

Radboud University



university of  
 groningen

**ASTRON**

Netherlands Institute for Radio Astronomy

**Study of the cosmic ray  
radio emission pattern  
at ground level with LOFAR**

**Laura Rossetto**

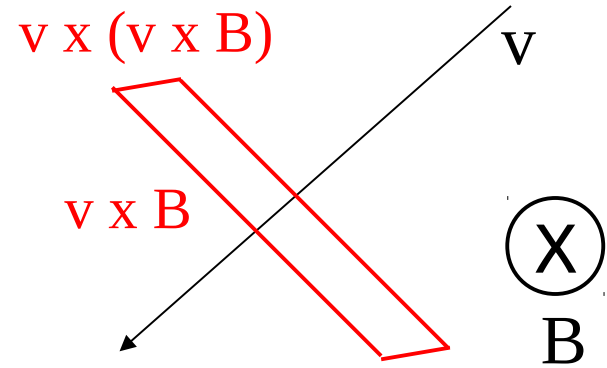
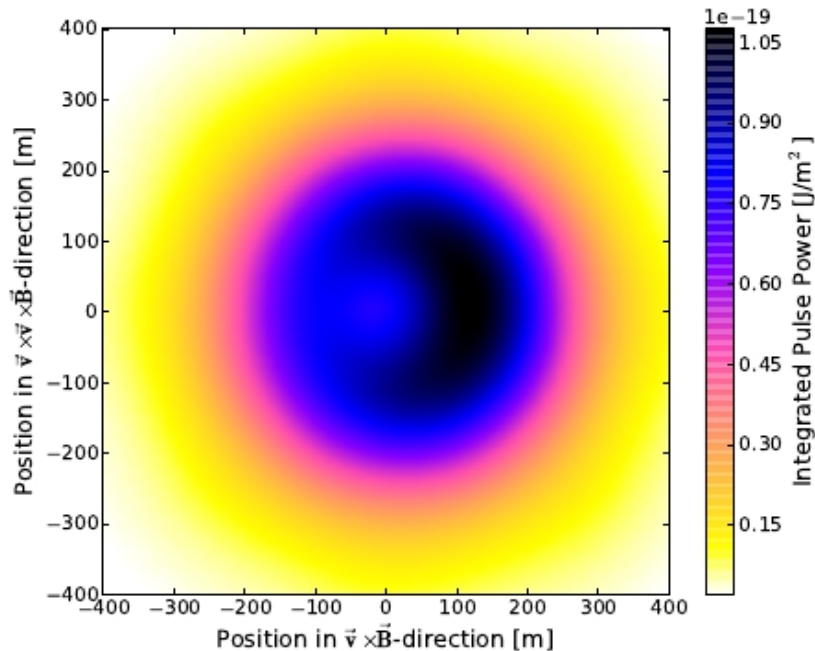
**on behalf of the Cosmic Ray Key Science Project**

*LOFAR Science Workshop, Assen, June 2<sup>nd</sup> – 3<sup>rd</sup> 2015*

# The footprint of the radio signal

*A. Nelles et al., 2014, Astroparticle Physics, 60, 13-24*

*A. Nelles et al., 2015, Journal of Cosmology and Astroparticle Physics 05, 018*

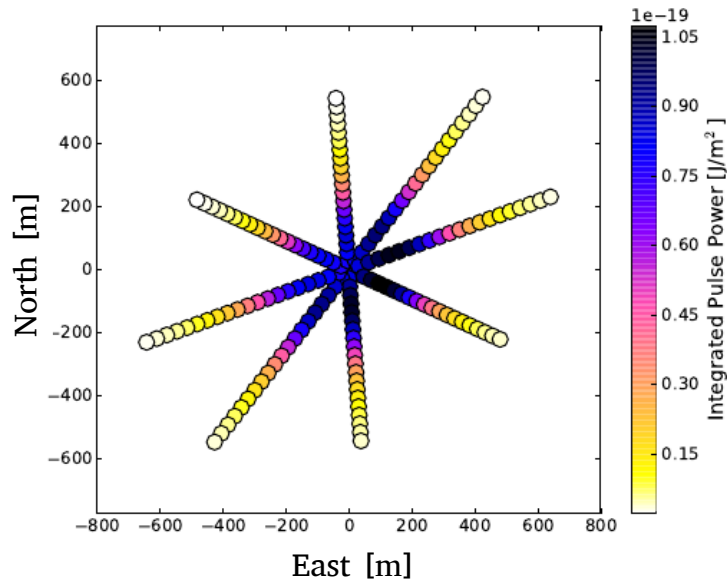


$\vec{v}$  = shower axis

$\vec{B}$  = Earth magnetic field vector

## What is the shape of the radio signal footprint?

# Simulation study



← On the ground level

Particle production

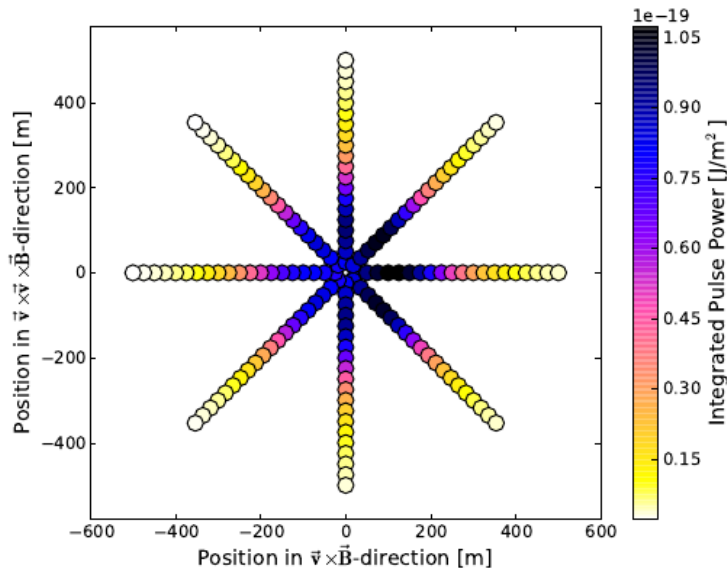
→ CORSIKA 7.400

with FLUKA 2011.2b and QGSJETII.04

Radio emission → CoREAS

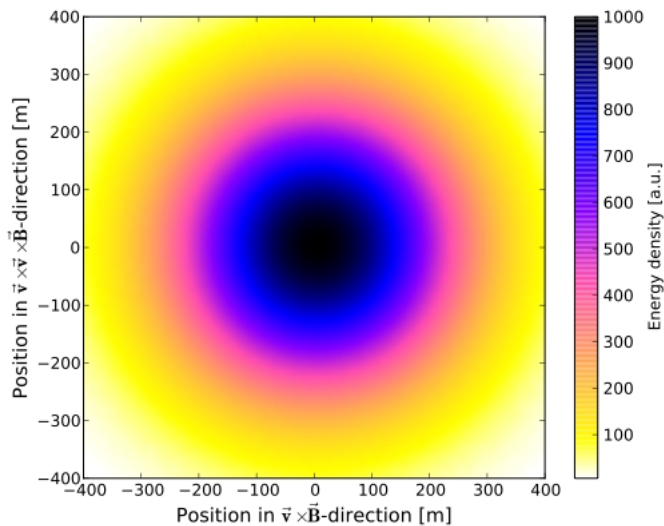
Energy =  $10^{16} - 10^{18.8}$  eV, zenith =  $3^\circ - 55^\circ$

