

# LOFAR SCIENCE

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**LOFAR**

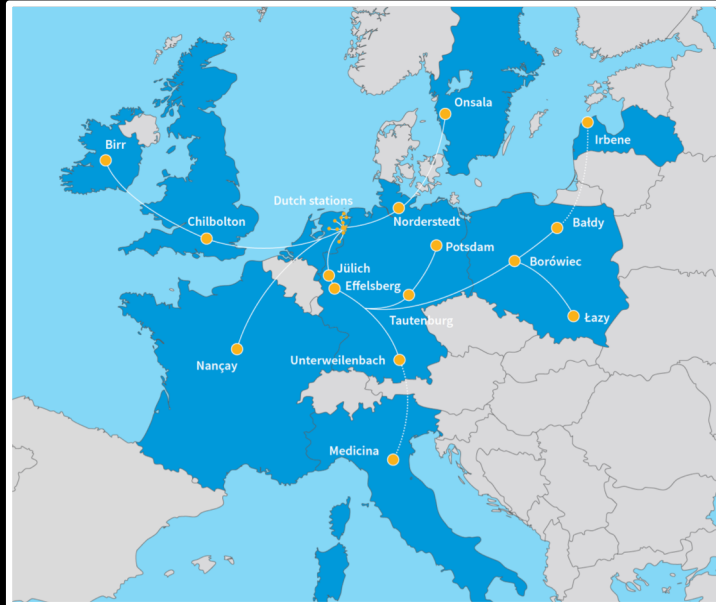
# OUTLINE

An aerial photograph of the LOFAR radio telescope array. The array consists of numerous small, dark, rectangular antenna elements arranged in a grid pattern across a large, flat, green field. A winding river flows through the landscape, curving around the central part of the array. The background shows a vast, flat landscape with more fields and a distant treeline under a clear sky.

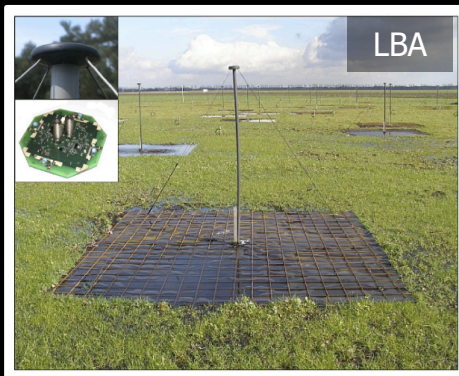
## ➤ The Low Frequency Array

- Key facts
- The science drivers of LOFAR
- Recent science highlights
- Science output

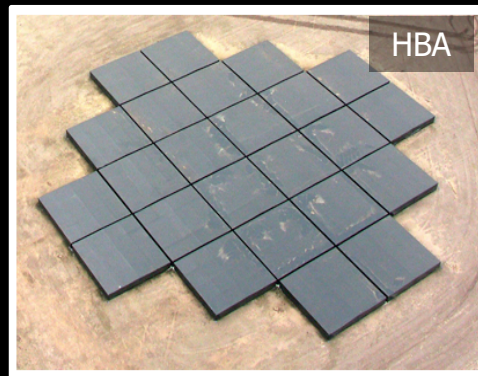
# LOFAR: KEY FACTS



- Array of 51 dipole antenna stations **distributed across EU**
- **10-250 MHz**
- Low band antenna (LBA; 10-90 MHz); High Band Antenna (HBA; 110-250 MHz)
- **Several observing modes** (imaging, BF, BF+IM, TBB)
- **Responsive telescope**
- **96 MHz bandwidth** (multi-beam option)
- **Big data: important technological pathfinder for next-gen facilities and data intensive astronomy**



LBA



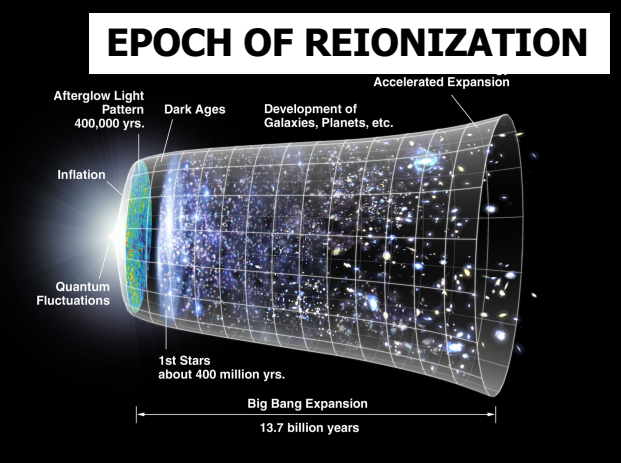
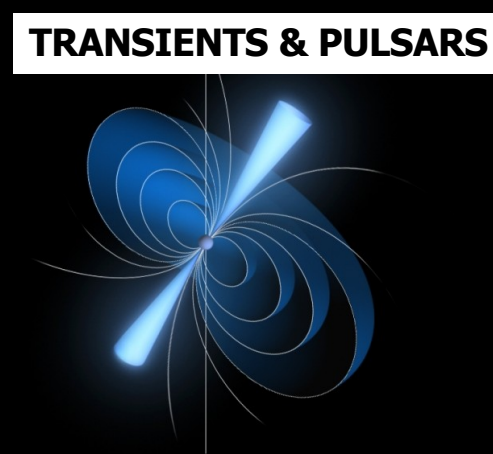
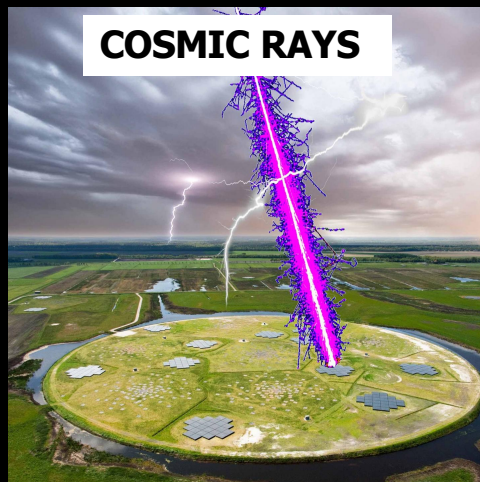
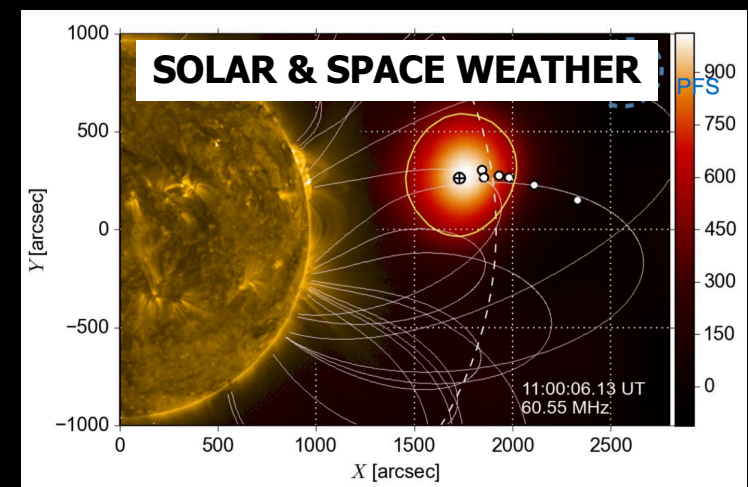
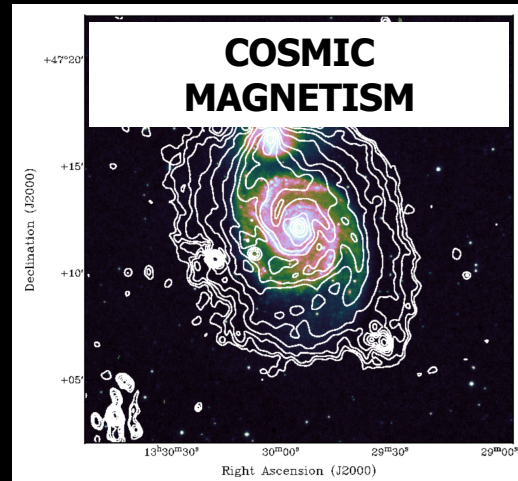
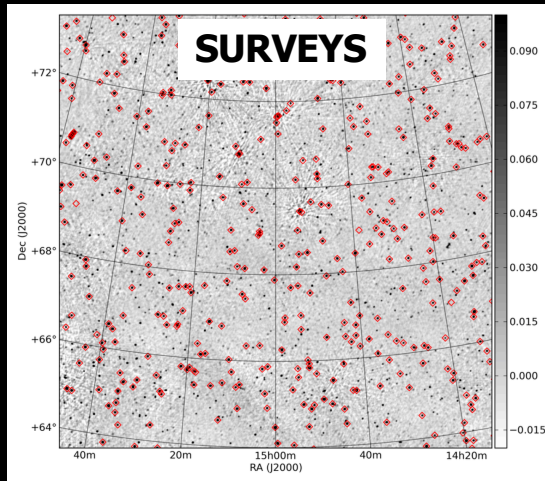
HBA



