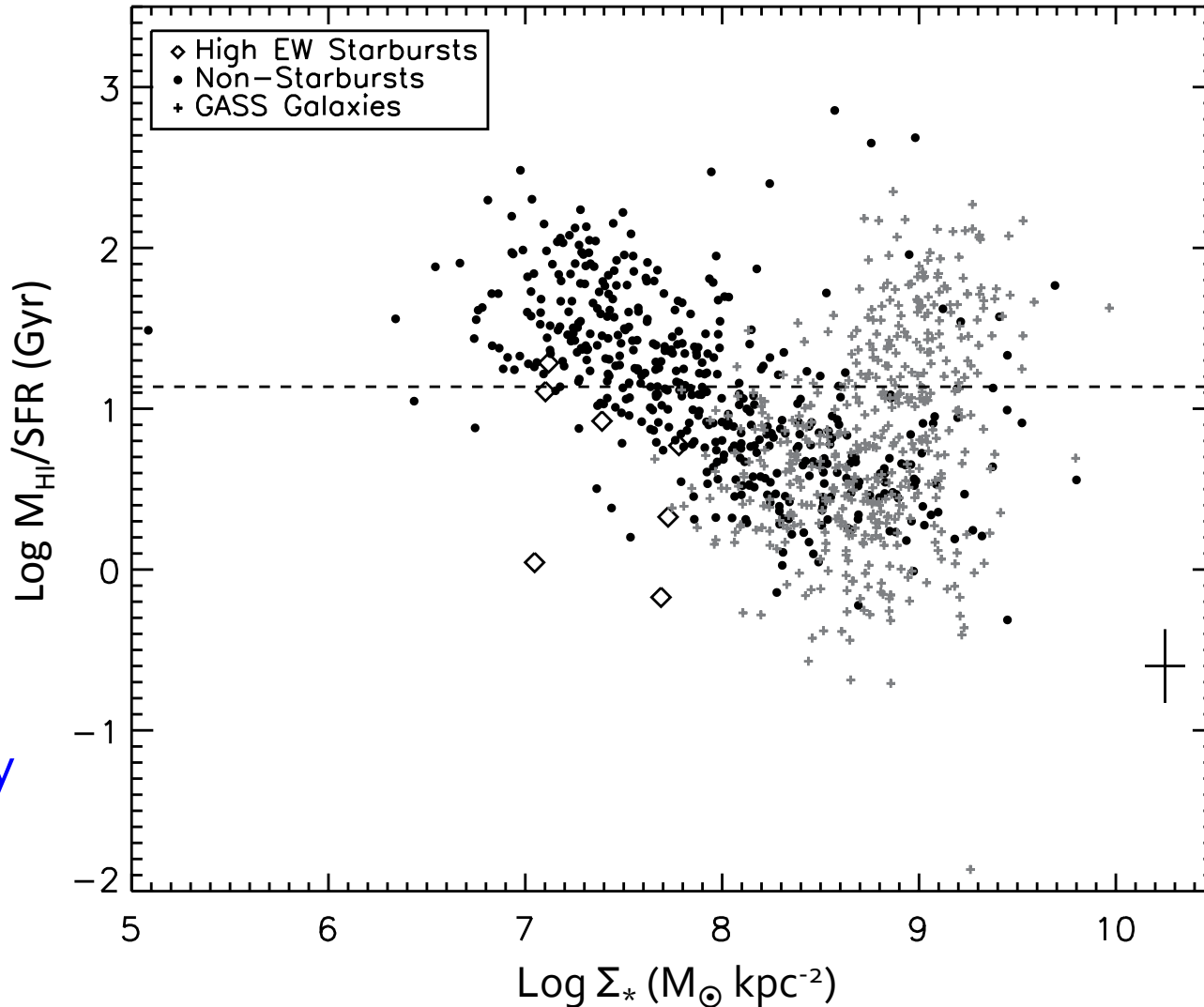
- Mass-metallicity relations
 - *e.g., Mannucci et al. 2010; Davé et al. 2011; Lilly et al. 2013*
 - Gas consumption
 - Metal retention
 - Dilution from accreted HI
- Shielding of H₂ (*e.g., Krumholz et al. 2009; Bolatto et al. 2011*)