Antenna layout → star-shape pattern in the shower plane



← Plane perpendicular to the shower axis

# Simulation study



Two-dimensional Gaussian distribution

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right)$$

$P$  → total power of the integrated signal

$x'$  → coordinate along the  $\mathbf{v} \times \mathbf{B}$  direction

$y'$  → coordinate along the  $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$  direction

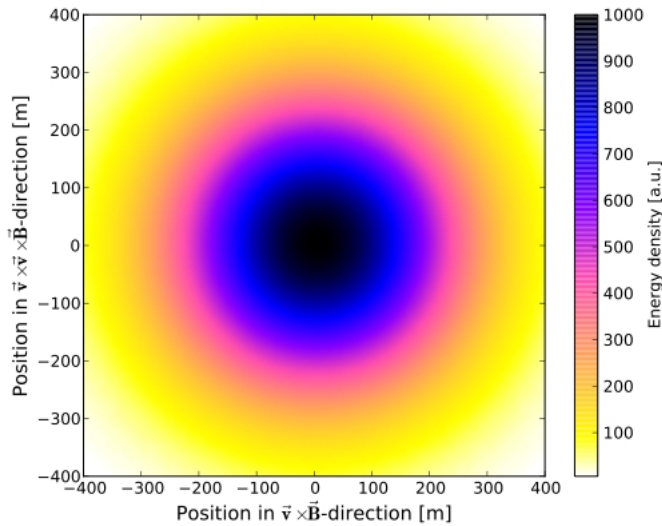
$X_+$  → location parameter along  $\mathbf{v} \times \mathbf{B}$

$Y_+$  → location parameter along  $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$

$\sigma_+$  → width parameter

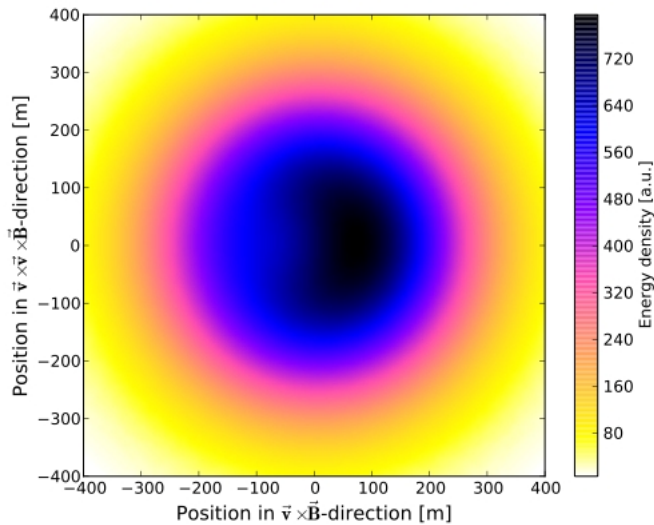
$A_+$  → amplitude parameter

# Simulation study



Two-dimensional Gaussian distribution

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right)$$



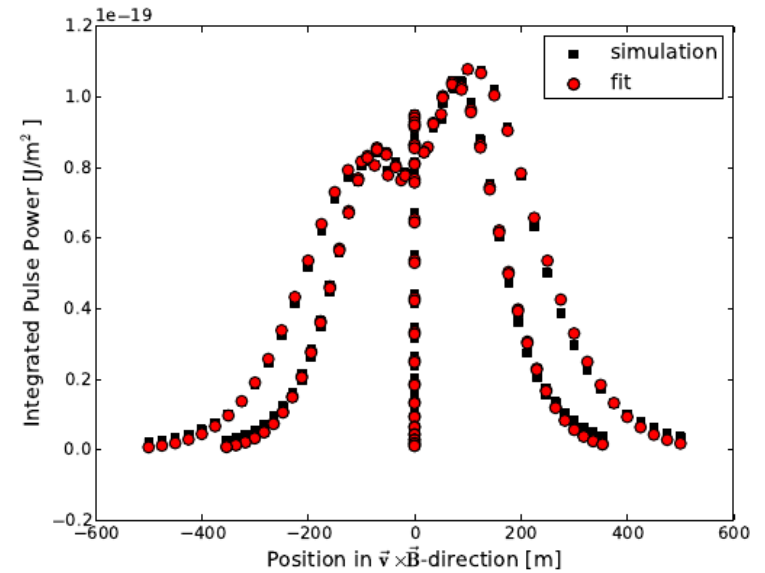
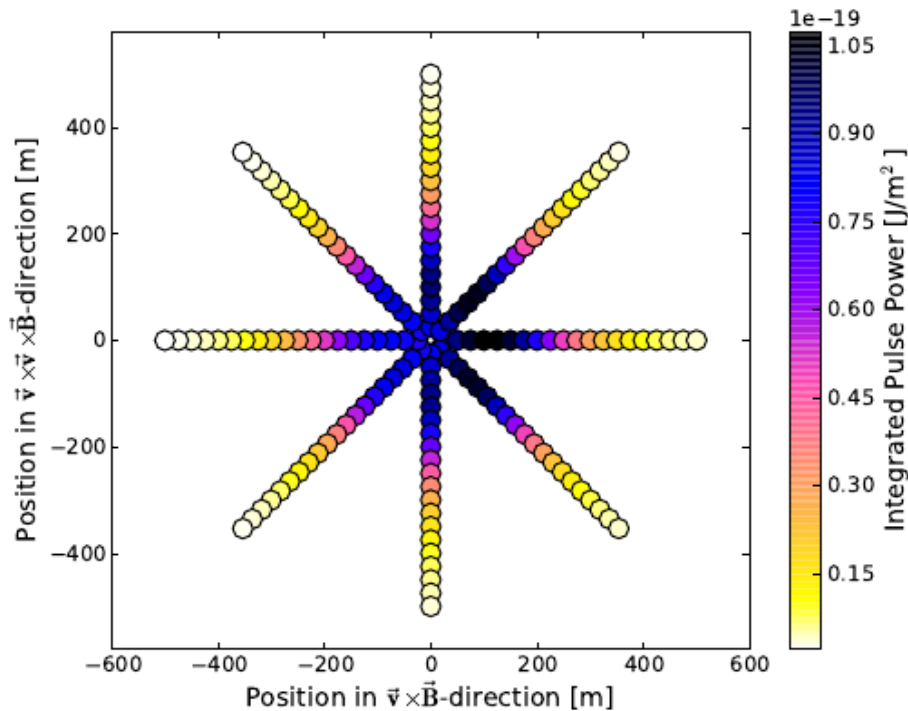
“Bean shape” → combination of two-dimensional Gaussian distributions

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right) - A_- \cdot \exp\left(\frac{-[(x' - X_-)^2 + (y' - Y_-)^2]}{\sigma_-^2}\right) + O$$

- The two Gaussians are shifted and subtracted from each others ( $A_+ > A_-$ )
- 9 free parameters

