



# Amplified Radio Emission from Air Showers in Thunderstorms

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# Outline

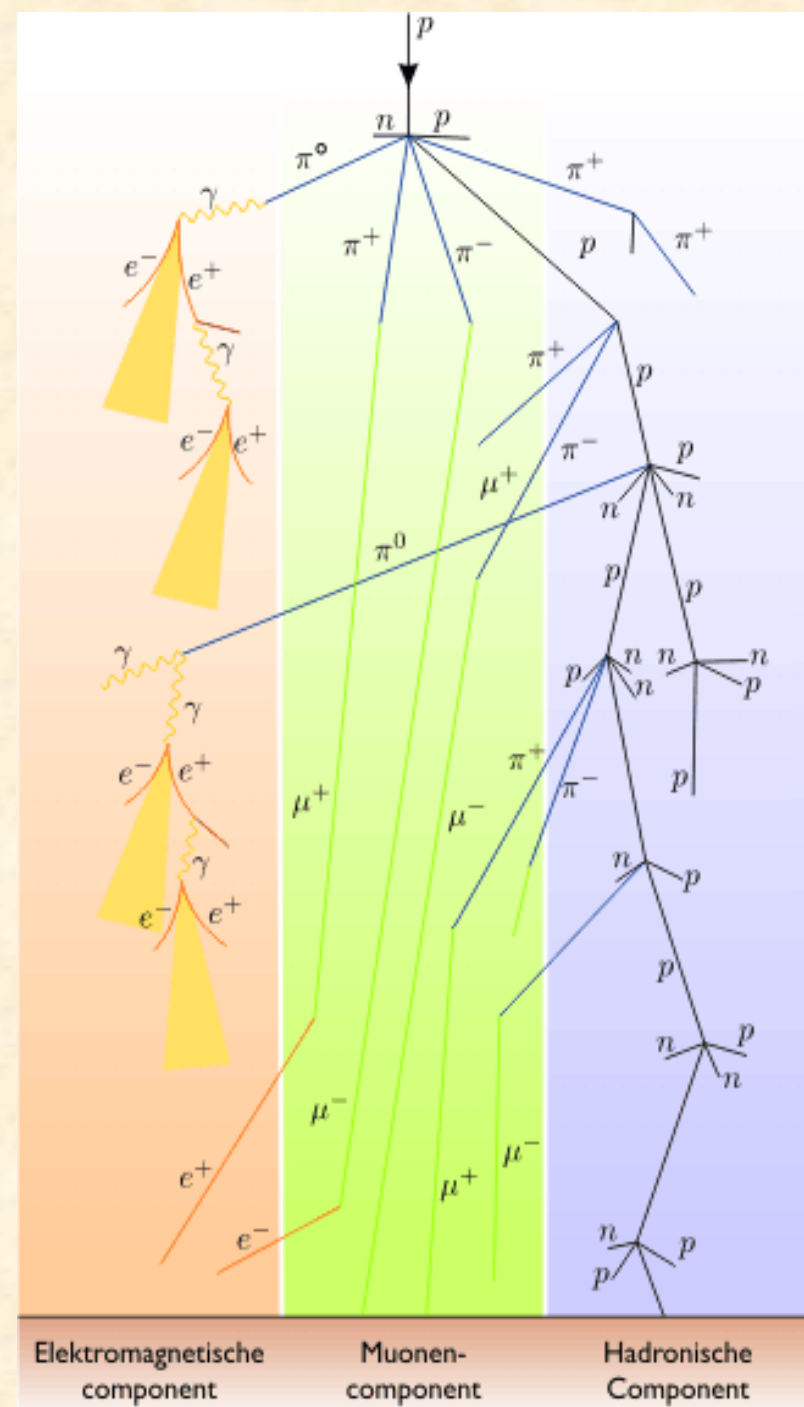
- Part I: Emission mechanism
- Part II: LOPES thunderstorm results
- Part III: Other thunderstorm research

# Part I: Emission Mechanism

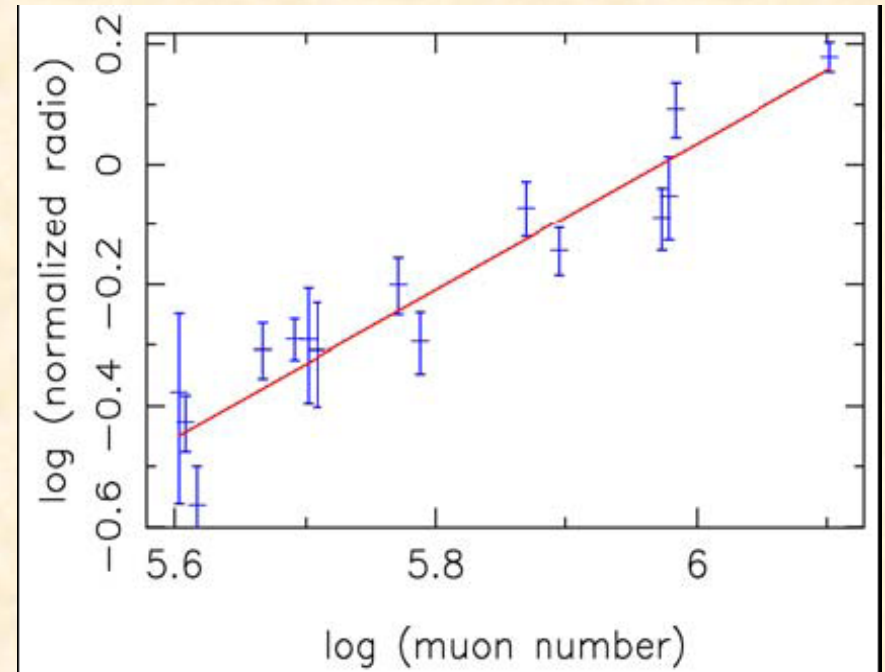
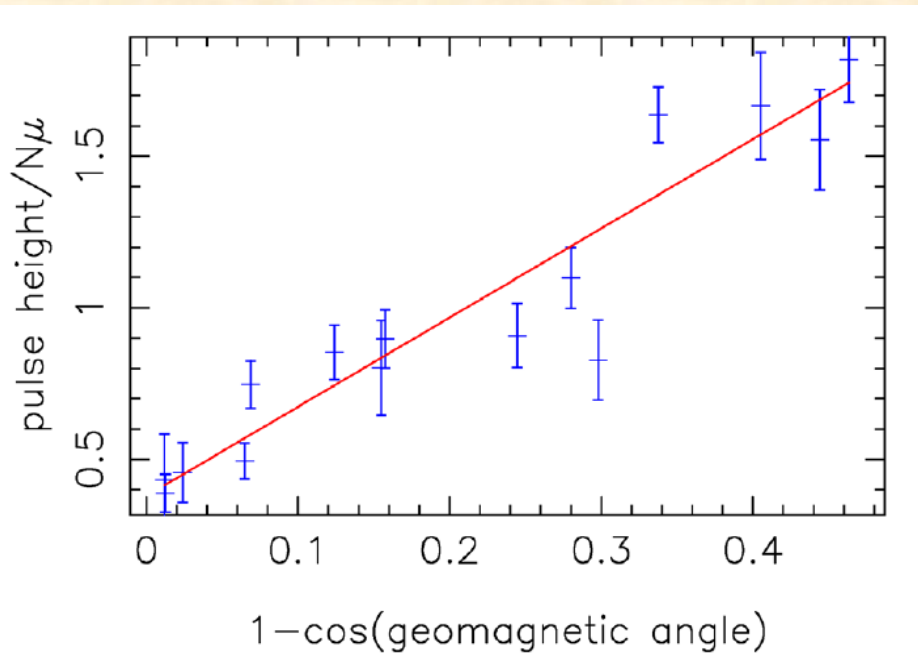


