



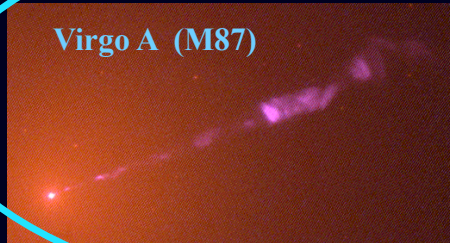
**LOFAR CONSTRAINTS ON WEAKLY
ACCRETING BLACK HOLE JETS**

Sera Markoff

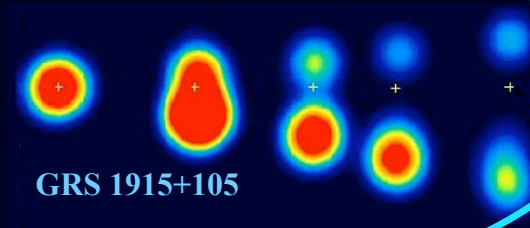
Astronomical Institute "A. Pannekoek"
University of Amsterdam

Jets appear in many astronomical objects

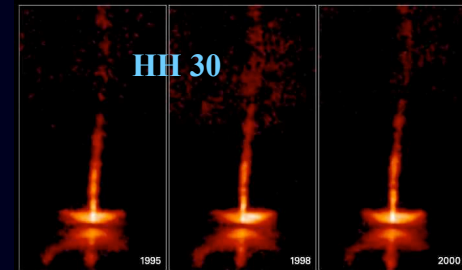
SUPERMASSIVE BLACK HOLES



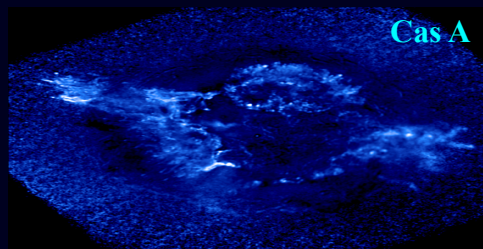
STELLAR-MASS BLACK HOLES



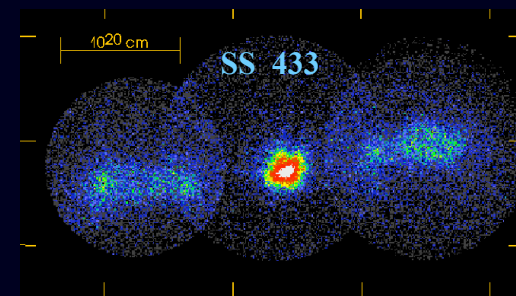
FORMING STARS



DYING STARS



DEAD STARS



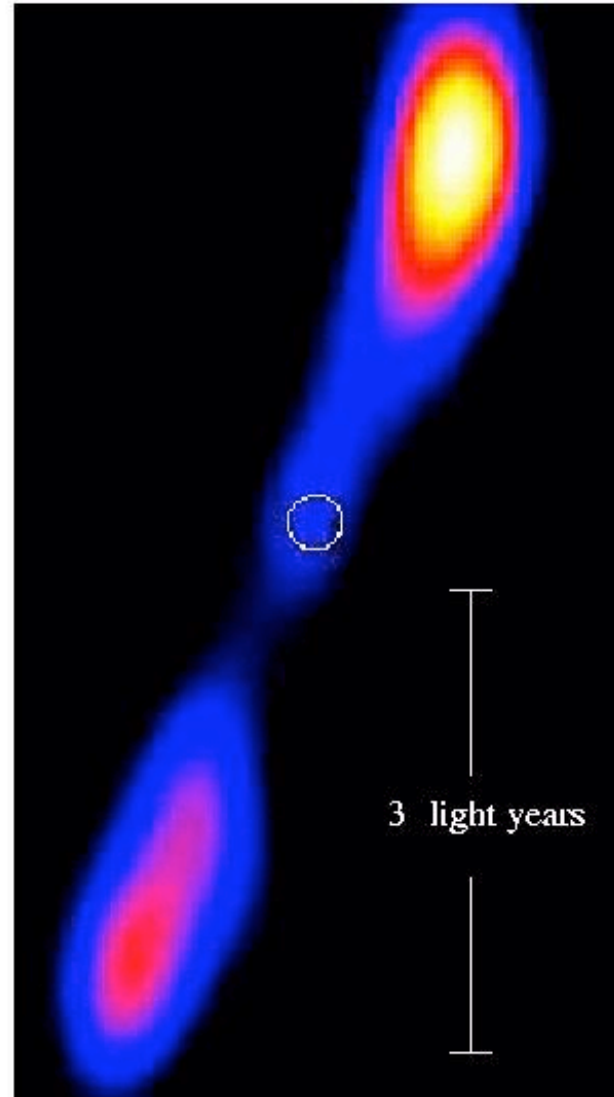
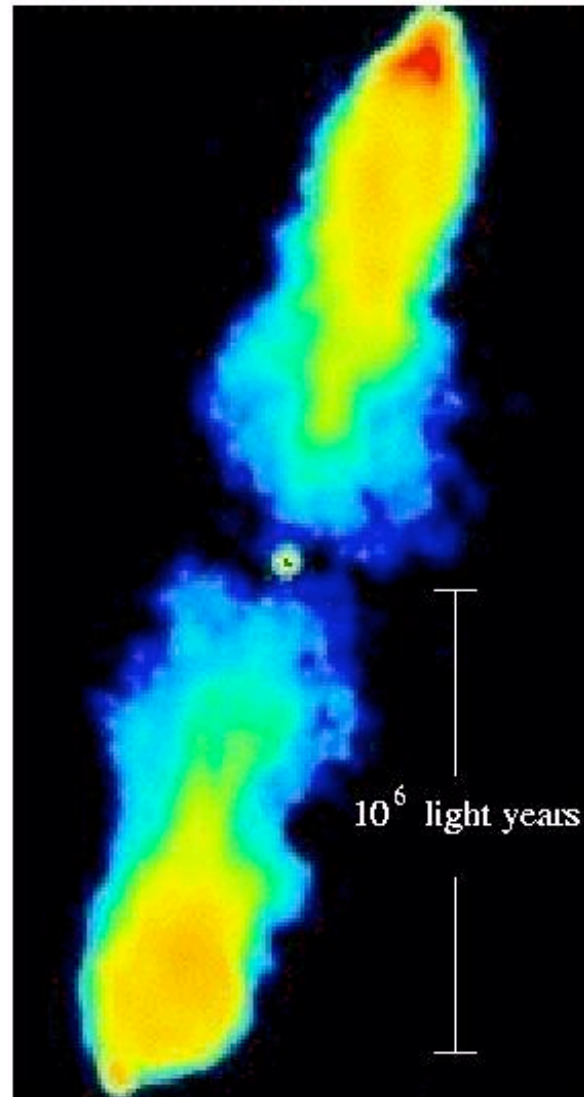
Fundamental Questions

- ★ Why do jets form? What are the conditions near the black hole leading to jet creation, collimation etc.? Does the black hole play a special role or could it be anything in there?
- ★ Jet structure/geometry/bulk velocity
- ★ Jet intrinsic physics: Matter (e^-p vs e^+e^-) vs. Poynting flux dominated?
- ★ Is BH physics generic? If yes, should scale predictably between stellar/galactic systems

Accreting BHs with Jets - Mass (Size) Scales

QUASAR (AGN)

MICROQUASAR (XRB)