HI Consumption Timescales

Higher P =
Higher H₂/HI
e.g., Blitz &
Rosolowsky
2006

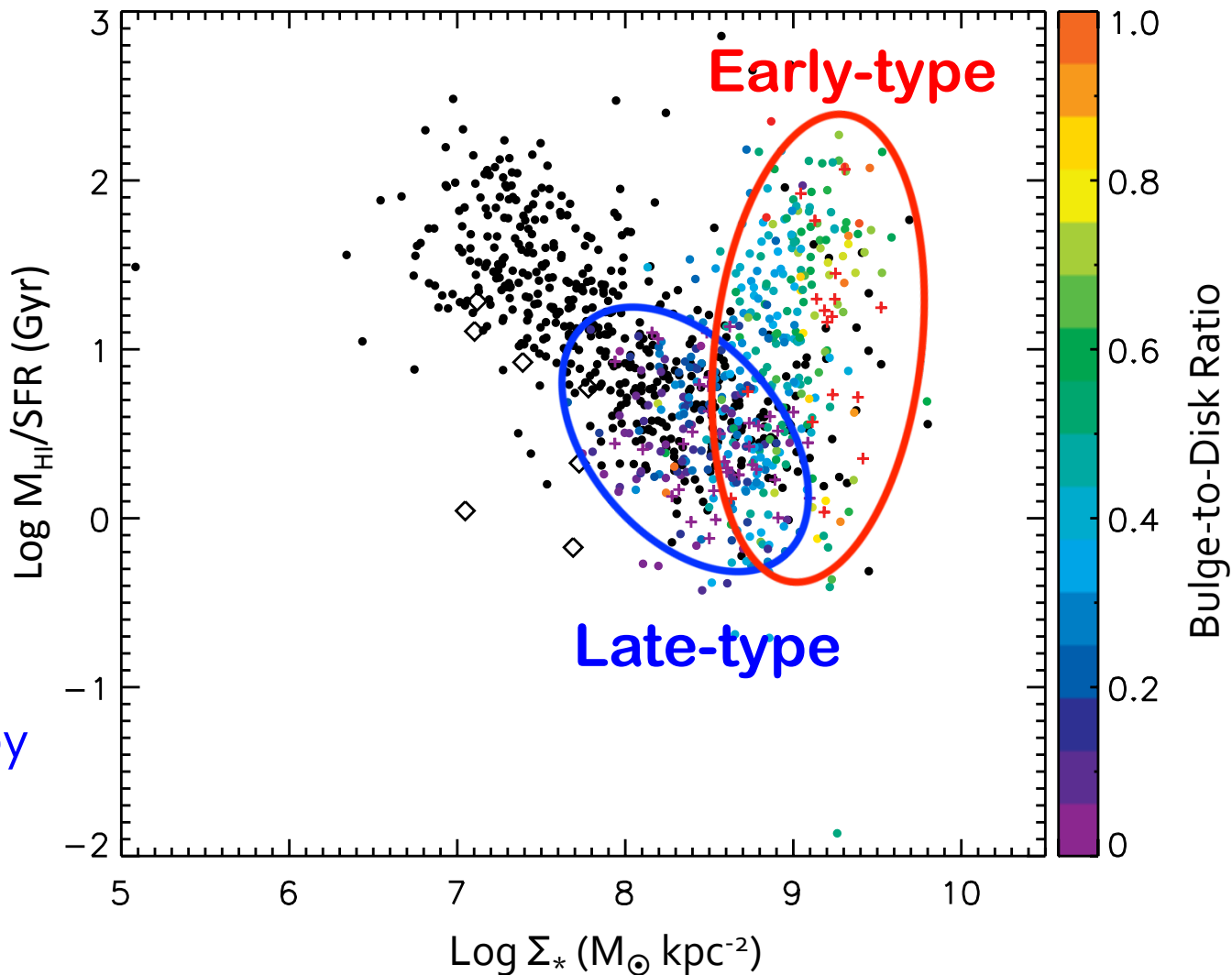
See talks by
Hughes,
Schruba,
Bonnell



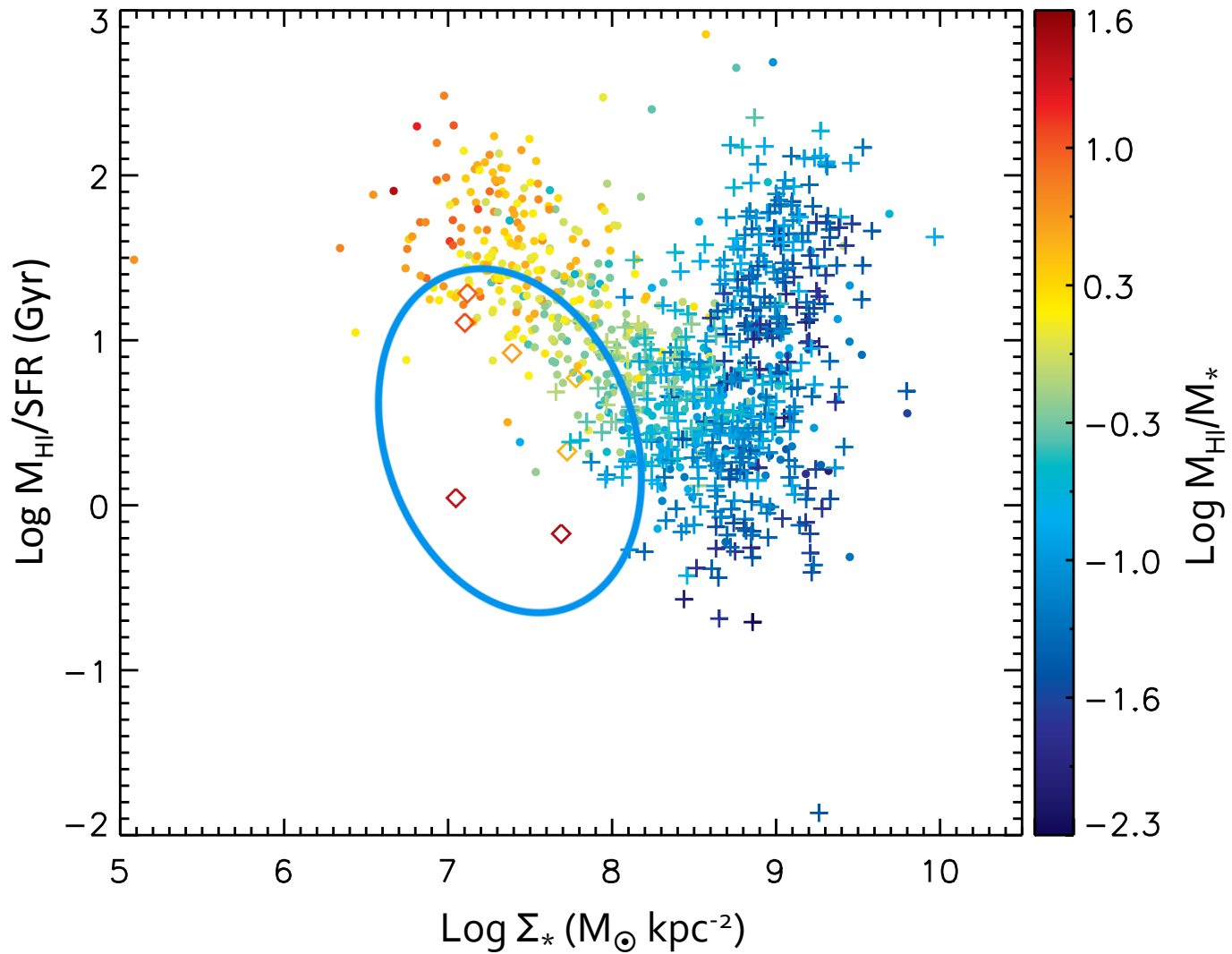
