

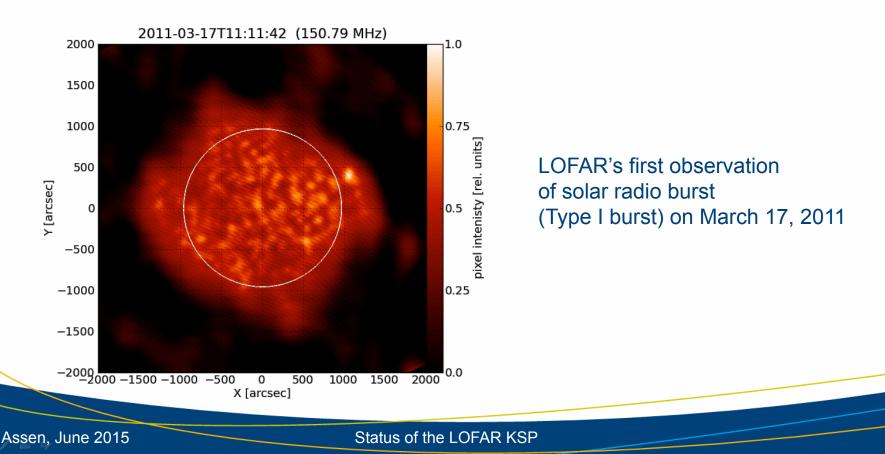
Status of the LOFAR KSP

Solar Physics and Space Weather with LOFAR



Bundesministerium für Bildung und Forschung

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KSP Workshops



1st workshop in Potsdam Oct. 5/6, 2006 2nd workshop in Potsdam June 24/25, 2009 July 5/6, 2010 3rd workshop in Potsdam LOFAR splinter meeting in Graz Sept. 14, 2010 4th workshop in Potsdam Nov. 8/9, 2010 June 28/30, 2011 5th workshop in Aberysthwyth LOFAR splinter meeting in Graz June 5, 2012 6th workshop in Potsdam June 28/29, 2012 7th workshop in Dublin April 24-26, 2013 8th workshop in Dwingeloo May 6-7, 2014 9th workshop in Brussels May 26/27, 2015

http://www.aip.de/groups/osra/german/de_lofar.html

Assen, June 2015

Status of the LOFAR KSP



Management Structure



	Name	Affiliation	Country	
Ы	Prof. Dr. Gottfried Mann	AIP	Germany	
core members	Dr. Mario Mark Bisi Dr. Richard Fallows Dr. Peter Gallagher Dr. Alain Kerdraon Dr. Alexander Konovalenko Dr. Jasmina Magdalenic Dr. Alec McKinnon	Univ. Aberystwyth ASTRON Trinity College Dublin Obs. de Paris-Meudon Institute of Radio Astronomy Royal Obs. of Belgium Univ. Glasgow	N Netherlands Ireland Paris-Meudon France of Radio Astronomy Ukraine bs. of Belgium Belgium asgow UK	
(project manager)	Prof. Dr. Helmut Rucker Prof. Dr. Bo Thide Dr. Christian Vocks	IWF Graz Univ. Uppsala AIP	Austria Sweden Germany	
ordinary members	Dr. Jens Berdermann Frank Breitling Eoin Carley Dr. Harry Enke Dr. Norbert Jakowski Dr. Eduard Kontar Diana Morosan Dr. Jürgen Rendtel Dr. Hamish Reid Pietro Zucca	DLR NeustrelitzGermanyAIPGermanyGrinity College DublinIrelandAIPGermanyDLR NeustrelitzGermanyDLR NeustrelitzGermanyJniv. GlasgowUKTrinity College DublinIrelandAIPGermanyJniv. GlasgowUKFrinity College DublinIrelandAIPGermanyJniv. GlasgowUKFrinity College DublinIreland		32 1(
associated memb.	Dr. Philippa Browning Dr. Bartosz Dabrowski Prof. Dr. Carsten Denker Dr. Lyndsay Fletcher Prof. Dr. Arnold Hanslmeier Dr. Matthias Hoeft Dr. Karl-Ludwig Klein Dr. Christophe Marqué Dr. Hanna Rothkaehl Dr. Astrid Veronig Dr. Alexander Warmuth	Univ. Manchester Royal Obs. of Belgium AIP Univ. Glasgow Univ. Graz Thür. LSW Tautenburg Obs. de Paris-Meudon Royal Obs. of Belgium SRC Warsaw Univ. Graz AIP	UK Belgium Germany UK Austria Germany France Belgium Poland Austria Germany	