station



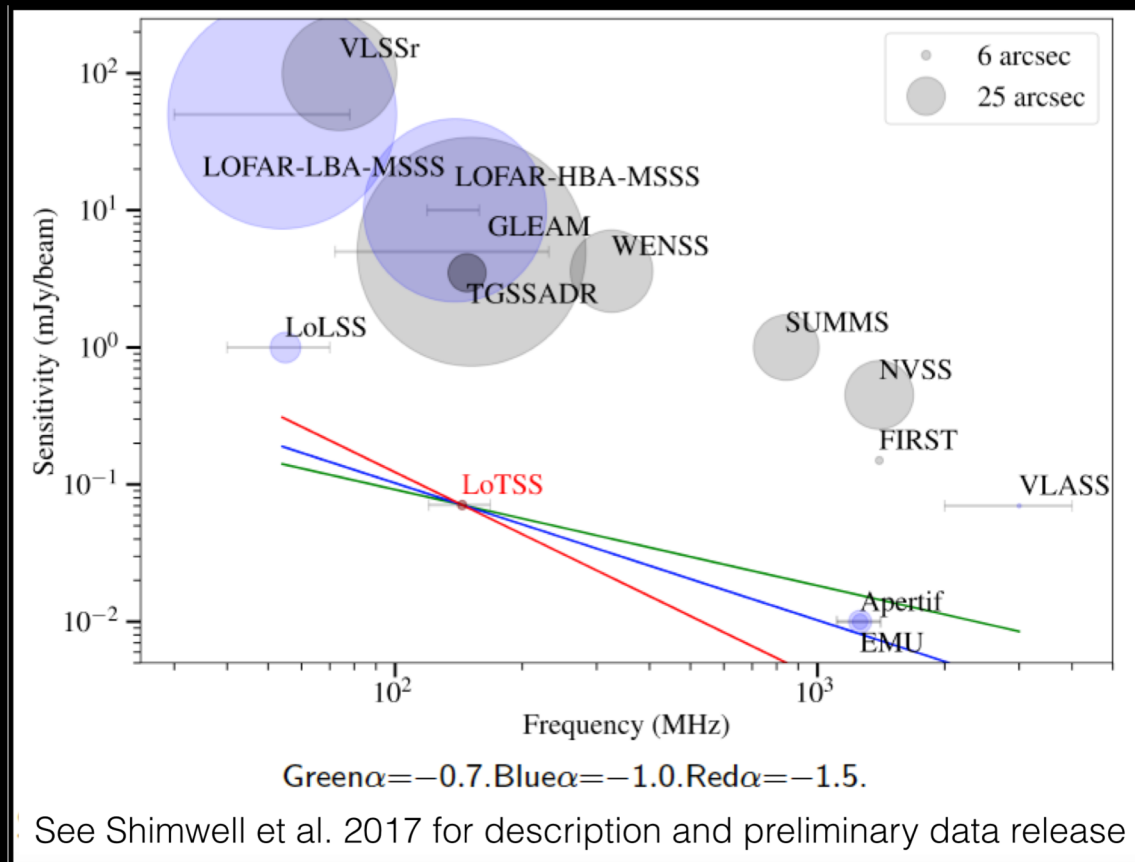
multi-beam



**International membership from countries all over the World**

**Contribute development and commissioning resources**

# SURVEYS: THE LOFAR TWO-METRE SKY SURVEY (LoTSS)



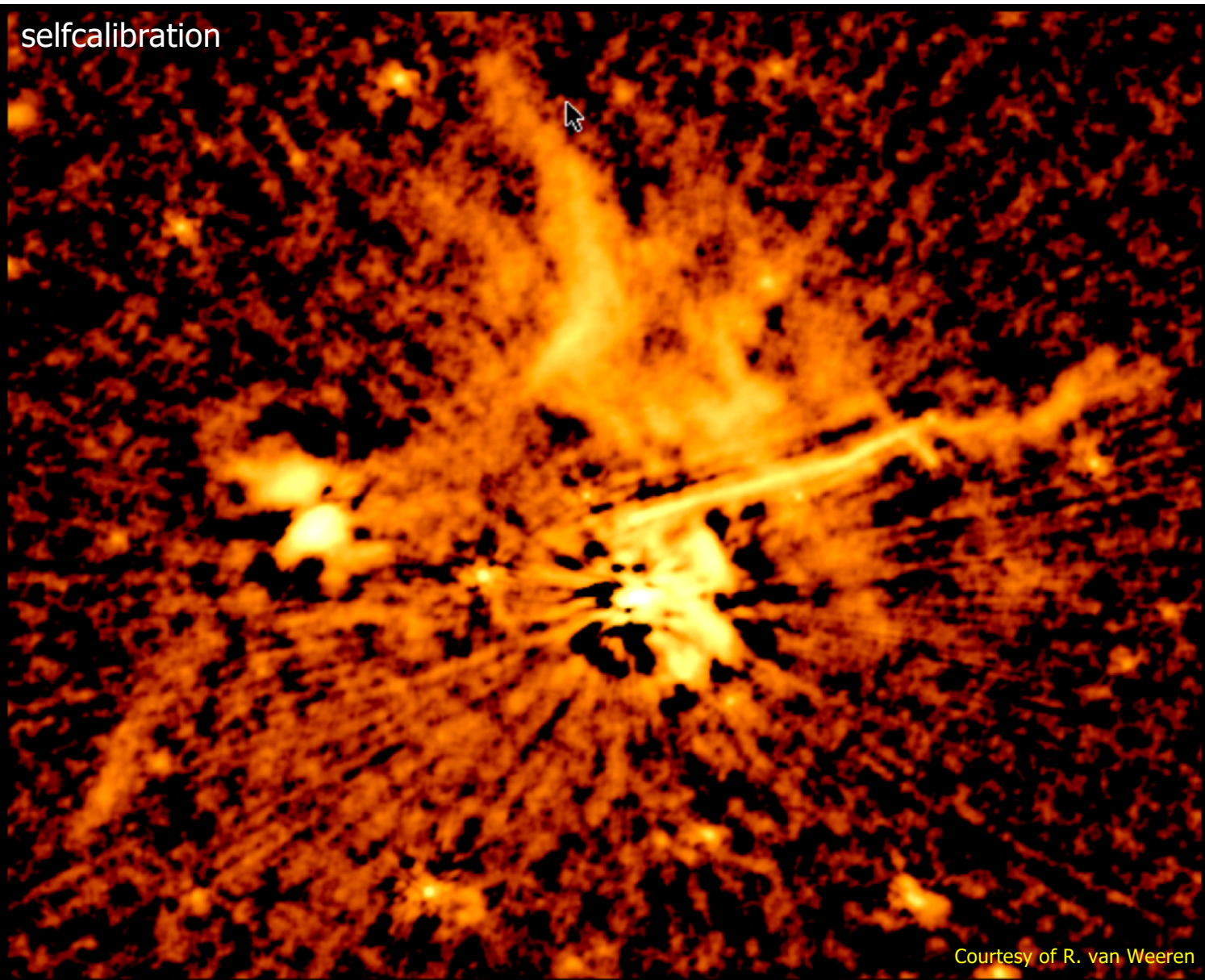
See Shimwell et al. 2017 for description and preliminary data release

- **High-redshift radio galaxies:** formation and evolution of massive galaxies, rich clusters and massive black holes
- **Galaxy clusters:** origin and evolution of magnetic fields and relativistic electrons
- Determining the **cosmic star-formation** history of the Universe
- **serendipitous** discoveries
- How? Produce **high fidelity images** of the entire Northern sky with a resolution of 5" and sensitivity of 100 μJy/beam at most declinations.
- **Will not be surpassed as a northern sky survey for the foreseeable future**

# SURVEYS

selfcalibration

more in lecture L7 by  
Shimwell, van  
Weeren, Mevius  
(Wednesday)



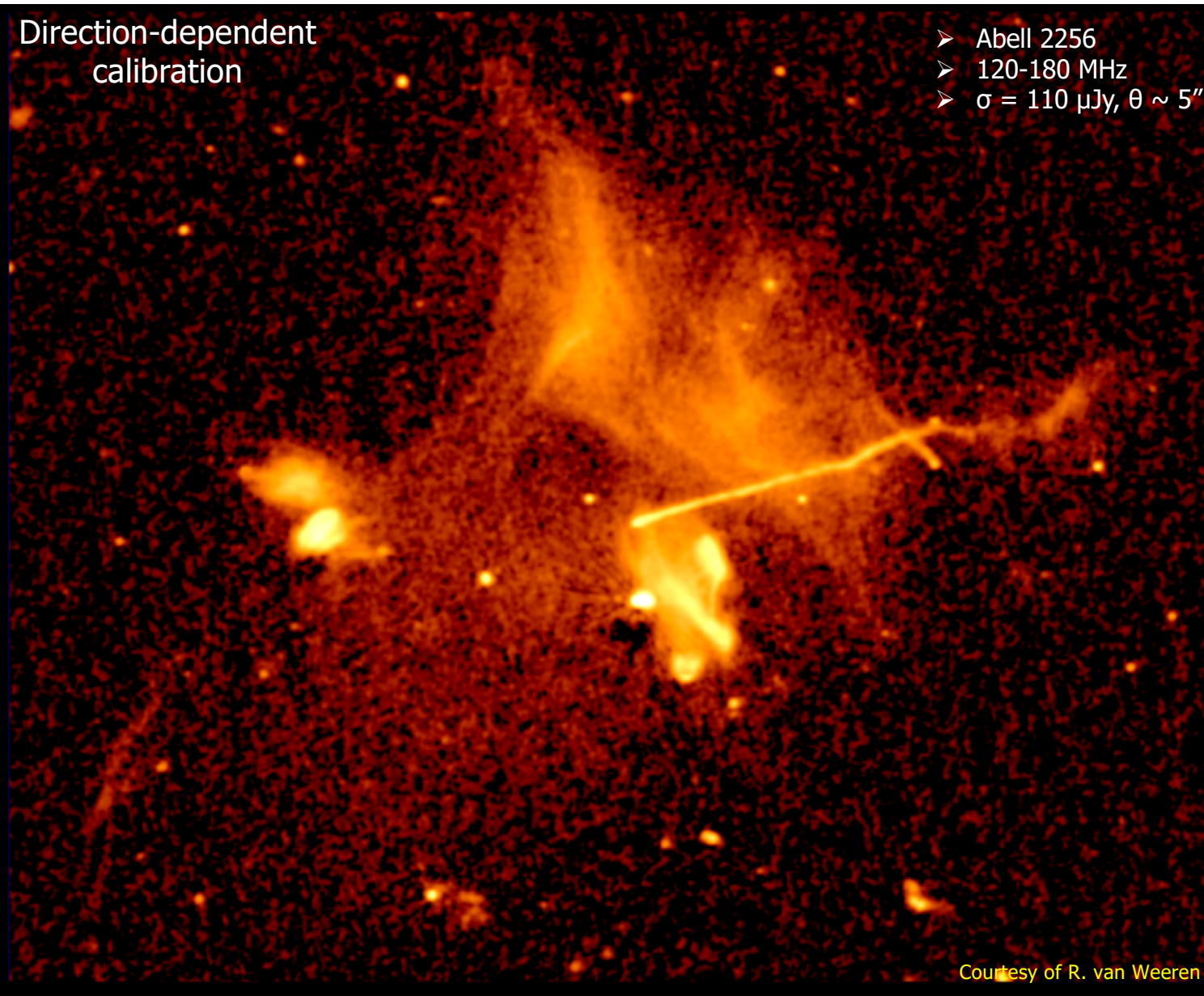
Courtesy of R. van Weeren

# SURVEYS

Direction-dependent  
calibration

- Abell 2256
- 120-180 MHz
- $\sigma = 110 \mu\text{Jy}$ ,  $\theta \sim 5''$

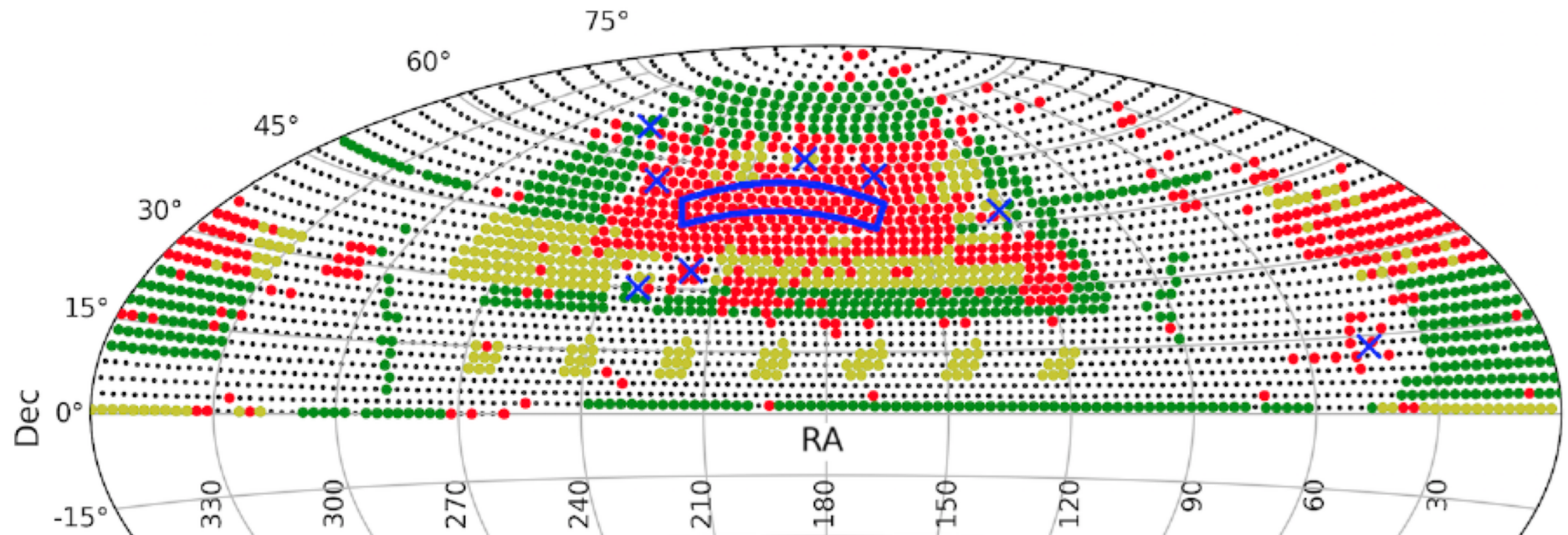
You will also do it  
by Friday!