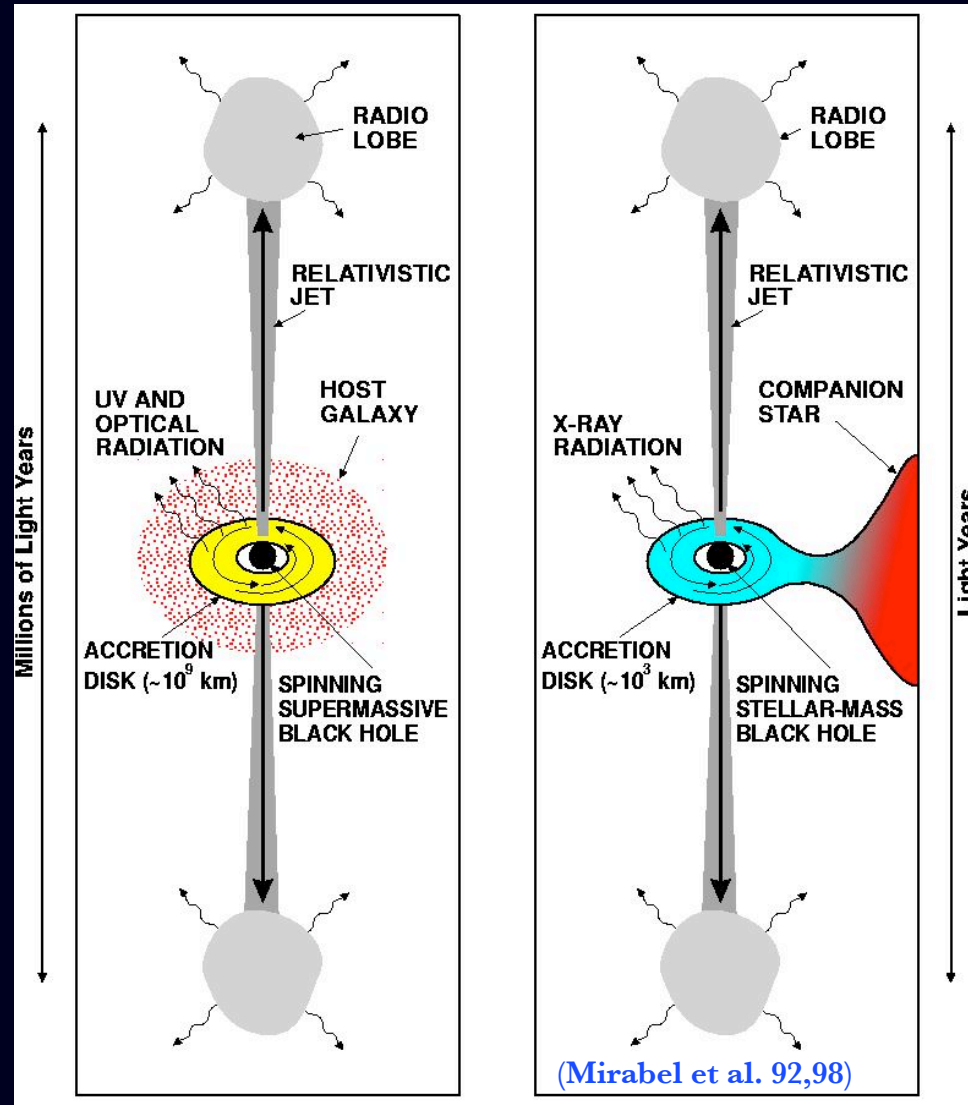
(Mirabel et al. 92,98)

Comparing accretion across the mass scale

QUASAR (AGN)

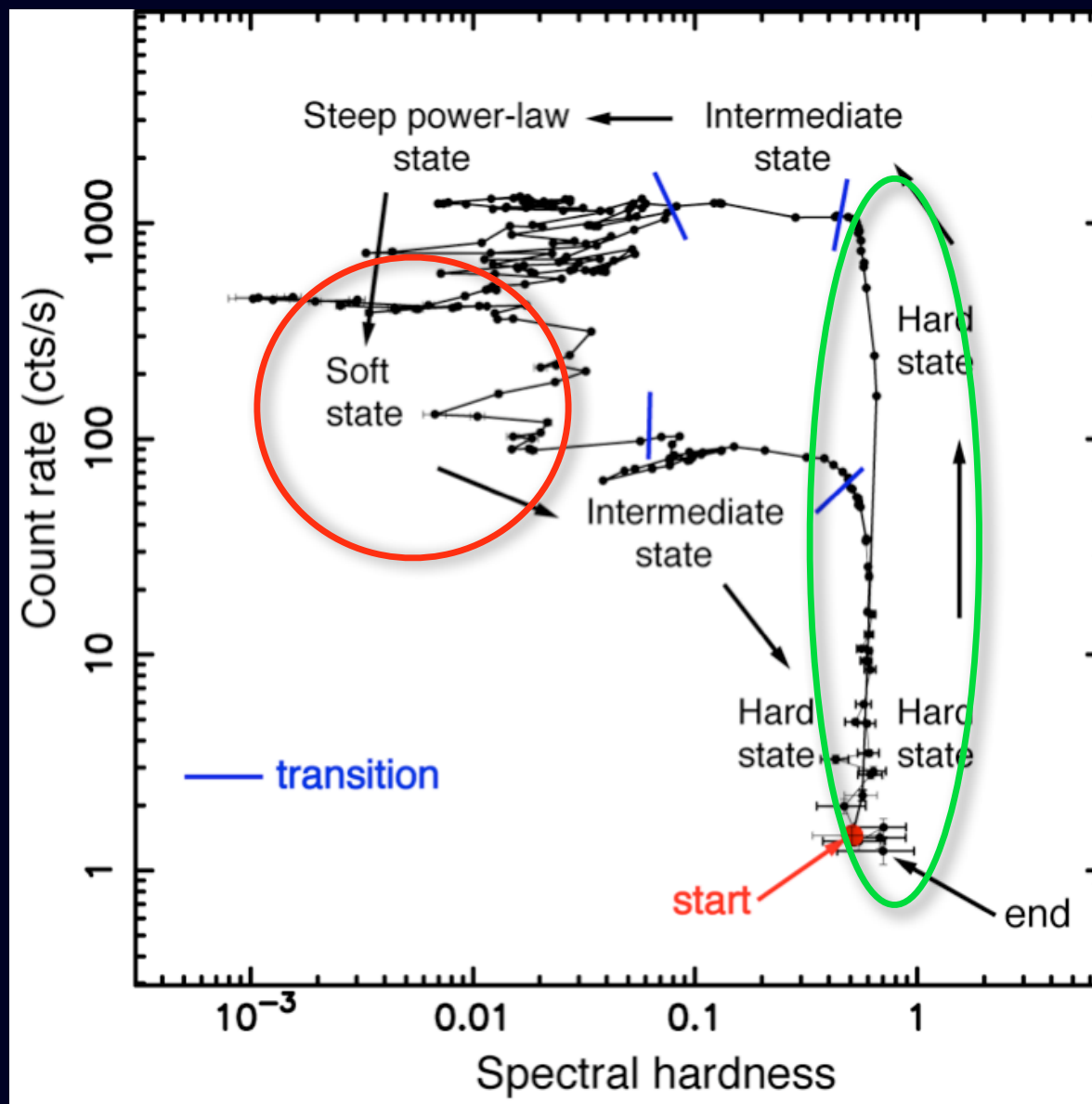
MICROQUASAR (XRB)

10^{4-5} yrs!



1 day

XRB accretion states

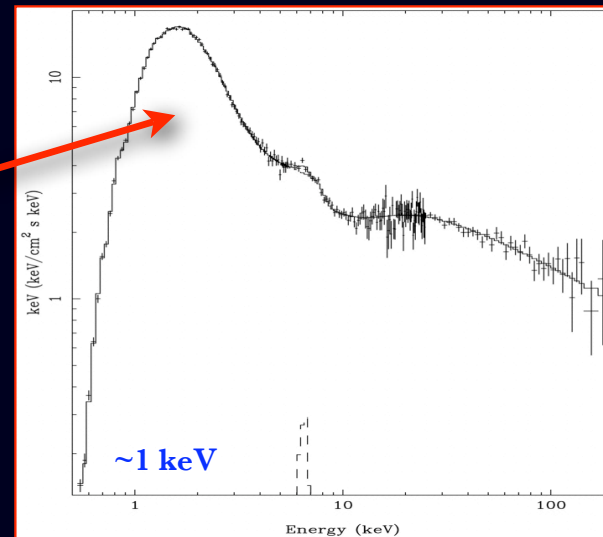
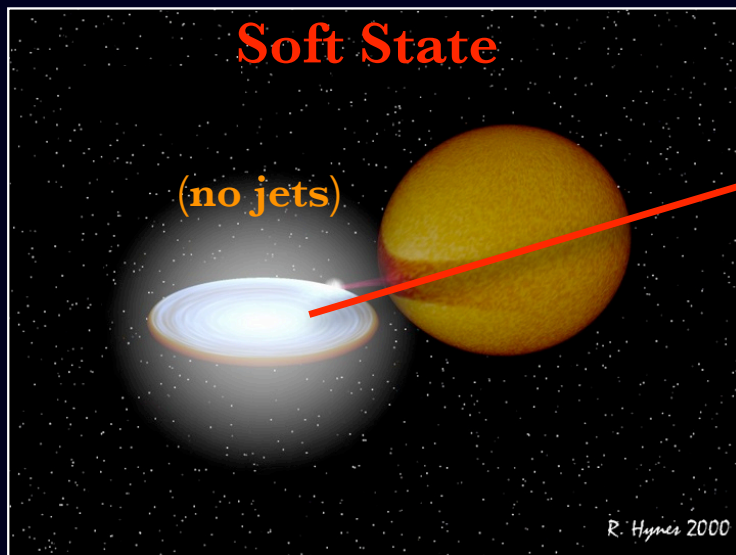


(Homan & Belloni 2004)

