Factor pipeline

Facet Calibration for LOFAR (T6)

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Outline

Lecture

DDE calibration/imaging & Factor in a nutshell The Factor parset & directions file The selfcal strategy Running Factor

Tutorial handout

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Tutorial handout

Introduction: DDE cal/img

Direction-Dependent Effects in LOFAR data are primarily caused by:

- The ionosphere (see D2,T1) \rightarrow mostly phase effects (vary quickly in time)
- The LOFAR beam (see L3,L4) \rightarrow mostly amplitude effects (vary slowly in time)

Direction-Dependent calibration attempts to correct for these effects

Introduction: Factor

Factor performs Direction-Dependent calibration and imaging

- Based on the facet-calibration scheme of van Weeren et al. (2016)
- Supports multi-epoch datasets (interleaved or multiple nights)
- Designed to operate on HBA data only
- Recommended for targeted exposures
- Sub-optimal for:
 - full-field of view processing possible
 - very extended (e.g. >=1 deg) target sources (i.e. >1 facet needed)

Introduction: Factor

Factor performs Direction-Dependent calibration and imaging

- Uses LOFAR generic pipeline framework as backend:
 - Allows distribution over cluster nodes
 - Allows resuming of interrupted jobs
- Available from GitHub at <u>https://github.com/lofar-astron/factor</u>
- Relatively modest resources required: 2 CPUs, 32 GB memory, 1TB of disk space



Introduction: Factor

- subtract best model of the sky (pre-Factor Initial-Subtract, see T2)
- divide field into facets based on bright DD calibrators
- cycle over facets:
 - add back the calibrator source
 - self calibrate the calibrator sources (facetselfcal)
 - add back the rest of the facet and apply calibration solutions
 - improve the subtraction with new model and calibration (facetsub)
- image the facets (facetimage)
- make a mosaic of all facets & correct for the primary beam attenuation (fieldmosaic)



Singularity lds-img-cep3.sif:/data/scratch/iacobelli/FactorLDS> checkfactor factor2.0.parset /opt/lofarsoft/lib/python2.7/site-packages/lofarpipe/support/utilities.py : Using default subprocess module!

/usr/local/lib/python2.7/dist-packages/matplotlib/cbook/deprecation.py:107:

MatplotlibDeprecationWarning: The mpl_toolkits.axes_grid module was deprecated in version 2.1. Use mpl_toolkits.axes_grid1 and mpl_toolkits.axisartist provies the same functionality instead. warnings.warn(message, molDeprecation, stacklevel=1)

- INFO factor:progress Plotting directions...
- INFO factor:progress Left-click on a direction to select it and see its current state
- INFO factor:progress Right-click on a direction to deselect it
- INFO factor:progress (In both cases, pan/zoom mode must be off)
- INFO factor:progress Press "c" to display calibrator selfcal images for selected direction
- INFO factor:progress Press "i" to display facet image for selected direction
- $\ensuremath{\text{INFO}}$ factor:progress Press "v" to display facet verify image for selected direction
- INFO factor:progress Press "t" to display TEC solutions for selected direction
- INFO factor:progress Press "g" to display Gain solutions for selected direction
- $\ensuremath{\text{INFO}}$ factor:progress Press "u" to update display (display is updated automatically every minute)
- INFO factor:progress Press "h" to repeat these instructions on this terminal

Introduction: Factor

Factor workflow can be monitored interactively with checkfactor:

\$ checkfactor <parset>

Workflow of a Factor run

Using Factor on a dataset generally involves the following basic steps:



Data preparation

First calibrate data via pre-Factor pipelines https://github.com/lofar-astron/prefactor

 input MS files (uv data) consist of concatenated bands of e.g. 10 subbands each (i.e. = 2 MHz); typical observations will have 24 bands (= 48 MHz total)

After calibration, the Initial-subtract pre-Factor pipeline must be run to:

- image the field at medium and low resolution to make initial sky models
- subtract these models of the sources from the uv data



Setting up the Factor pipeline

Make a directory in your area to hold the parset and Factor output:

\$ mkdir /your/FactorTutorial/dir

For this tutorial, copy the tutorial parset to the directory you made above:

\$ cd /your/FactorTutorial/dir

Edit your copy of the parset:

\$ emacs /your/FactorTutorial/dir/factor.parset

The parset is divided into sections to control and specify the workflow:

- See <u>http://www.astron.nl/citt/facet-doc/parset.html</u> for a full description.
- An example parset that you can use as a basis for your reduction is available at <u>https://www.astron.nl/citt/facet-doc/parset.html</u>
- A customized parset for this tutorial is also shared on the Slack channel. For most parameters, the default values will be fine.