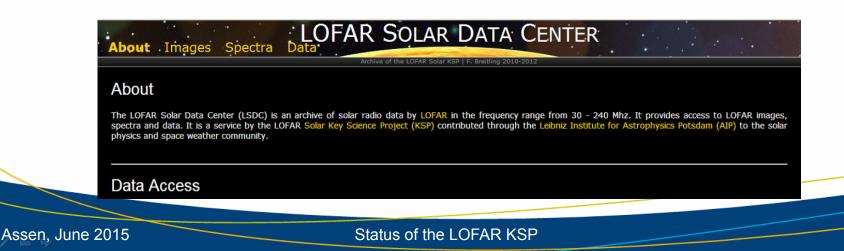
32 members of 10 countries

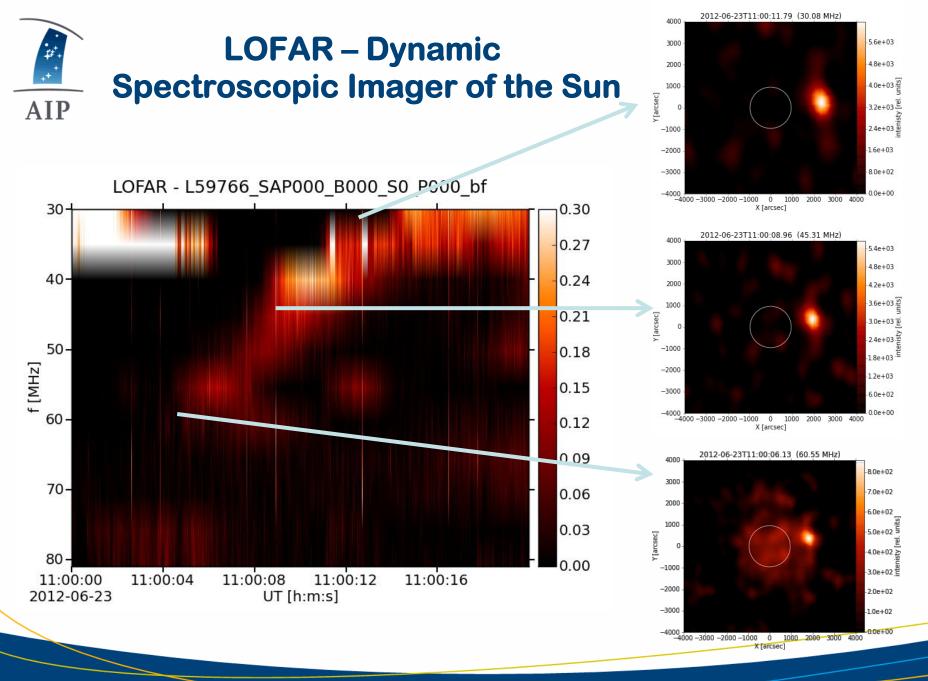






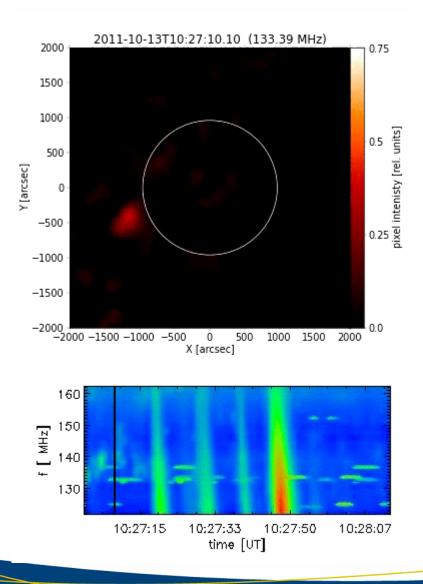
- development of the solar imaging pipeline
 - with self-calibration
 - with an external calibrator
- simultaneous dynamic radio spectra from imaging data
- studying of type III burst
- studying of the ionospheric influences
- submission of regular proposals (C 3 & 4) with 1 envelope sheets
- establishment of the LOFAR Solar Data Center (LSDC) at AIP: http://lsdc.aip.de







Type III Bursts (HBA Range)



LOFAR can really work as a dynamic spectroscopic radio imager of the Sun.



Cycle 3 & 4



Cycle 3: LC3_002, PI R. Fallows, "Monitoring Ionospheric Scintillation above LOFAR", 1600 hours (piggy-back) requested, awarded time in gaps between regular LOFAR observations LC3 005: PI D. Morosan, "Imaging Coronal Mass Ejections with LOFAR", 0 hours LC3 012: PI E. Kontar, "Quantitative diagnostics of electron beams in the solar corona", 20 hours LC3 021, PI G. Mann, "Radio emission from coronal shock waves", 48 hours Cycle 4: LC4 001: PI C. Vocks, "Quiet