# Air Showers

- Created by high energy cosmic rays
- Electromagnetic component contains electron-positrons pairs
- Radio emission by synchrotron radiation



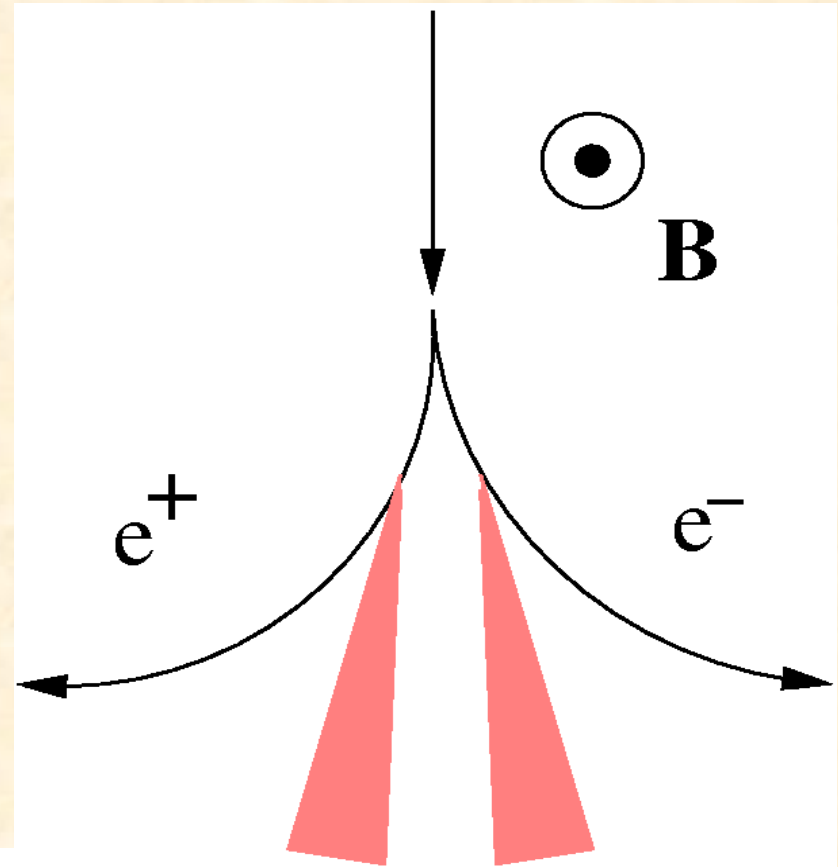
# Radio Emission from Air Showers



- Emission driven by Earth's magnetic field
- ...how about atmospheric electric fields?

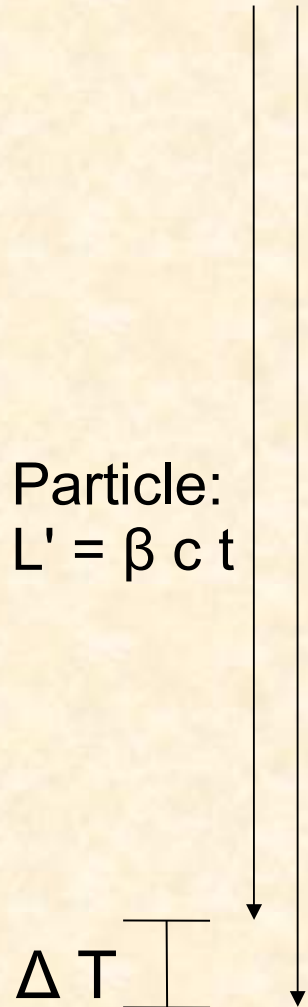
# Emission mechanism

- Geosynchrotron
- Same sign for  $e^+/e^-$
- Pulse height:  
 $E \sim \gamma^3 F_{\perp}$



$$\mathbf{E}(\mathbf{x}, t) = \frac{e}{c} \left[ \frac{\mathbf{n} \times [(\mathbf{n} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}]}{(1 - \boldsymbol{\beta} \cdot \mathbf{n})^3 R} \right]_{\text{ret}}$$