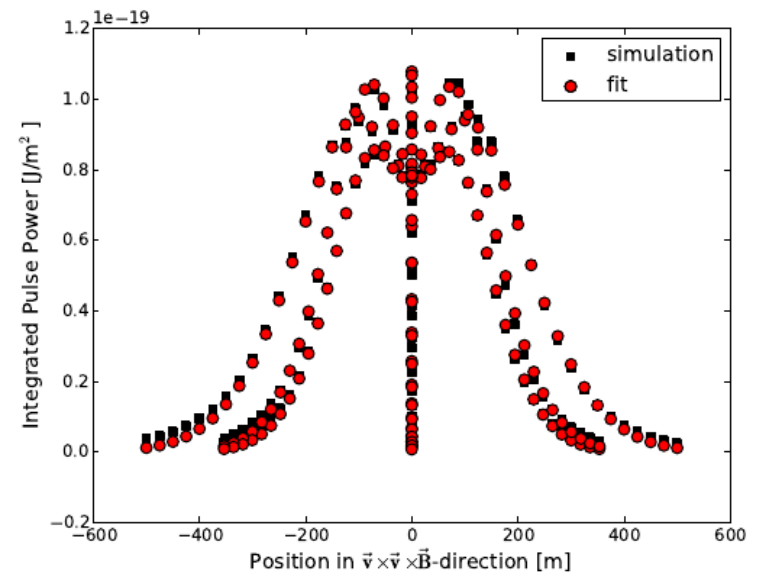
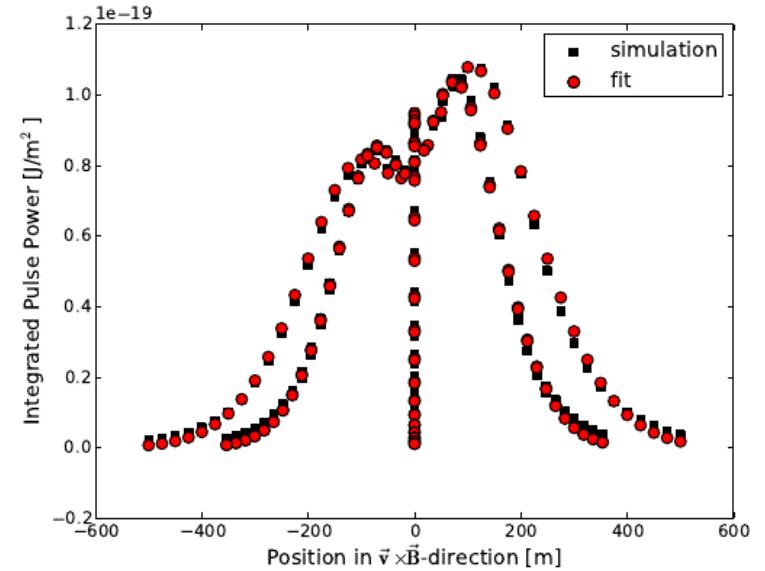
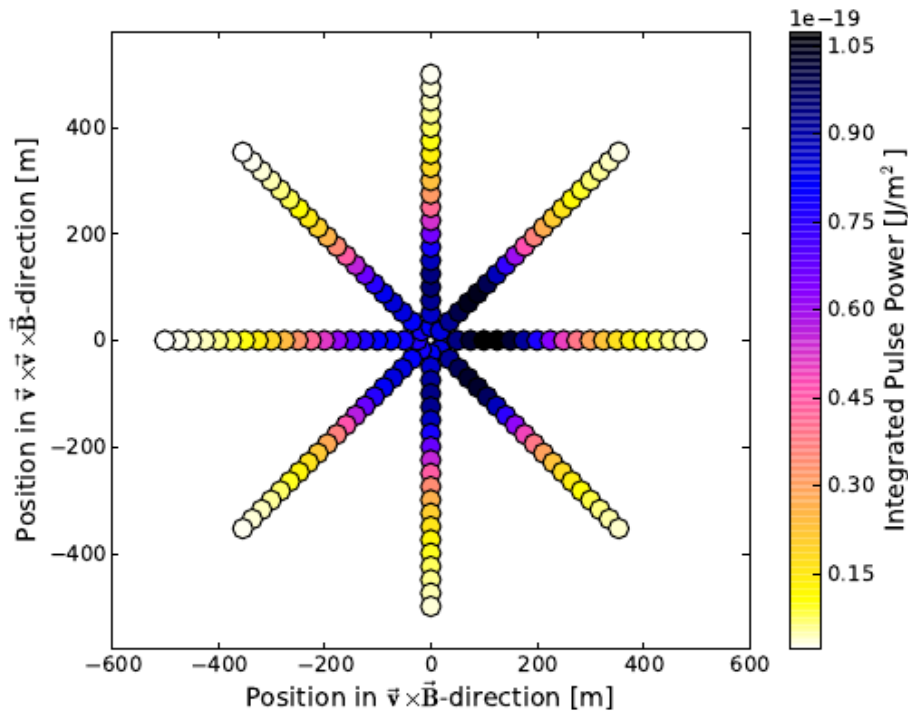
# Simulation study

## Fit results for a single simulated shower



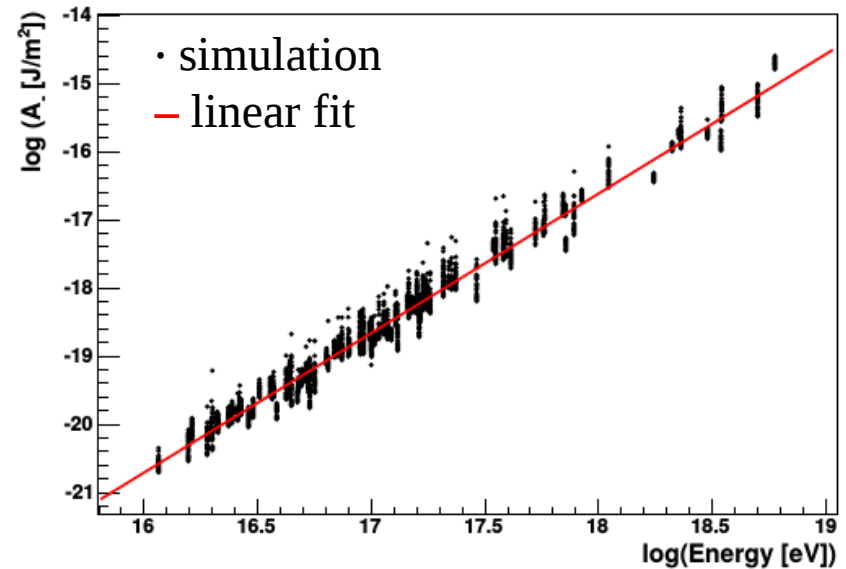
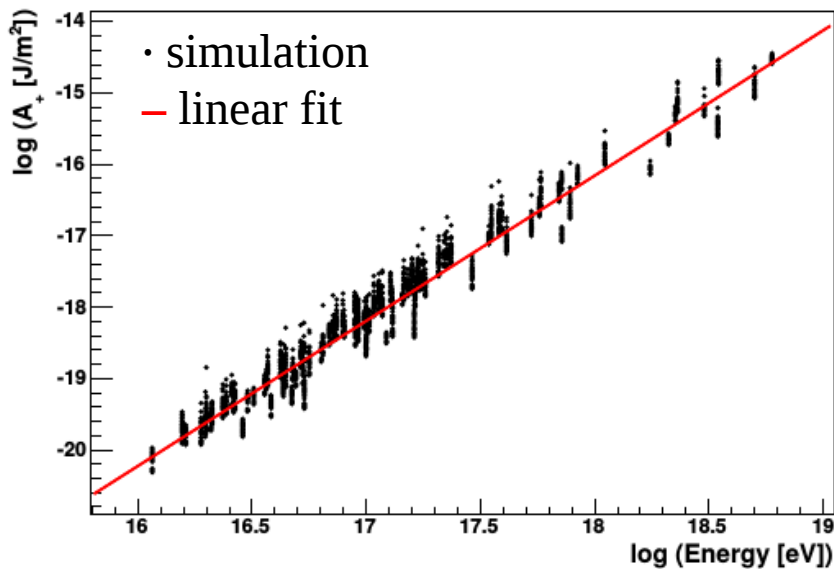
# Simulation study

