



Astronomy and Space Science

Kinematical misalignments in the EAGLE simulations

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with ATLAS^{3D} and EAGLE teams

Atlas^{3D}: 260 ETGs, $M_{\star} > 10^{10} M_{\odot}$

IFU survey

WHT/SAURON
stars and ionised gas
 $r < R_e$

Cappellari+ 11
Davis+ 11

CO survey

IRAM+CARMA
 $M(\text{H}_2) \sim 3 \times 10^7 M_{\odot}$

Young+ 11
Alatalo+ 13

HI survey

WSRT
 $M(\text{HI}) \sim 10^7 M_{\odot}$
 $N(\text{HI}) \sim 3 \times 10^{19} \text{ cm}^{-2}$

Serra+ 12, 14

*$\text{H}_2 + \text{HI}$ detection rate $\sim 50\%$
neutral gas mass $\sim 10^7 - 10^9 M_{\odot}$*

mostly gas discs and rings, $\sim 50\%$ misaligned from stars

HI - stars kinematical misalignments

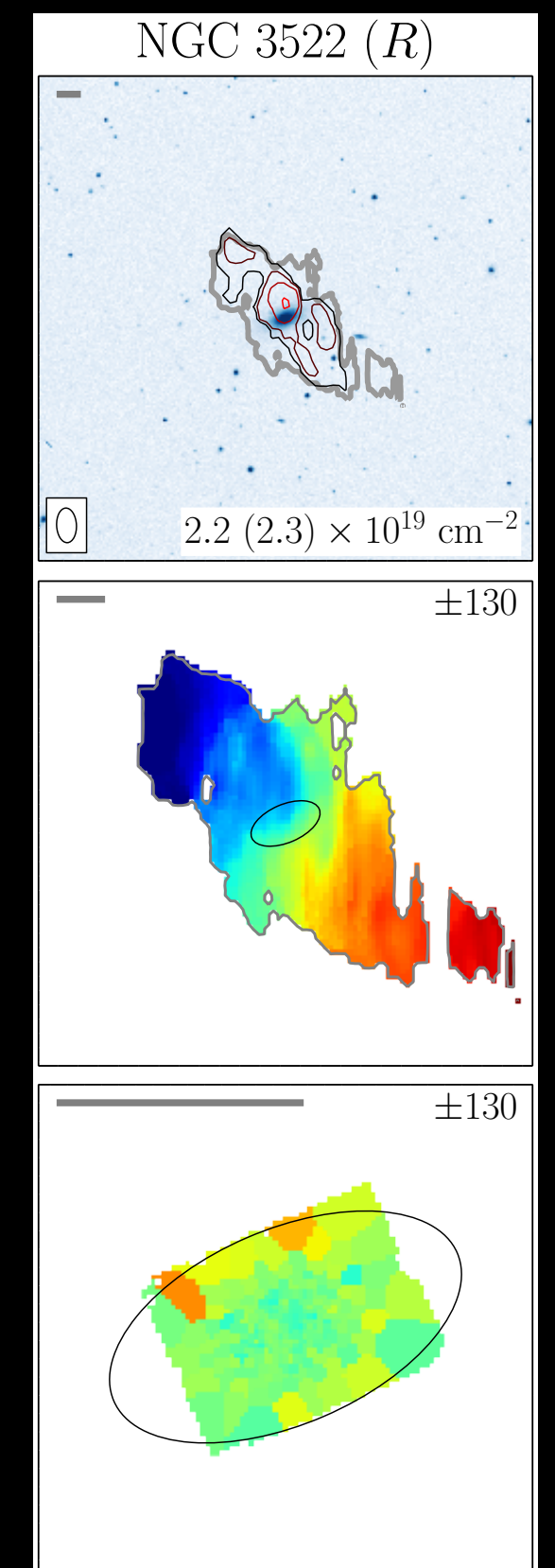
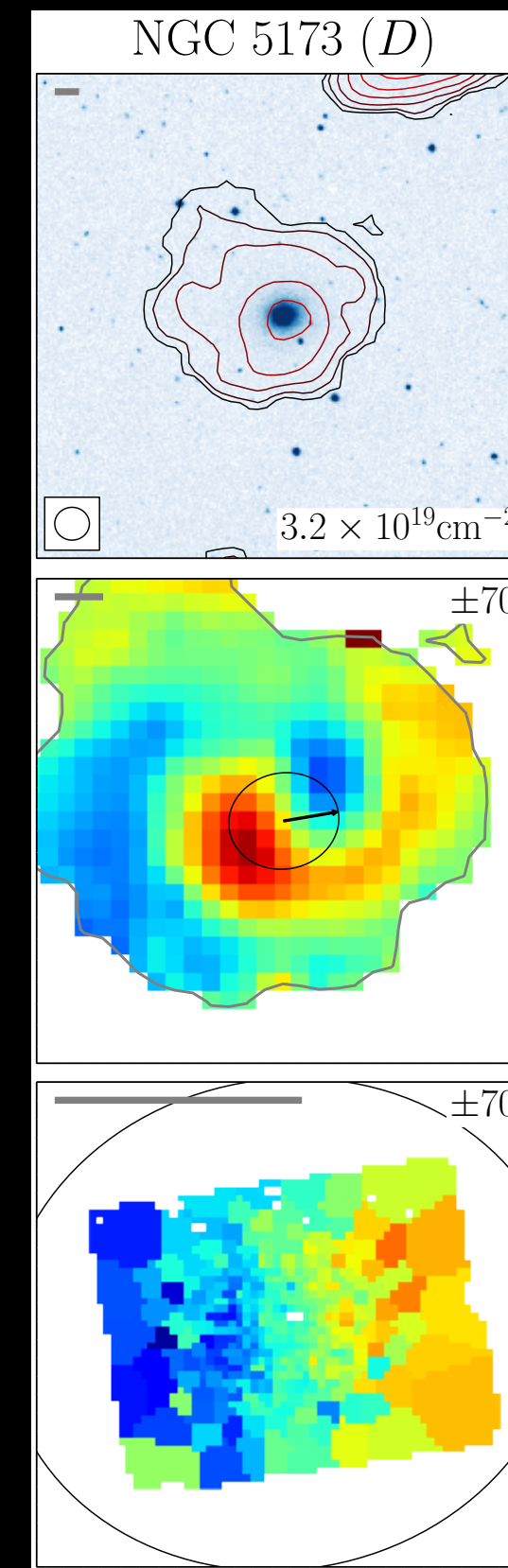
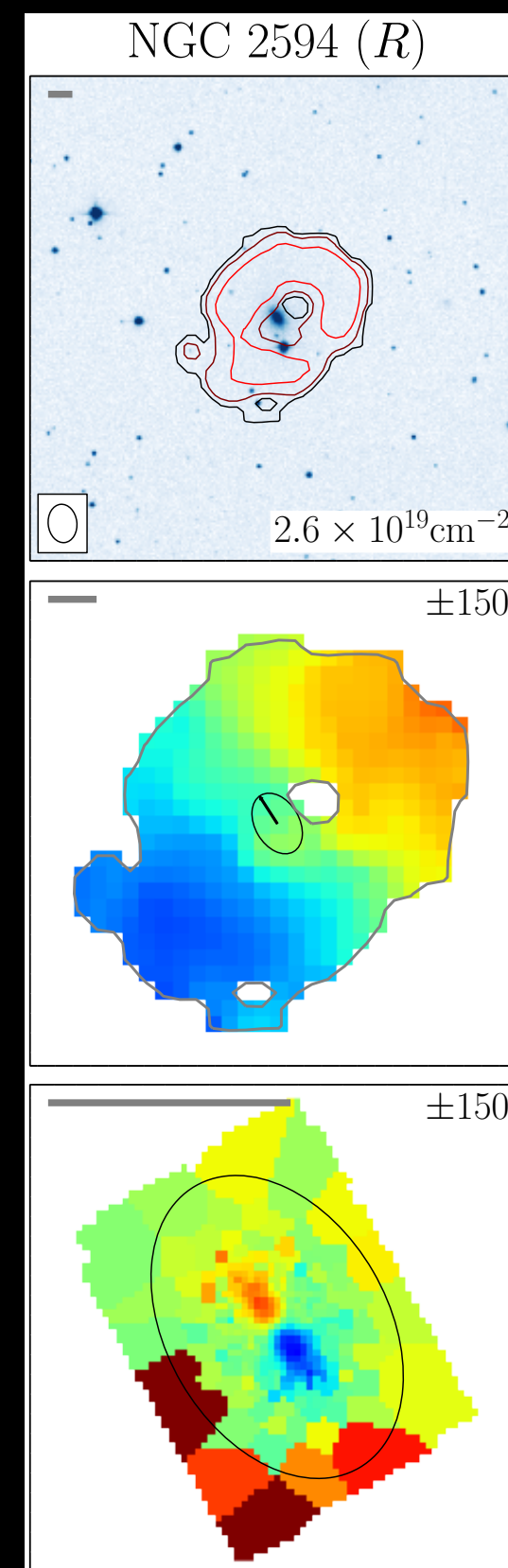
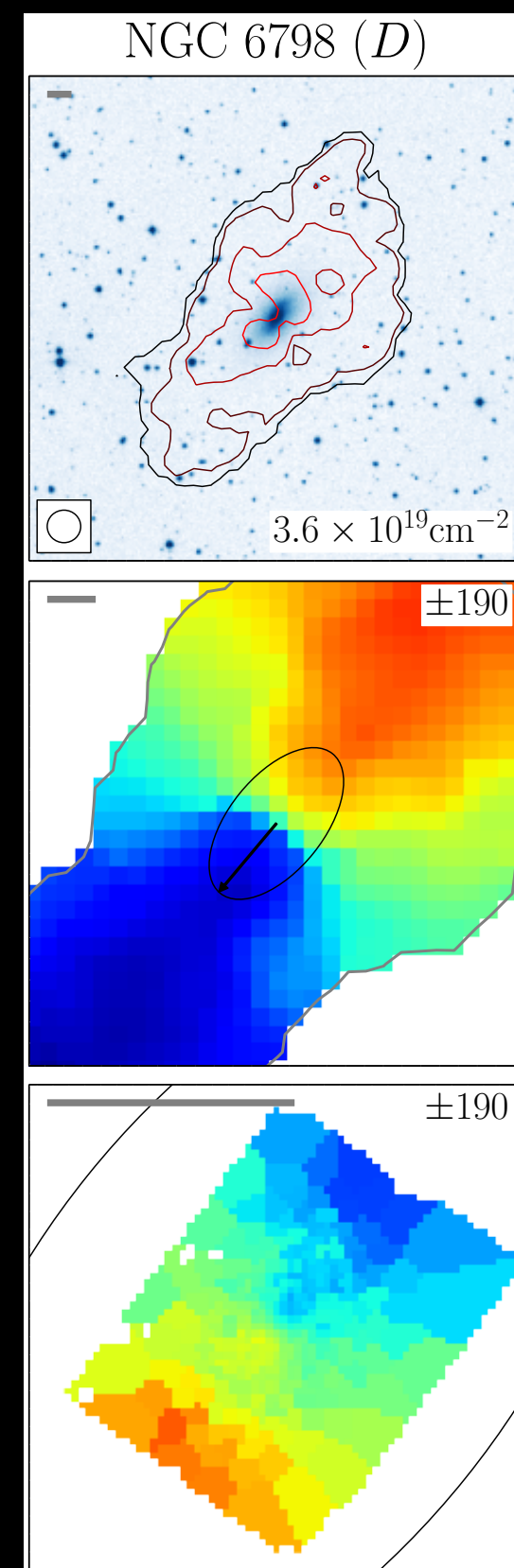
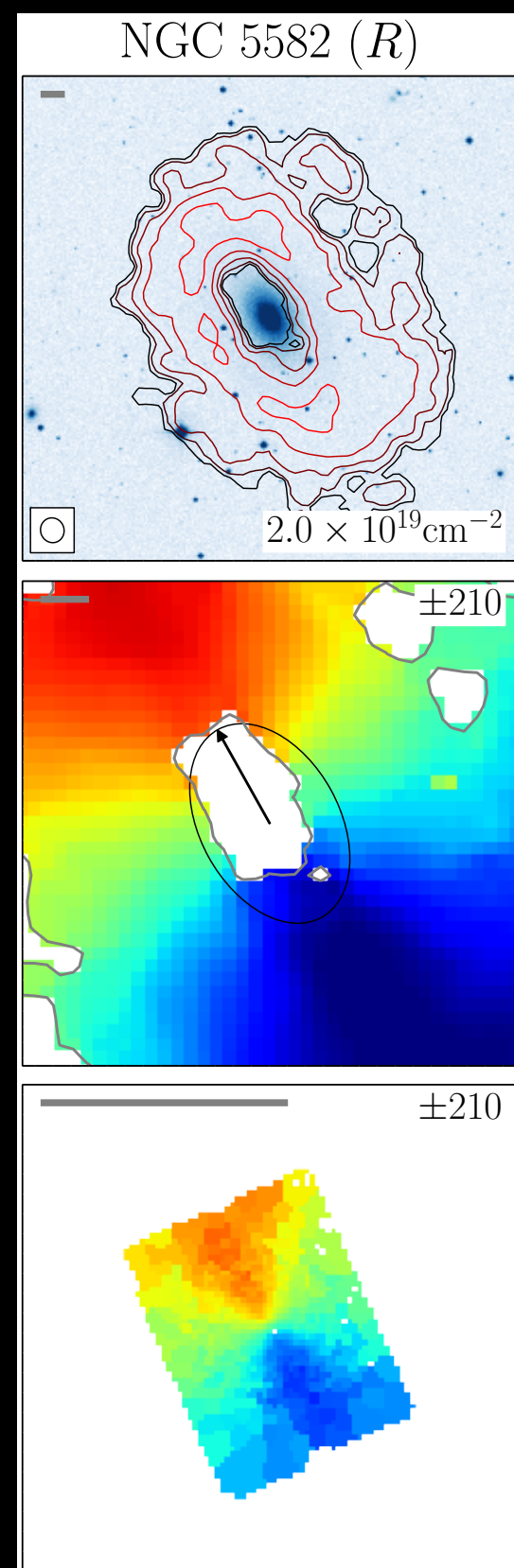
co-rotation