Sun studies with LOFAR", 24 hours LC4 002: PI G. Mann, "Radio emission from coronal shock waves and nascent CMEs", 24 hours LC4 006: PI G. Mann, "Energetic electrons generated in solar flares", 17.5 hours LC4 016: PI E. Kontar, "Fine structures in solar radio emission", 40 hours



Problems Which Should be Discussed



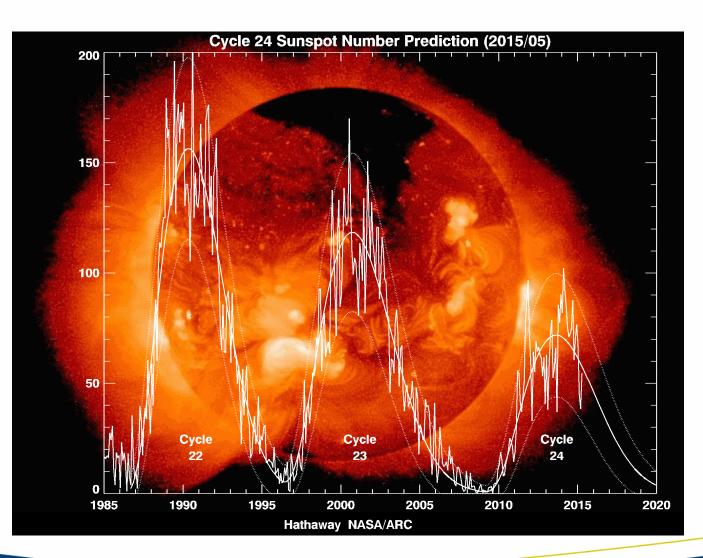
- LOFAR observations of the Sun
 - The Sun is a special target !!!
 - regularly observations of the Sun by spacecraft Hinode, RHESSI, and SDO
 - * We must wait up to the Sun provides us an event, which we can scientificly study.
 - \rightarrow We ask for a long term proposal.
 - external triggering (\rightarrow solar activity)
 - * alert time to ASTRON (3 days or 1 hour) ???
 - * Could a single LOFAR station be used only for solar observations i.e. as a spectrometer ?
 - * potential triggers !!!
 - Would it be possible to record regularly an image of the Sun
 - e.g. every minute ?

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Status of the LOFAR KSP



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Thank you for your attention!