## Fit results for a single simulated shower



# Simulation study

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right) - A_- \cdot \exp\left(\frac{-[(x' - X_-)^2 + (y' - Y_-)^2]}{\sigma_-^2}\right)$$



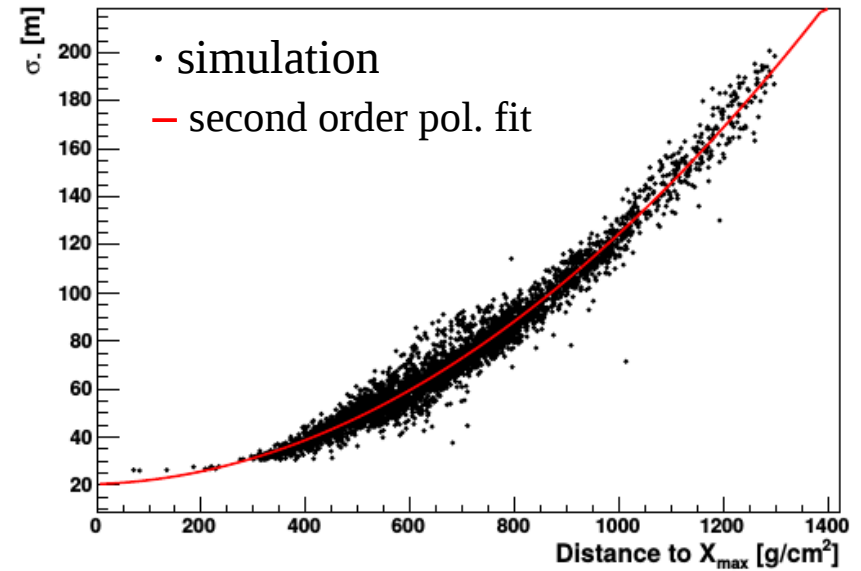
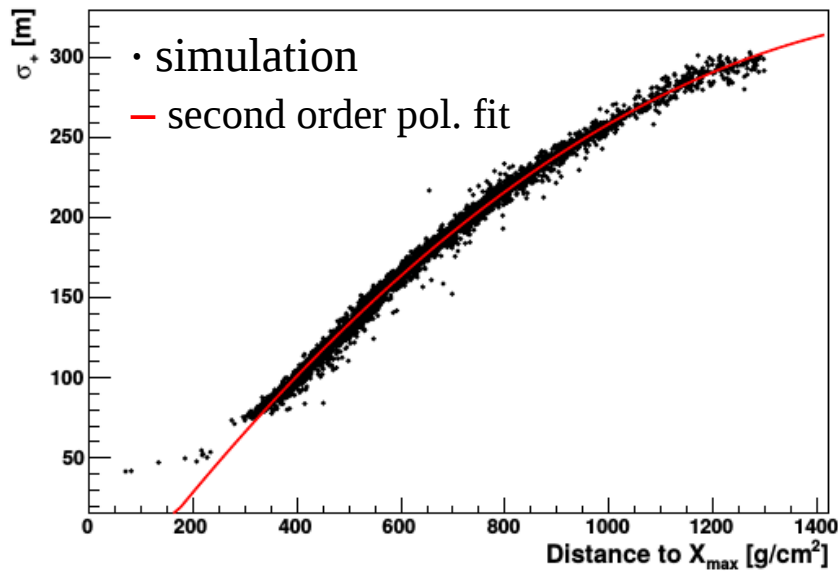
$$A_{\pm} \propto n_{\text{charged}} \propto E^2$$

For coherent emission



# Simulation study

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right) - A_- \cdot \exp\left(\frac{-[(x' - X_-)^2 + (y' - Y_-)^2]}{\sigma_-^2}\right)$$

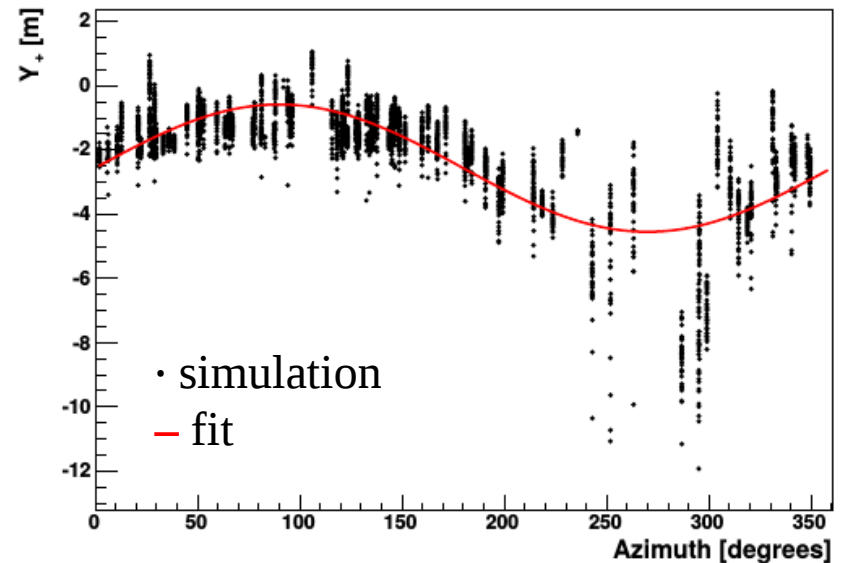
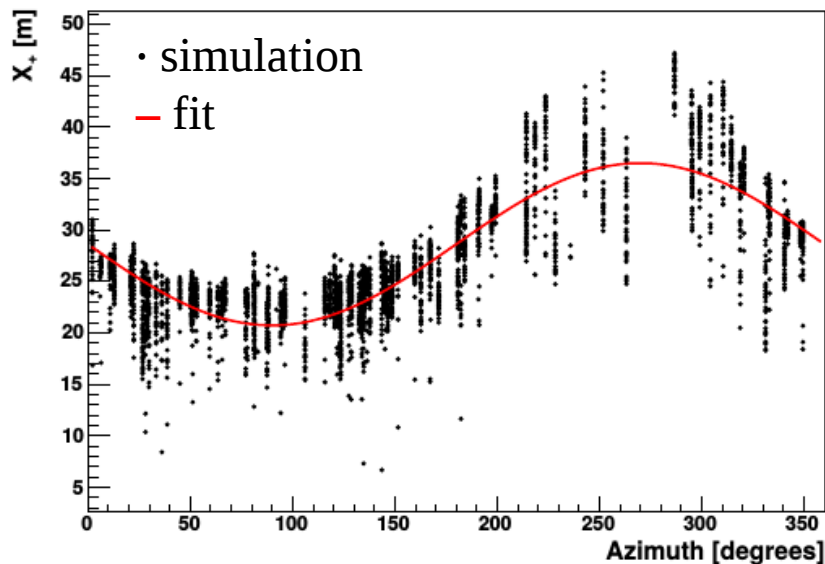