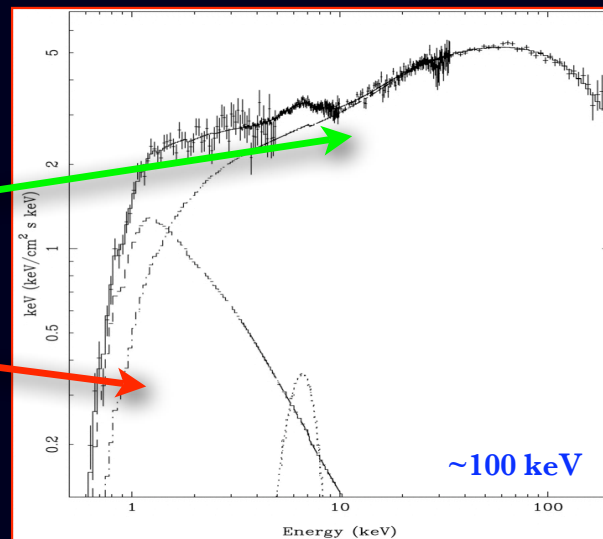
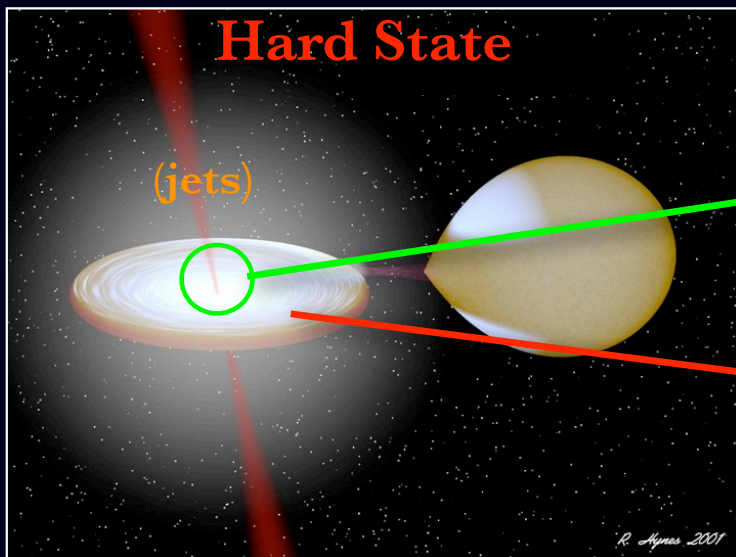
(BH) XRB Accretion States

Luminosity

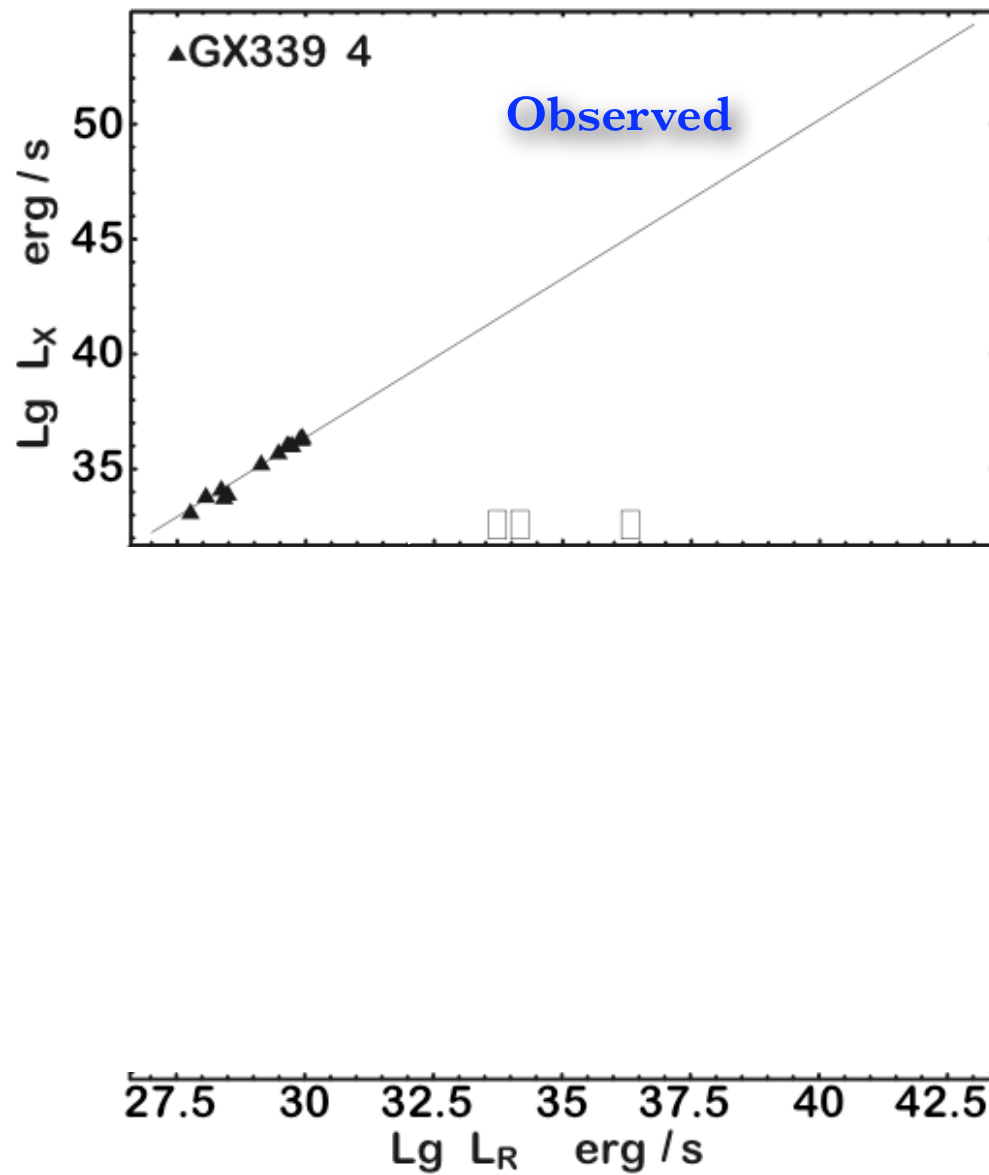
$\approx L_{\text{Edd}}$



$\ll L_{\text{Edd}}$

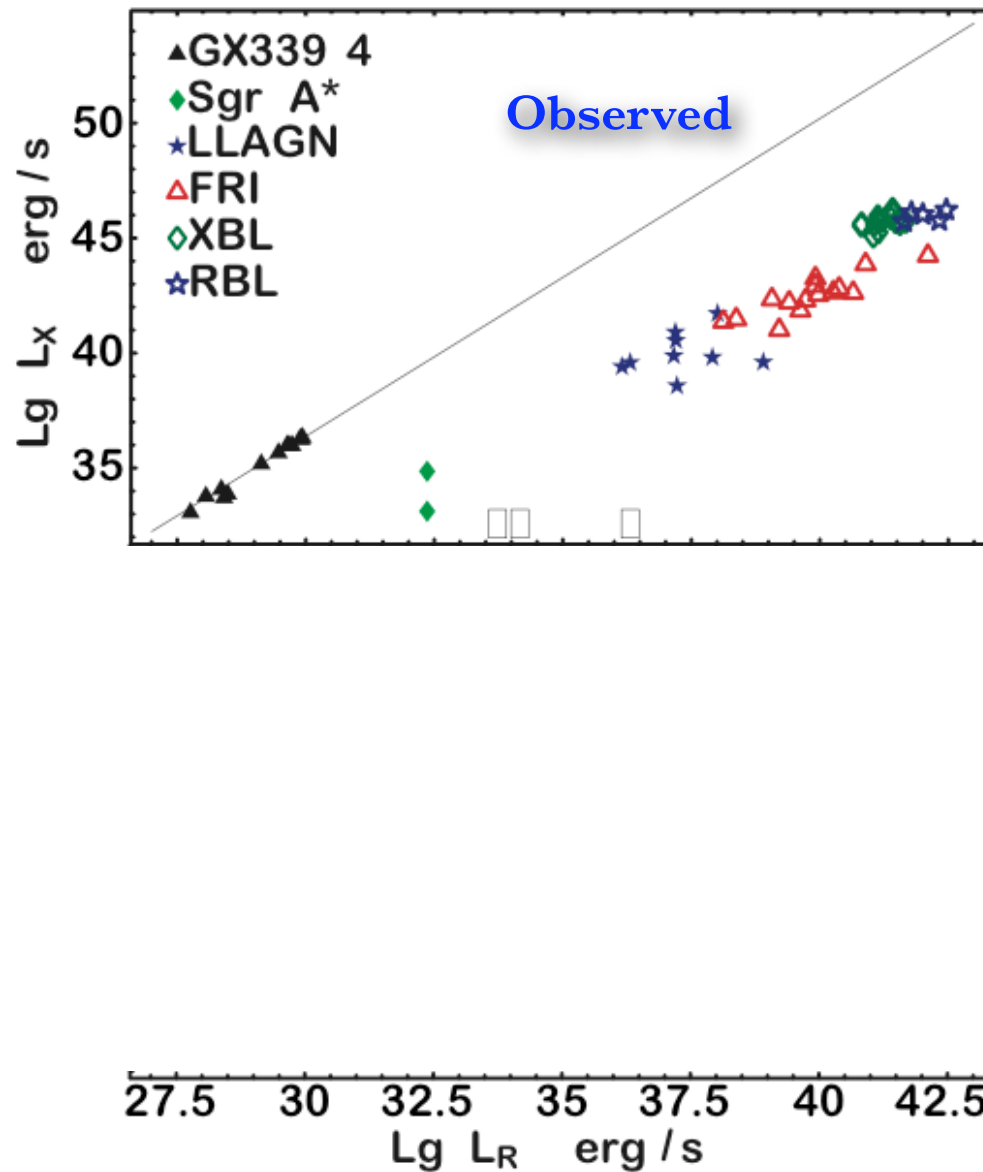


Fundamental plane of BH accretion!