Courtesy of R. van Weeren

# SURVEYS

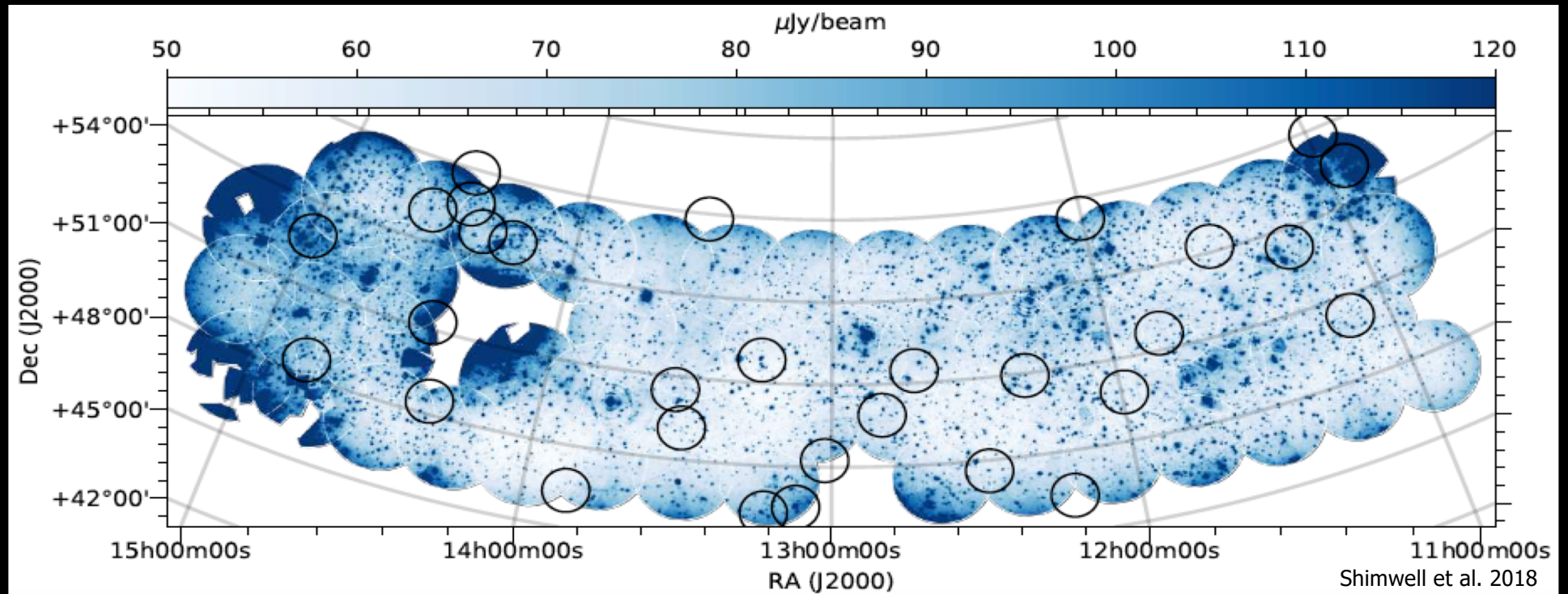
Since Cycle 2: 30% of the Northern Sky observed -> 4000 observing hours.



**Red** observed, **yellow** observed in next 6 months, **green** observed in next 2 years.



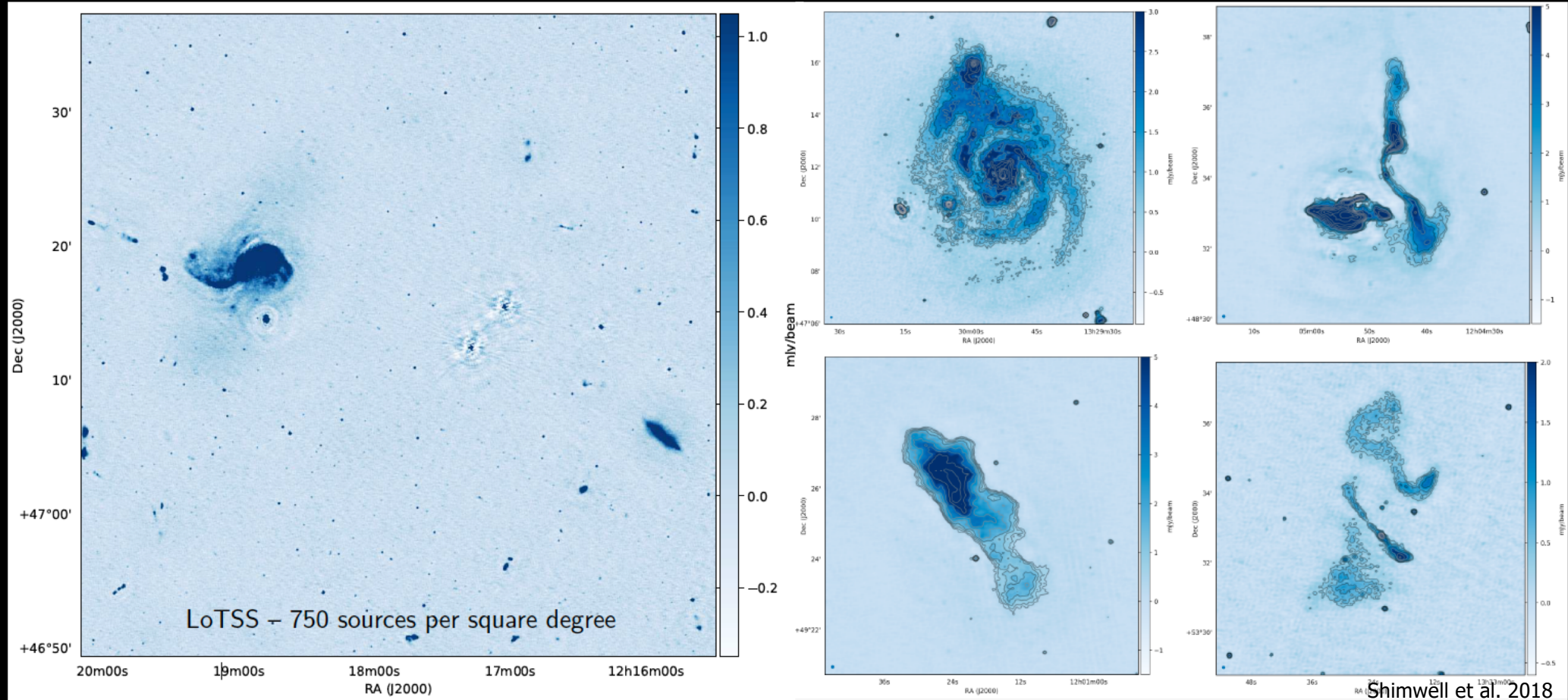
# SURVEYS



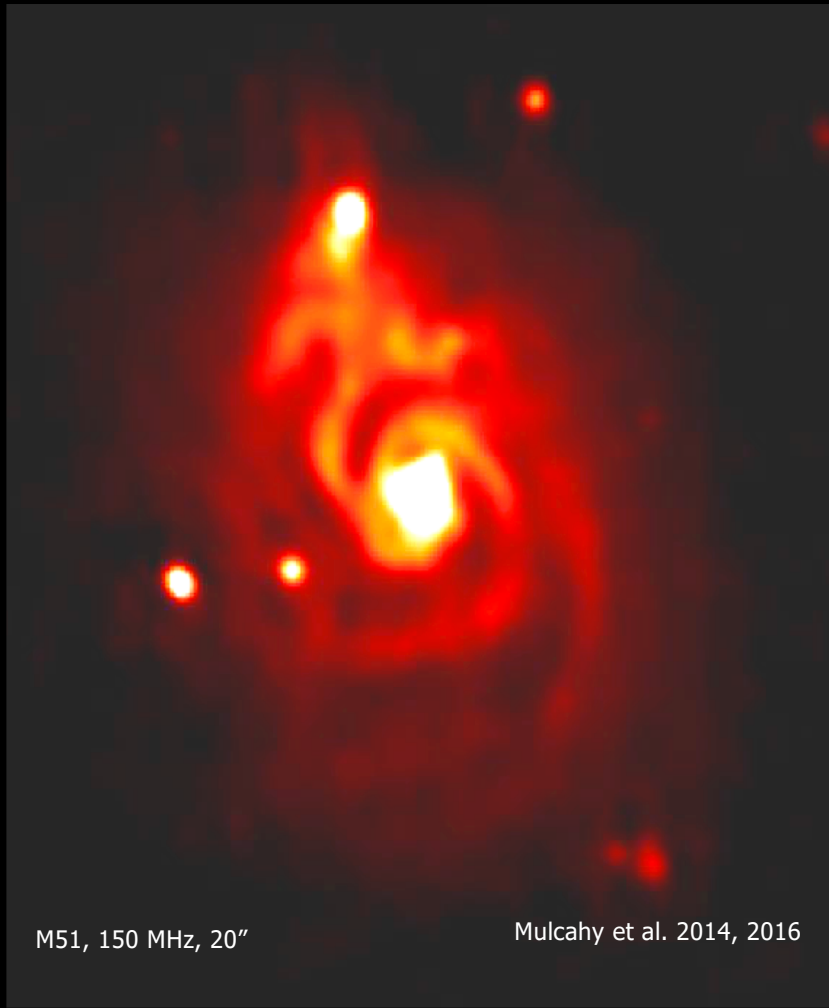
First full quality **data release**: 6" resolution images with a sensitivity of  $71\mu\text{Jy}/\text{beam}$  that covers 424 square degrees in the HETDEX Spring Field region.

The catalogue contains 325,694 radio sources (Shimwell et al 2018, in press), 225,457 optical identifications (Williams+submitted) and 158,284 photometric redshift estimates (Duncan+submitted)

# SURVEYS



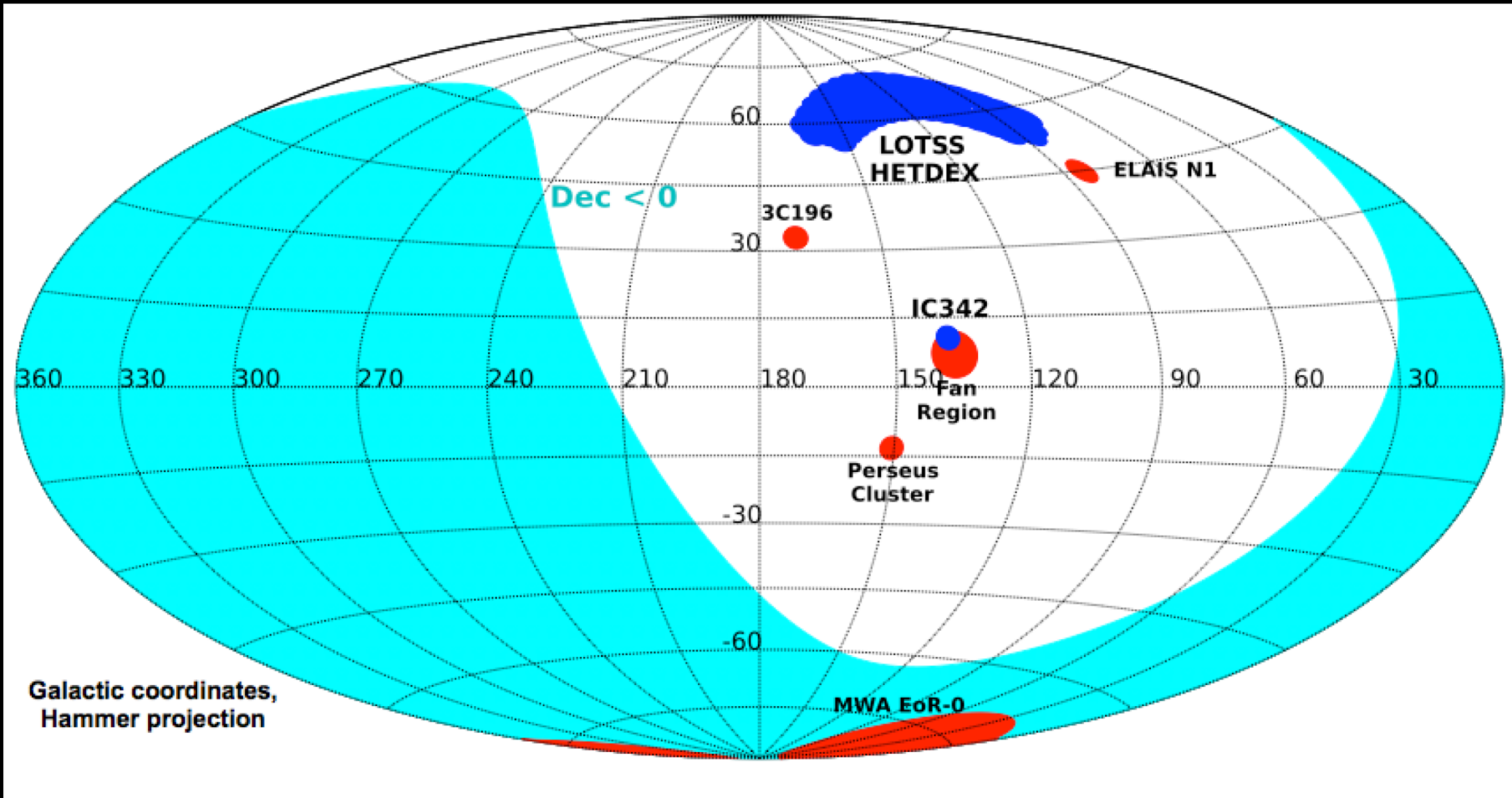
# COSMIC MAGNETISM



- Magnetic fields are ubiquitous in the Universe – **their origin, evolution and structure still remain open fundamental problems**
- Due to its wide bandwidth at low frequency, LOFAR can:
  - Provide info on **spectral properties of the synchrotron radiation**
  - **Trace magnetic fields** far away from CR acceleration sites
  - Trace weak magnetic fields through **Faraday rotation studies** (RM synthesis)