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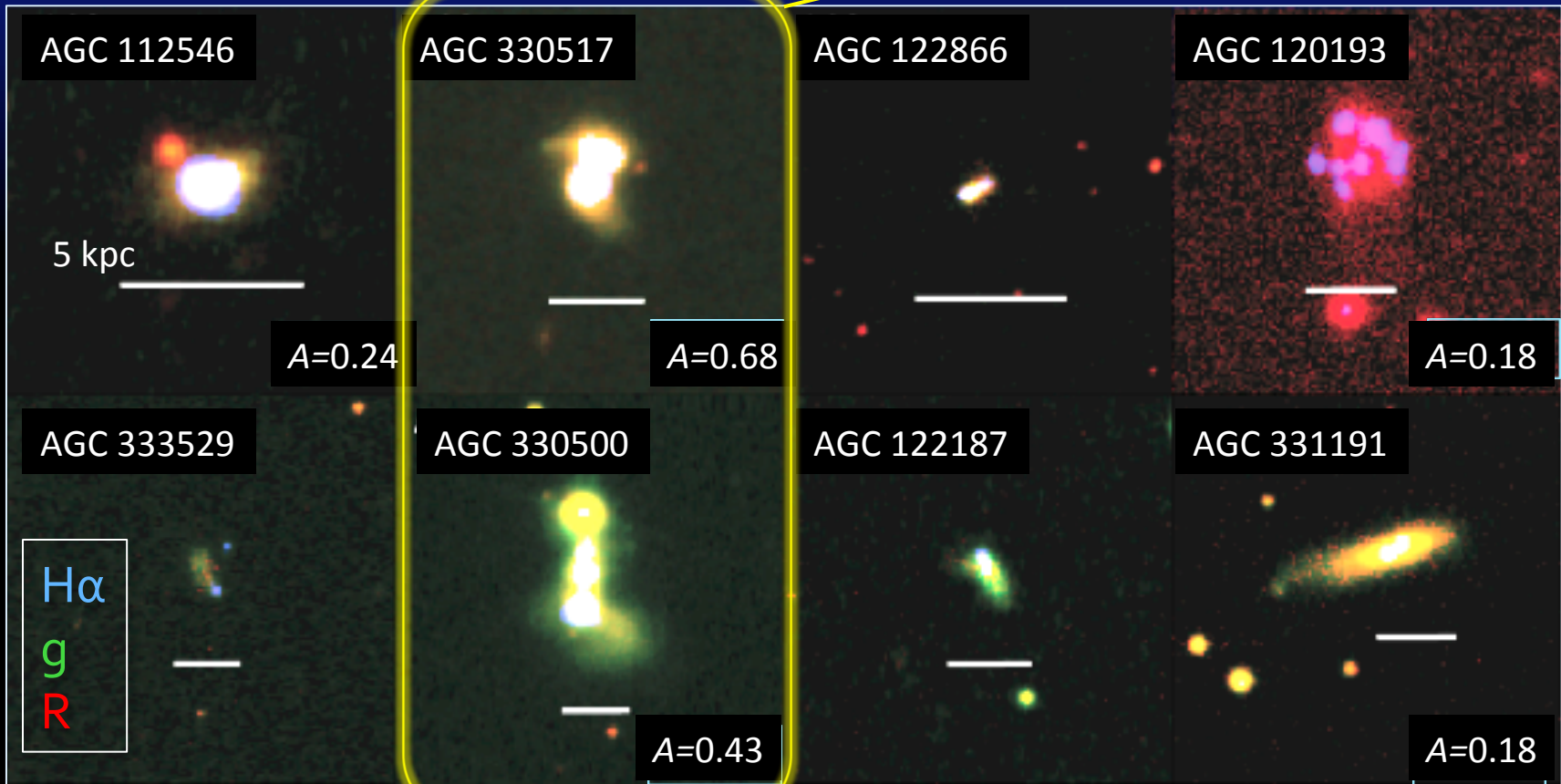


