

Netherlands Institute for Radio Astronomy



LOFAR Users meeting 4 April 2016

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ASTRON is part of the Netherlands Organisation for Scientific Research (NWO)

# **CITT project overview**



#### CITT 1

Sept. 2013 – Sept. 2015

Work streams:

- DPPP calibration
- AWImager
- Generic pipeline framework
- Selfcal pipeline

CITT 2

Oct. 2013 – Sept. 2017

Work streams:

- bring results of CITT 1 to the user
- DPPP calibration
- AWImager / GPU Gridder
- Factor / HBA calibration

- LBA calibration

+ maintenance of DPPP / BBS / AWImager / PyBDSM

Advisory group for feedback on priorities



- October 2015: Start of project
- 9 11 December 2015: Factor Busy Week 23
- 25 29 January 2016: Surveys Workshop on Facet Calibration
- 22 25 February 2016: Factor Busy Week 24
- 29 Feb 4 March 2015: LBA expert workshop at Leiden
- Before summer 2016 Factor Busy Week 25 (to be planned)

# **DPPP developments**

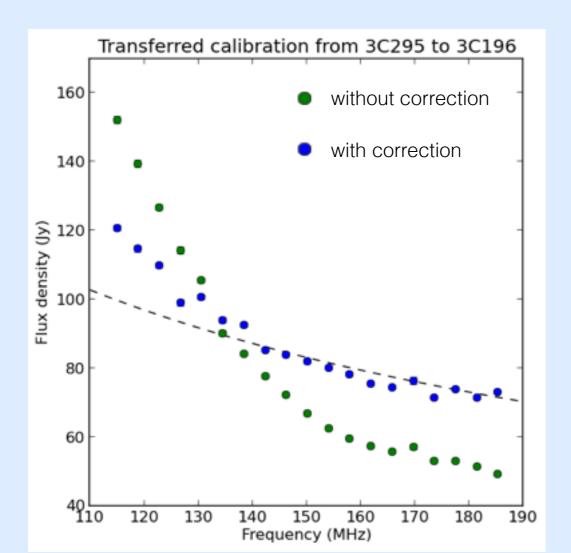


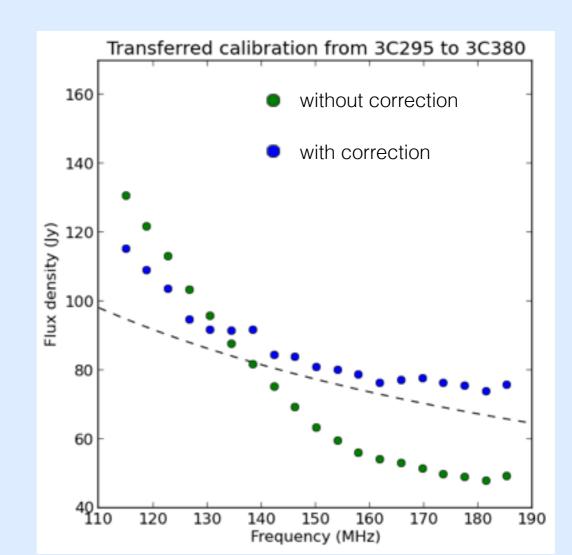
- Plan: replace BBS calibration with DPPP calibration where possible
- Now in DPPP:
  - GainCal
    - Full Jones / diagonal / phase only / scalar phase
    - With or without beam model
  - ApplyCal
  - ApplyBeam (in phase center)
  - Predict / subtract / add
  - Work on DPPP now:
    - Solve per group of channels
    - Fitting phase slope

### **Beam normalization**



- Updated HBA beam model was made by George Heald & TJD
- Incorporates normalization for power over the hemisphere
- Currently (still) in testing, <u>not</u> implemented in the production beam model (DPPP, BBS, AWImager, WSClean, ...)





# **AWImager and Image Domain Gridder**

Development on AWImager has focused on Image Domain Gridder

**AST**(RON

- Image Domain Gridder:
  - Speedup of  $\sim 10 \times$  in gridding and degridding
  - Uses GPUs
  - Works well with A-term / phase screens
- AWImager
  - Uses Image Domain Gridder
  - Uses CASA for deconvolution
  - Building on top of CASA code is not supported, hard to maintain (but works)
- Alternative currently investigated:
  - WSClean + Image Domain Gridder



# Facet Calibration paper by Van Weeren et al: <u>http://arxiv.org/abs/1601.05422</u>, accepted for publication in ApJS

See talk by Andreas Horneffer tomorrow!

Prefactor	Initial subtract	Prepare facets	Selfcal per facet	Mosaic
Calibration: - Flux calibration - Clock / TEC - Flagging soln's - Diagnostic plots	Image at high resolution Subtract high resolution model Image at low resolution Subtract low resolution model Merge low- and high-res models	Find calibrators, make tesselation	Add calibrators to facet, do selfcal on full bandwidth with heavy averaging - 2x phase only - 2x amp+phase Add all facet sources Image at high resolution (1.5") Subtract facet sources	

#### **Prospects HBA**



- Hardest bit in Factor is the facet self-calibration.
- The direction independent calibration is more or less stable.

The CITT thinks direction independent calibration ('pre-factor') and imaging ('initial subtract') should become available on CEP4.

Requires adjustments on CEP4: "generic pipeline pilot"



Experts now all together in Leiden (Huib + students, Franceso, Leah, ...)

CITT facilitates and collaborates where possible - e.g. by implementing DPPP features

One calibration option for LBA is Factor



- Factor 'almost there', will push to observatory / CEP4
- LBA: work in progress
- Imager: lots of good stuff, not there yet
- DPPP: done for main scenarios, working on more scenarios