# Coherency



Radiowave:  
 $L = (c/n) t$

Particle:  
 $L' = \beta c t$

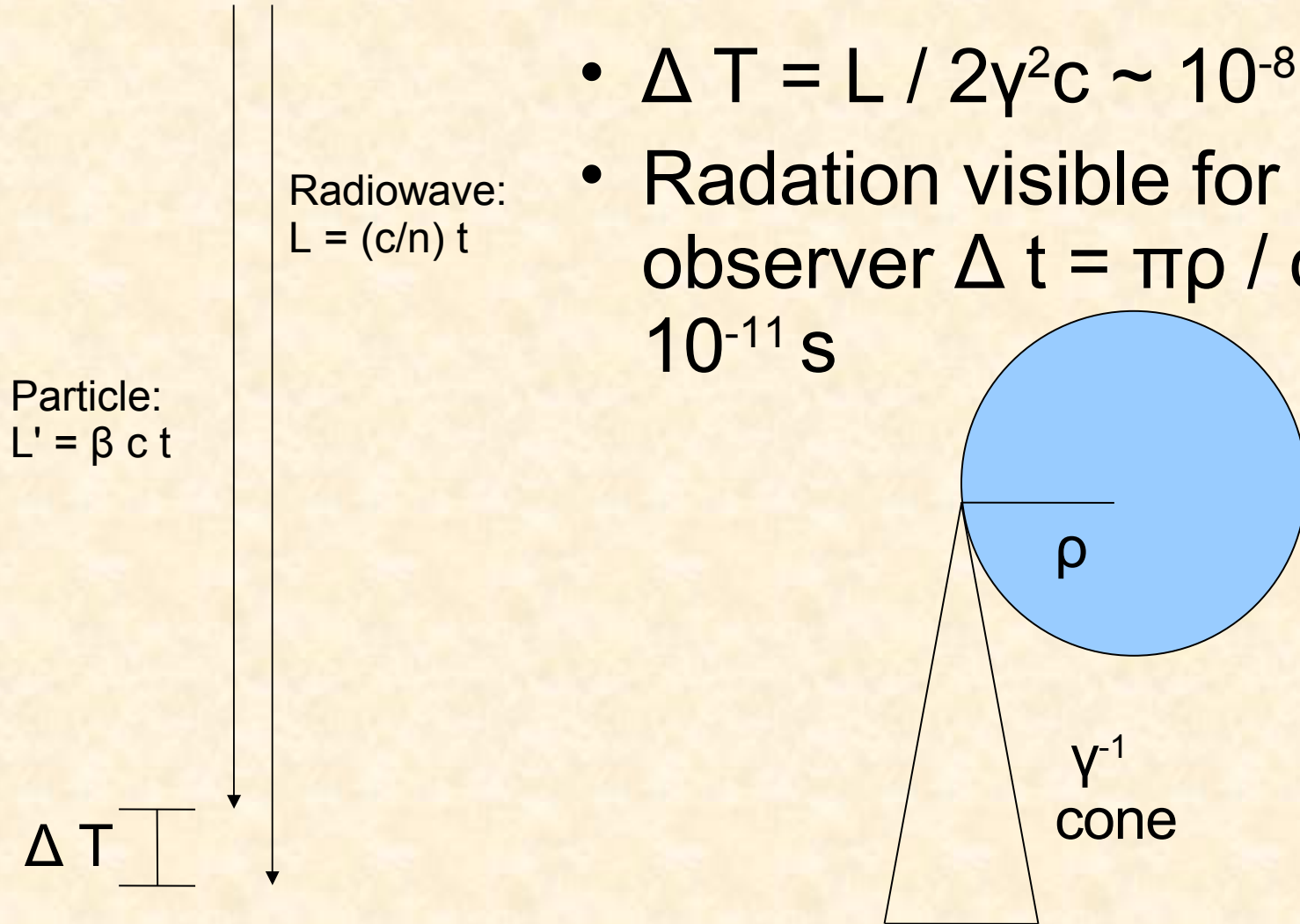
- Total shower length  $L$  (few kilometers)

- $\Delta x = (c/n) t - \beta c t$   
 $\approx L / 2\gamma^2$

- $\Delta T = L / 2\gamma^2 c \sim 10^{-8} \text{ s}$

# Coherency

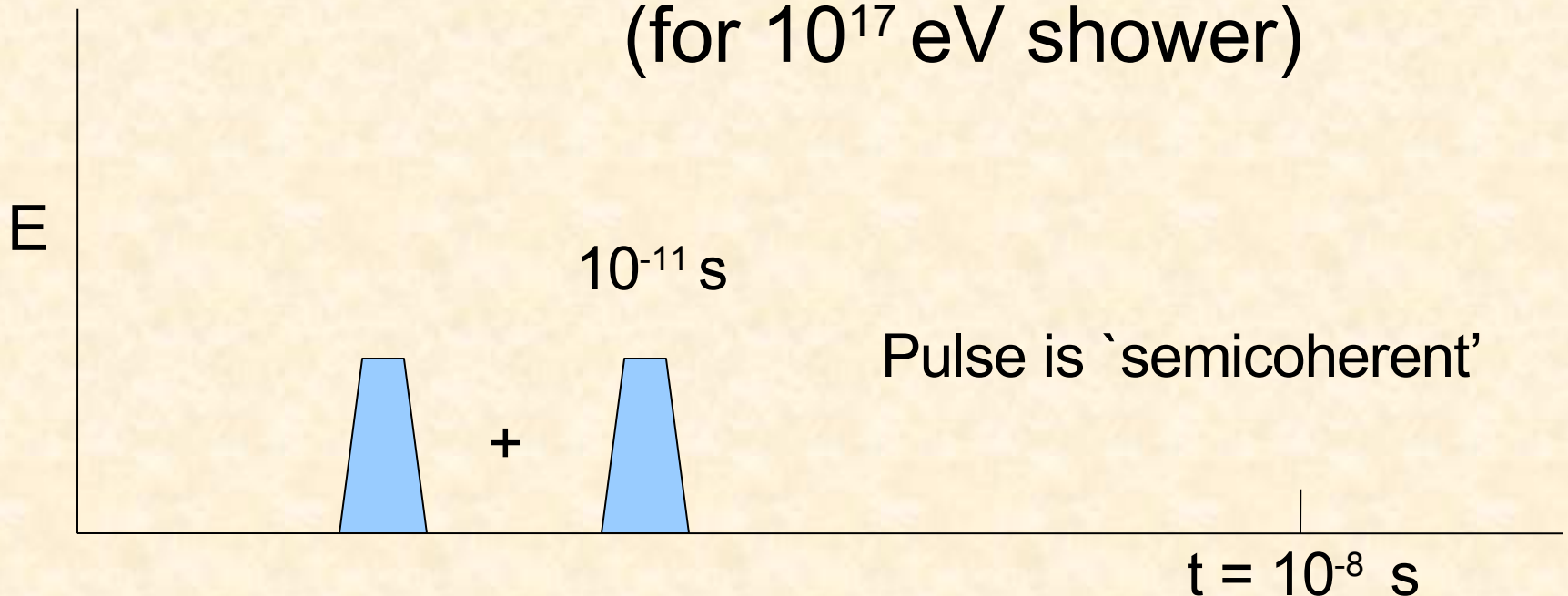
- $\Delta T = L / 2\gamma^2 c \sim 10^{-8} \text{ s}$
- Radiation visible for observer  $\Delta t = \pi\rho / c\gamma^3 \sim 10^{-11} \text{ s}$