counter-rotation

polar

warp/other

slow rotators

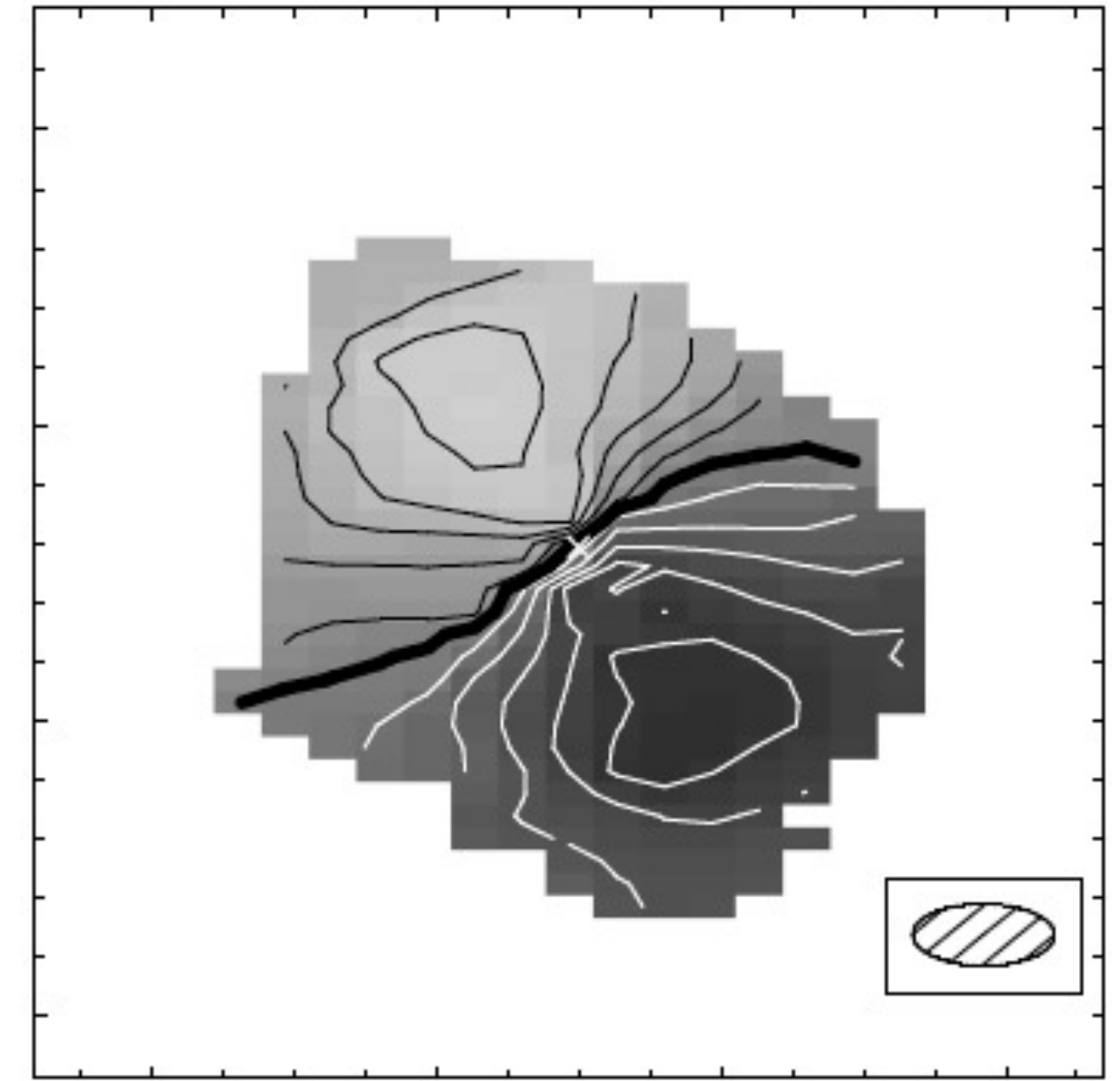
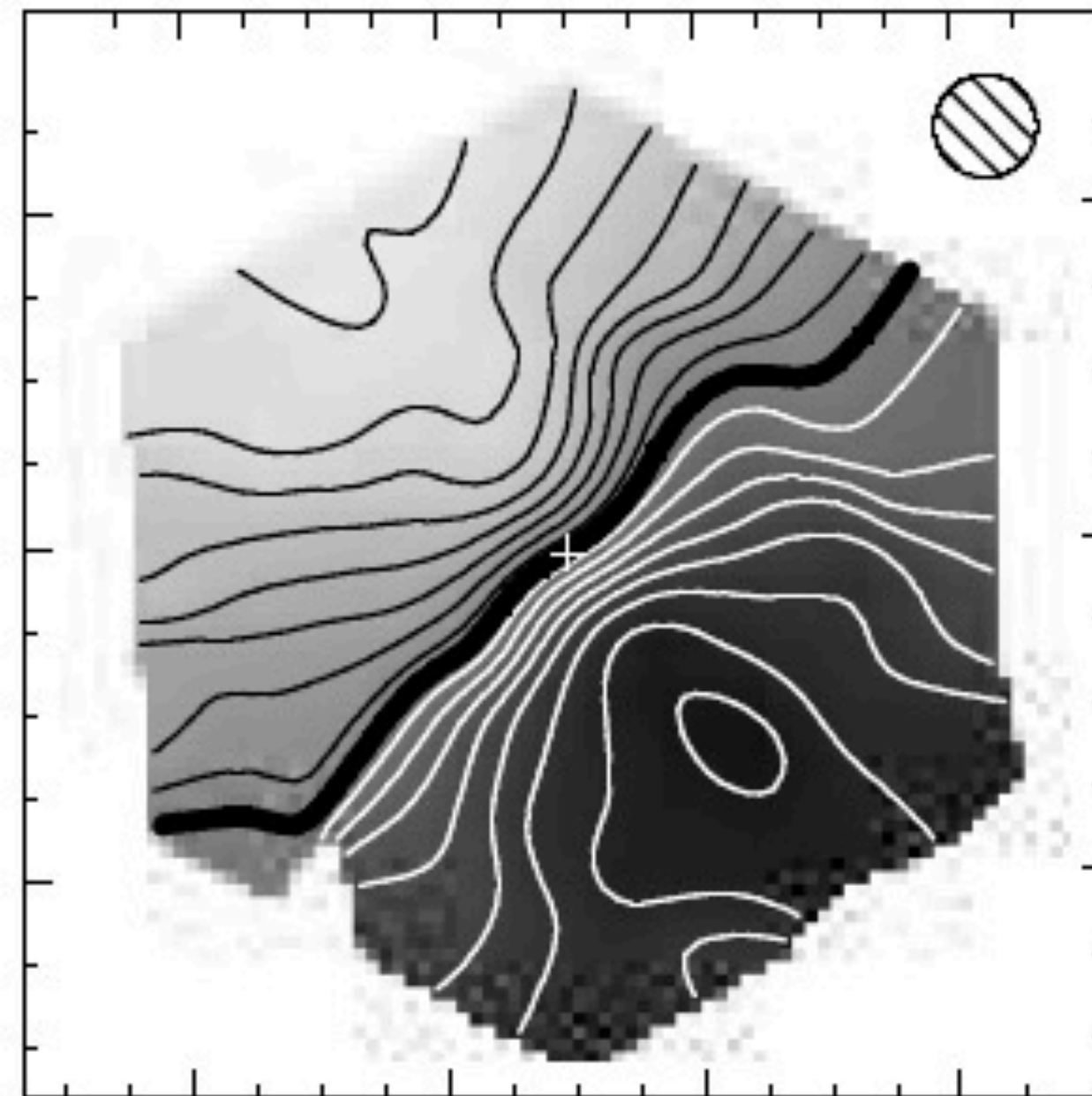
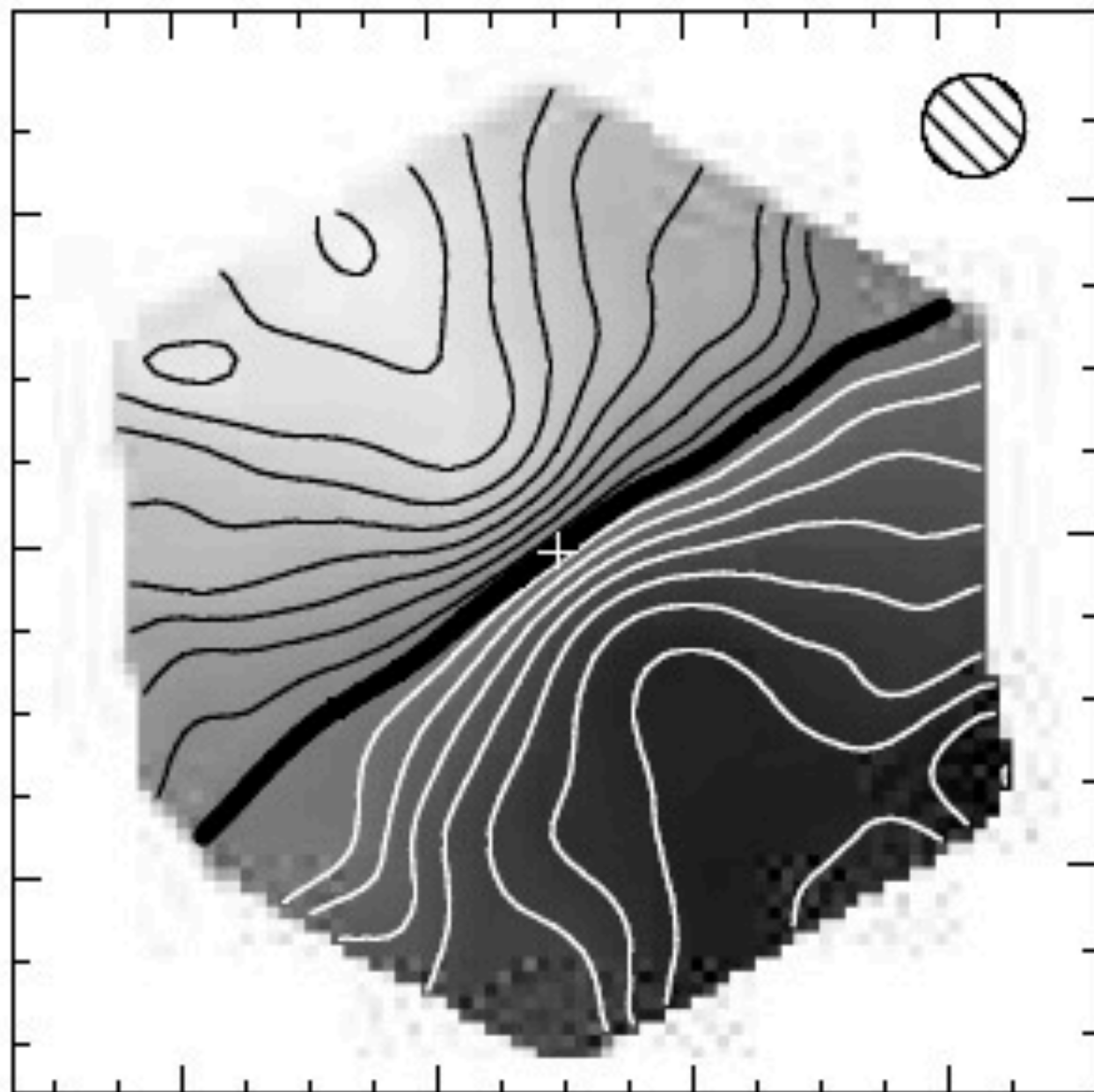