HI Consumption Timescales



Are Mergers Enhancing SFE?

Most HI-rich for their mass

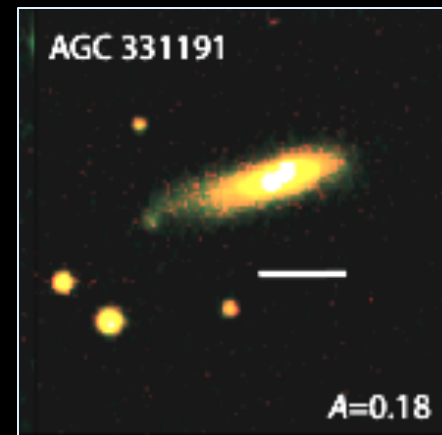


- Asymmetry (Conselice et al. 2000)
- Major mergers: $A > 0.35$

- Full sample median: $A=0.14$
- Asymmetry peaks before peak SF (Lotz et al. 2010)

HI Regulation

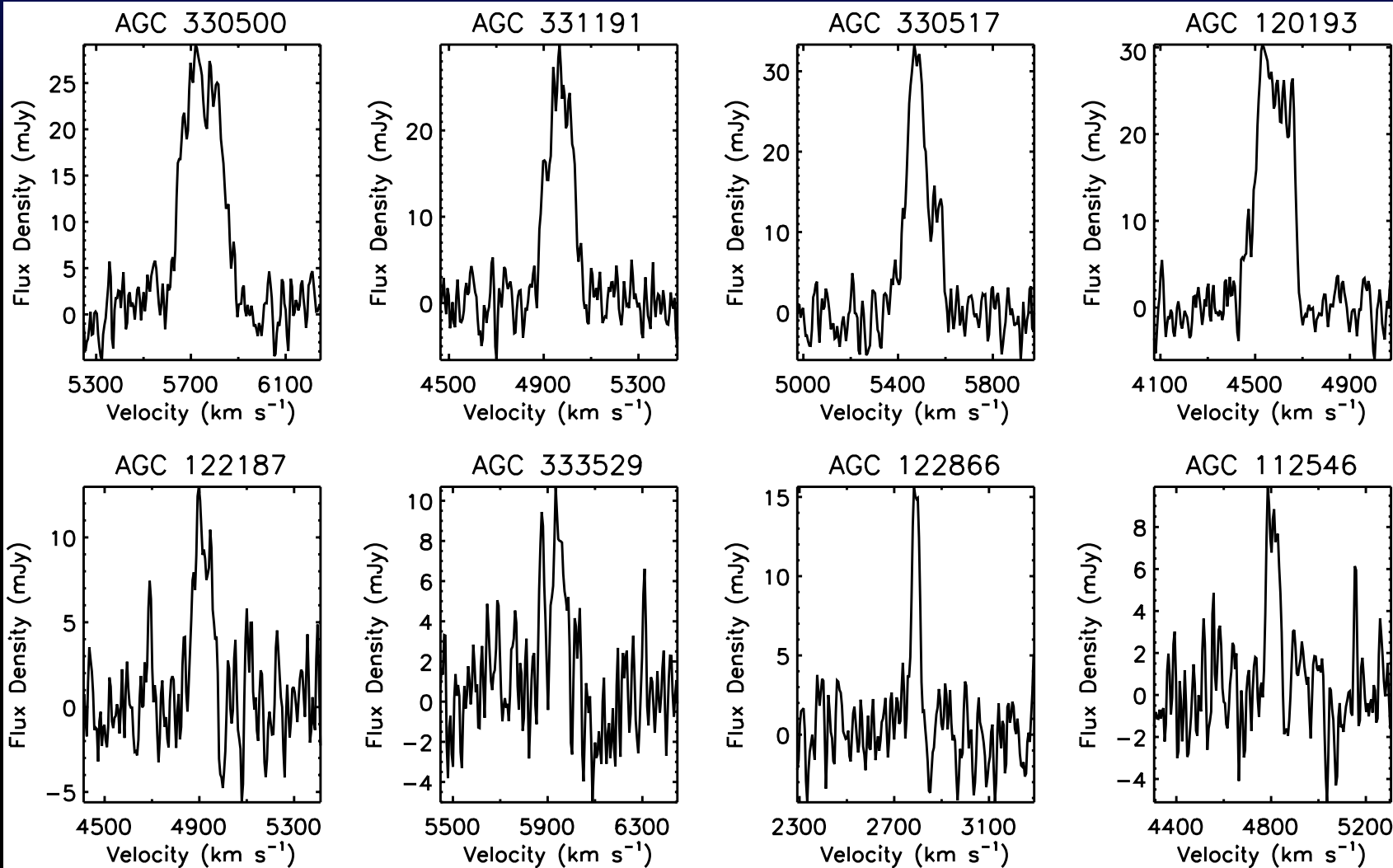
- Starbursts maintain moderate HI gas fractions
 - HI inflow and H₂ formation as merger progresses
 - In starburst regions, photodissociation replenishes HI (*e.g.*, *Stacey et al. 1991*)
 - HI fuel available for future bursts
- (*see also: Stierwalt et al. 2014; Ellison et al. 2015*)