See lecture L14 by M. Iacobelli on Friday

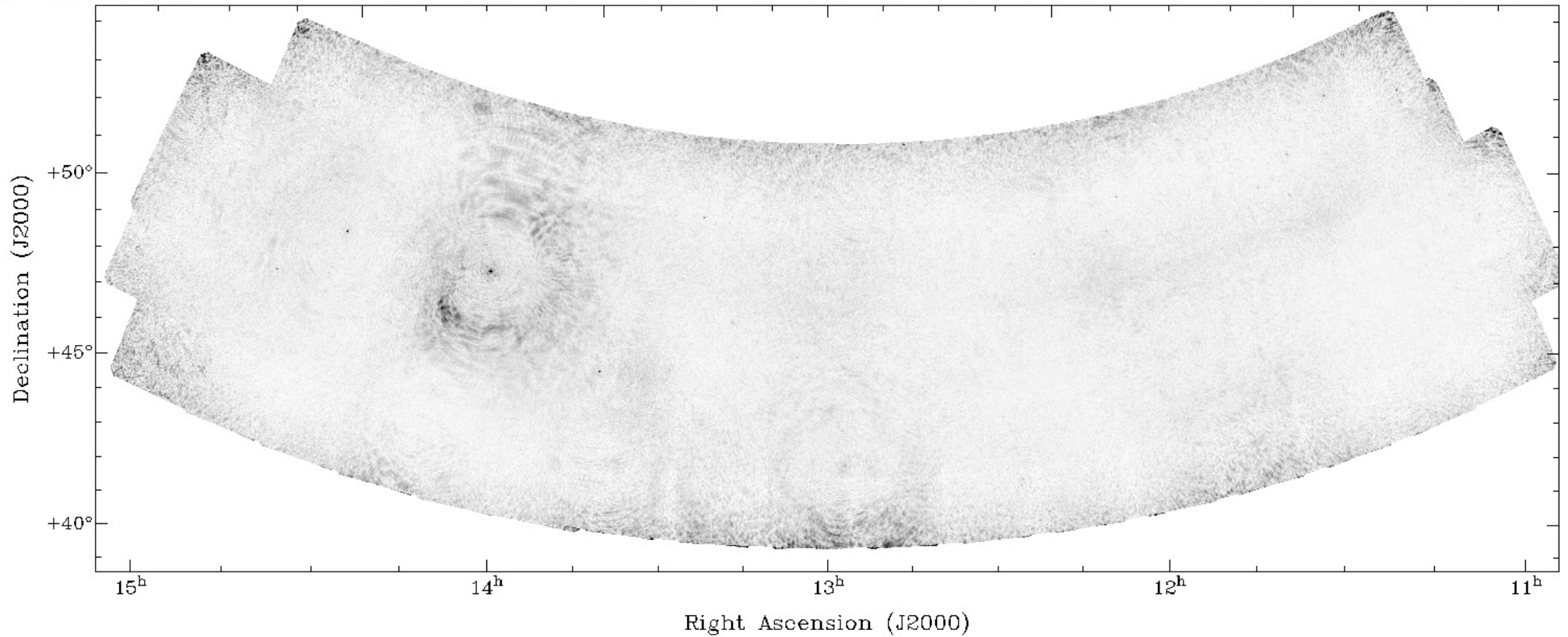
# COSMIC MAGNETISM



Courtesy of C. Van Eck

# COSMIC MAGNETISM

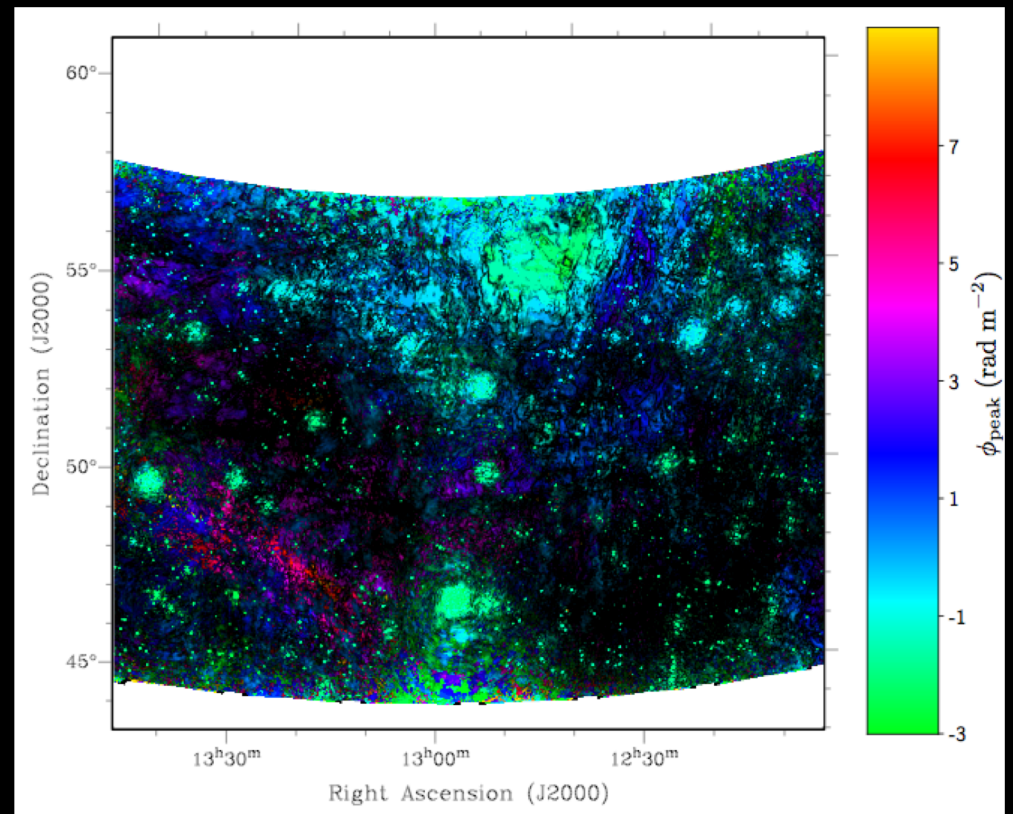
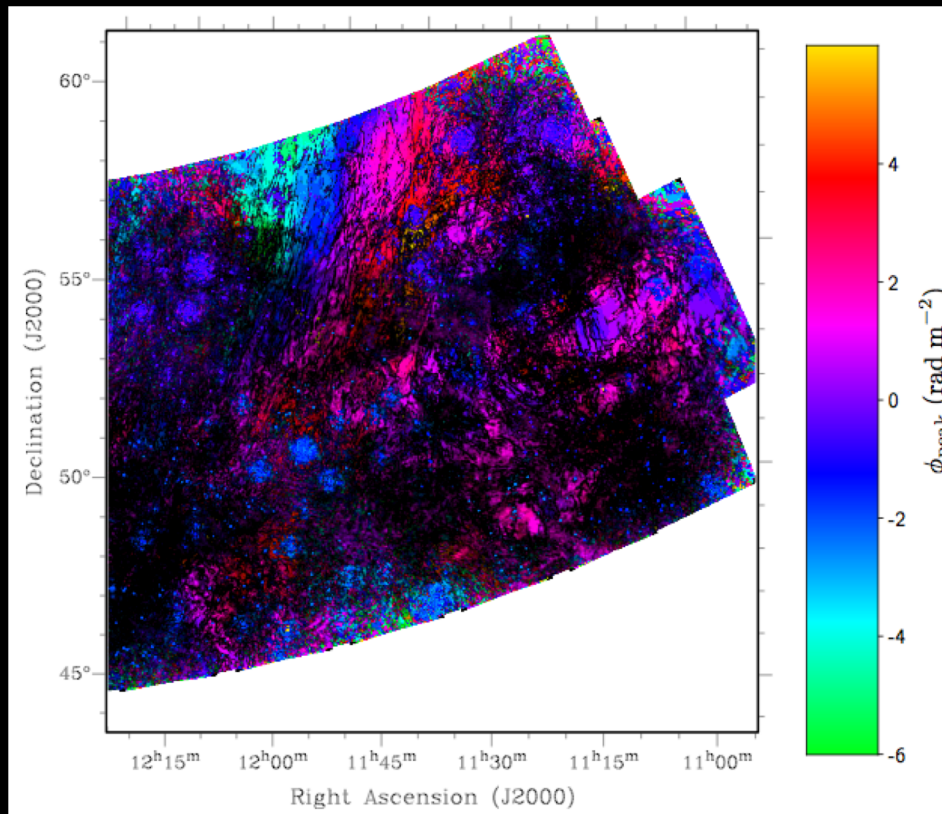
Phi:  $-1.000000e+01$