Kinematical misalignments in spirals

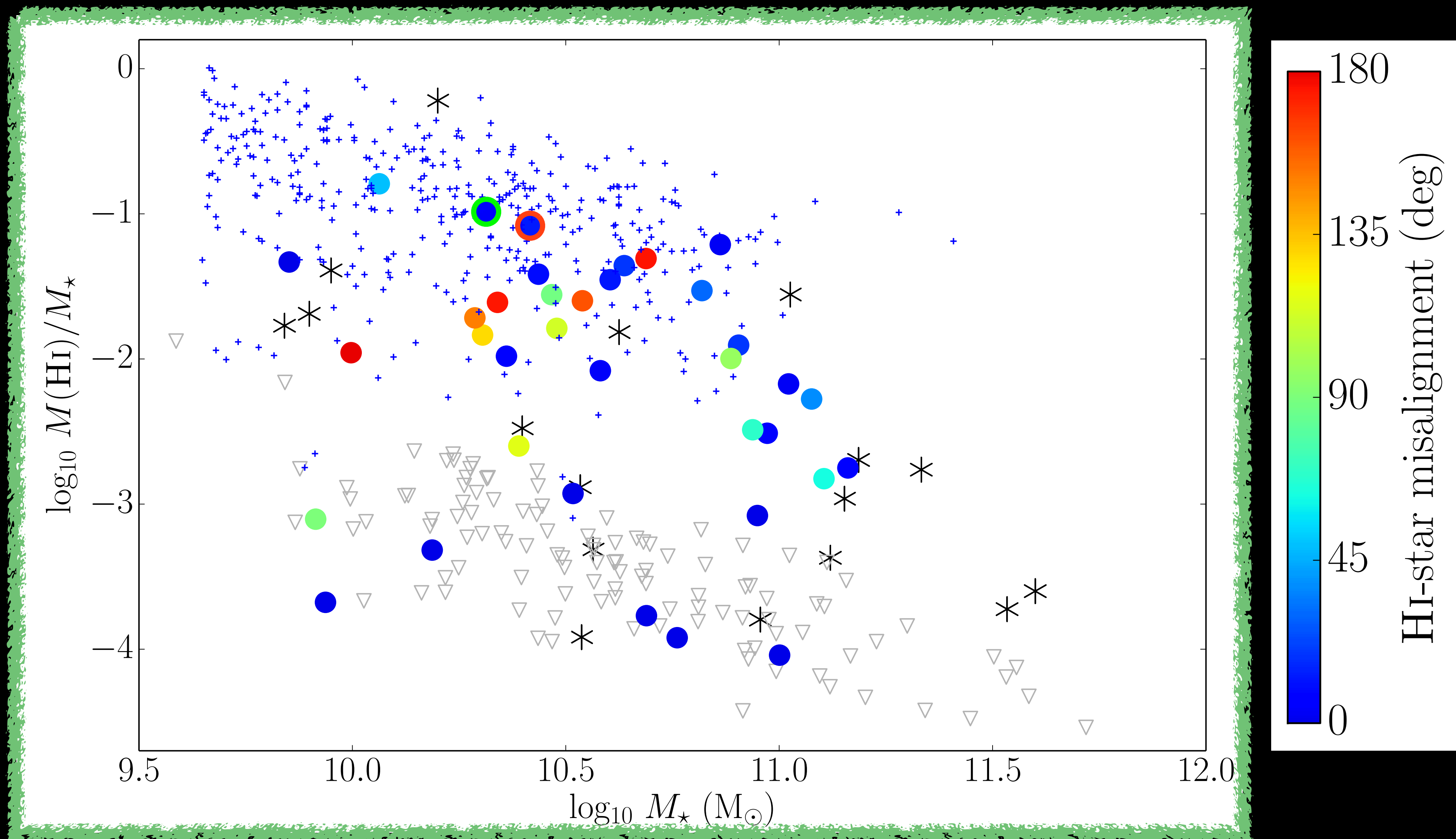
stars

OIII

HI



Ganda et al. (2006); Martinsson et al. (2013a,b);
Barrera-Ballesteros et al. (2014); Bryant et al. (in prep.)
TOTAL \sim 200 spirals; misalignment rate \sim 10%



- ETG HI disc/ring
- * ETG HI unsettled
- + spirals
- ▽ ETG HI undetected

EAGLE simulations

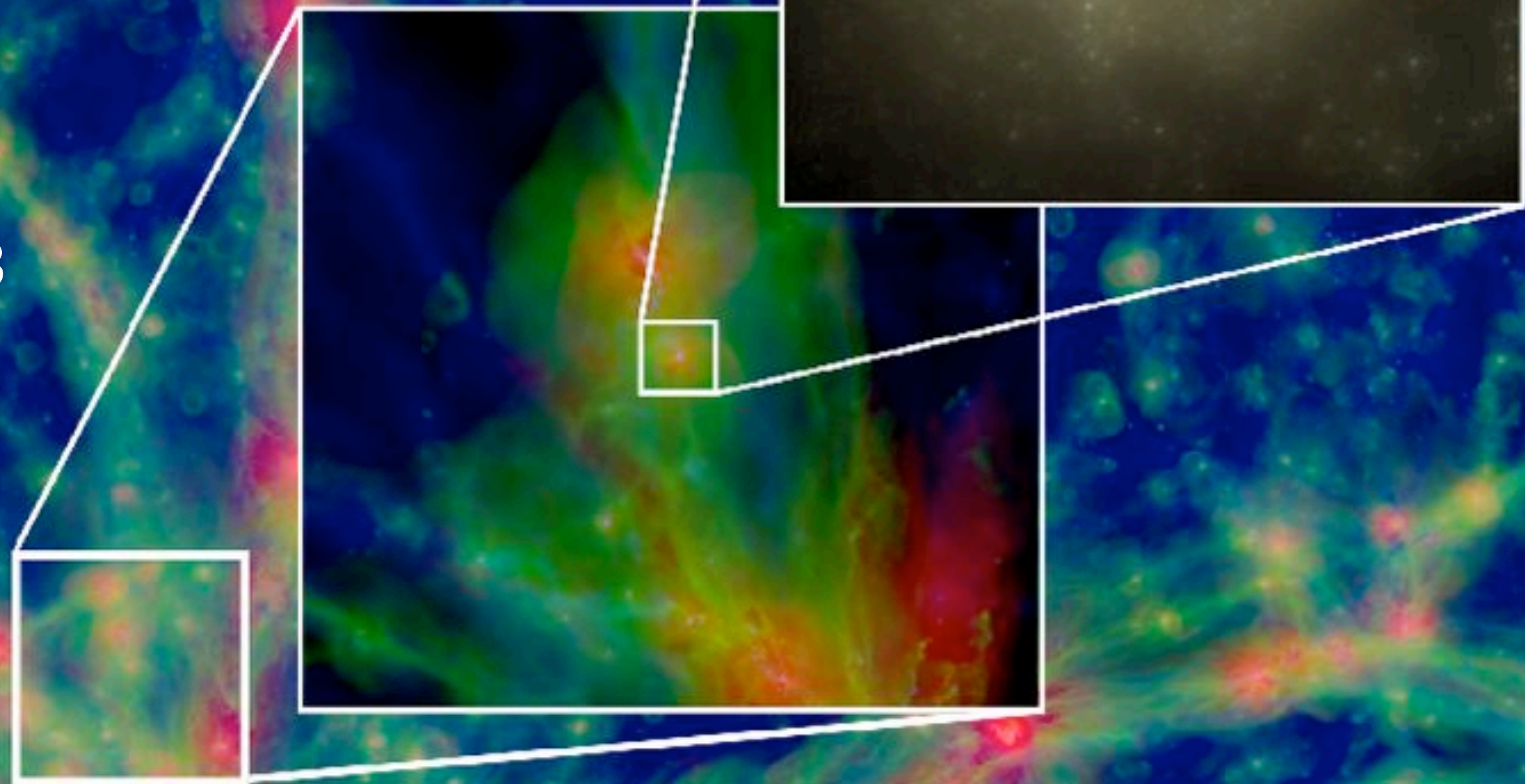
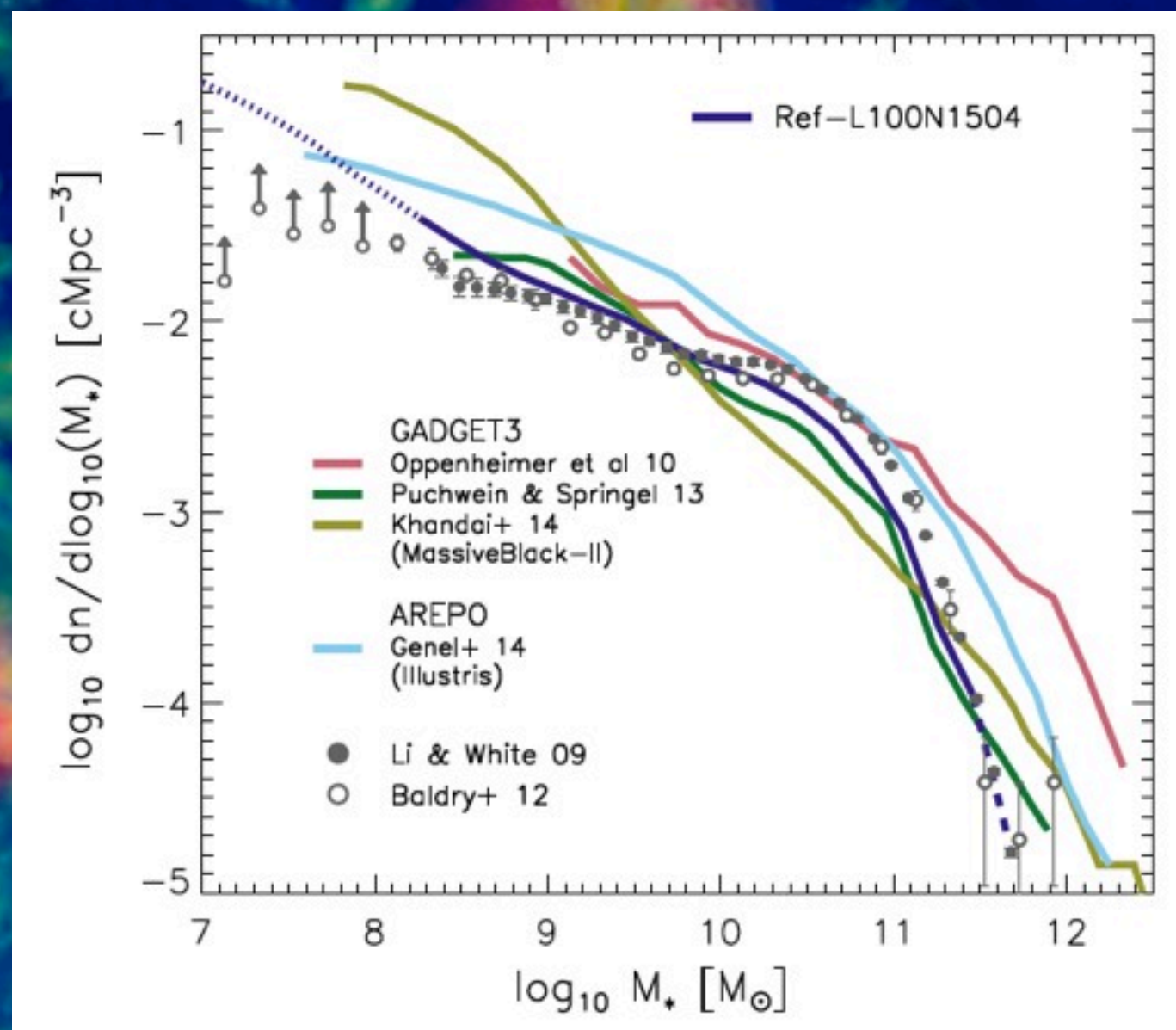
Schaye et al. (2015)