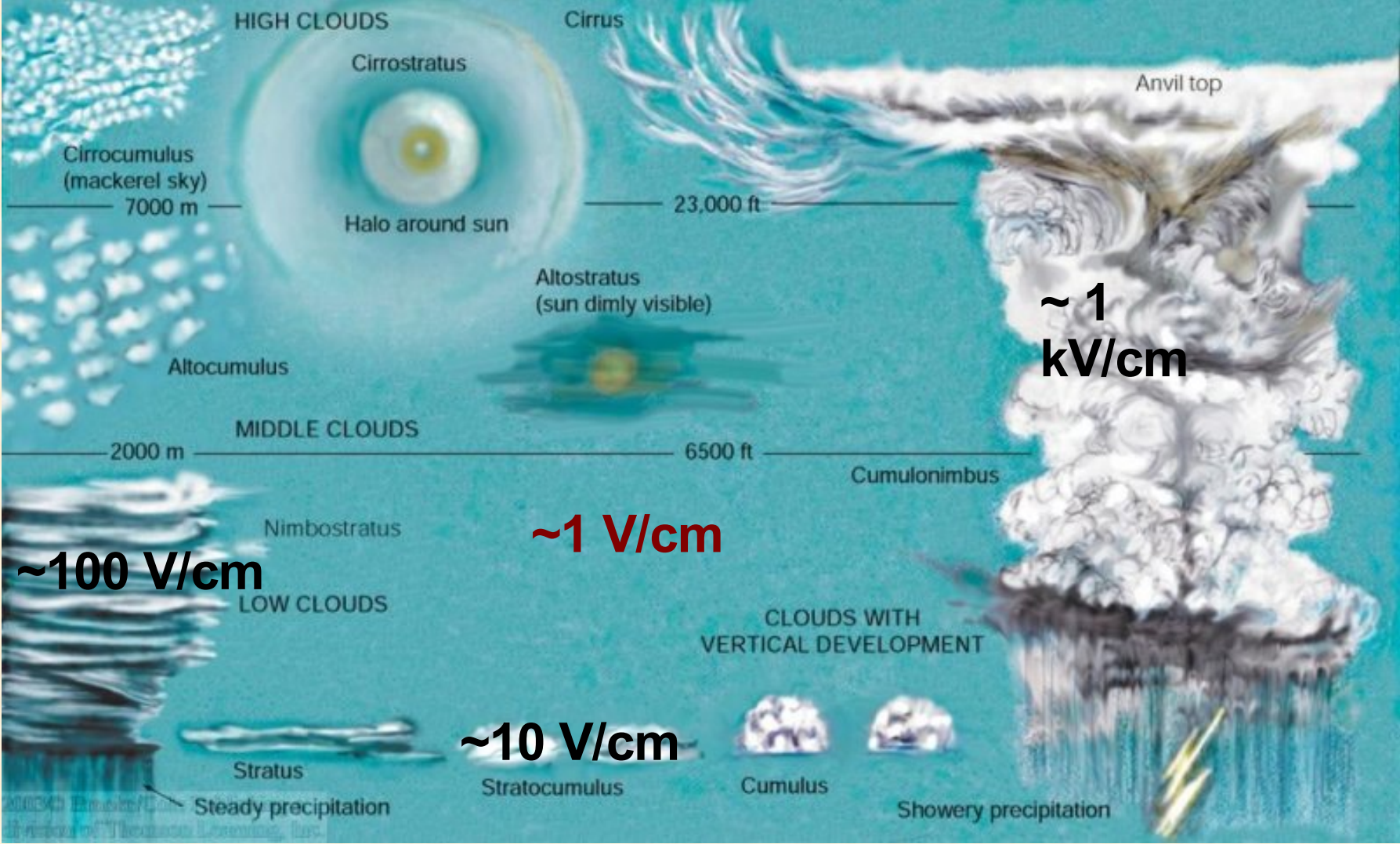


# Coherency

- $\Delta T = L / 2\gamma^2 c \sim 10^{-8} \text{ s}$
- $\Delta t = \pi\rho / c\gamma^3 \sim 10^{-11} \text{ s}$
- Number of pairs:  $N \sim 10^8$   
(for  $10^{17} \text{ eV}$  shower)



# Electric fields in atmosphere



# Electric field

- *Case 1: Perpendicular*

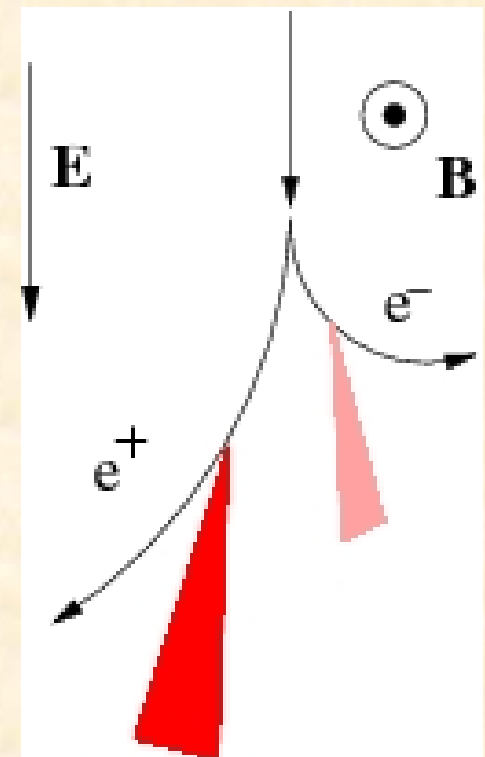
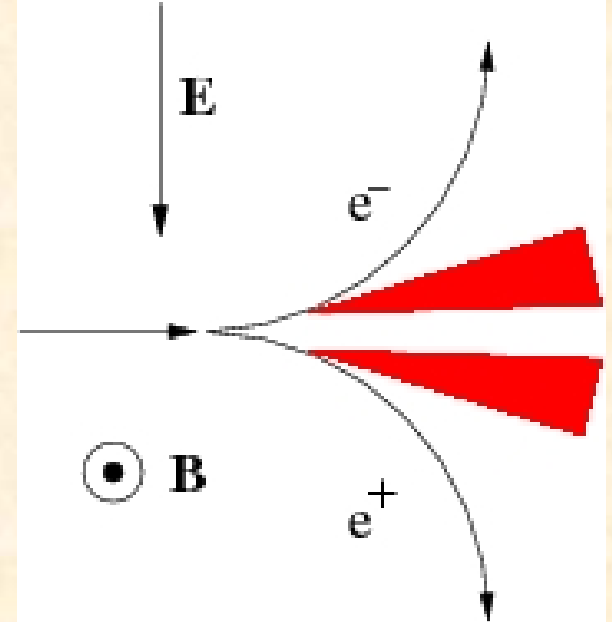
Increase in  $F_{\perp}$