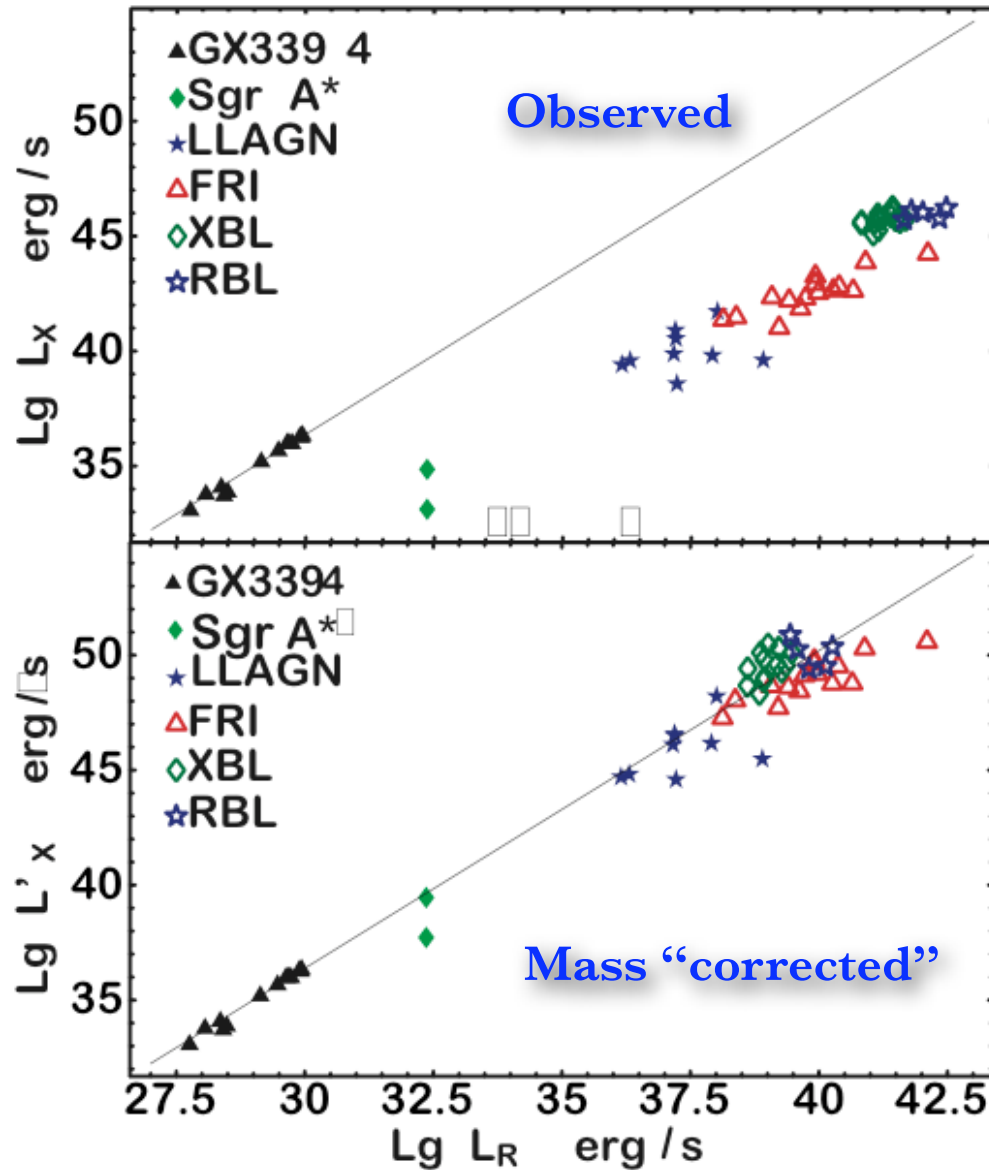
(Falcke, Körding & Markoff 2004
Merloni, Heinz & diMatteo 2003)

Fundamental plane of BH accretion!



(Falcke, Körding & Markoff 2004
Merloni, Heinz & diMatteo 2003)

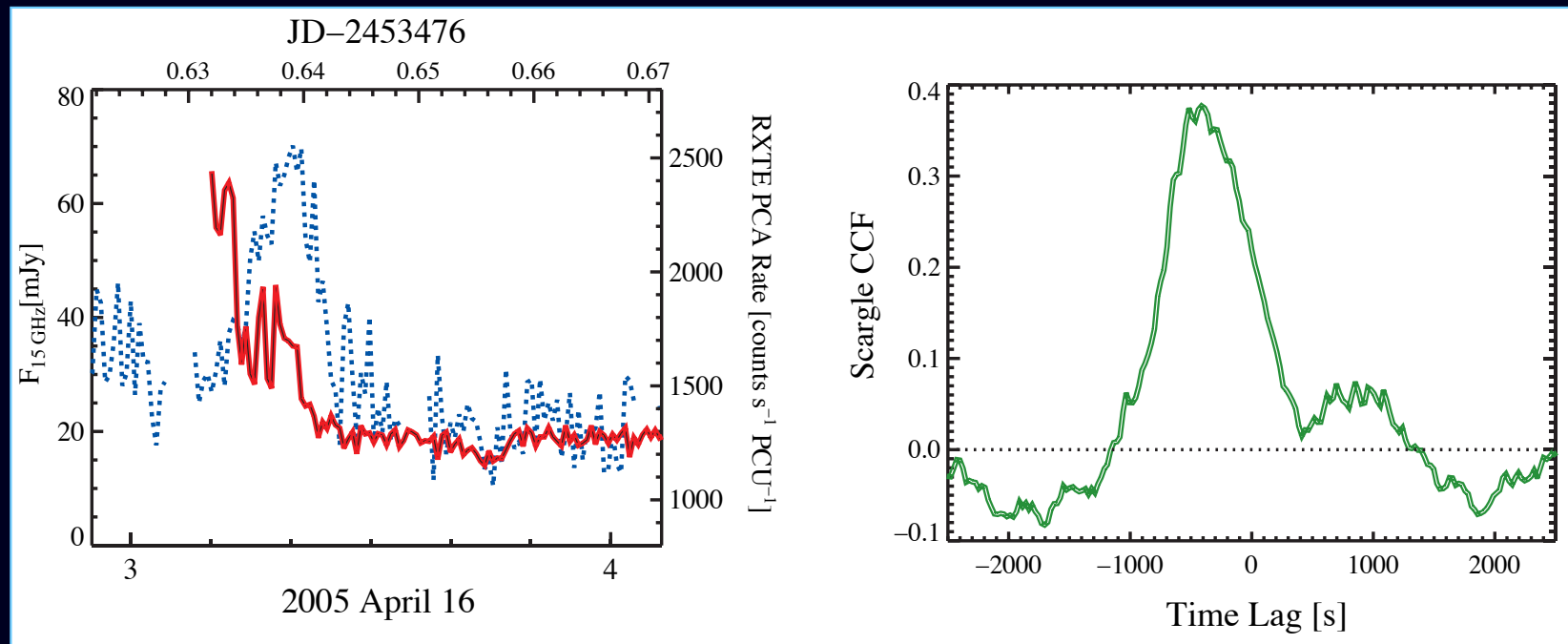
Fundamental plane of BH accretion!