This work was done in collaboration with the solar KSP and LOFAR/ASTRON team

Status of the LOFAR KSP



Allocation Time I



8740 h (365 x 24 h) are available per year

- further commissioning (33 %)
- maintenance (7 % = 612 h)
- director's discretionary time (5 % = 437 h)
- available for science (88 %)

for the first semester

- open skies access (5.5 %)
- NL reserved access (31.9 %)
- GLOW reserved access (9.35 % = 818 h) \rightarrow 409 h per semester \rightarrow 81 h per station
- France, Sweden, UK (2.75 %)

in reality: 48 h solar observations in the first semester



Allocation Time



in future

	NL	DE	PL	FR, IT, SE, UK	Open
				each	Skies
Year 1	0.49	0.14	0.09	0.04	0.12
Year 2	0.44	0.13	0.08	0.04	0.19
Year 3	0.36	0.10	0.07	0.03	0.35
Year 4	0.27	0.08	0.05	0.02	0.52
Year 5	0.19	0.06	0.04	0.02	0.63

365 d x 24 h = 8760 h per year (4380 h per semester)

German time: DE Potsdam time: DE/5

example: DE = $0.1 \rightarrow 88$ h per semester should be sponsored time for the solar KSP



Funding



- by German government: D-LOFAR I + II
 - 400 k€ for 2 x 3 years (Frank Breitling) 07/2009 - 06/2014
 - D-LOFAR III
 230 k€ for 3 years (Frank Breitling)
 07/2014 06/2017
- by Irish government:
- Diana Morosan
- Pietro Zucca
- by UK government:
- Hamish Reid
- by AIP operation costs:
- ≈ 80 k€ for ASTRON
- ≈ 35 k€ for costs
- <u>≈ 15 k€ data link</u>
- ≈ 120 k€ in total per year
- 1 staff position (Christian Vocks LOFAR scientist at AIP)



Cycle 0 & 1



Cycle 0:

LC0_027, PI G. Mann, "Solar activity studies with LOFAR", 72 h requested, 47 h awarded LC0_030, PI G. Mann, "LOFAR studies of the evolution of coronal mass ejections in the heliosphere", 24 h requested, 24 h awarded

Cycle 1:

LC1_012, PI P. Gallagher, "LOFAR Imaging and Spectroscopy of Quiet Sun Variability", 24 h requested LC1_022, PI G. Mann, "Energetic electron production in solar flares", 36 h requested LC1_025, PI G. Mann, "Coronal mass ejections and type II radio bursts", 24 h requested

LC1_044, PI. J. Magdalenic, "Propagation of Electron Beams in the Solar Corona and Associated Type III Radio Emission as Seen by LOFAR", 36 h requested

Awarded time for LC1_022, LC1_025, and LC1_044 under one envelope sheet is 48 h

LC1_057, PI. R. Fallows, "Observing the Interplanetary Magnetic Field with LOFAR", 110 h requested, 39 h awarded







Cycle 2: LC2_001, PI R. Fallows, "Investigating Refraction Through the Solar Wind using Intensity and Phase Scintillation", 5 h LC2_002, PI R. Fallows, "Probing a Coronal Mass Ejection with Scintillation Arcs", 12 h LC2_003, PI R. Fallows, "Monitoring Ionospheric Scintillation above LOFAR", 1600 h (piggyback mode) LC2_027, PI G. Mann, "Energetic electron propagation in solar flares", 14.5 h LC2_028, PI G. Mann, "Solar coronal mass ejections", 35 h LC2_031, PI C. Marque, "Solar noise storms in the decametric and metric range: a study with LOFAR and the Nançay Radioheliograph", 22 h LC2_033, PI J. Magdalenic, "Observations of solar type II radio bursts by LOFAR", 16 h LC2_035, PI H. Reid, "The LOFAR Quiescent Sun", 16 h

total: 120.5 h (1600 h piggyback mode)