- *Case 2: Linear*

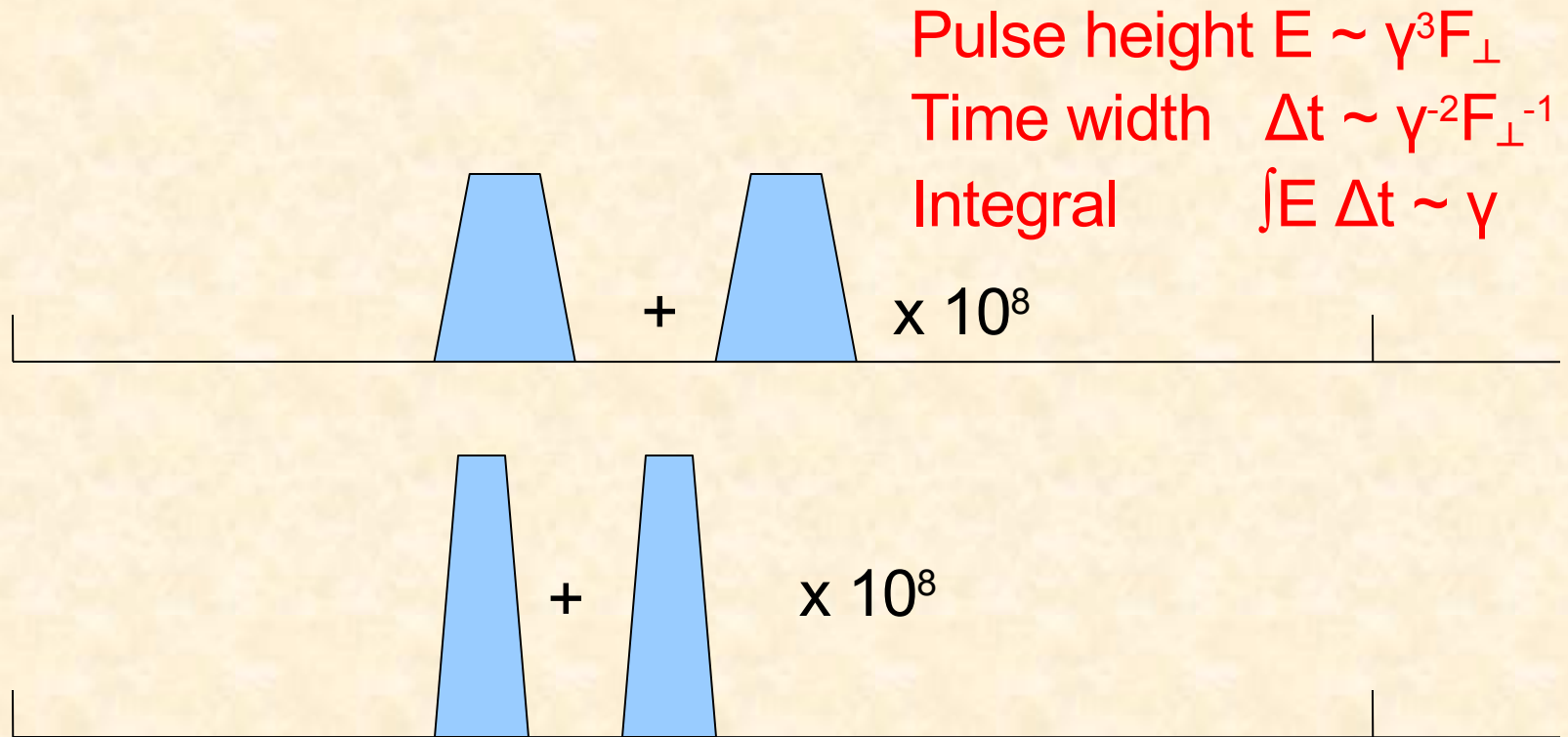
$e^{+}$  : Increase in  $\gamma$

$e^{-}$  : Decrease in  $\gamma$

(Radiation by  $F_{\parallel}$  is suppressed by a factor  $\gamma^2$ )

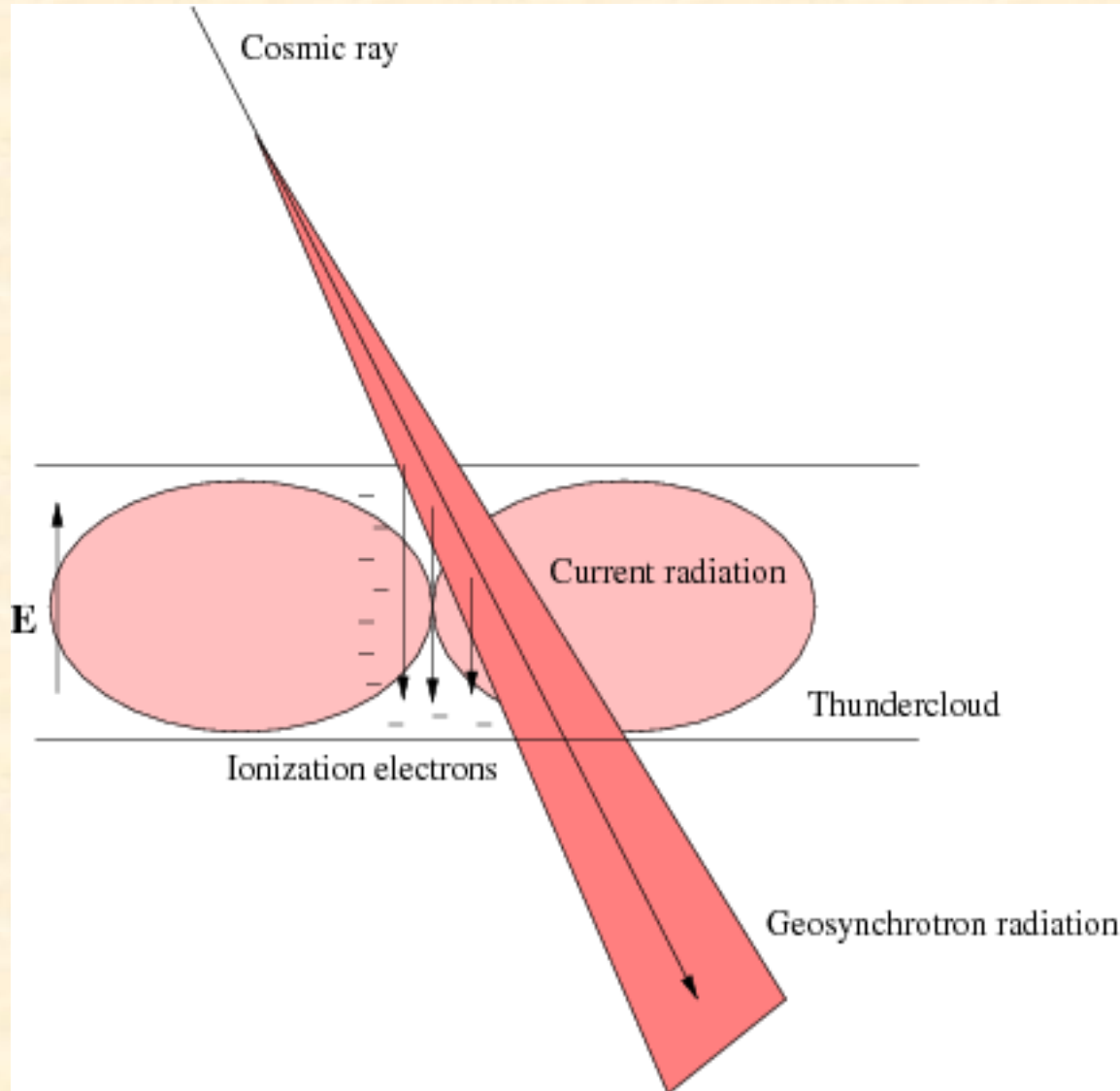


# Amplification of shower ?



- Amplification by linear acceleration more efficient
- Shower development crucial
- Monte Carlo simulation in progress