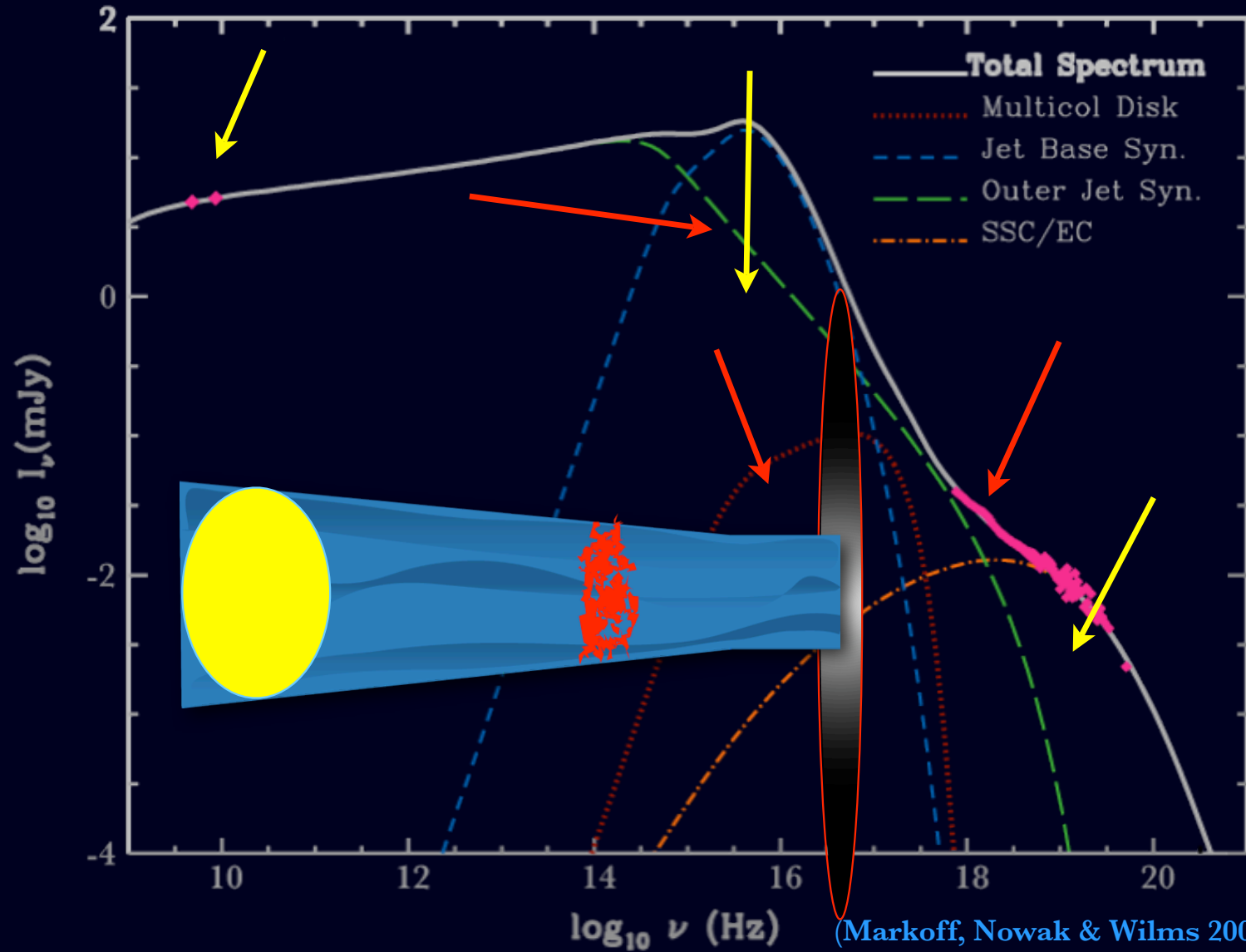
(Falcke, Körding & Markoff 2004
Merloni, Heinz & diMatteo 2003)

Jet structure/plasma flow

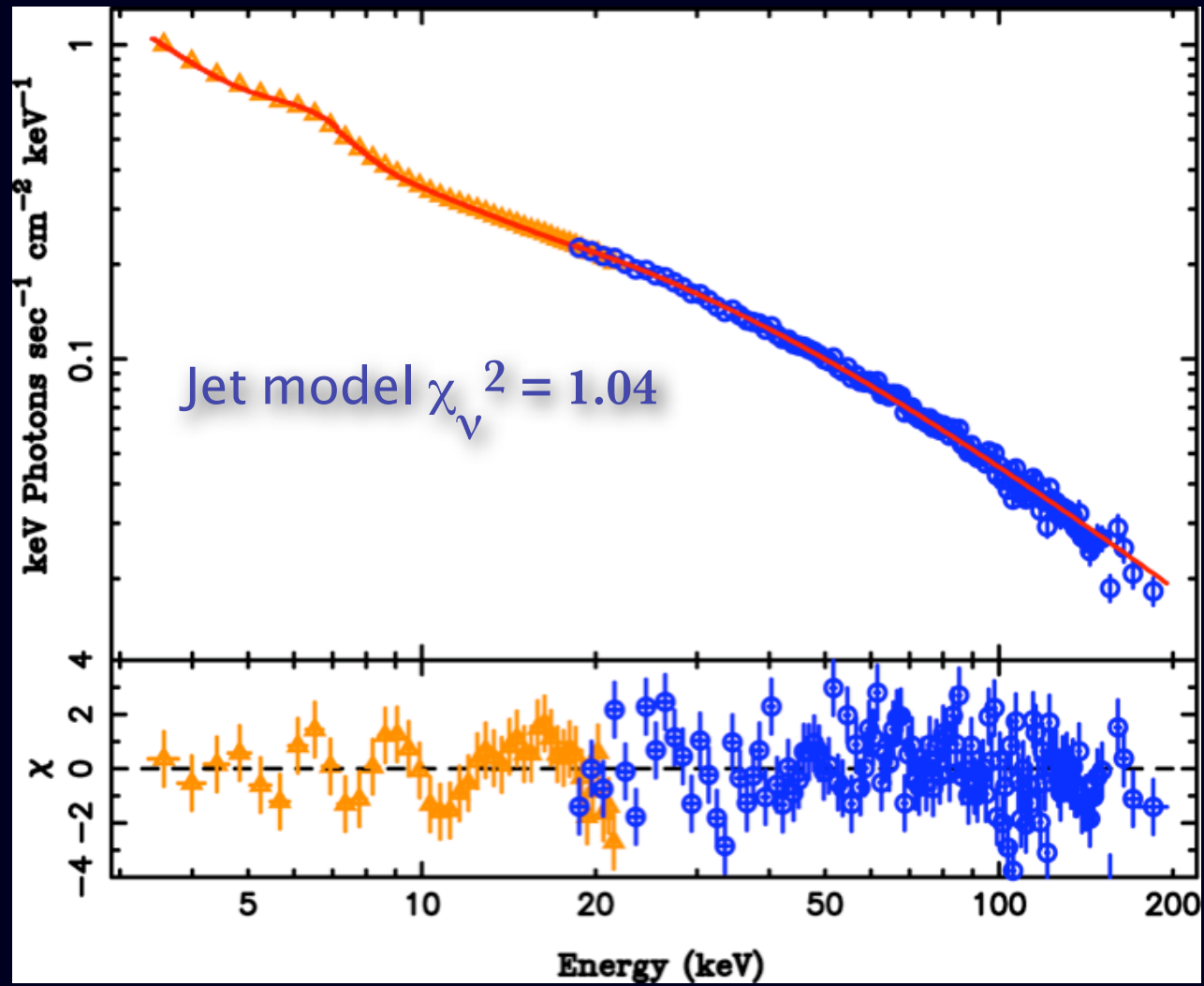
- * Measuring lags between frequencies, and amplitude/shape decay/evolution of flare events
- ▶ Gives information about plasma velocities and internal physics, e.g., Cyg X-1: (Wilms et al. 2007)