$M_{\text{res}} \sim 10^6 M_{\odot}$, $\epsilon \sim 0.7$ kpc

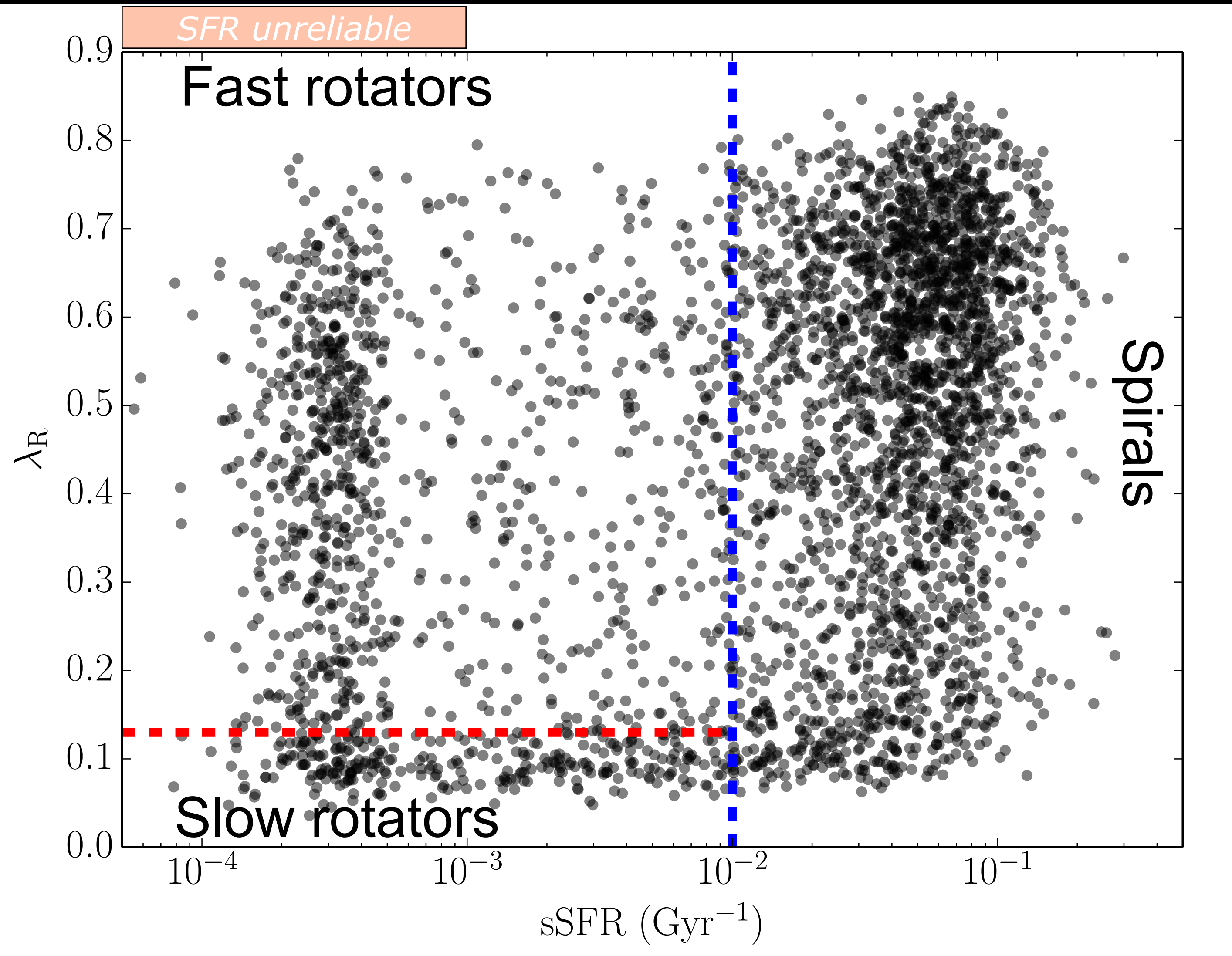
100 Mpc volume

~ 3500 galaxies with $M_{\star} > 10^{10} M_{\odot}$

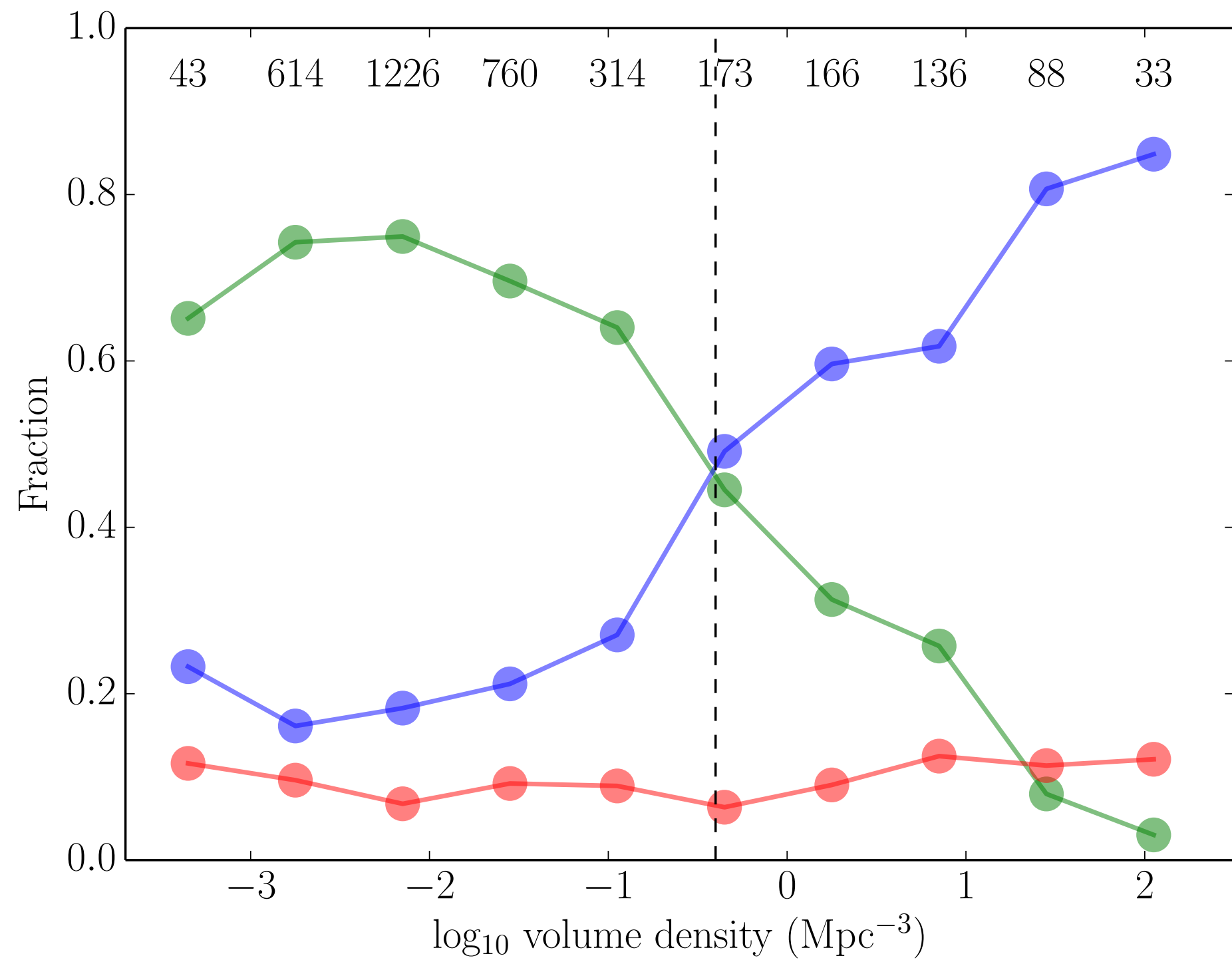