60 LOTS observations – 570 deg<sup>2</sup>

Courtesy of C. Van Eck

# COSMIC MAGNETISM

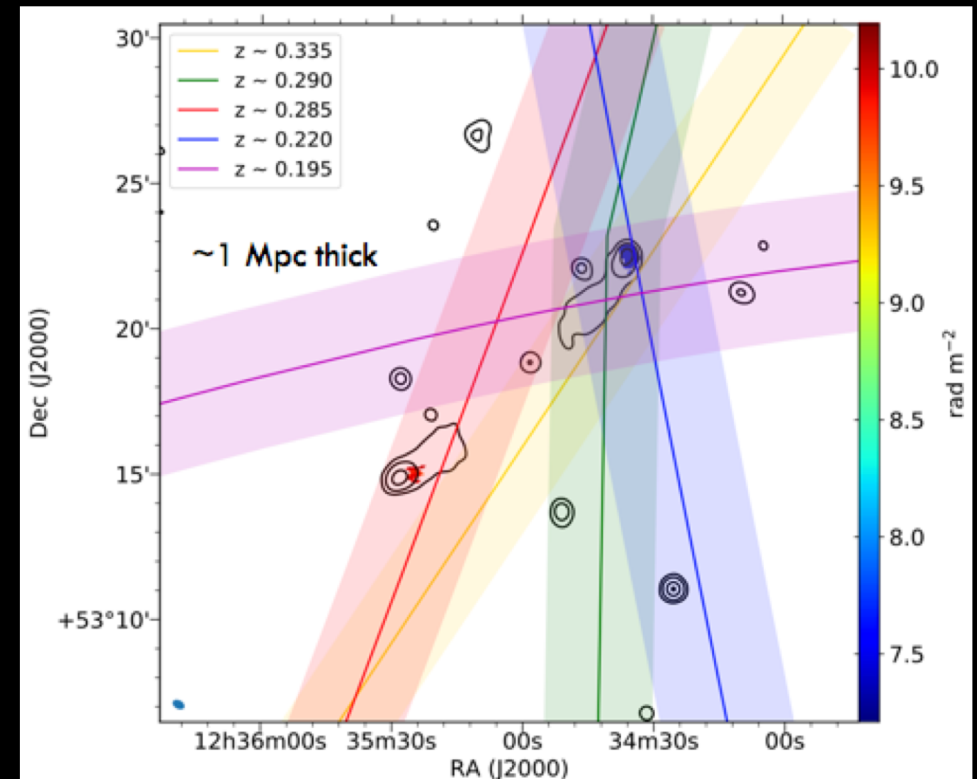
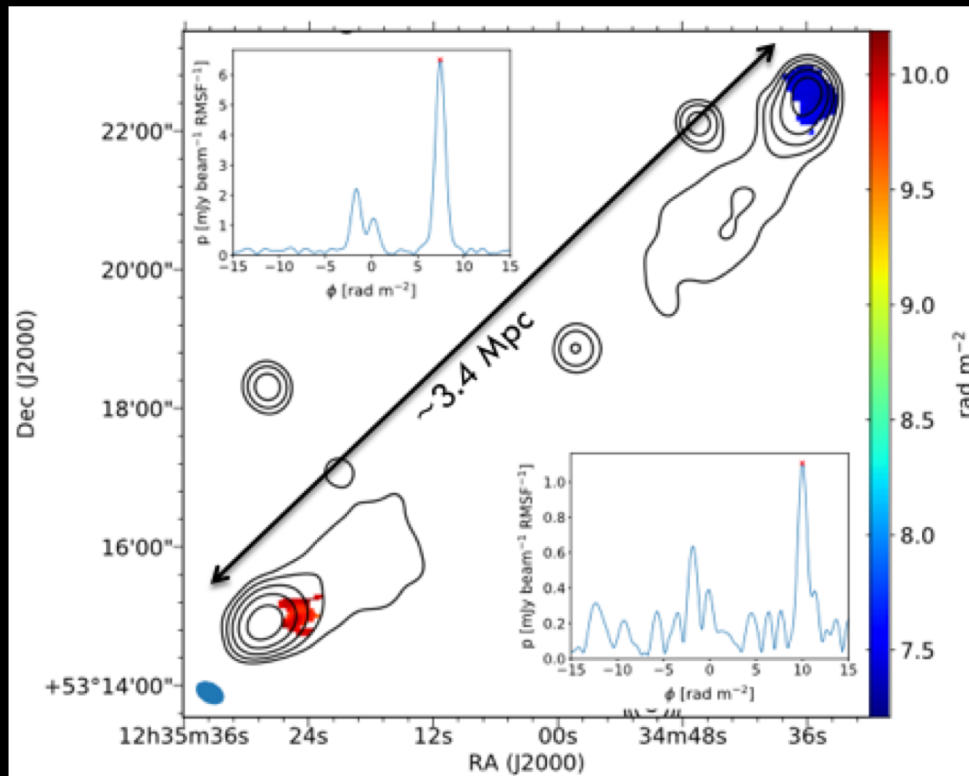


Color -> Faraday Depth of the peak  
Brightness -> Polarized intensity

Courtesy of C. Van Eck

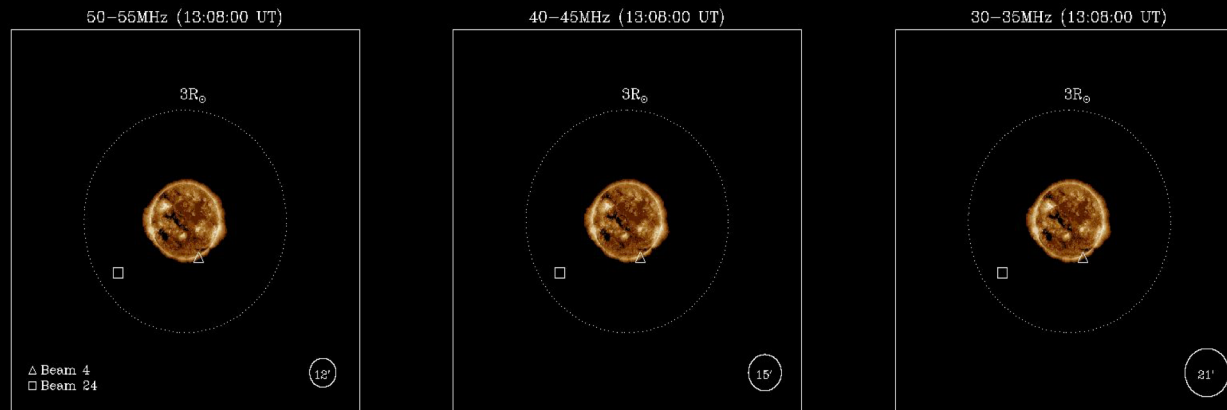
# COSMIC MAGNETISM

Probing magnetic fields in **intergalactic filaments**: excess of  $2.5 \text{ rad/m}^2$  on  $3.4 \text{ Mpc}$  scales in Giant Radio Galaxy

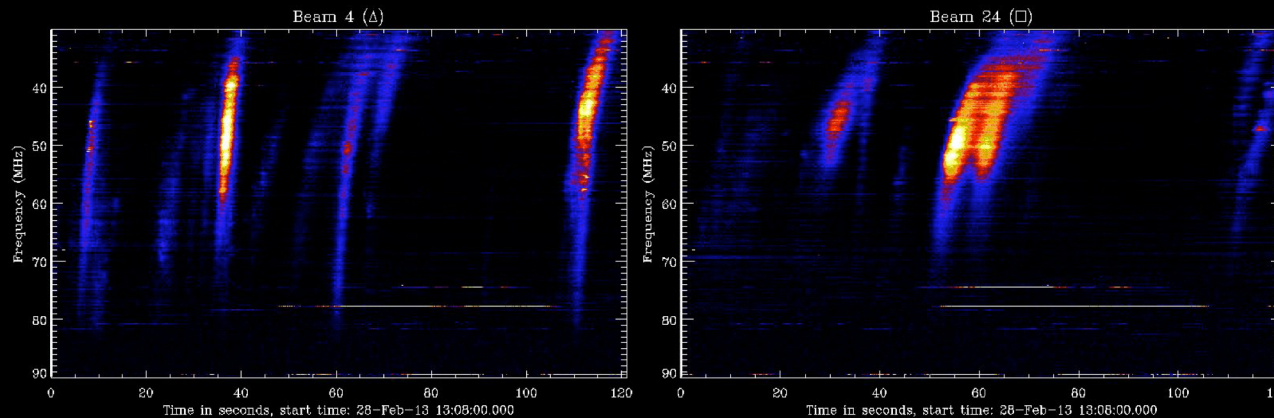


Courtesy of S. O' Sullivan

# SOLAR PHYSICS AND SPACE WEATHER



- Thermal radiation of the quiet Sun interspersed with intense **radio bursts (flares and CMEs)**
- LOFAR: dynamic spectroscopic radio imager
- Several type I-II-III radio bursts detected since 2011