Model Components

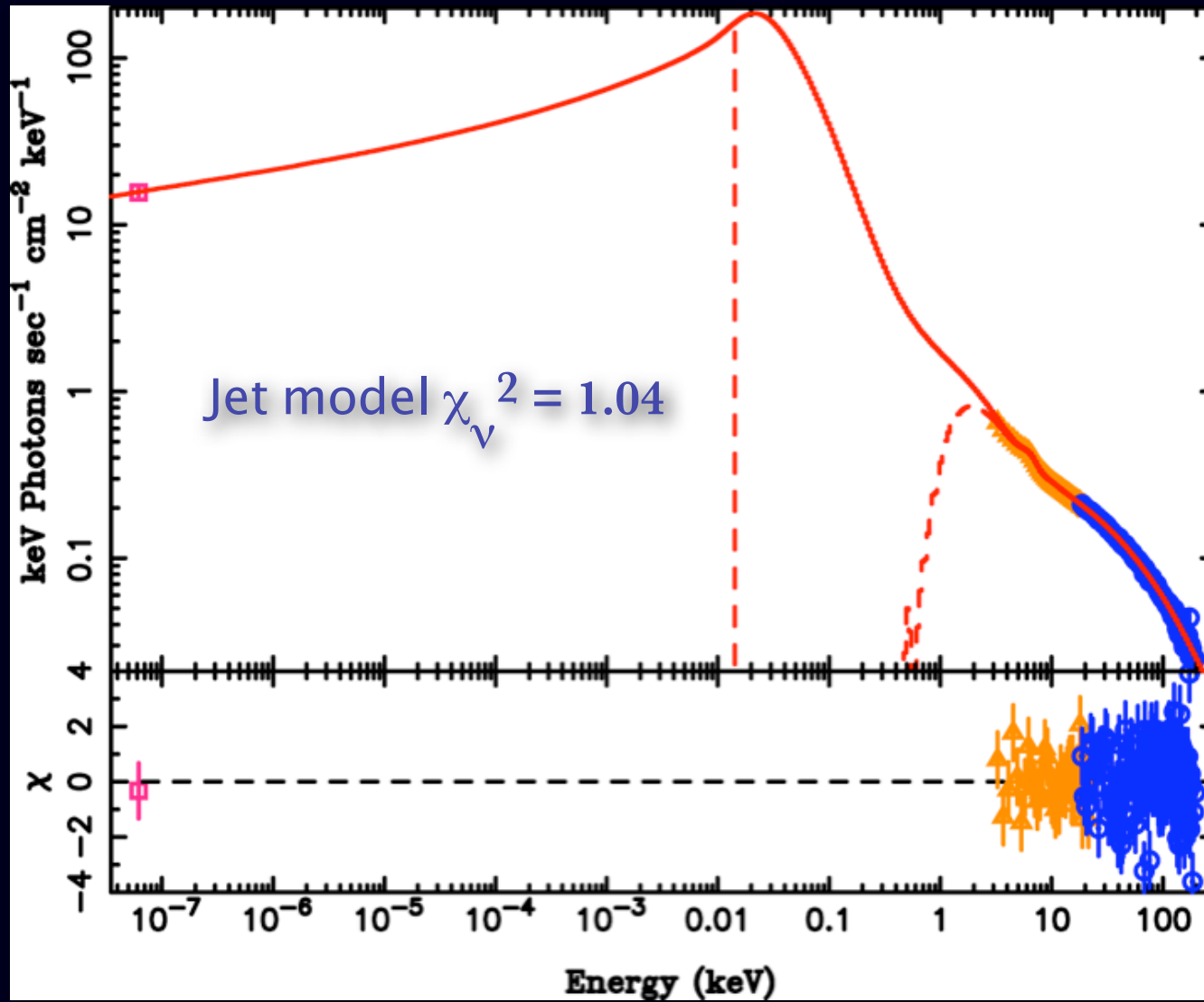


Radio/Xray only: Cyg X-1 spectrum



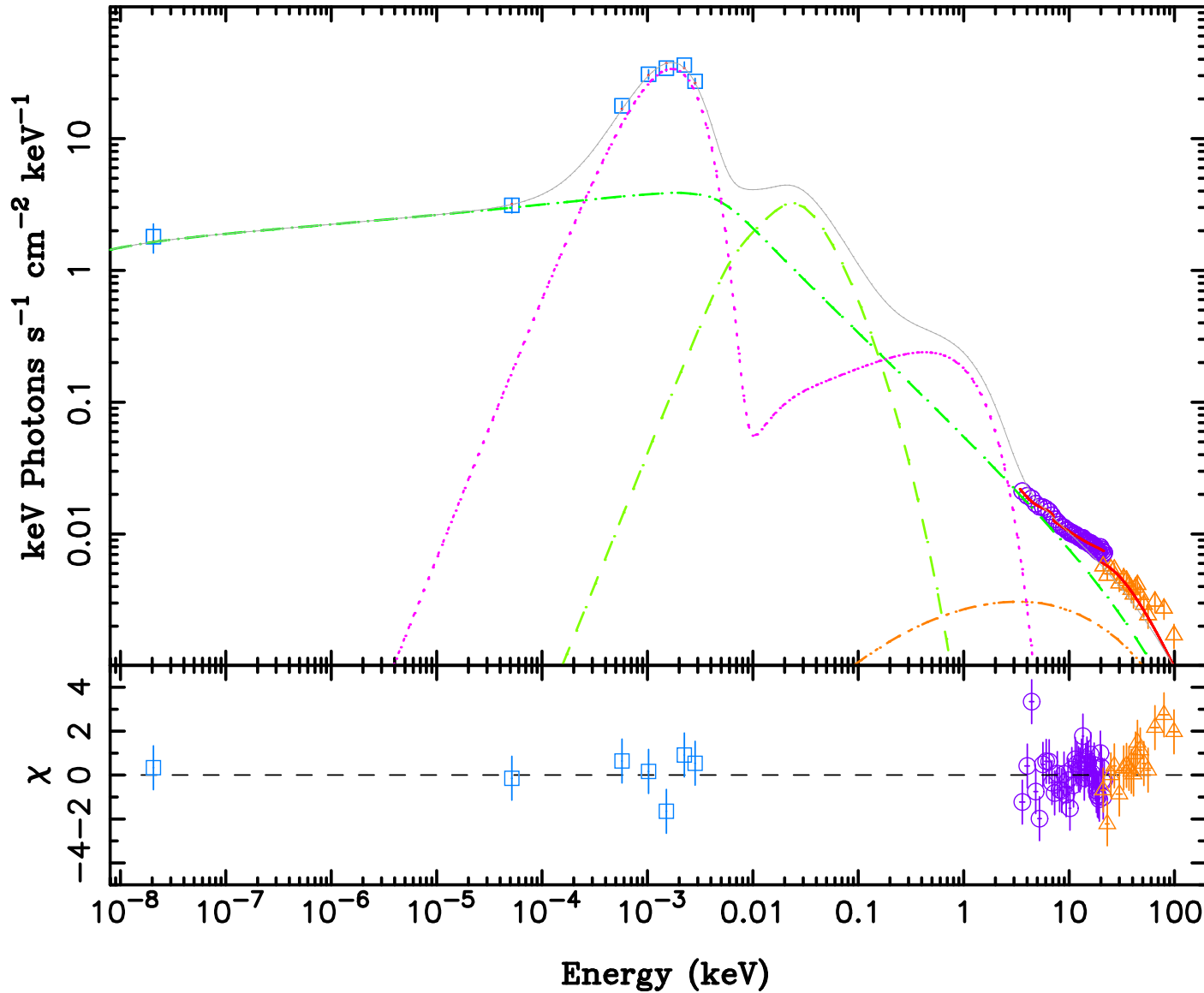
(Markoff & Nowak 2004; Markoff, Nowak & Wilms 2005)

Radio/Xray only: Cyg X-1 spectrum



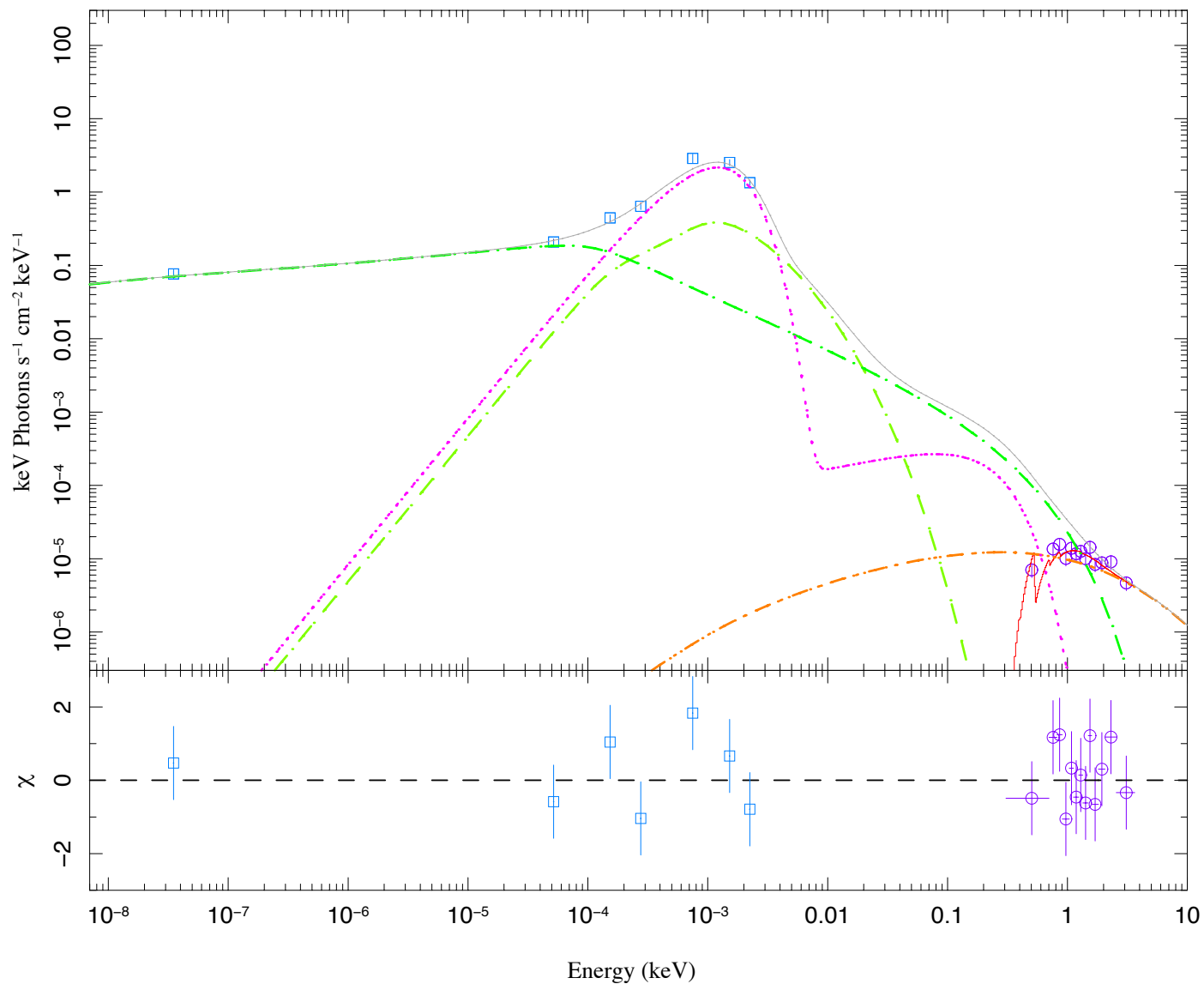
(Markoff & Nowak 2004; Markoff, Nowak & Wilms 2005)