Width of the distribution  $\sigma_{\pm}$  is proportional to the distance of the shower maximum from ground level

$$D(X_{\max})[\text{g}/\text{cm}^2] = X_{\text{atm}}[\text{g}/\text{cm}^2] / \cos(\theta) - X_{\max}[\text{g}/\text{cm}^2]$$

# Simulation study

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_+)^2 + (y' - Y_+)^2]}{\sigma_+^2}\right) - A_- \cdot \exp\left(\frac{-[(x' - X_-)^2 + (y' - Y_-)^2]}{\sigma_-^2}\right)$$



Shift of the positive and negative Gaussian w.r.t. the shower core is proportional to the azimuth  
→ combination of the geomagnetic and charge excess effect in the radio emission

# Data analysis: reduction of free parameters

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_c)^2 + (y' - Y_c)^2]}{\sigma_+^2}\right) - C_0 \cdot A_+ \cdot \exp\left(\frac{-[(x' - (X_c + x_-))^2 + (y' - Y_c)^2]}{(C_3 \cdot e^{C_1 + C_2 \cdot \sigma_+})^2}\right)$$

The number of free parameters can be reduced by considering correlations between parameters

→ the maximum reduction is obtained by using only 4 free parameters

$X_c$  → location parameter along  $v \times B$

$Y_c$  → location parameter along  $v \times (v \times B)$

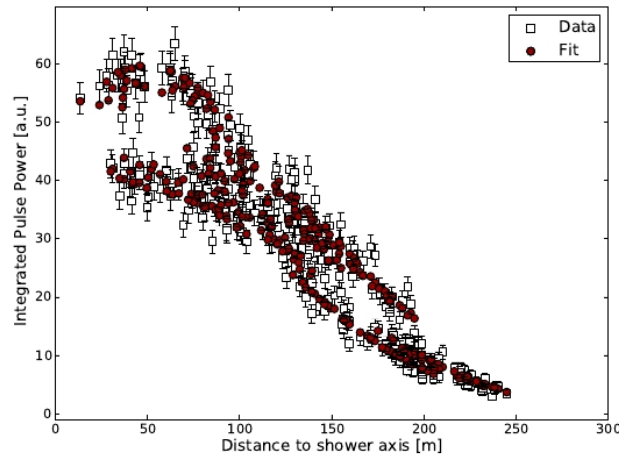
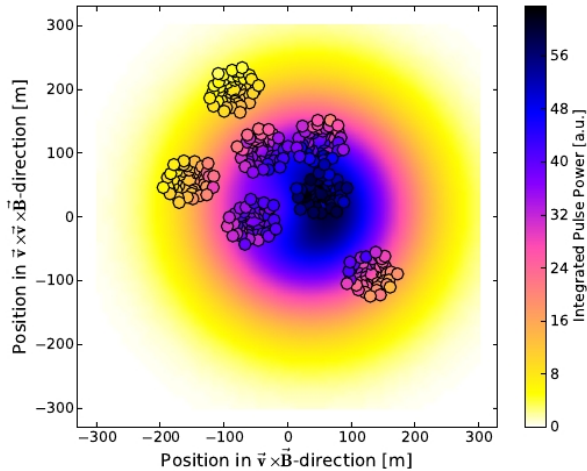
$\sigma_+$  → width parameter

$A_+$  → amplitude parameter

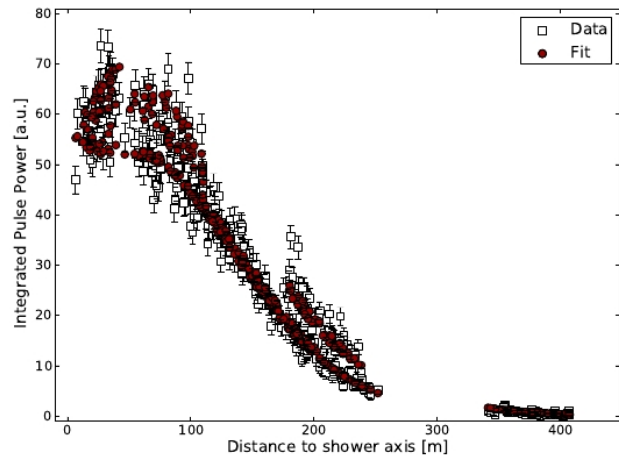
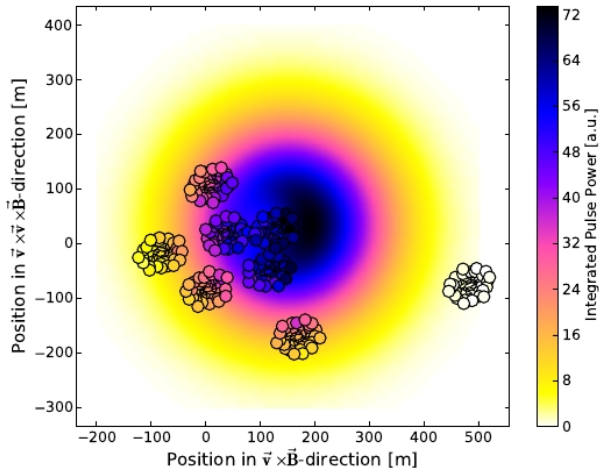
→  $C_0$ ,  $C_1$  and  $C_2$  constant