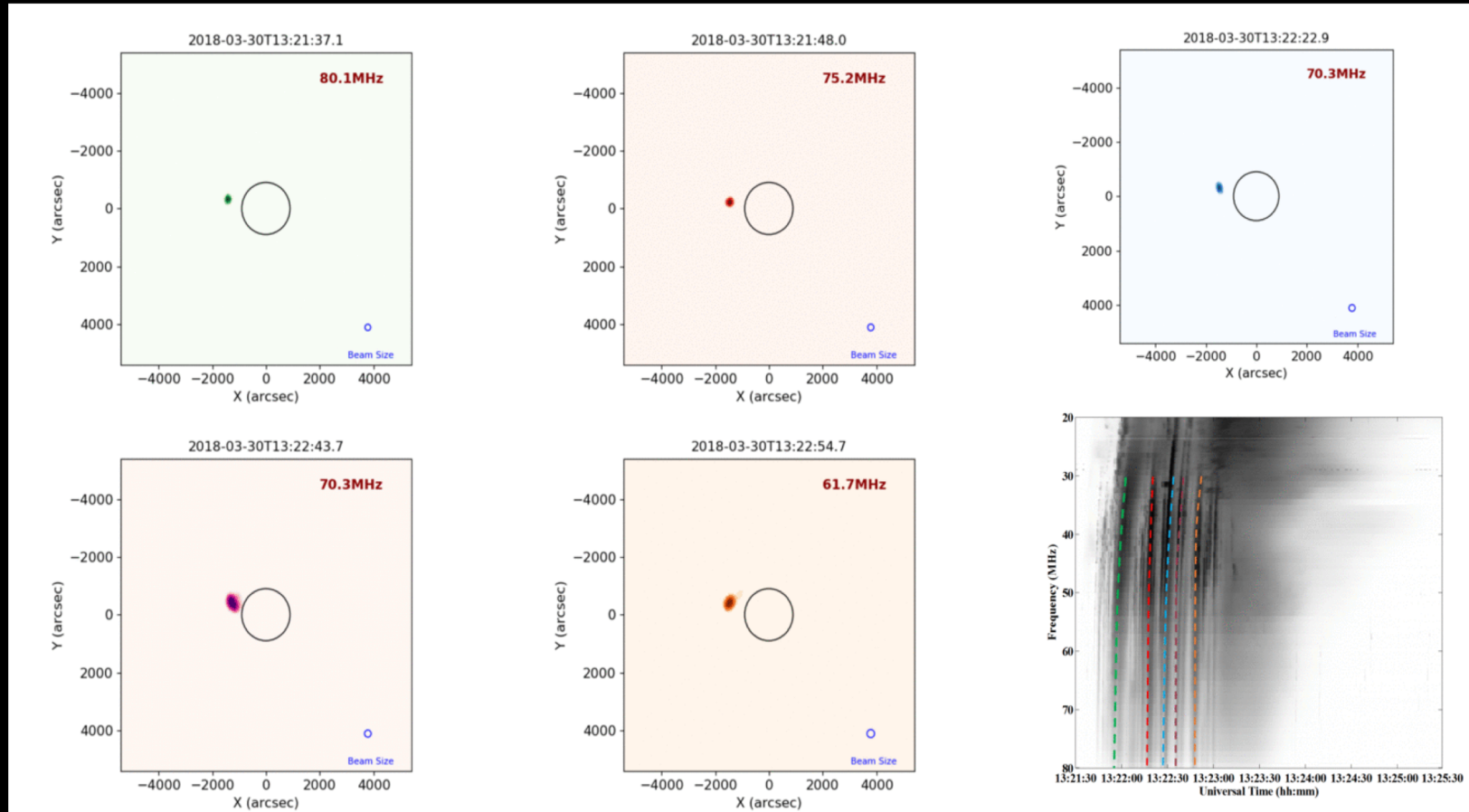
*See lecture L15 by Zucca on Friday*

LOFAR tied array imaging of the Sun: electrons escaping the solar corona

Courtesy Morosan, Gallagher, Zucca, Fallows, Carley and the Solar KSP

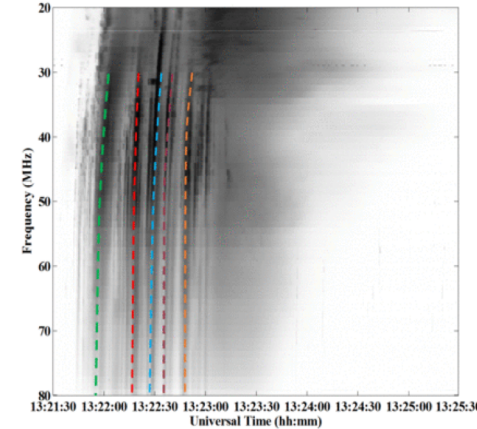


# SOLAR PHYSICS AND SPACE WEATHER



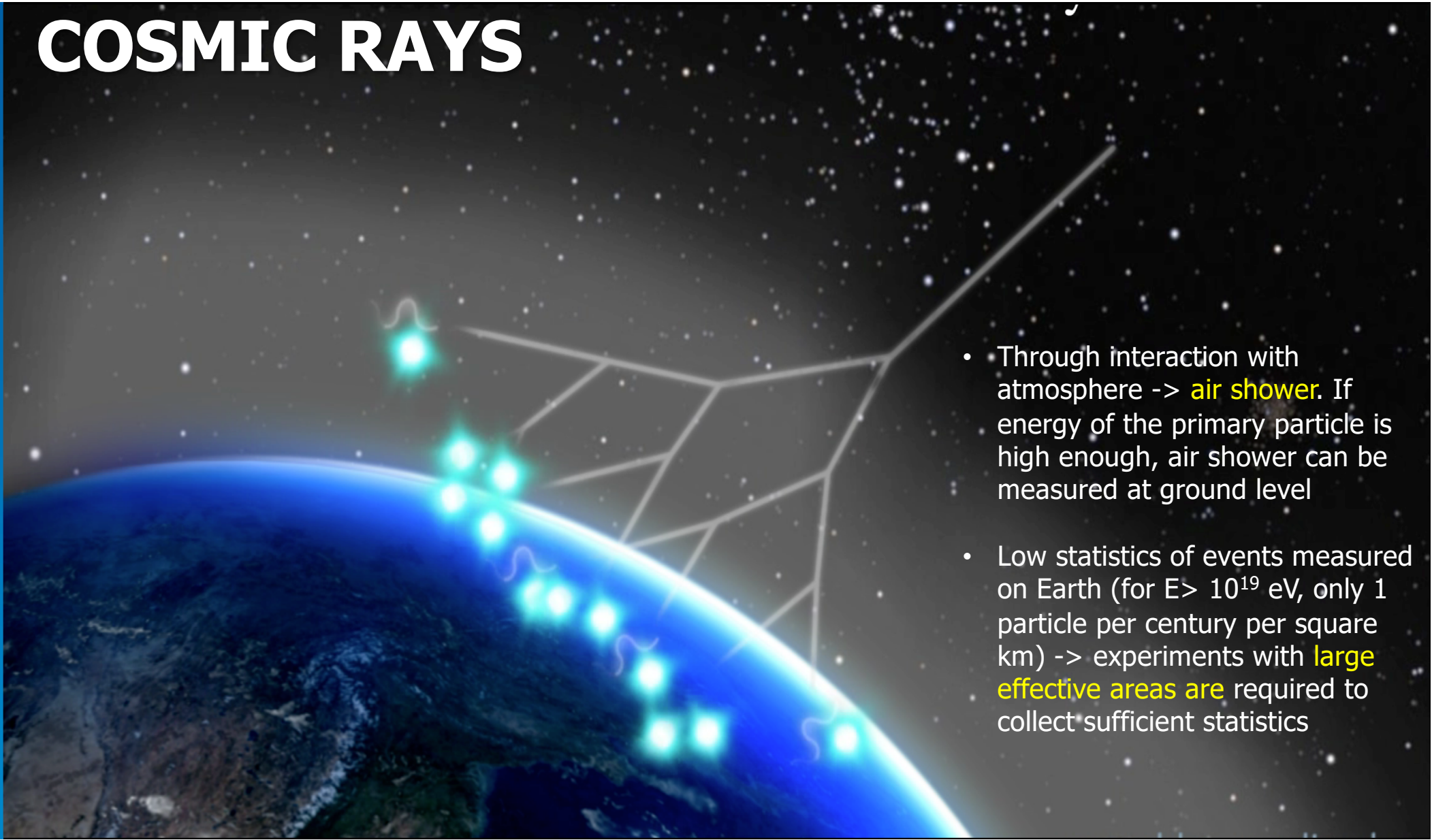
Interferometric images of type III bursts

*See lecture L15 by Zucca on Friday*



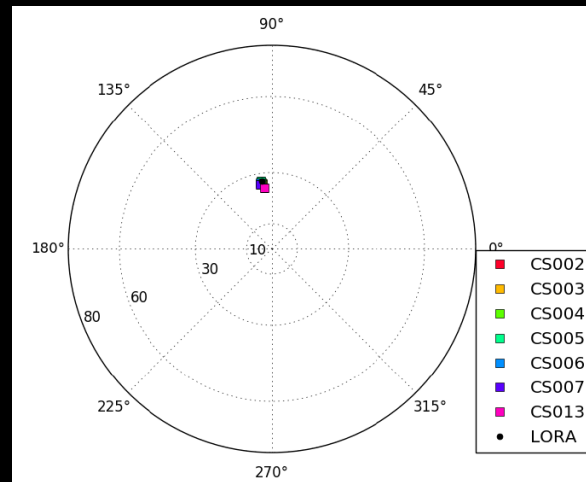
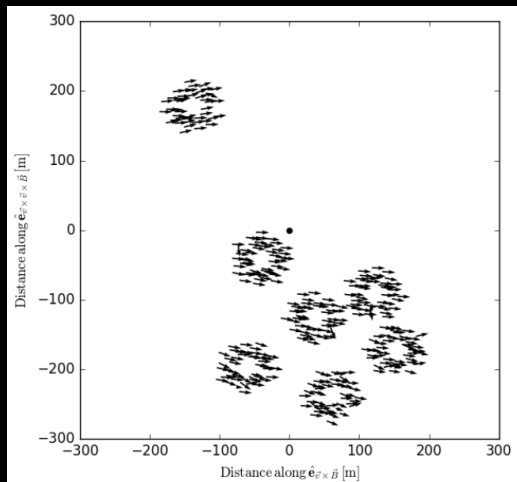
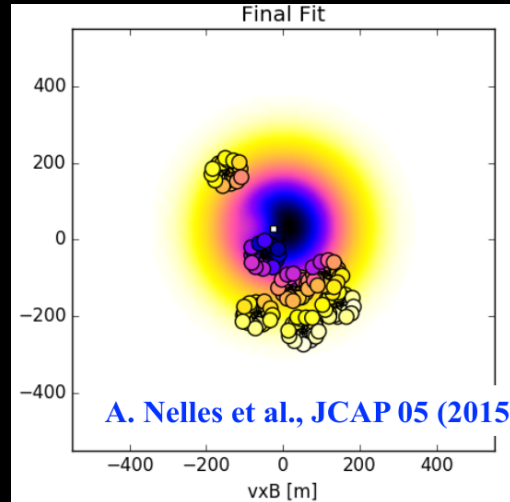
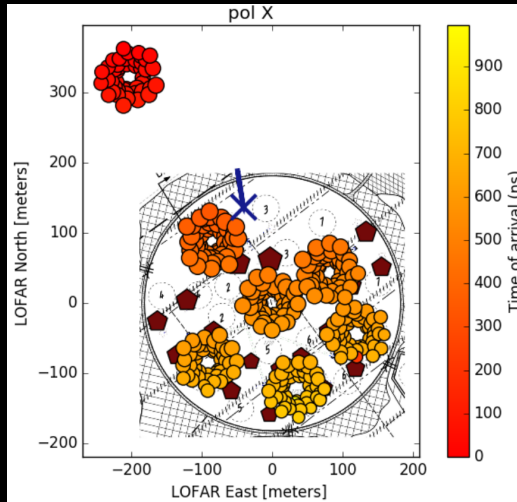
Kumari & Zucca in prep.

# COSMIC RAYS



- Through interaction with atmosphere -> **air shower**. If energy of the primary particle is high enough, air shower can be measured at ground level
- Low statistics of events measured on Earth (for  $E > 10^{19}$  eV, only 1 particle per century per square km) -> experiments with **large effective areas** are required to collect sufficient statistics

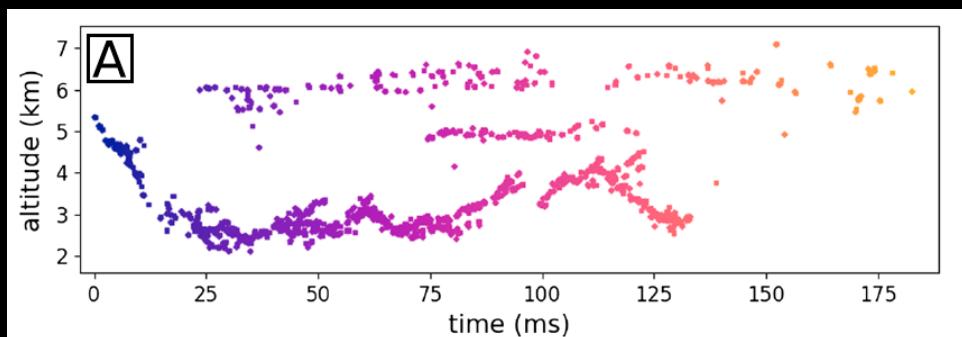
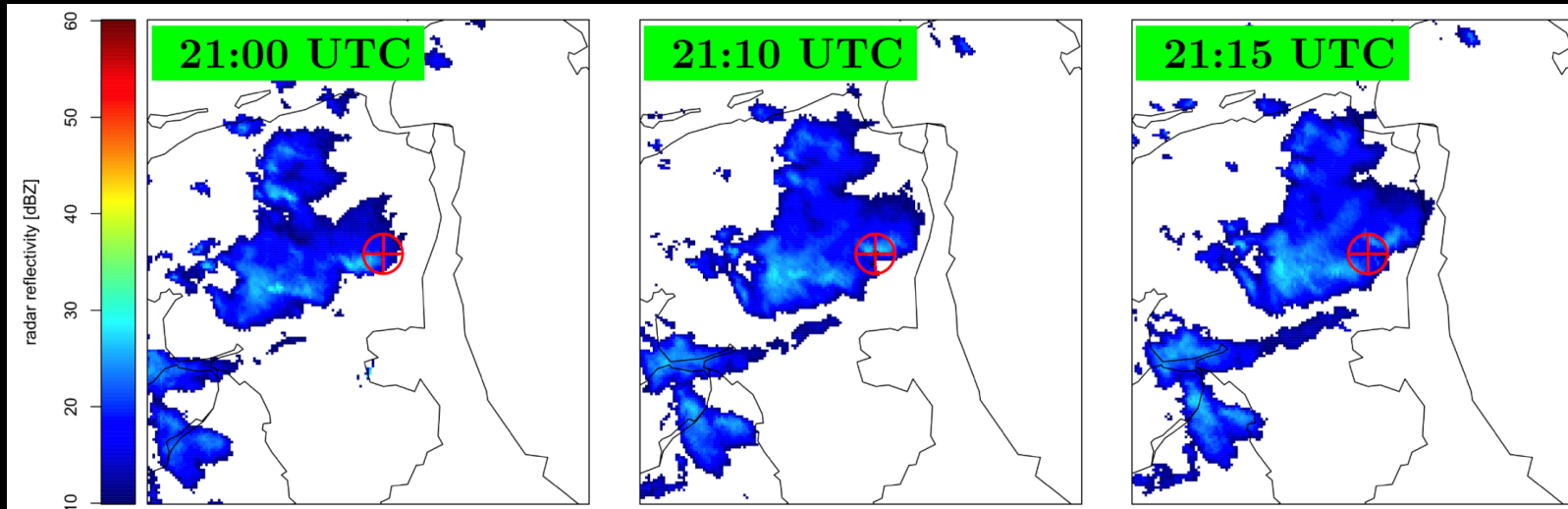
# COSMIC RAYS



- Cosmic-ray events reconstruction between  $10^{16} - 10^{18}$  eV -> acceleration and propagation mechanisms
- Distribution of radio footprint allows to reconstruct arrival direction, energy and mass composition of the primary particle.
- Best and most precise CR composition measurements to date (Buitink et al., Nature, 2016).
- Good reconstruction of polarization direction

Courtesy of L. Rossetto

# COSMIC RAYS



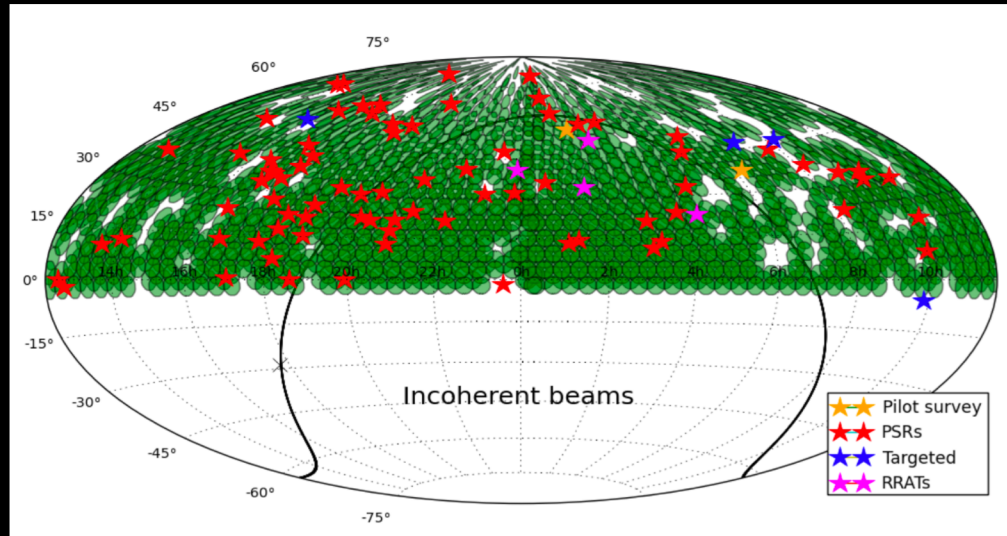
B. Hare et al. Journal of Geophysical Research (2018) 123  
T.N.G. Trinh et al. Submitted to Journal of Geophysical Research (2018)