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
global	dir_working	/your/FactorTutorial/dir	
global	dir_ms	/your/FactorTutorial/dir	
global	min_fraction_per_band	0.8	A large value set to maximise the S/N
global	use_compression	True	For less storage usage & faster processing

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
calibration	max_selfcal_loops target_max_selfcal_loops	5 5	For faster processing
calibration	preaverage_flux_Jy	1.0	To maximise the S/N
calibration	multires_selfcal	True	To improve convergence of self-cal

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
imaging	wsclean_nchannels_factor	2	For faster processing
imaging	fractional_bandwidth_selfc al_facet_image	0.75	Large bandwidth for a better subtraction
imaging	wsclean_use_idg idg_mode	True cpu	For faster processing

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
directions	max_radius_deg	3	For faster processing
directions	flux_min_for_merging_Jy size_max_arcmin separation_max_arcmin flux_min_Jy	0.4 1.0 7.0 0.4	Constraints to select calibrators automatically (rule of thumb: bright and compact)
directions	minimize_nonuniformity	True	To make size of facets more uniform in size

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
directions	faceting_radius_deg	0.8	For faster processing
directions	target_ra target_dec target_radius_arcmin target_has_own_facet	03h35m23.407 +55d04m07.124 4.0 True	To specify a target area to ensure that a facet edge does not cross it. For this tutorial, we specify a region around a faint extended source
directions	minimize_nonuniformity	True	To make size of facets more uniform in size

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
cluster	ncpu	all	Tune it based on your machine features !
cluster	wsclean_fmem	0.9	Tune it based on your machine features !
cluster	ndir_per_node	1	Tune it based on your machine features !

The parset is divided into sections to control and specify the workflow:

Section	Option	Value	Notes
checkfactor	facet_viewer	ds9	
checkfactor	ds9_load_regions	False	Set to True for inspection
checkfactor	image_display	eog	

The directions (facets) file

Factor records selected DDE calibrators in a file called factor directions.txt

\$ more factor directions.txt

name position atrous_do mscale_field_do cal_imsize solint_ph solint_amp dynamic_range region_selfcal region_facet
peel_skymodel outlier_source cal_size_deg cal_flux_mJy

LSMTool history:

• • •

facet_patch_41 3h27m31.1093s,55d21m30.5782s empty empty 0 0 0 LD empty empty False 0.1050687533 36442.853868 facet_patch_35 3h29m52.7506s,53d32m43.7471s empty empty 0 0 0 LD empty empty False 0.0128485275297 10059.9407256 facet_patch_22 3h36m28.7034s,53d38m56.1925s empty empty 0 0 0 LD empty empty empty False 0.0112901740997 8669.60507632 facet_patch_40 3h28m11.8884s,55d09m08.5092s empty empty 0 0 0 LD empty empty empty False 0.0110088789021 7315.00279903





Factor will divide the field into facets (one for each DDE calibrator) inside of faceting_radius_deg (under the [directions] section of the parset). Outside of this radius, boxes are used instead (to speed up processing).

The directions (facets) file: editing

Directions can be:

- added or removed to improve facet layout
- reordered (e.g., in order to process only a subset of directions)
 Other, advanced options can be activated on a per-direction basis (see online doc)

Warning

 after the first direction has completed for a run, changes to the facet layout will result in incorrect results!



Verification of DDE calibrators

To check pre-selected calibrators and facets use auxiliary files saved in /your/FactorTutorial/dir/regions:

- load a wide-field image of the field in ds9 (e.g., one made by the Initial Subtract pre-Factor pipeline)
- load (under the "Region" menu) files named
 - o calimages_ds9.reg
 - o facets_ds9.reg

Warning

• To check & edit DDE calibrators list set in the [global] section of the parset interactive=True

The selfcal strategy

Selfcal is done on each calibrator to improve its model and DDE corrections

Factor minimizes the number of free parameters solved for during selfcal in order to avoid overfitting:

- Fast phases (TEC): one solution every ~10-20 seconds and ~10 MHz to track rapid changes due to ionosphere
- Slow gains (amp+phase): one solution every ~10-20 minutes and ~2 MHz to correct for beam effects



The selfcal strategy

TEC (from fast-phases): correction for ionospheric errors

- Tracking of phase changes has different S/N for different DDE calibrators as well as stations
 - The brightest and compact the DDE calibrator is the better results expected
- Quicker phase changes expected for stations further away from the reference one
- Noisy periods or jumps indicate problems !



The selfcal strategy

Amplitudes (from slow-gains): correction for beam effects

- Tracking of amplitude changes has different S/N for different DDE calibrators as well as stations
 - The brightest and compact the DDE calibrator is the better results expected
- Amplitude correction values for XX and YY should be close to 1; systematic deviations suspicious:
 - resolved source structure,
 - Instrumental systematics









Gain TEC +



image32





The selfcal strategy

Results of iterative selfcal (DIE, TEC, amplitude and TEC) corrections on the image plane are provided as useful diagnostic per DDE calibrator (see next section)

A run of Factor is started via the command

```
$ runfactor -v <your factor>.parset
```

Because the expected pipeline runtime is >2 days run the pipeline in a screen session!