→  $C_3$  binned for zenith angle

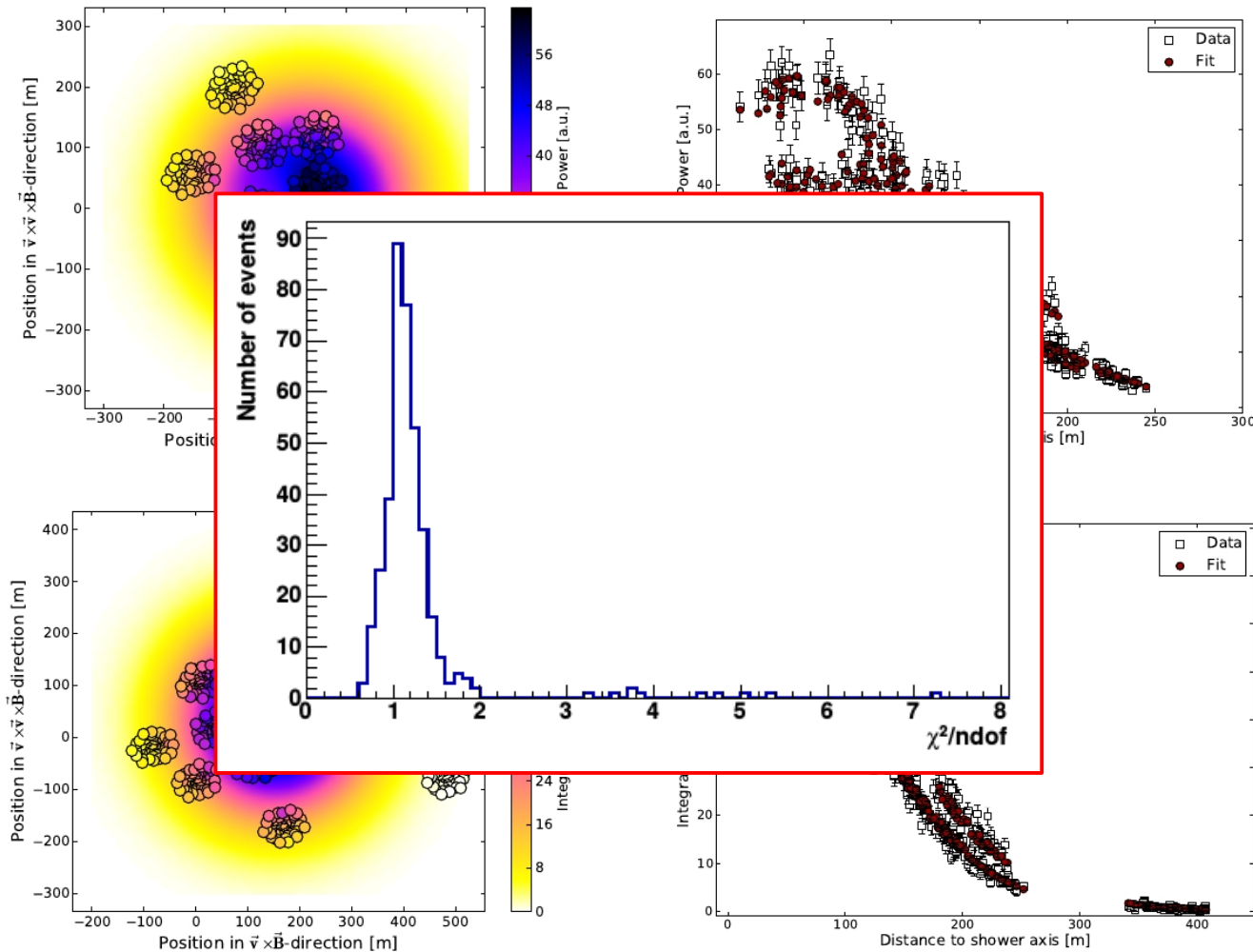
# Parametrization applied to DATA



→ Fit applied to ALL LBA LOFAR data



# Parametrization applied to DATA



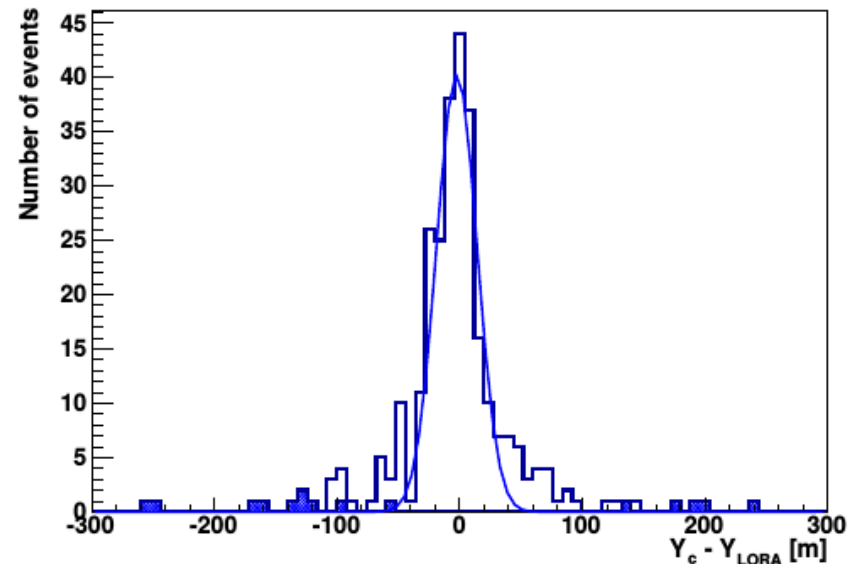
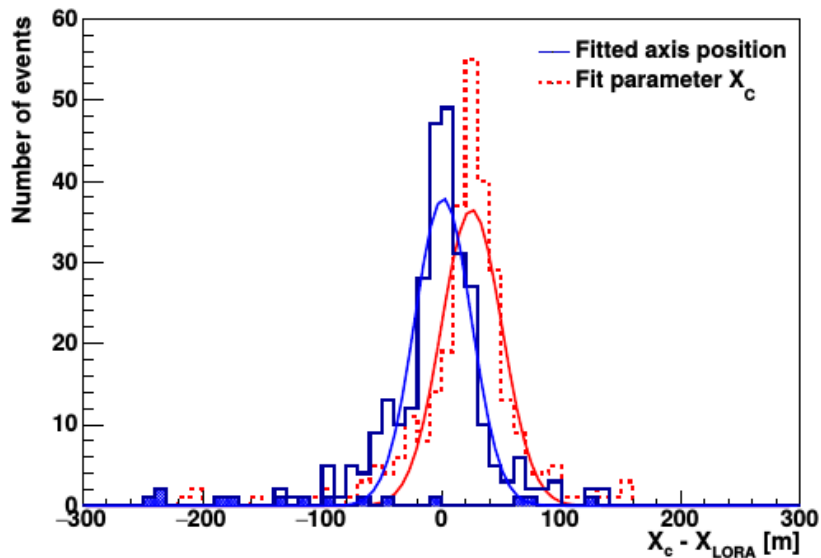
- Fit applied to ALL LBA LOFAR data
- Good quality fit
- events with  $\theta < 15^\circ$  show slightly less good fit

# Reconstruction of shower parameters

## Shower core position

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_c)^2 + (y' - Y_c)^2]}{\sigma_+^2}\right) - C_0 \cdot A_+ \cdot \exp\left(\frac{-[(x' - (X_c + x_-))^2 + (y' - Y_c)^2]}{(C_3 \cdot e^{C_1 + C_2 \cdot \sigma_+})^2}\right)$$