Mergers and HI Kinematics

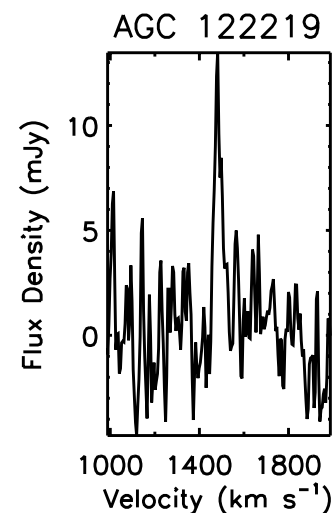
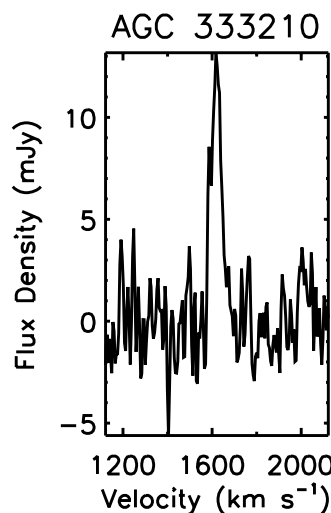
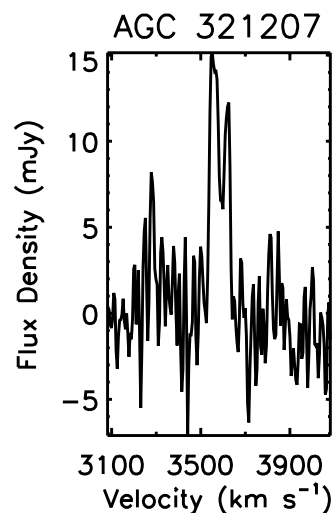
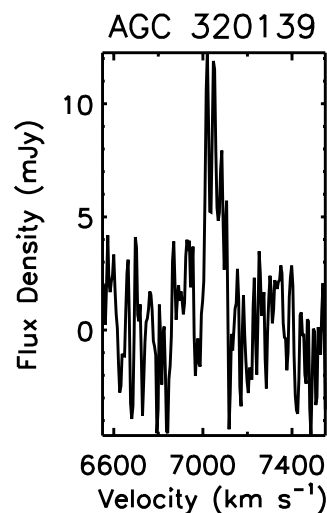
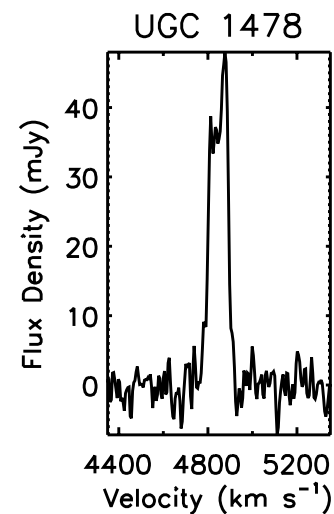
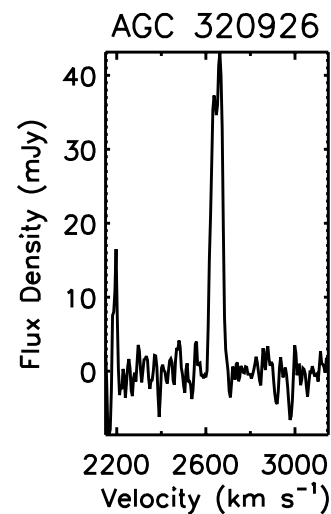
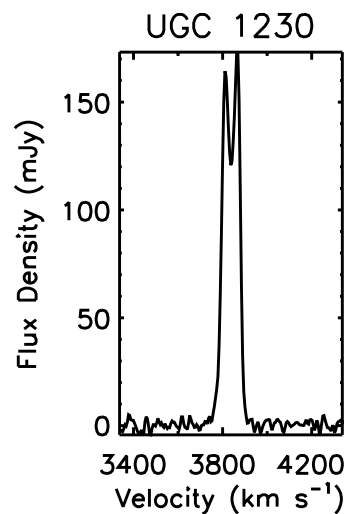
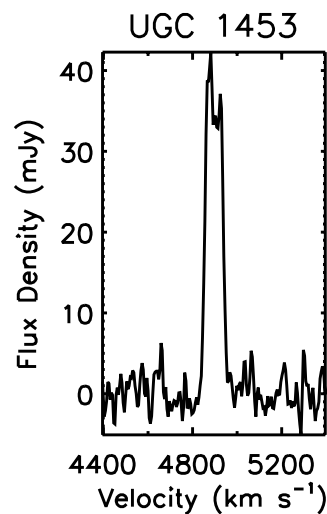
Starbursts