Routine for mapping thunderstorm and lightning events:

- reconstruction of the **on-sky position of the electric discharge**
- mapping the **electric fields within clouds during thunderstorms**, and characterizing their influence on cosmic-rays radio emission

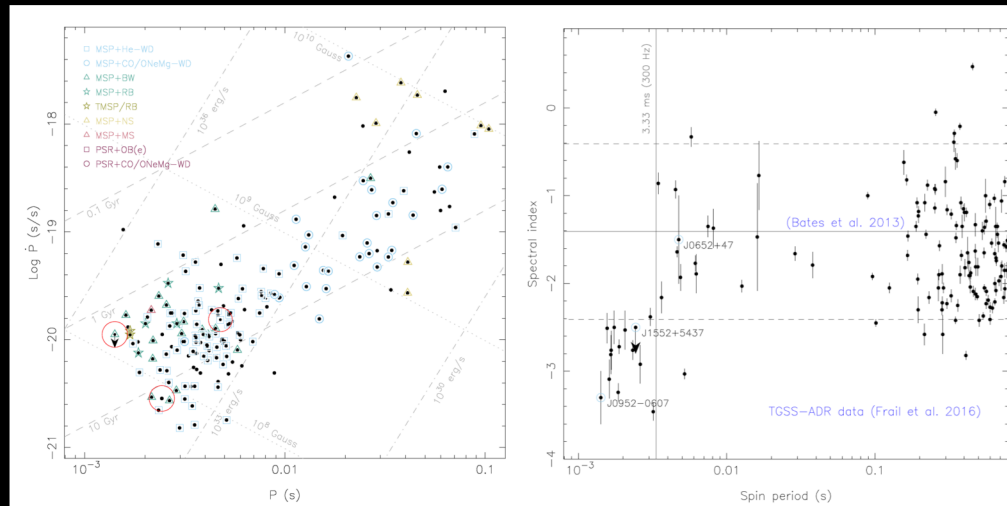
Courtesy of L. Rossetto

# TRANSIENTS & PULSARS



➤ **LOTAAS** – LOFAR Tied Array All-Sky Survey - deepest low frequency pulsars survey ever performed:

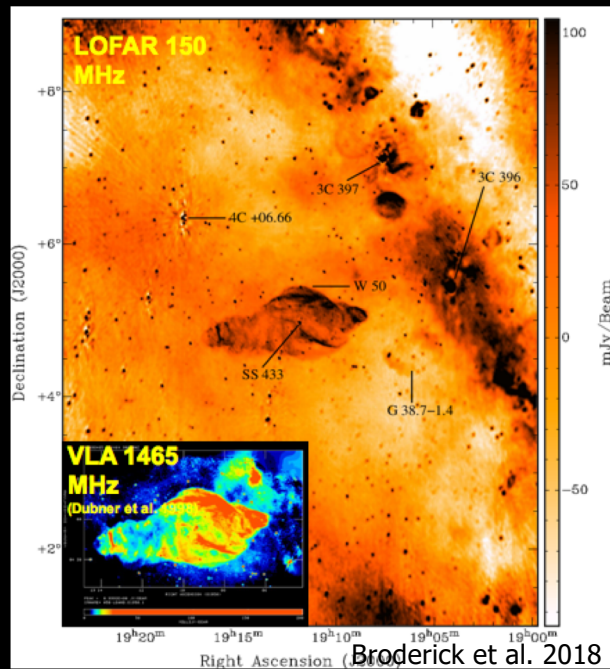
- Discover exotic pulsar systems to **test gravity**, constrain the physics of dense matter, and probe the **pulsar emission mechanism**
- Characterize the low-frequency transient radio sky on sub-second timescales
- Almost completed!
- **85 pulsars discovered** so far
- One of the most successful pulsars surveys in the last decade



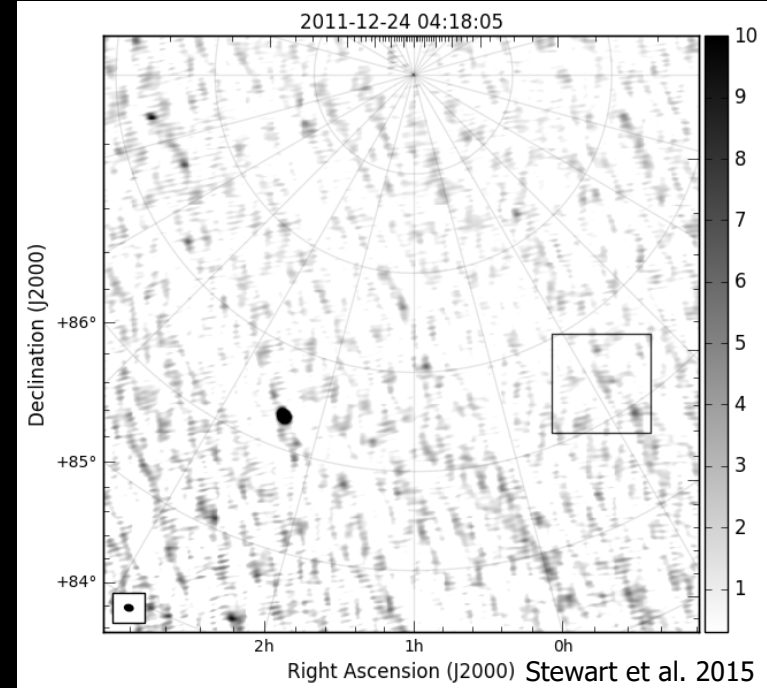
- Discovery of 3 MSPs with LOFAR, including the **fastest MSP in the Galactic field (PSRJ0952-0607)**
- J0250+58: **slowest PSR ever**: 23.5 s period!

Courtesy of C. Bassa, V. Kondratiev and C. M. Tan  
*See beamformed lectures on Thursday*

# TRANSIENTS & PULSARS

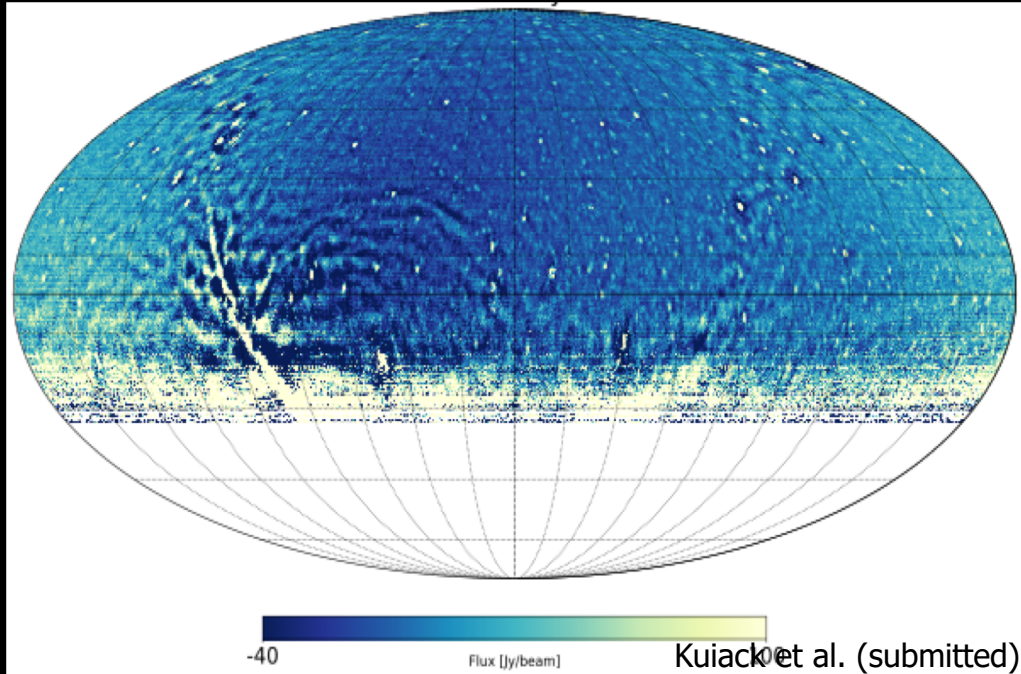


- W 50 morphology in excellent agreement with higher-frequency maps. 150-MHz integrated flux  $\sim 210$  Jy.
- Most complete detection of radio shell of SNR G 38.7-1.4.
- SS 433 marginal variability at 150 MHz; rise corresponds to extended flaring activity at GHz frequencies.



- Detection of **first LOFAR transient** event (LBA)
- **400 h monitoring** data of NCP (single LBA sub-band in MSSS).
- ILT J225347+862146:  $\sim 20$  Jy at 60 MHz. Estimated time-scale of event  $\sim 4$  min
- Flare star? (Scattered) FRB with unusually steep radio spectrum ( $\alpha < -4.7$ )?

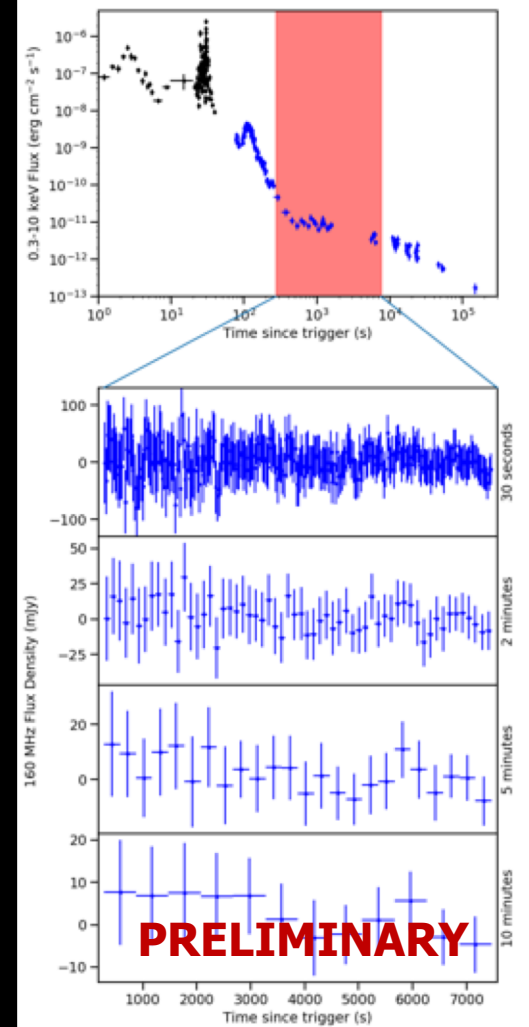
# TRANSIENTS & PULSARS



- **AARTFAAC**: Amsterdam-ASTRON Radio Transients Facility And Analysis Center
- New AARTFAAC source catalogue at 60 MHz