environment from voids to \sim Coma/3



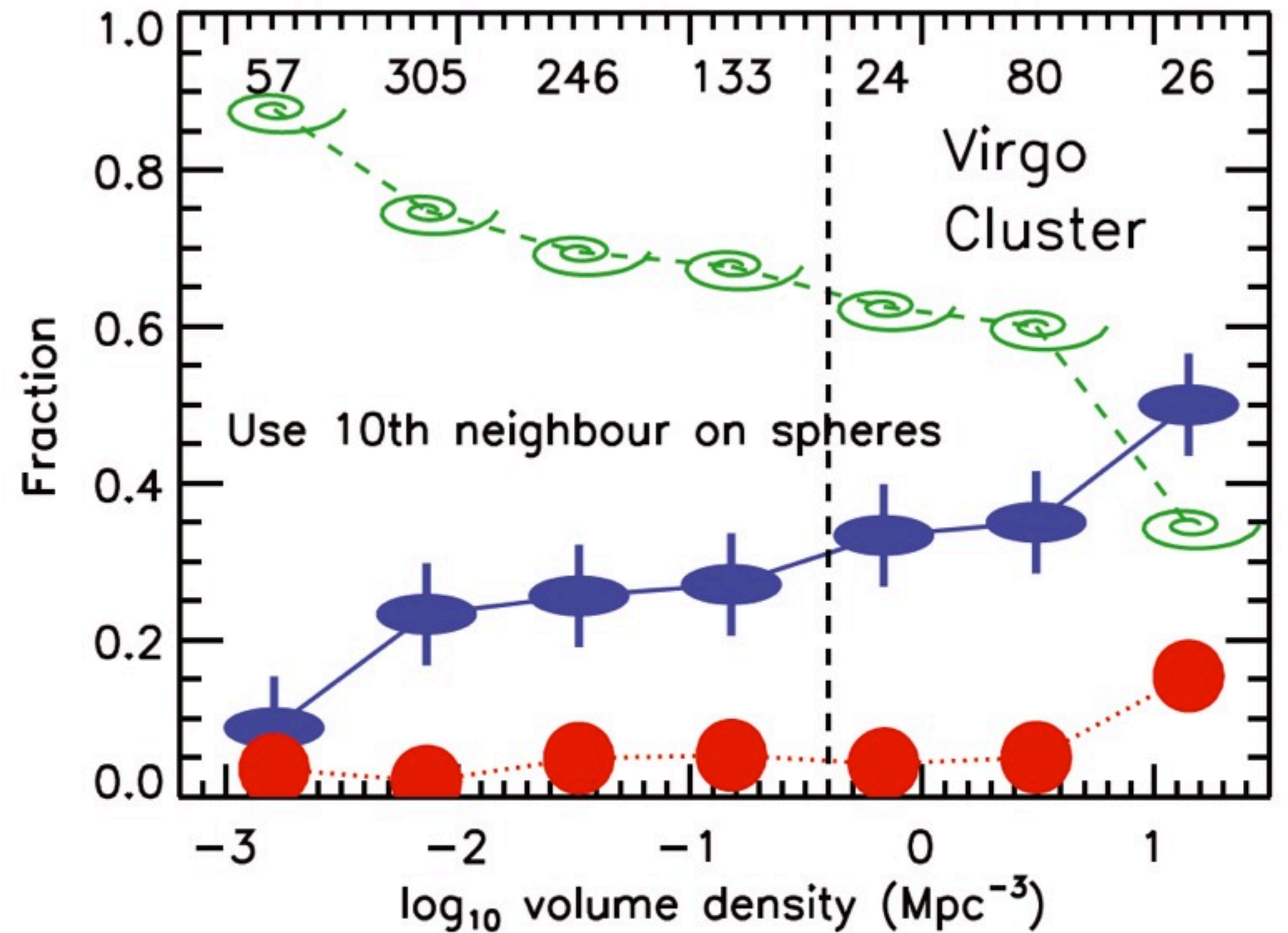
Selection of Spirals, FRs, SRs



EAGLE volume (env. dens. courtesy of Marasco)



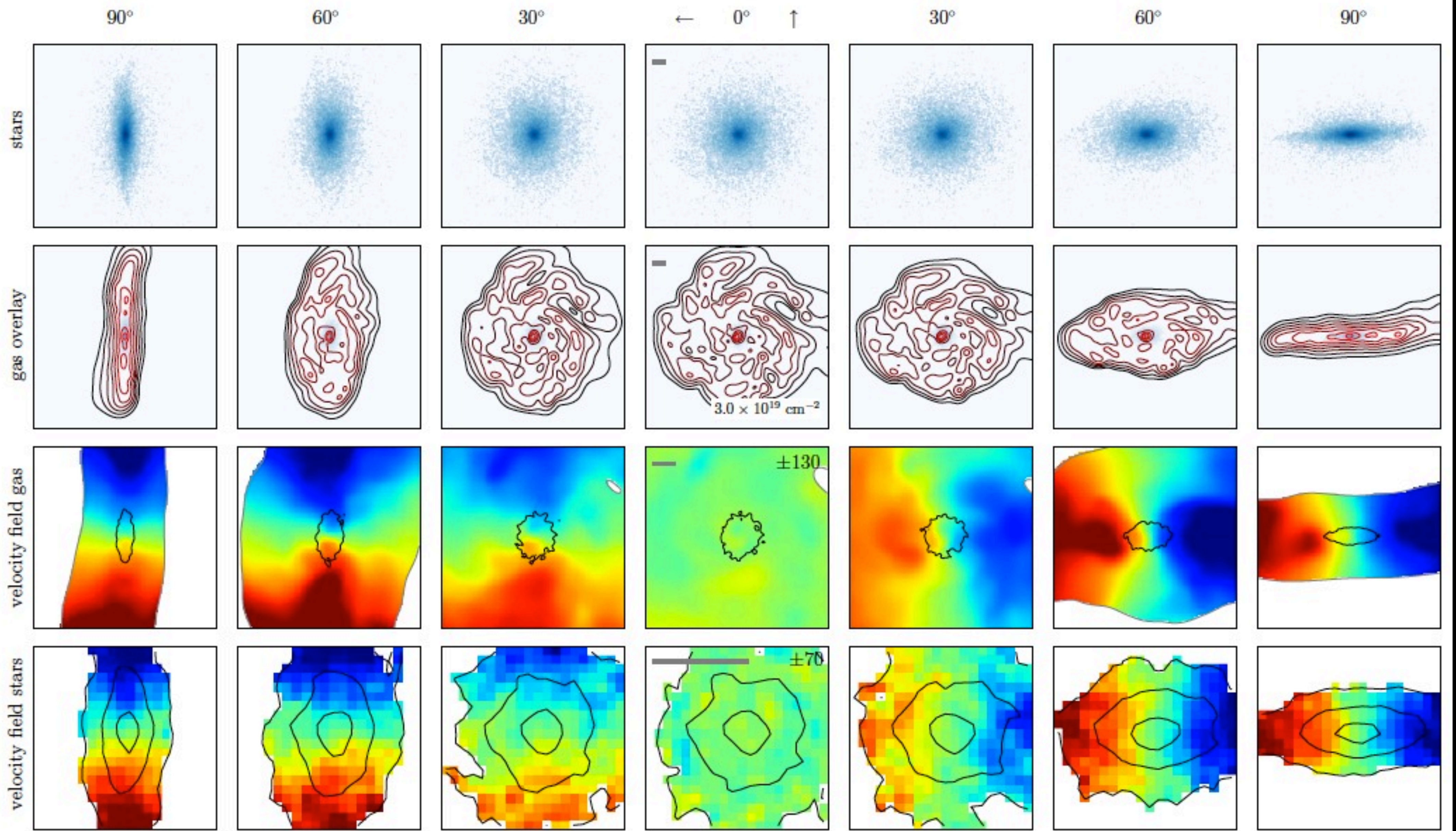
Atlas^{3D} volume (Cappellari et al. 2011)