New constraints from IR/Optical I: GRO J1655-40



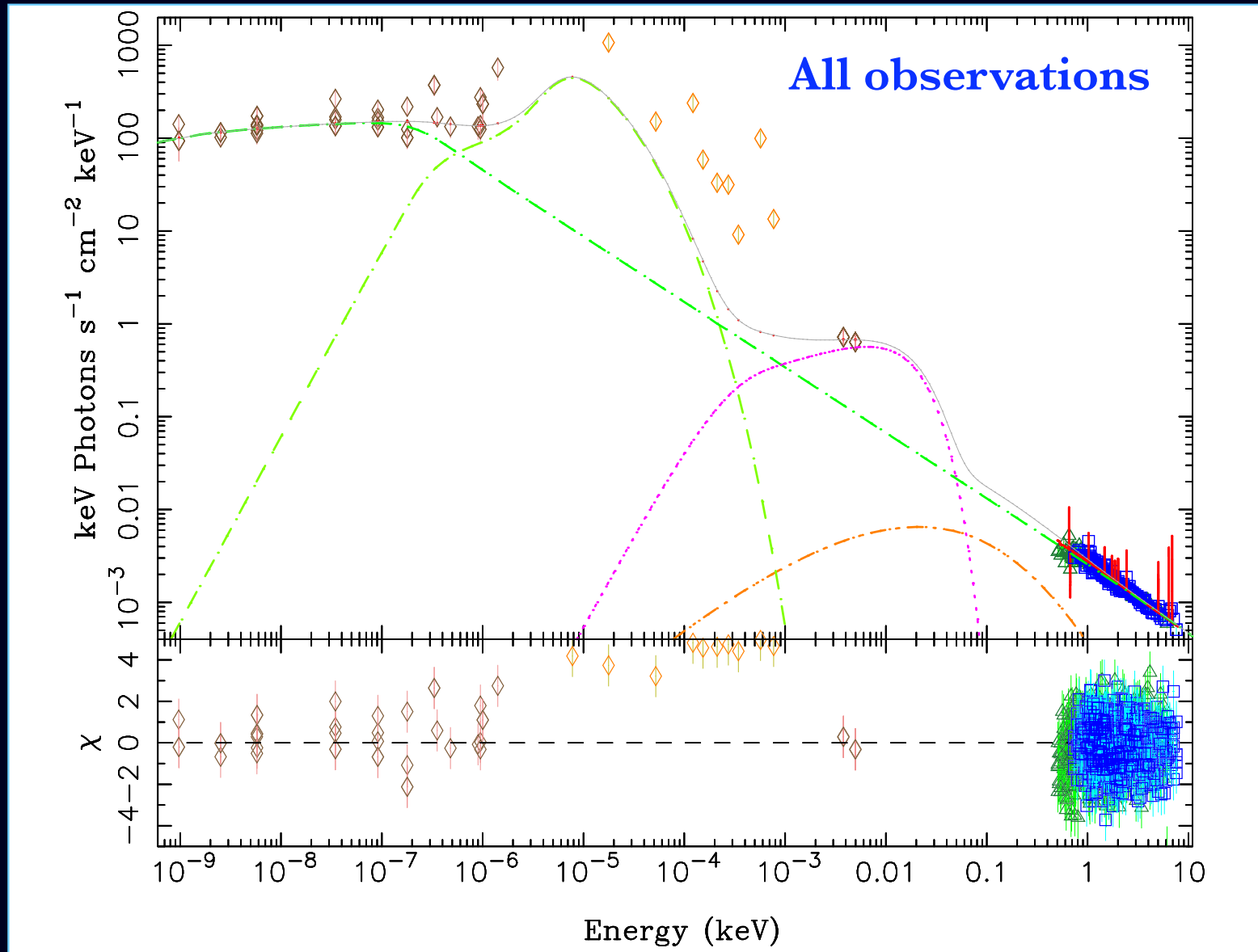
(Migliari et al. 2007)

New constraints from IR/Optical II: A0620-00



(Gallo et al. 2007)

M81* simultaneous campaign

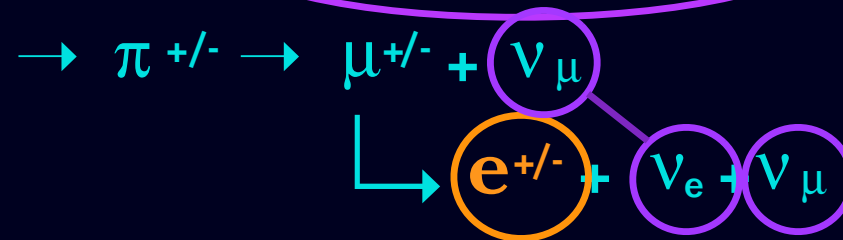


(Markoff et al. 2007)

“Multimessenger” = multiwavelength

- ★ If protons accelerated in the jets, additional contribution from hadronic interactions

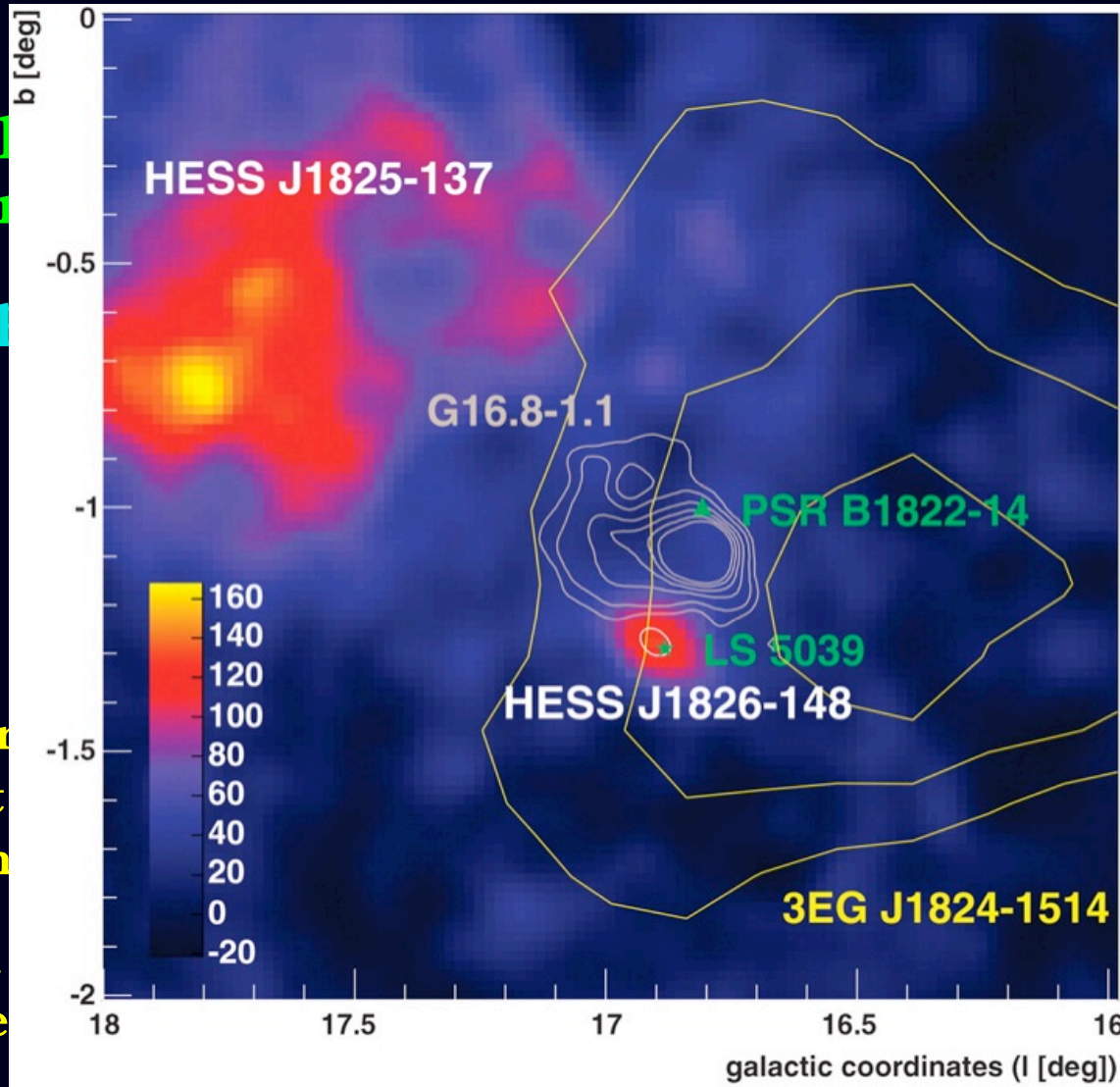
pp or $p\gamma \rightarrow \pi^0 \rightarrow 2\gamma$ (~ 70 MeV CM frame)