Once Factor has started, you can check the progress of a run with the checkfactor tool:

\$ checkfactor <your factor>.parset

INFO - factor:progress - Plotting directions...
INFO - factor:progress - Left-click on a direction to select it and see its current state
INFO - factor:progress - Right-click on a direction to deselect it
INFO - factor:progress - (In both cases, pan/zoom mode must be off)
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INFO - factor:progress - Press "u" to update display (display is updated automatically every minute)
INFO - factor:progress - Press "h" to repeat these instructions on this terminal

After a short time, a window should appear showing the facet layout

Click on a facet

- yellow: the one being processed
- green: the one completed

Hit the "c" key to see the calibrator images made during selfcal



X Figure 1

Overview of Factor run in /data/scratch/iacobelli/FactorLDS

RΔ

Summary of processing for selected facet

Click on a facet

- yellow: the one being processed
- green: the one completed

Hit the "c" key to see the calibrator images made during selfcal

Image before any direction dependent calibration



Click on a facet

- yellow: the one being processed
- green: the one completed

Hit the "c" key to see the calibrator images made during selfcal

image02 Clean masks in red bir. Inder image12 image22 iter0 image22 iter1 Ш image32 image42 iter0 image42 iter1 TEC + Gain

Images after fast phases calibrations

Click on a facet

- yellow: the one being processed
- green: the one completed

Hit the "c" key to see the calibrator images made during selfcal

Images after slow gain calibrations



Click on a facet

- yellow: the one being processed
- green: the one completed

Hit the "c" key to see the calibrator images made during selfcal

The tutorial run setup was with max_selfcal_loops = 5 Factor stops after no more improvement is seen !



Click on a facet

yellow: the one being processedgreen: the one completed

Hit the "v" key to open the selfcal verification (i.e. the residual, source subtracted) images in ds9 (the facetselfcal operation must be finished first). This may take a few seconds..

Once selfcal has completed for all the facets, Factor re-images each facets using the full bandwidth.

• These are the final images, but must be corrected for the primary beam attenuation !

Once the reimaging is complete, Factor will mosaic the images together and correct them for the primary beam

For each facet Factor will make one or more images, as set by the WSClean imaging parameters in the [imaging] section of the parset, e.g.

facet cellsize arcsec = [1.5, 5.0]

facet_robust = [-0.5, -1.0]

facet_taper_arcsec = [0.0, 15.0]

facet_min_uv_lambda = [80.0, 80.0]

Select the "field" direction (in the top-left corner)

- Hit the "i" key to list the available field images
- Hit the "1" key to open the full- resolution image in ds9 (this may take a few seconds...)

Facets regions can be loaded:

- via the [checkfactor]section of the parset.
- via the ds9 "Region" menu (in Factor_output/regions/facets_ds9.reg)

Two mosaics are made for each set of imaging parameters:

- Non-primary-beam corrected ("flat noise"), loaded by checkfactor (for inspection) Factor_output/results/fieldmosaic/field/*.correct_mosaic. pbcut.fits
- Primary-beam corrected image needed for flux measurements: Factor_output/results/fieldmosaic/field/*.correct_mosaic. pbcor.fits

Due to problems with the LOFAR beam model, the flux scale can be off by up to 15%. A check against e.g. TGSS catalog can be done (see T11)

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Tutorial handout

To get started

- Make a directory in your area to hold the
 - parset,
 - Factor input data (10GB) & output (+10GB)
- Open in your web browser the online full documentation <u>https://www.astron.nl/citt/facet-doc/</u>

Input data & parset

- Collect input data in a single directory
 - Input data are 23 subbands (= 2 MHz) of a snapshot (10-min) HBA observation of the PSR0329+54
 - Averaged by preFactor to 0.05 MHz per channel and 8 seconds per timeslot
- Create & edit your parset file (see Lecture notes under "The Factor parset"):
 - \$ touch factor2.parset
 - \$ emacs factor2.parset

Starting Factor

- Enter a screen session
 - \$ screen [+enter]
- Initialise your Singularity environment, mounting your Factor directory
 - \$ singularity shell --bind /your/FactorTutorial/dir/, /your/Singularity.sif
 - \$ source /opt/lofarsoft/lofarinit.sh
- Run Factor in the screen session with the parset you made
 - \$ runfactor -v factor2.parset
- Exit from the screen (CTRL-A CTRL-D)

Check DDE calibrators & facets

- When the factor_directions.txt file is made, initialise your Singularity environment, mounting your Factor directory
 - \$ singularity shell --bind /your/FactorTutorial/dir/, /your/Singulatity.sif

```
$ source /opt/lofarsoft/lofarinit.sh
```

• Load in ds9 (1) the provided wide-field image of the field (obtained by the Initial Subtract preFactor pipeline) and (2) Factor auxiliary *.reg files (see Lecture notes under "Verification of DDE calibrators")

\$ ds9 L644395_SB000_uv_12BAF0A46t_163MHz.pre-cal.wsclean_high-image.fits

• Next check the progress of the run (see Lecture notes under "Running Factor")

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