 spirals

 slow-rotating ETGs

 fast-rotating ETGs



HI-stars kinematical misalignments

co-rotating

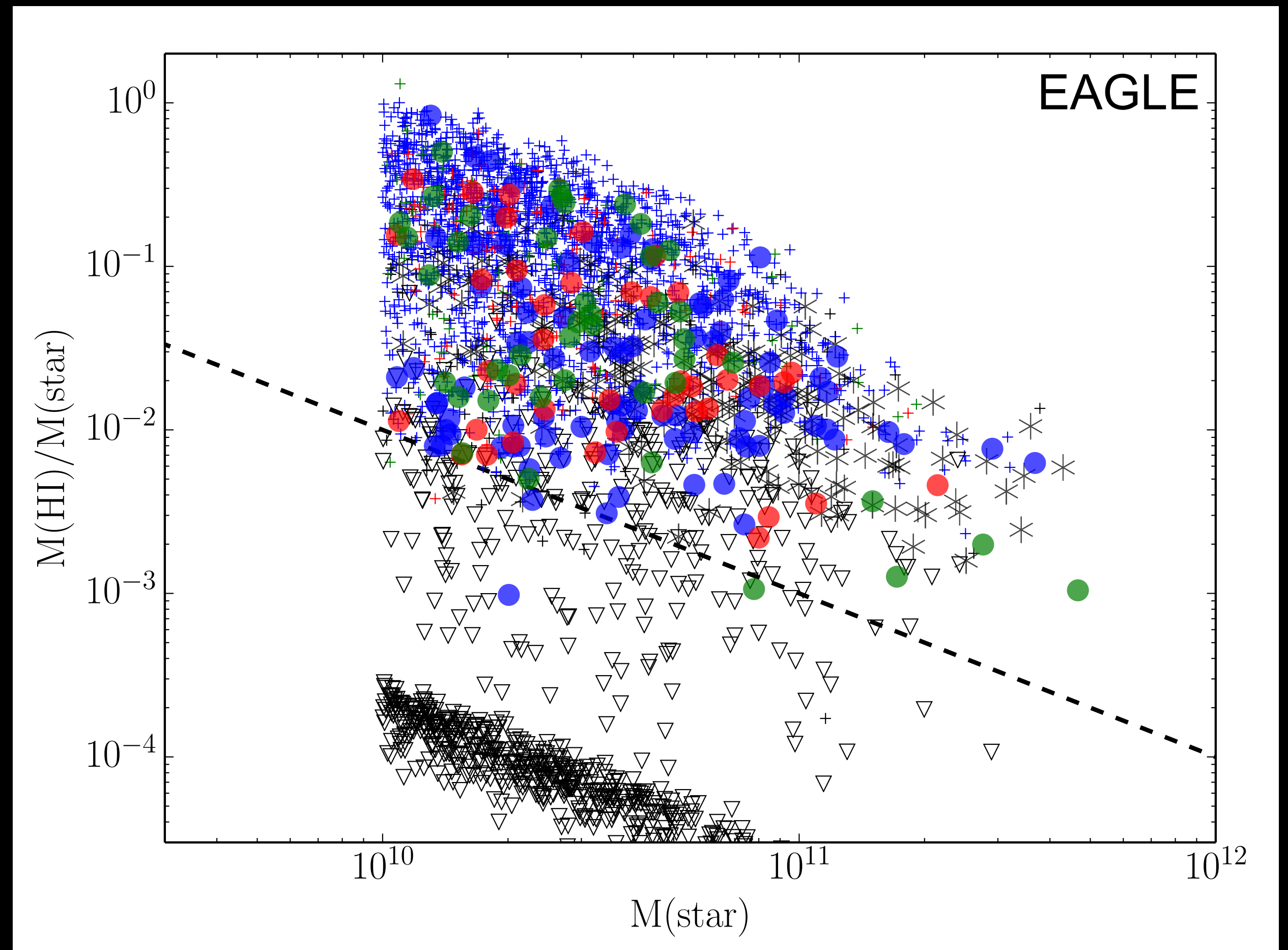
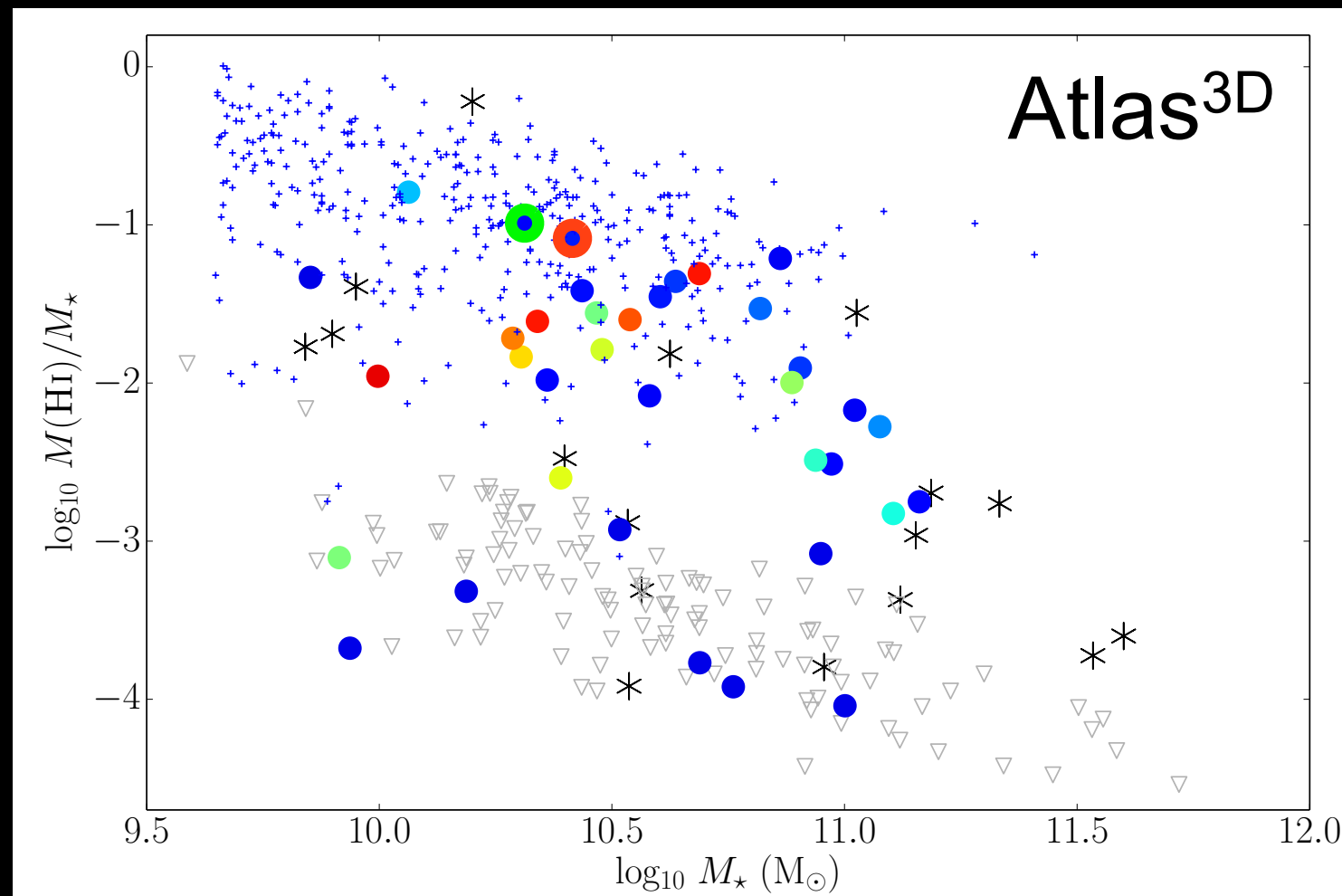
polar

counter-rotating

0°

90°

180°



- * ETG HI unsettled
- ▽ ETG HI undetected
- ETG HI disc/ring
- + spirals

Some numbers from *Atlas^{3D}* and *EAGLE* (outside clusters)