Cross-check with  
LORA detectors

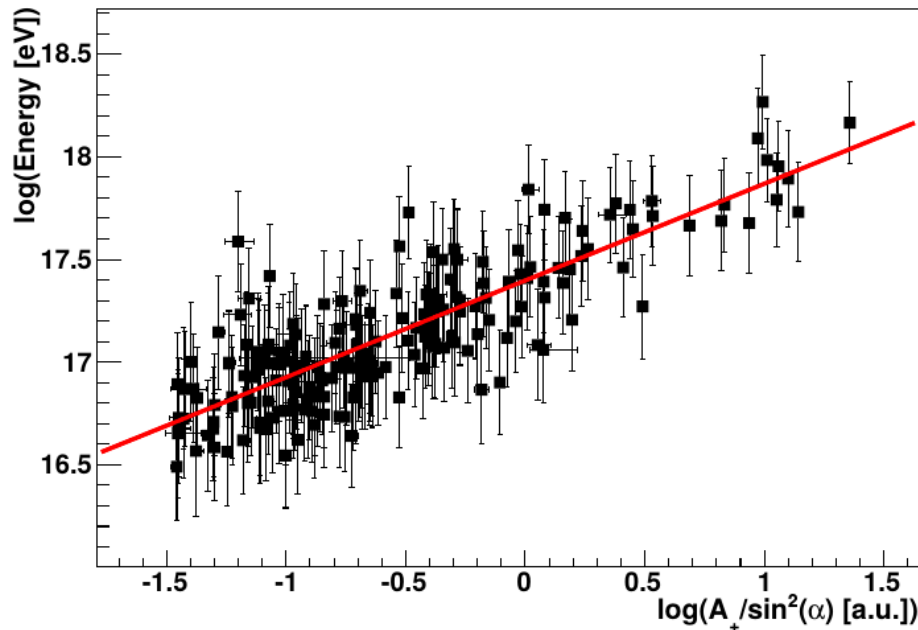


# Reconstruction of shower parameters

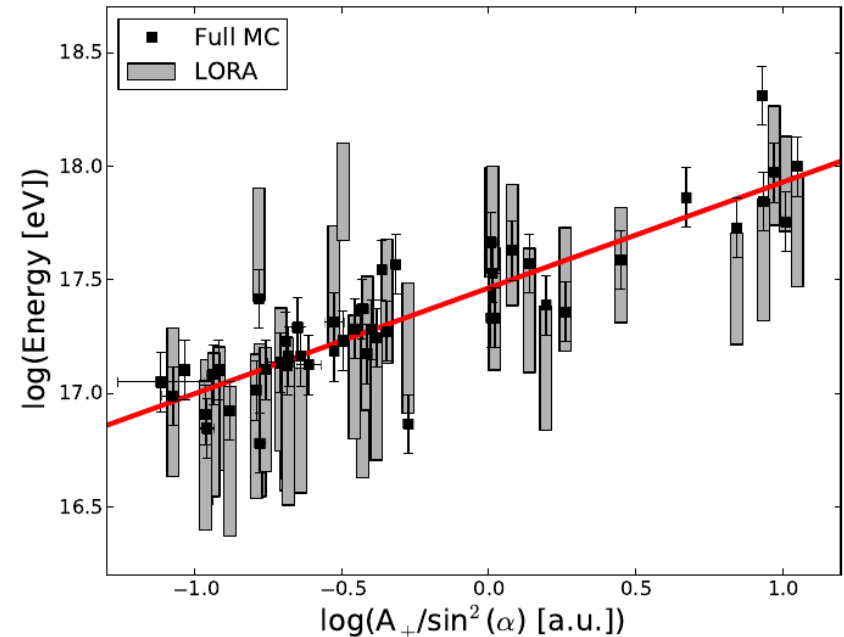
## Energy of primary particle

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_c)^2 + (y' - Y_c)^2]}{\sigma_+^2}\right) - C_0 \cdot A_+ \cdot \exp\left(\frac{-[(x' - (X_c + x_-))^2 + (y' - Y_c)^2]}{(C_3 \cdot e^{C_1 + C_2 \cdot \sigma_+})^2}\right)$$

Cross-check with  
LORA detectors



Cross-check with  
MC simulations



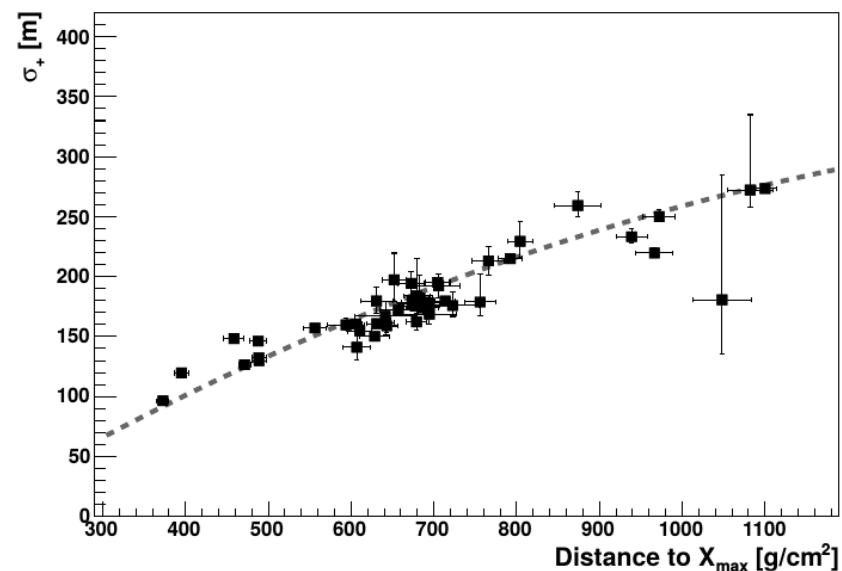
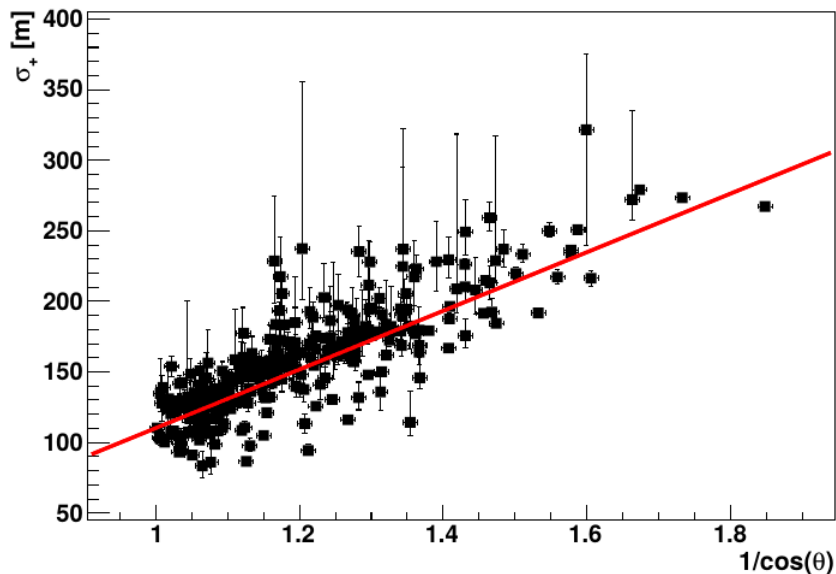
# Reconstruction of shower parameters

## Shower maximum

$$P(x', y') = A_+ \cdot \exp\left(\frac{-[(x' - X_c)^2 + (y' - Y_c)^2]}{\sigma_+^2}\right) - C_0 \cdot A_+ \cdot \exp\left(\frac{-[(x' - (X_c + x_-))^2 + (y' - Y_c)^2]}{(C_3 \cdot e^{C_1 + C_2 \cdot \sigma_+})^2}\right)$$