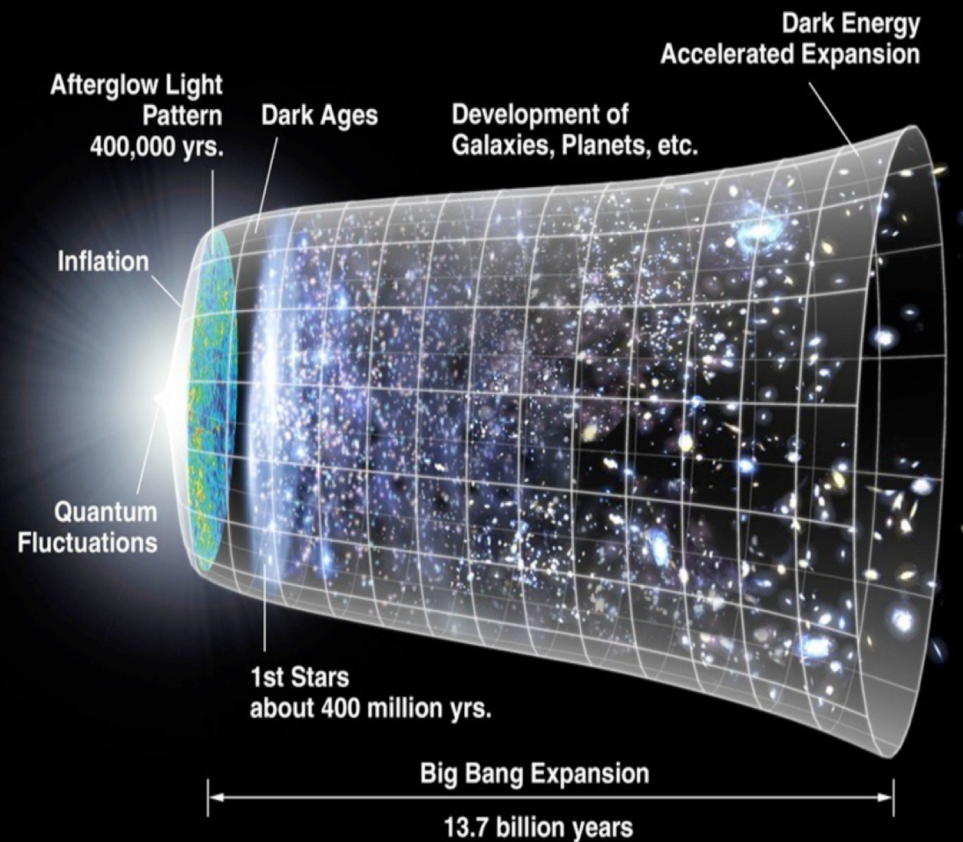
## LOFAR Rapid Response Observations of GRB 180706A

- Observations started within 5 minutes of the GRB
- top: X-ray light curve of the gamma-ray burst (GRB) detected by Swift Observatory. Red box: timescale of the LOFAR observations
- bottom: LOFAR light curve at the position of the GRB at 4 different timescales
- No emission was detected placing the **deepest limits on this to date**



Rowlinson, Gourdji et al. (in prep.)

# EPOCH OF REIONIZATION



- When was the Universe reionized ?
- How (fast) did reionization proceed ?
- Which objects were responsible ?  
stars/galaxies , QSOs, or ...

Redshifted HI to frequency mapping

$z = 6.7 \Rightarrow 185 \text{ MHz}$

$z = 8.5 \Rightarrow 150 \text{ MHz}$

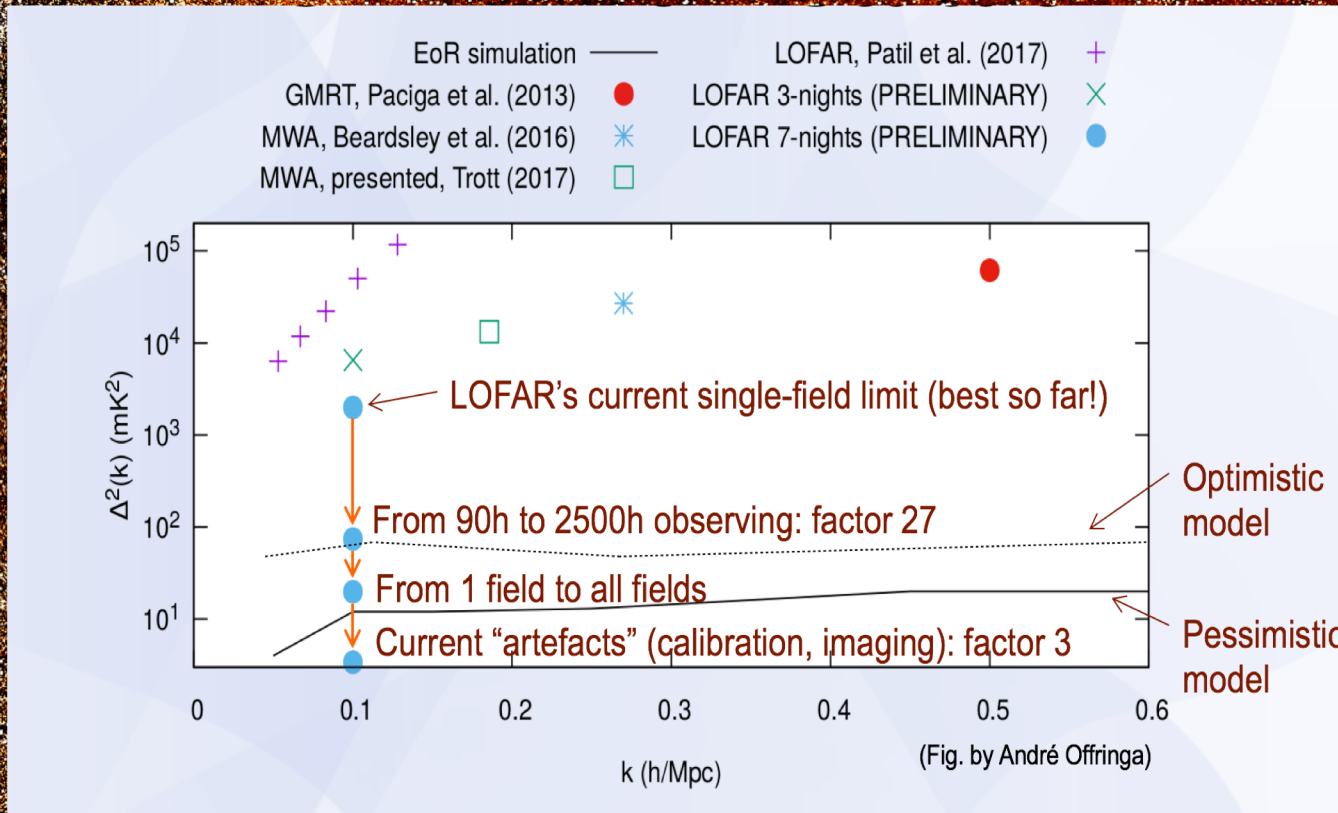
$z = 11.4 \Rightarrow 115 \text{ MHz}$

Goal: Detect cosmological 21cm signal ( $z \sim 6-10$ ) from the Epoch of Reionization

$\Rightarrow 1.5 \text{ Pbytes}$  and  $10^{21}-10^{22}$  FLOP to extract signal!



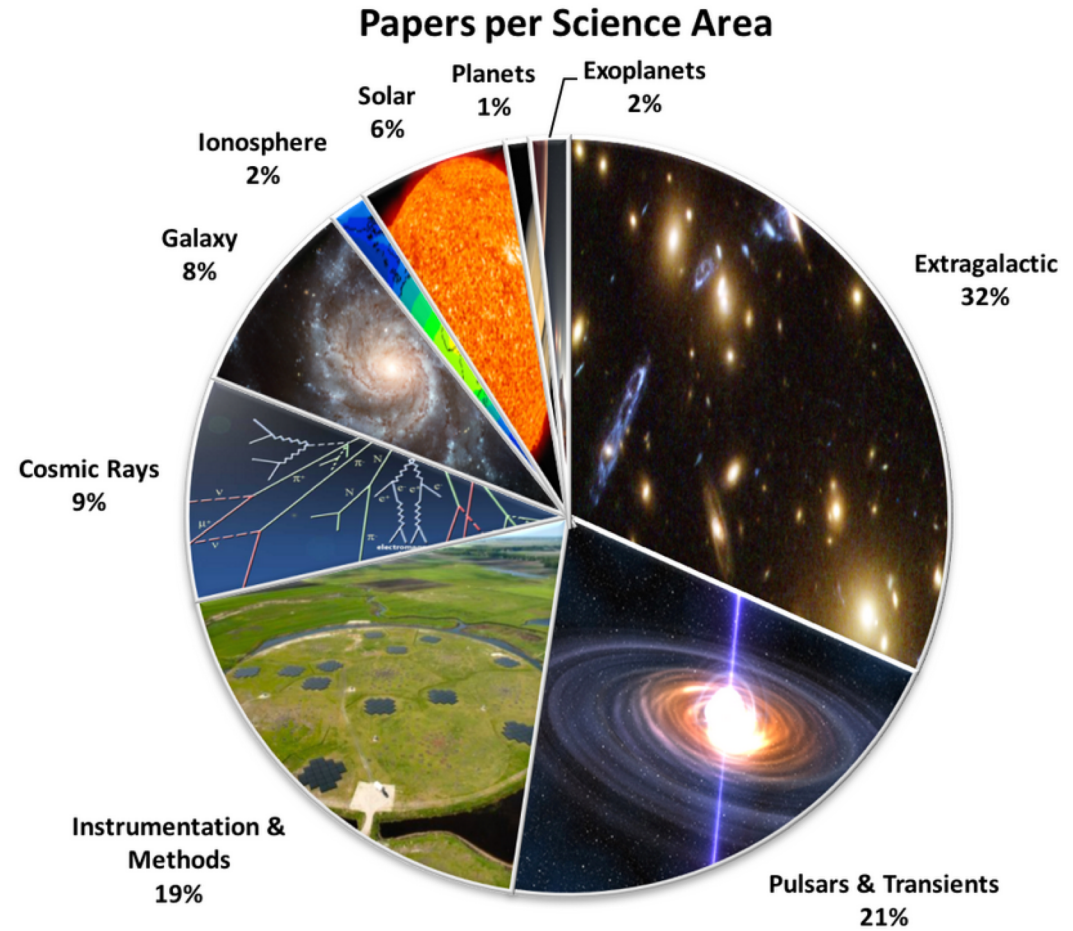
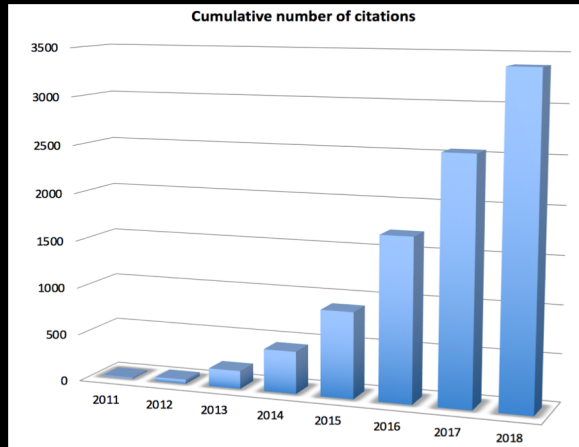
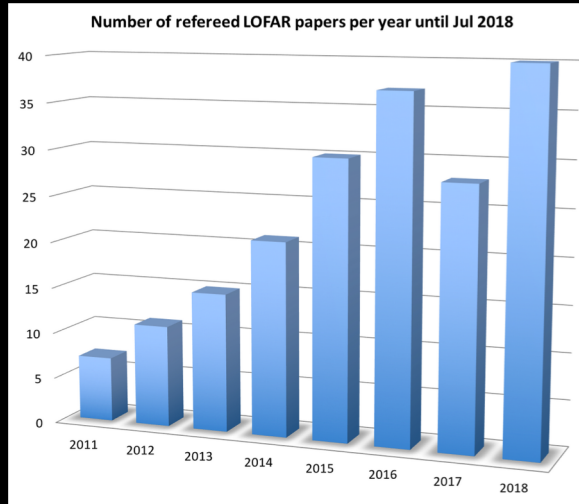
# EPOCH OF REIONIZATION



LOFAR is currently *the best telescope* to detect EoR signals:

- High sensitivity
- The only low-frequency telescope with high (arcsec) resolution
- The LOFAR team has published the current world-leading upper limit

# LOFAR SCIENCE OUTPUT



# SUMMARY

- **LOFAR is generating great scientific data**
- **LOFAR provides several unique scientific capabilities**
- **NEW EXCITING DISCOVERIES ARE THERE FOR YOU TO MAKE!!!**