- Submitted proposals to trigger MAGIC and IceCube with first bright LOFAR transients, as well as mutual monitoring of known sources
- Will directly address questions about particle acceleration and internal energetics/matter content

“Multimessenger” = multiwavelength

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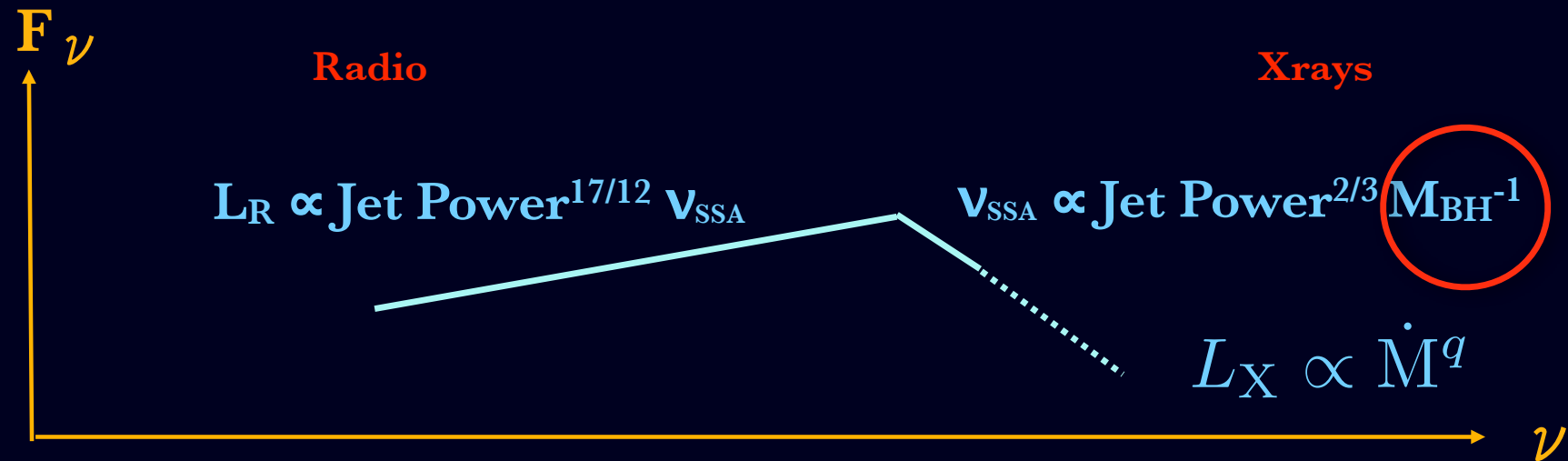
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Summary

- ★ **LOFAR promises significant progress in understanding jet physics in accreting black holes** ⇒ exploiting multi- λ and 8 orders of magnitude in mass/power scales!
 - ⇒ *RSM*: helping understand radio-dominated states
 - ⇒ *Low- v alone*: jet energetics, e^- distribution, prompt synchrotron
 - ⇒ *LOFAR in combination with multiwavelength* : new constraints on bulk velocity, geometry, emission mechanisms
 - ⇒ *LOFAR/ γ -ray*: expanding spectrum on both ends, determination of hadronic component? Eventually in combination with direct neutrino detections?
 - ⇒ *Spectral fitting*: constraints on geometry and plasma conditions very close to BH ⇒ clues about jet formation

A few extra slides

Predictions for radio/X-ray correlations



For objects with the *same* mass:

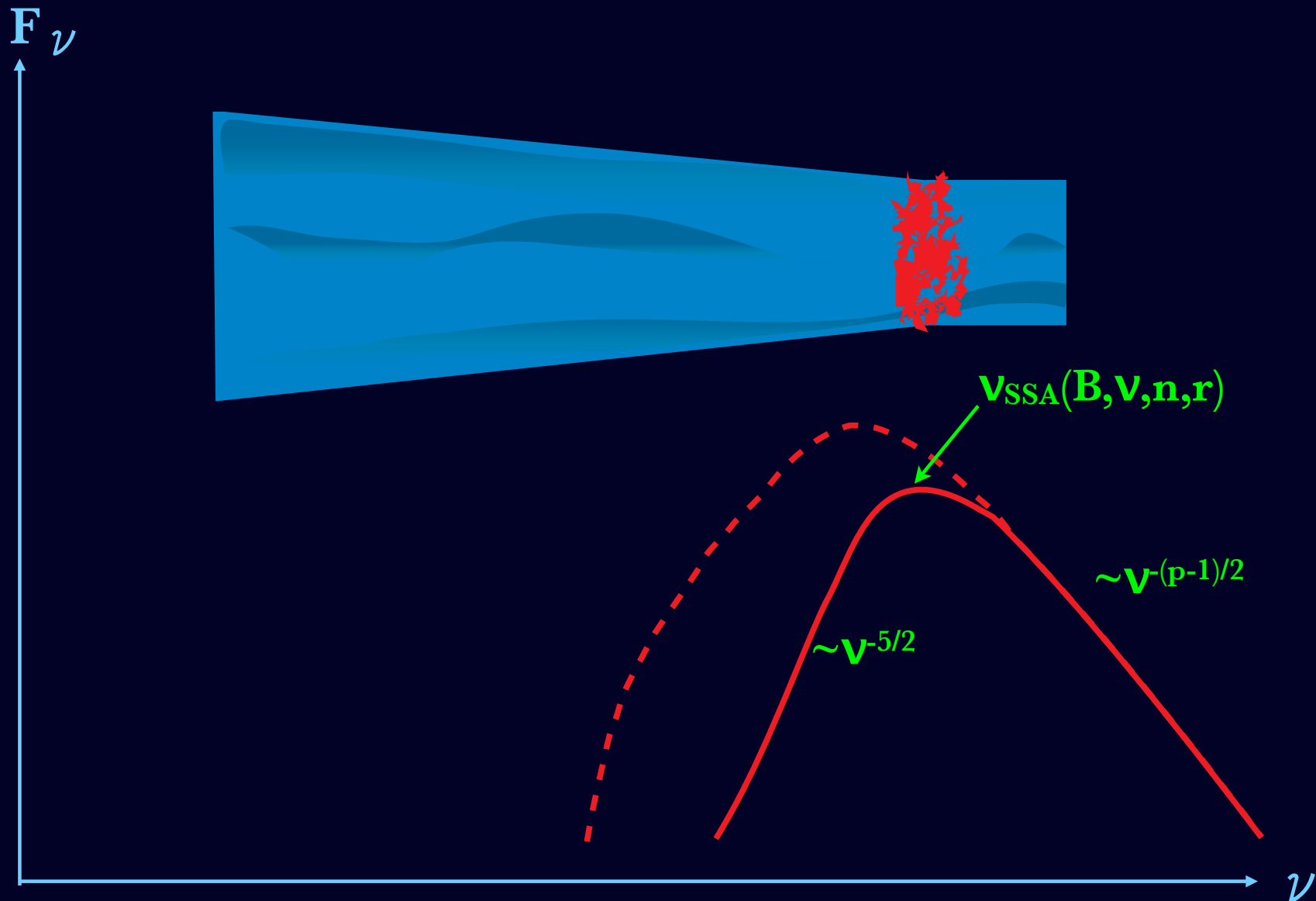
$$L_R \propto L_X^m \quad m = \frac{\frac{17}{12} - \frac{2}{3}\alpha_R}{q} \approx \frac{1.4}{q}$$

Synchrotron: $q=2$, ADAF/RIAF: $q=2-2.3$,

Radiatively efficient disk/corona: $q=1 \Rightarrow$ problematic

(Markoff et al. 2003, Merloni, Heinz & diMatteo 2003, Falcke, Körding & Markoff 2004)

Jet energetics: radiating particles



Jet energetics: radiating particles

F_ν