Width parameter  $\sigma_+$  proportional to  $1/\cos(\theta)$

$$D(X_{\max})[\text{g}/\text{cm}^2] = X_{\text{atm}}[\text{g}/\text{cm}^2] / \cos(\theta) - X_{\max}[\text{g}/\text{cm}^2]$$

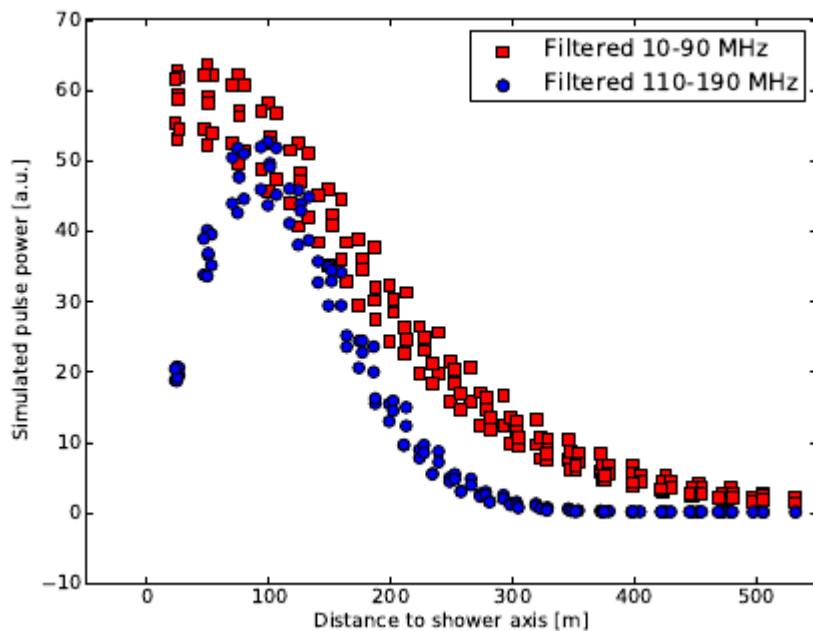




# Reconstruction of shower parameters

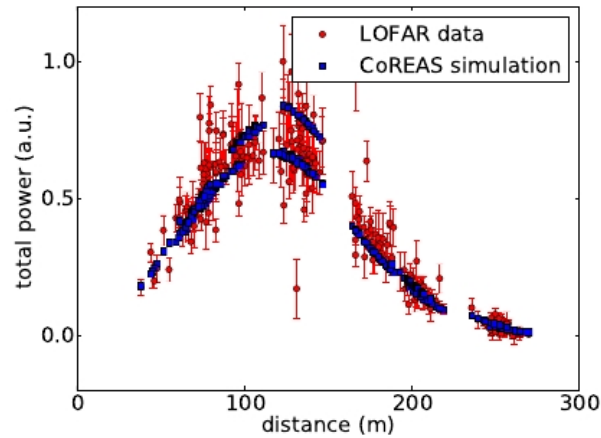
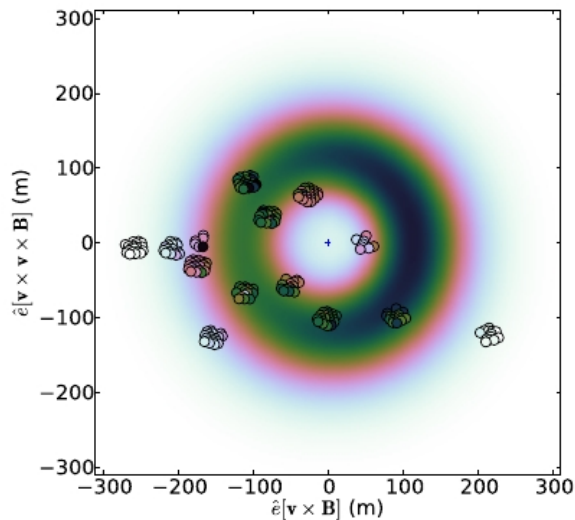
## What does it happen at frequencies around 100 MHz?

*A. Nelles et al., 2015, Astroparticle Physics, 65, 11-21*



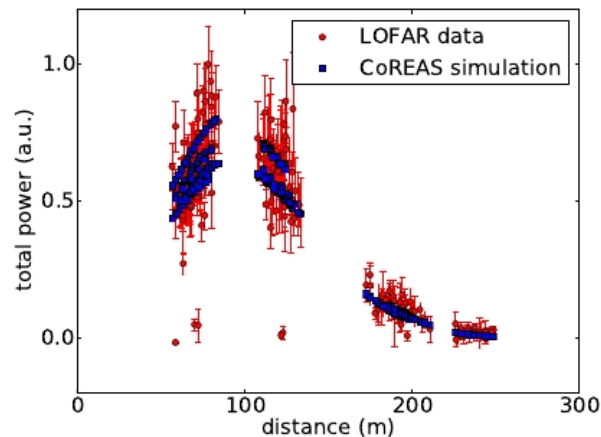
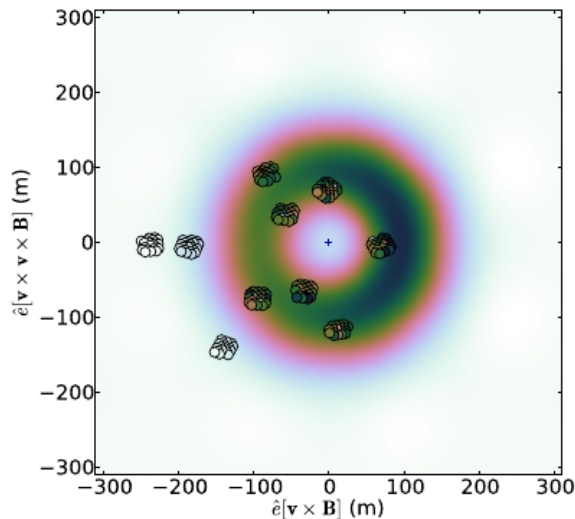
- Signature of a relativistic time compression
- the radio emission is amplified at a specific angle w.r.t. the shower axis
- “Cherenkov ring” has a diameter of about 100 m and is dominant at frequencies above 100 MHz

# Reconstruction of shower parameters



→ Fit applied to HBA data as well

→ amplified ring structure at about 100 m from shower axis



→ ring size sensitive to the depth of the shower maximum

# Conclusions

- Simulation study of the radio emission signal at ground level
  - double Gaussian distribution with minimum 4 free parameters
  - parameters related to shower properties
- The parametrization function has been used to fit ALL LOFAR data
  - good agreement for both LBA and HBA data
- Independent method for measuring the energy, depth of the shower maximum, position of the shower axis at ground level

*A. Nelles et al., 2014,  
Astroparticle Physics, 60, 13-24*

*A. Nelles et al., 2015,  
Astroparticle Physics, 65, 11-21*

*A. Nelles et al., 2015,  
Journal of Cosmology and Astroparticle Physics 05, 018*