# Runaway breakdown radiation



- Current Pulse Radiation
- Pulse width ~ 100-300 ns

(dependent on attachment time of electrons to oxygen)

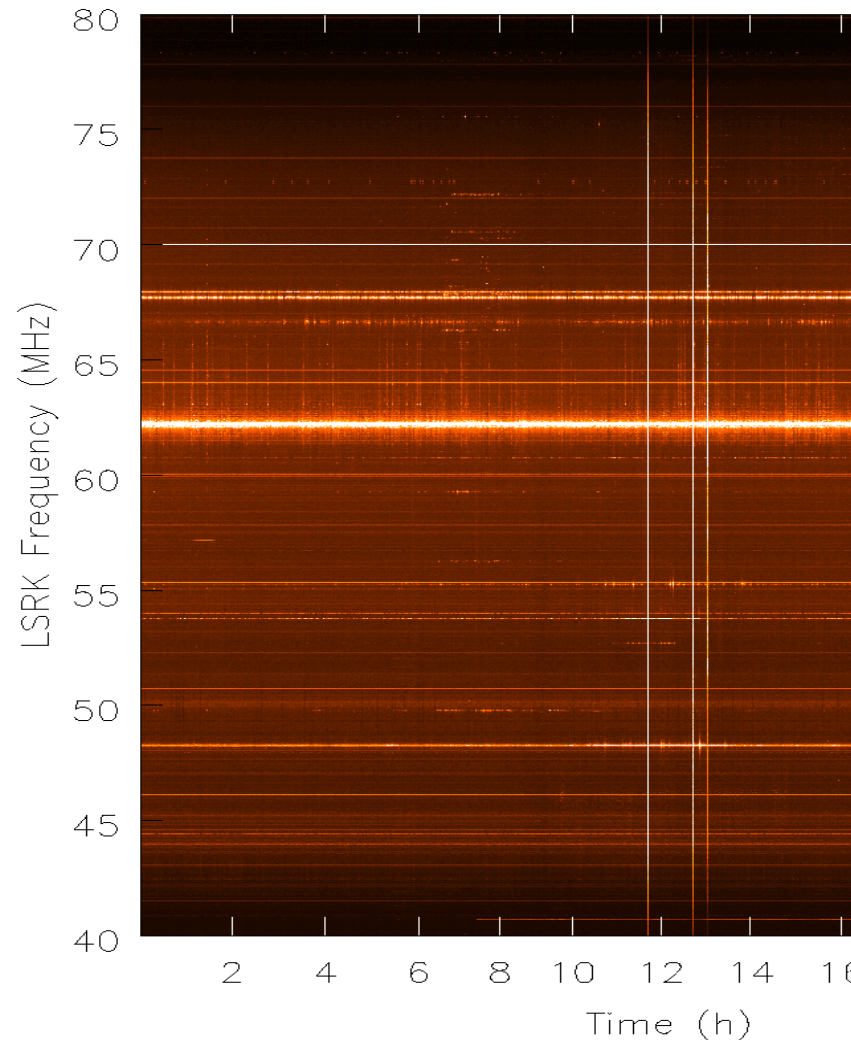


# Part II: LOPES thunderstorm results



# Data Analysis

- 3 selections of 2004 LOPES data
  - Cloudless
  - Nimbostratus
  - Thunderstorms