	Spirals	Early-type's	Fast rotators	Slow rotators
Nr galaxies	1936	827 127	583 111	244 16
HI det rate	99%	44% 39%	40% 37%	54% 50%
HI disc rate	94%	20% 25%	20% 25%	19% 25%
corotating fraction	89%	51% 45%	50% 50%	53% 0%

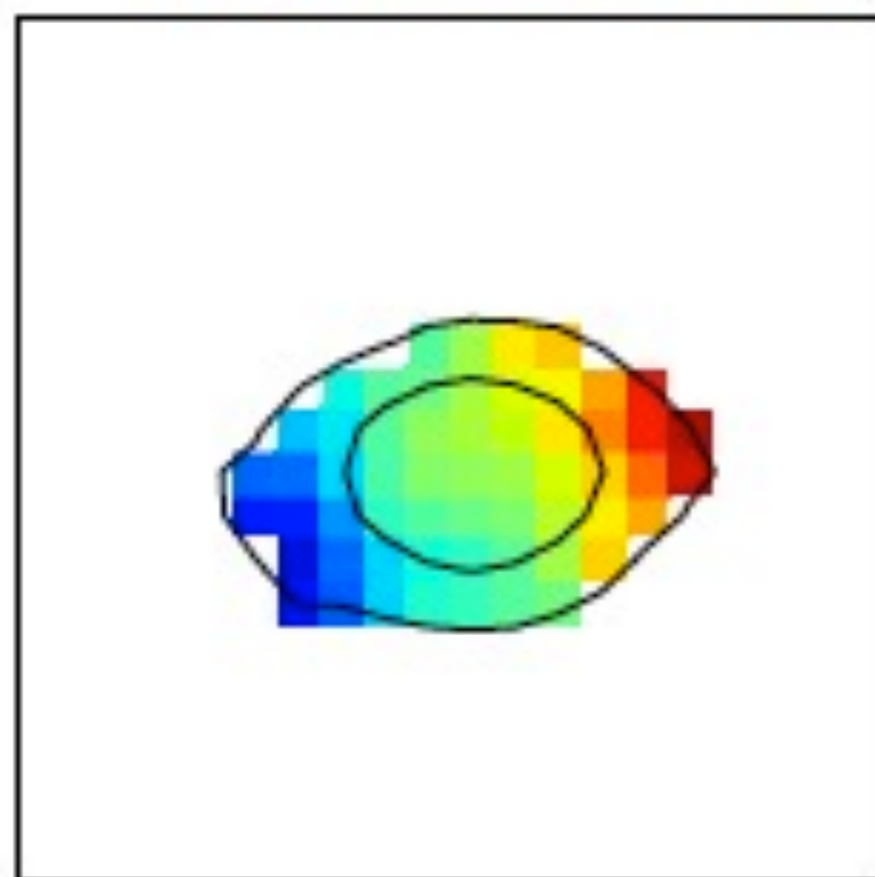
Possible problems with EAGLE

- Too many SRs
- Morphology-density relation not perfect
- Misaligned fraction of SRs

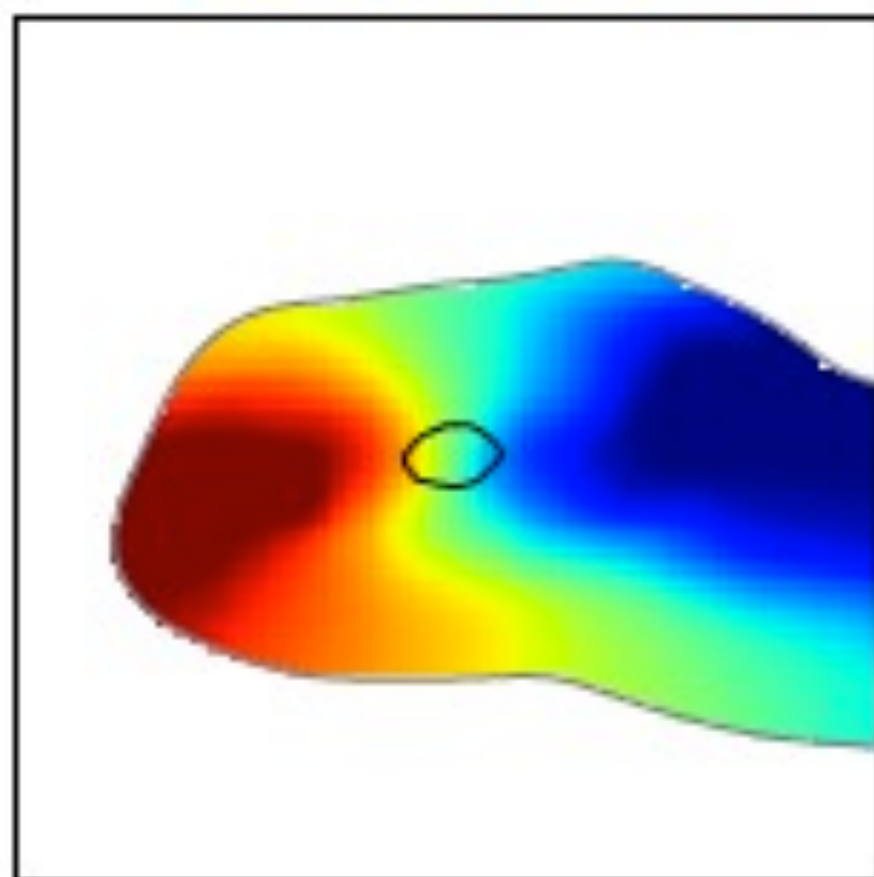
(Nearly) successes of EAGLE

- Can make gas-poor ETGs
- HI detection rate and mass of spirals, FRs and SRs
- HI disc rate of spirals, FRs and SRs
- Misaligned fraction of spirals and FRs
- (HI trends with environment about right)

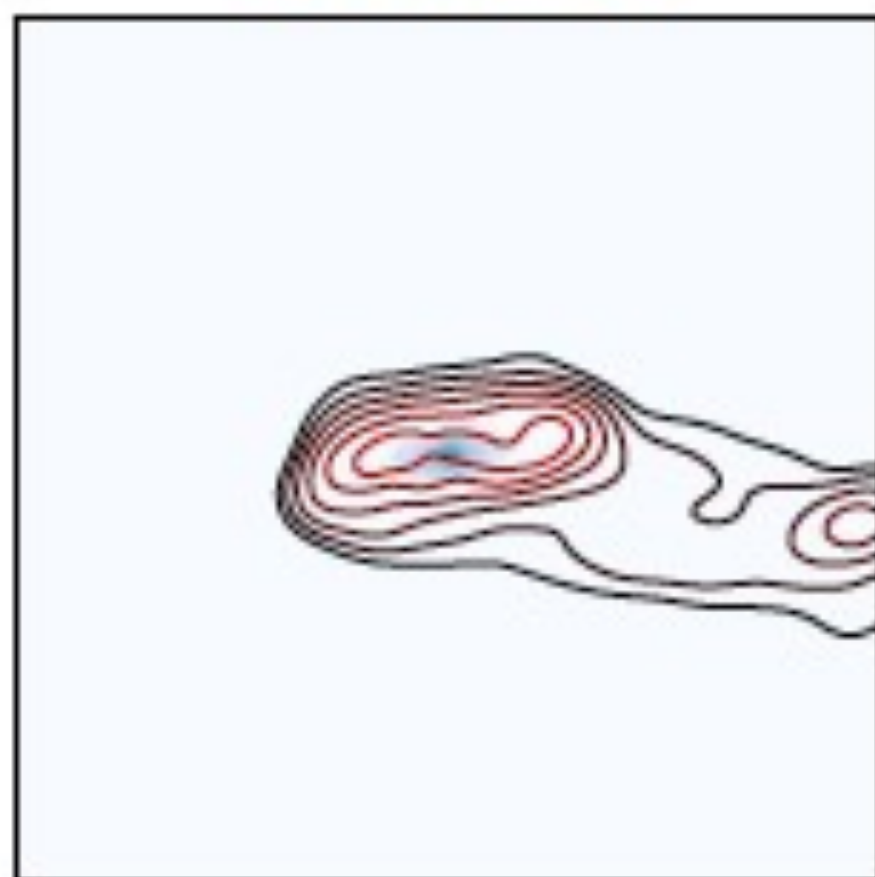
velocity field stars



velocity field gas



gas overlay



stars