Higher W_{20}/W_{50} indicates excess of high-velocity HI



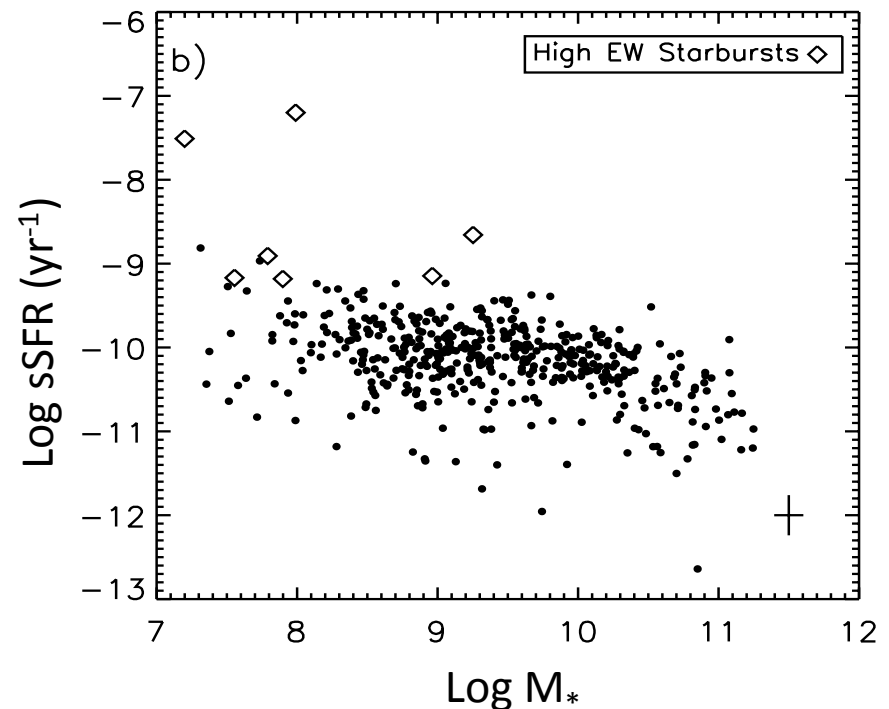
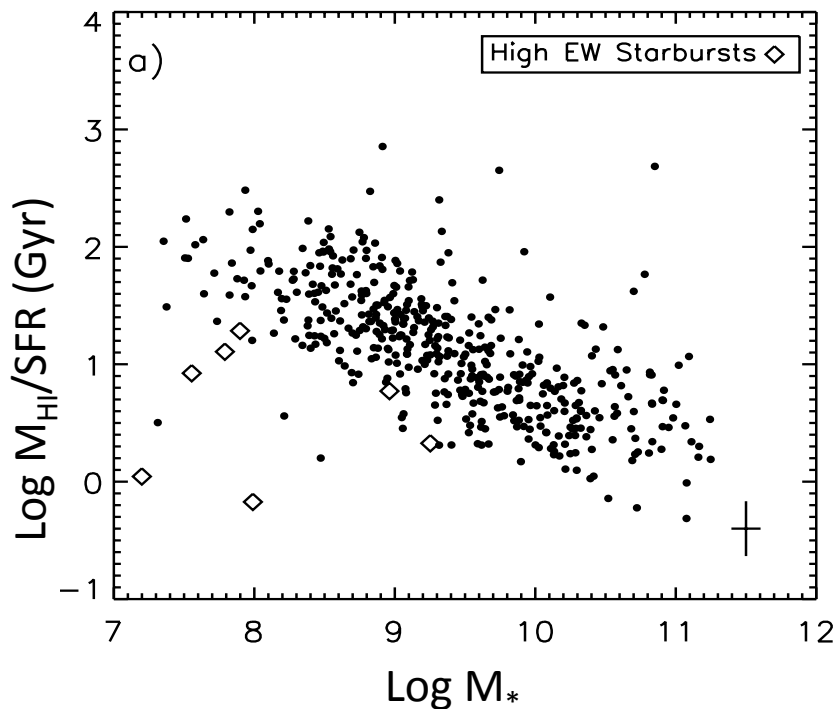
Mergers and HI Kinematics

Non-Starbursts



Low-Mass Starbursts

- Are the low-mass starbursts mergers?
- Periodic bursts due to feedback
 - *e.g., Lee et al. 2007; Verbeke et al. 2014; Hopkins et al. 2014*
- Initial disturbance may no longer be detectable
- Results in variable SFRs and delayed HI consumption



Summary

- HI weakly linked to sSFR, closely coupled with dust extinction
 - Gas cycles and metallicity
- HI/SFR correlates with stellar surface density for disks
 - Role of mid-plane pressure?
- Spheroids may differ in access to HI
- Starbursts may maintain approximately constant HI
- Bursts in low-mass galaxies may be cyclical