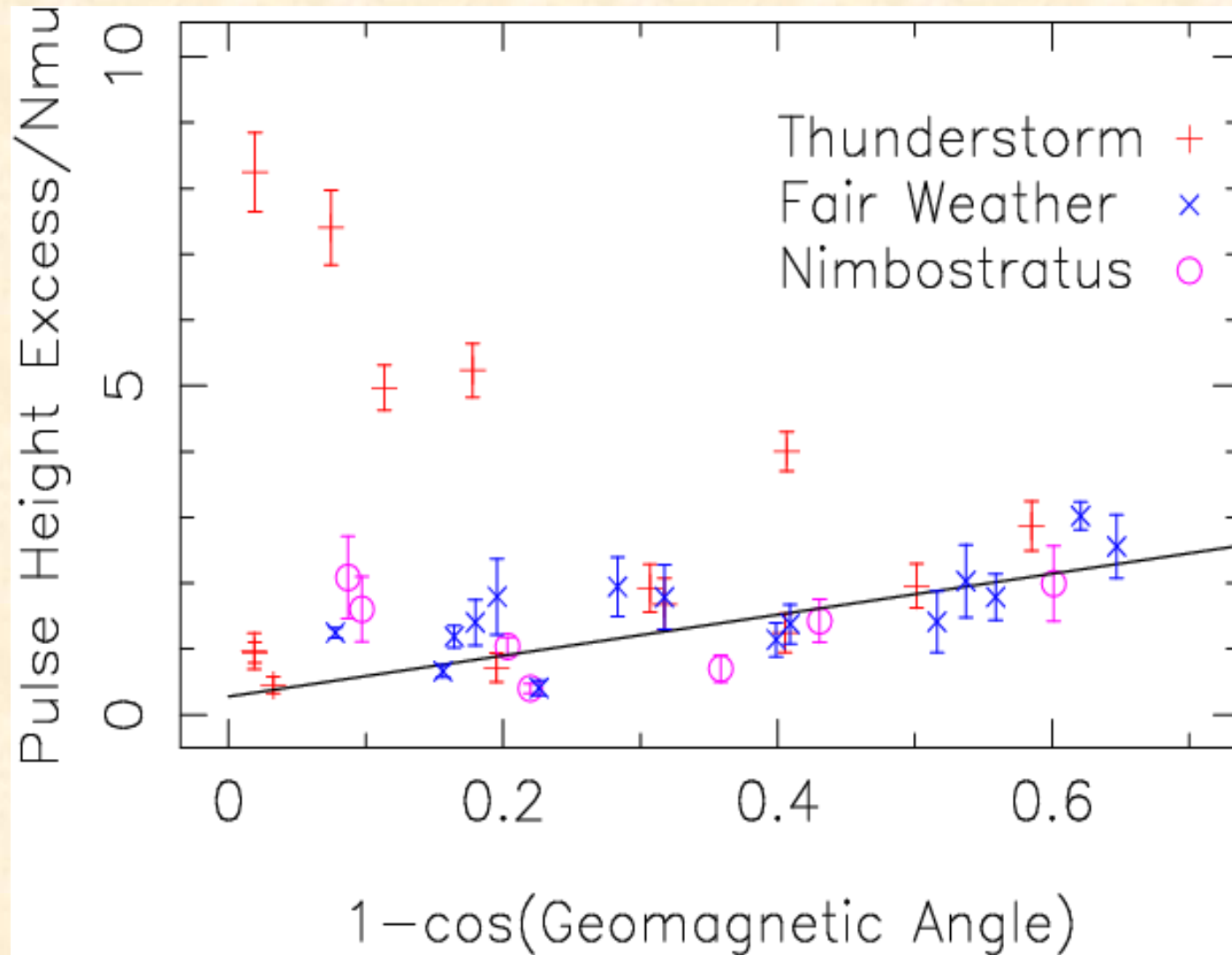


# Statistics

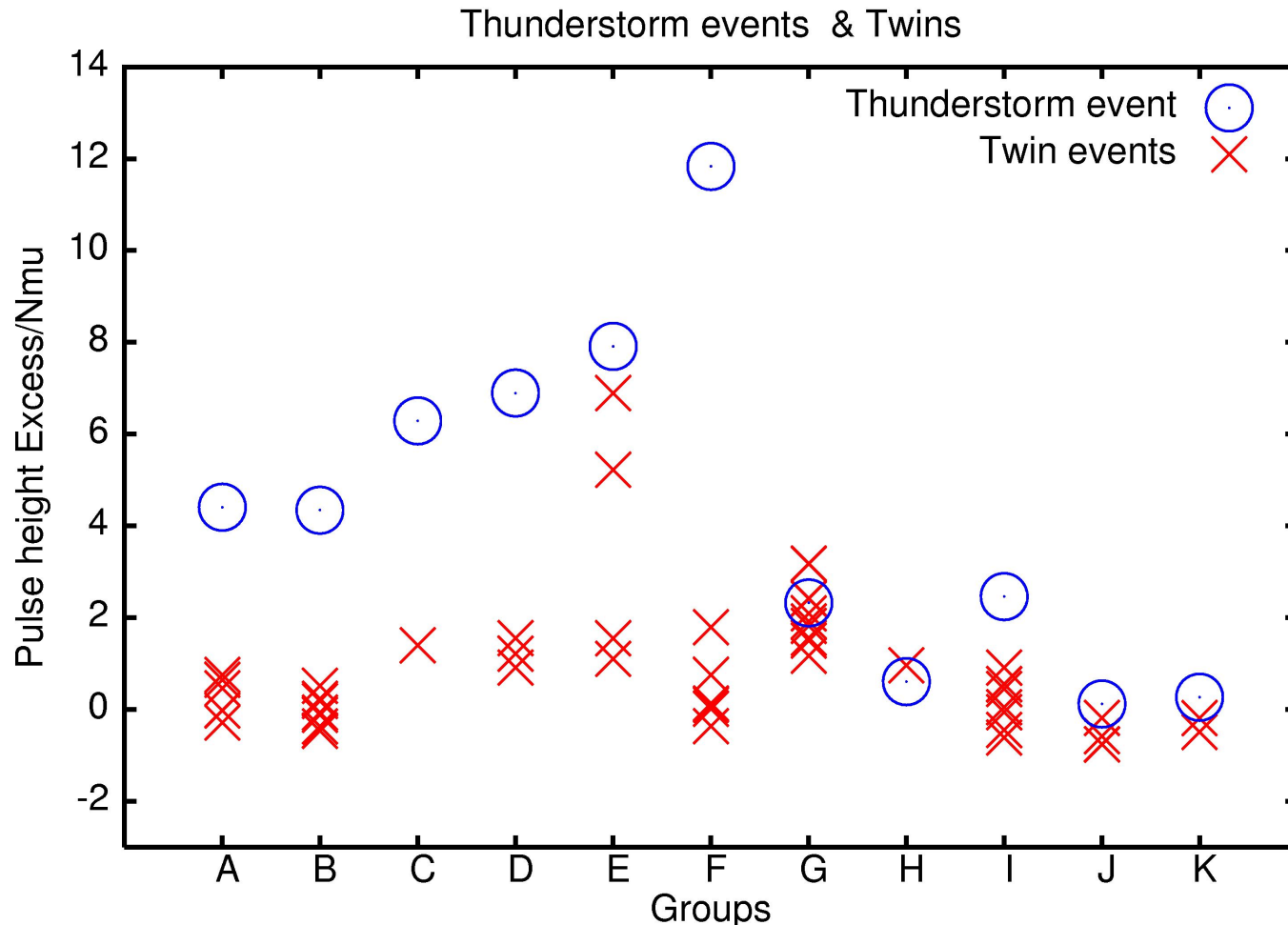
- No pre-selection on high energy  
→ low detection ratio !!
- Fair Weather: 9455 → 15 (0.16 %)
- Nimbostratus: 2659 → 7 (0.26%)
- Thunderstorm: 3510 → 14 (0.40%)



# Results



# Results: twin groups



# What causes amplification?

## Geosynchrotron + E-field

Emission is beamed forward

Emission is east-west  
polarised

Time width ~ 50 ns

## Ionization current

Emission is radiated in all  
directions

Emission is polarised in plane  
of current

Time width ~ 100-300 ns

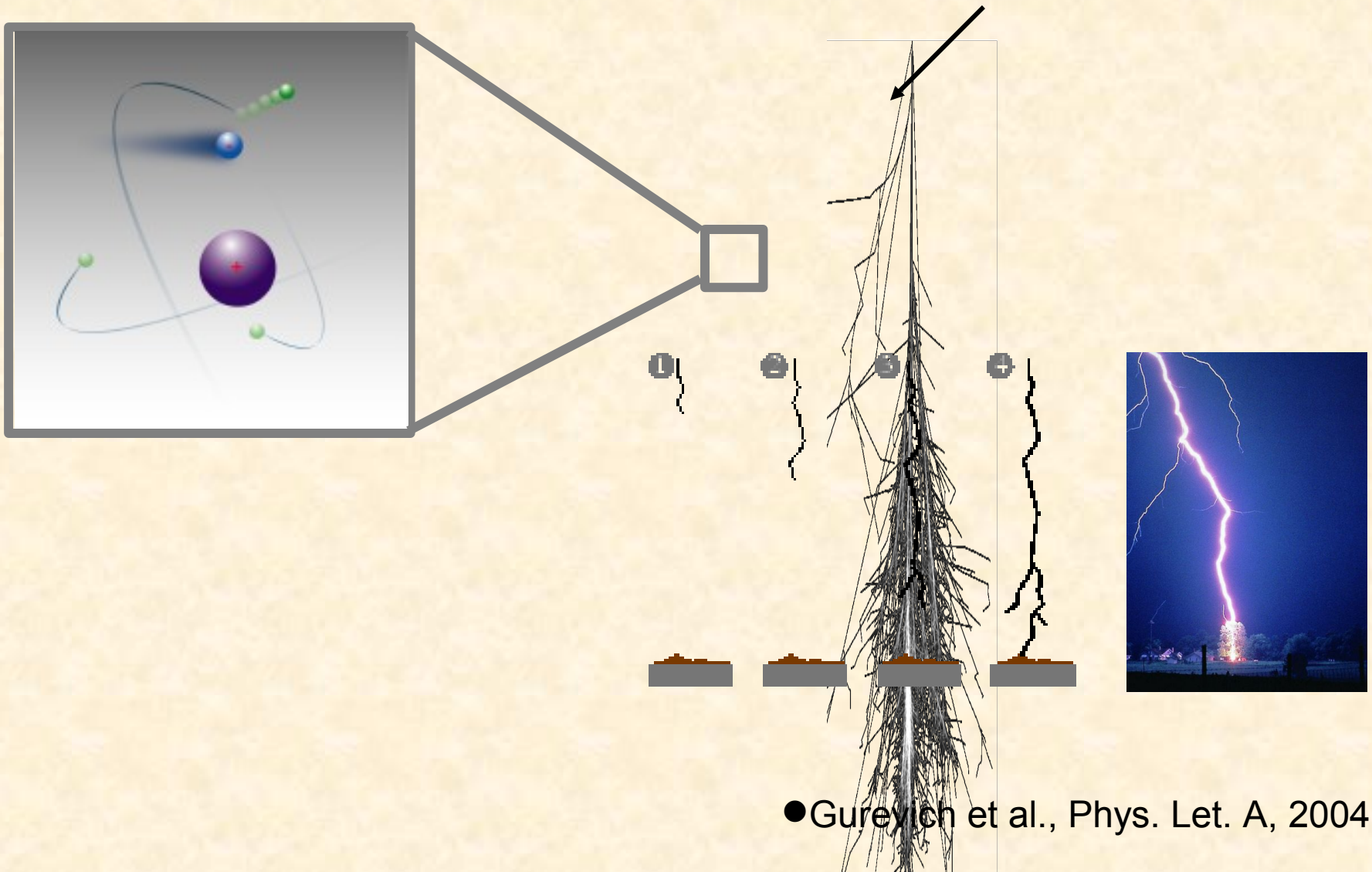
Observed width favours modified geosynchrotron emission

More information with LOFAR

# Part III: Other thunderstorm research



# Lightning triggering by air showers



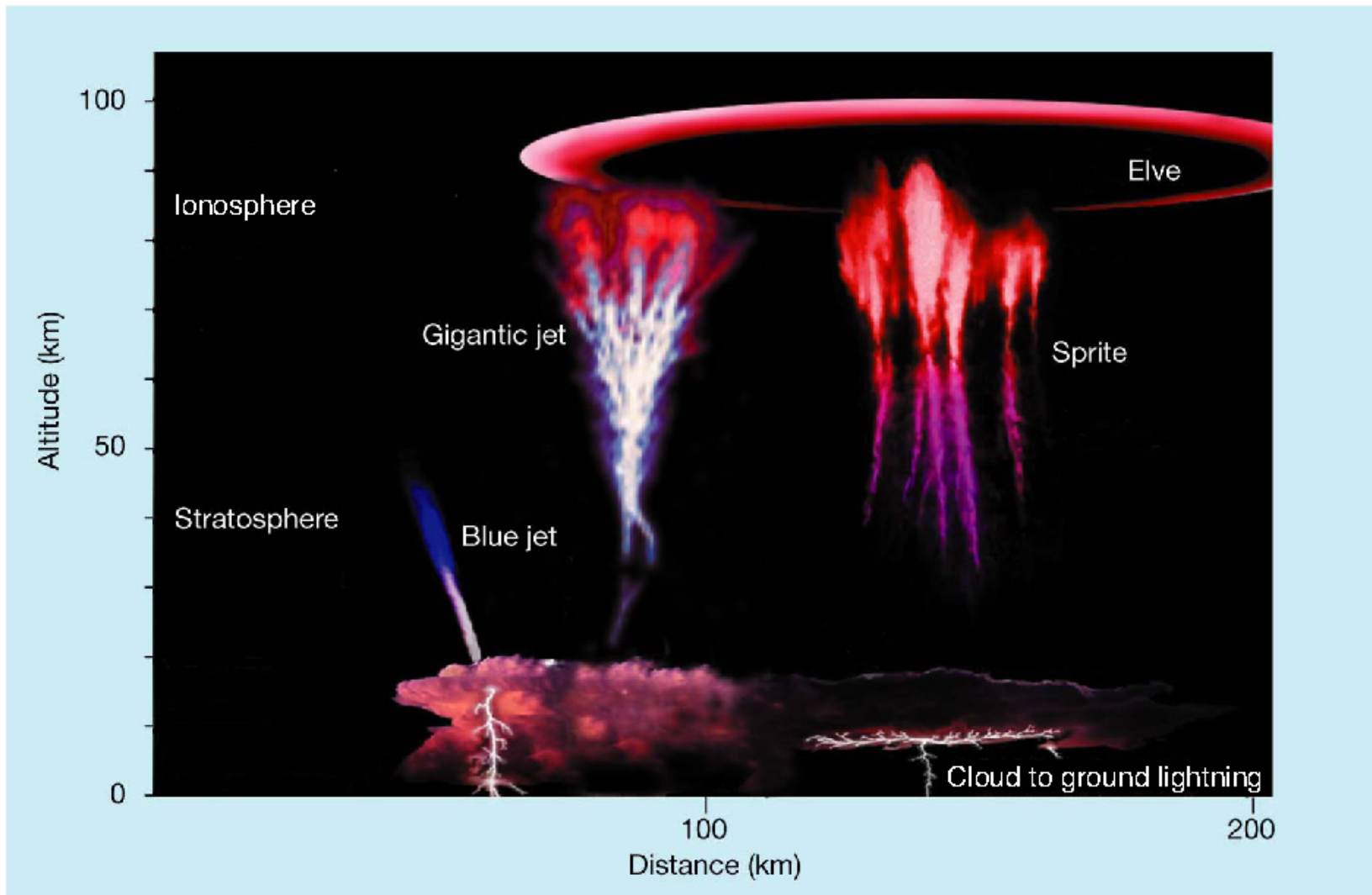
● Gureyich et al., Phys. Let. A, 2004

# Lightning imaging



Lars Baehren

# Lightning initiation / Sprites



[Pasko, Nature, 423, 927-929, 2003]



# Summary

- Atmospheric electric fields effect air shower emission ... but only strong fields
- This reduces duty cycle only slightly
- During thunderstorms a variety of interesting experiments can be performed

***Enjoy the weather !!***