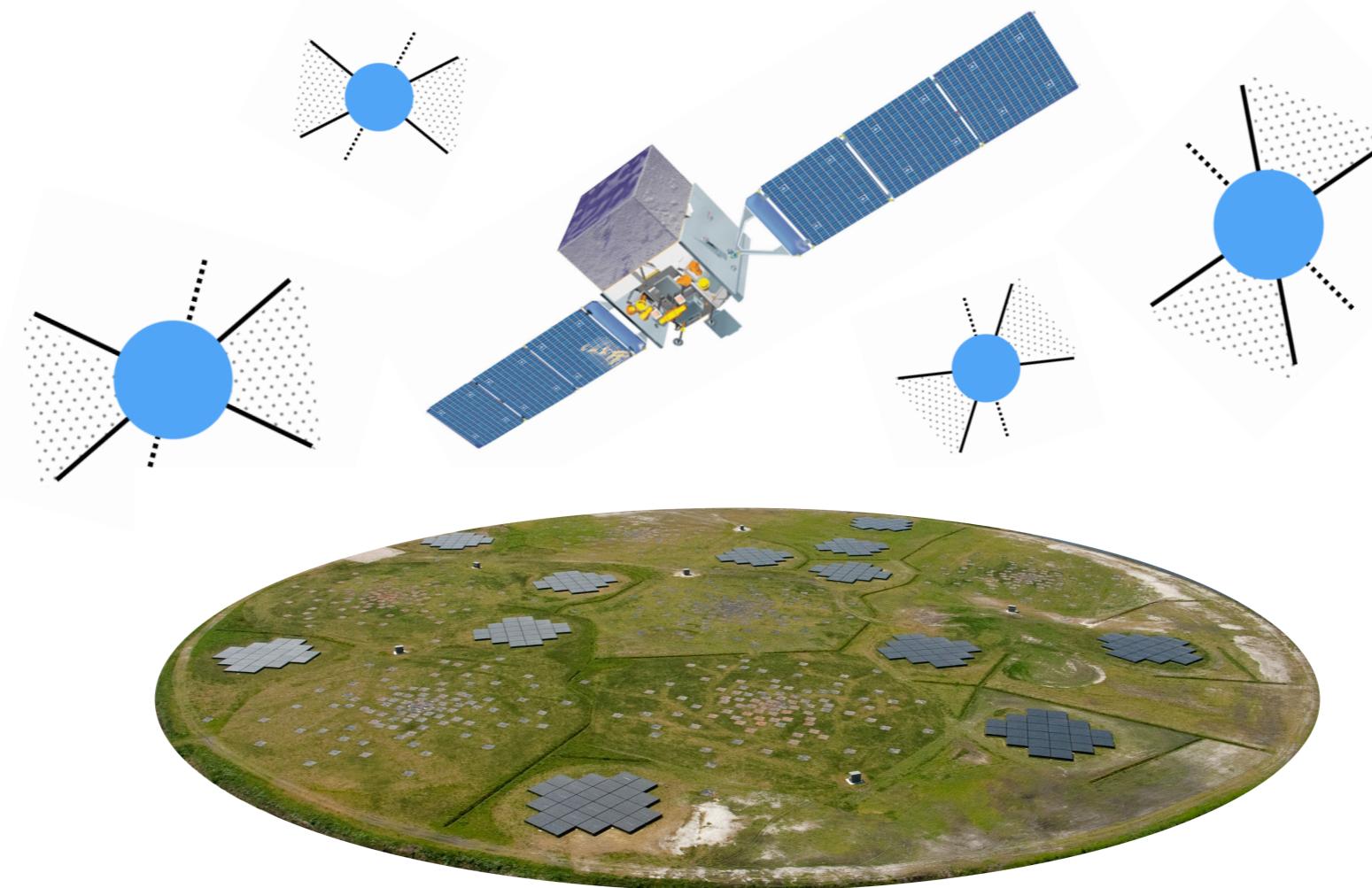


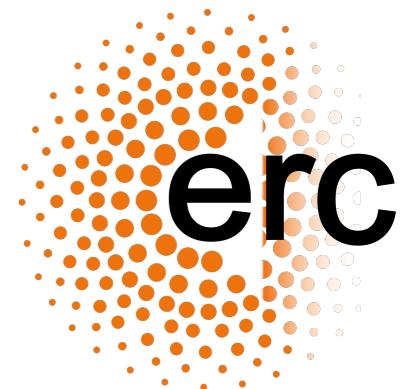
# LOFAR survey for millisecond pulsars in *Fermi* gamma-ray sources



**Ziggy Pleunis (API, UvA)**

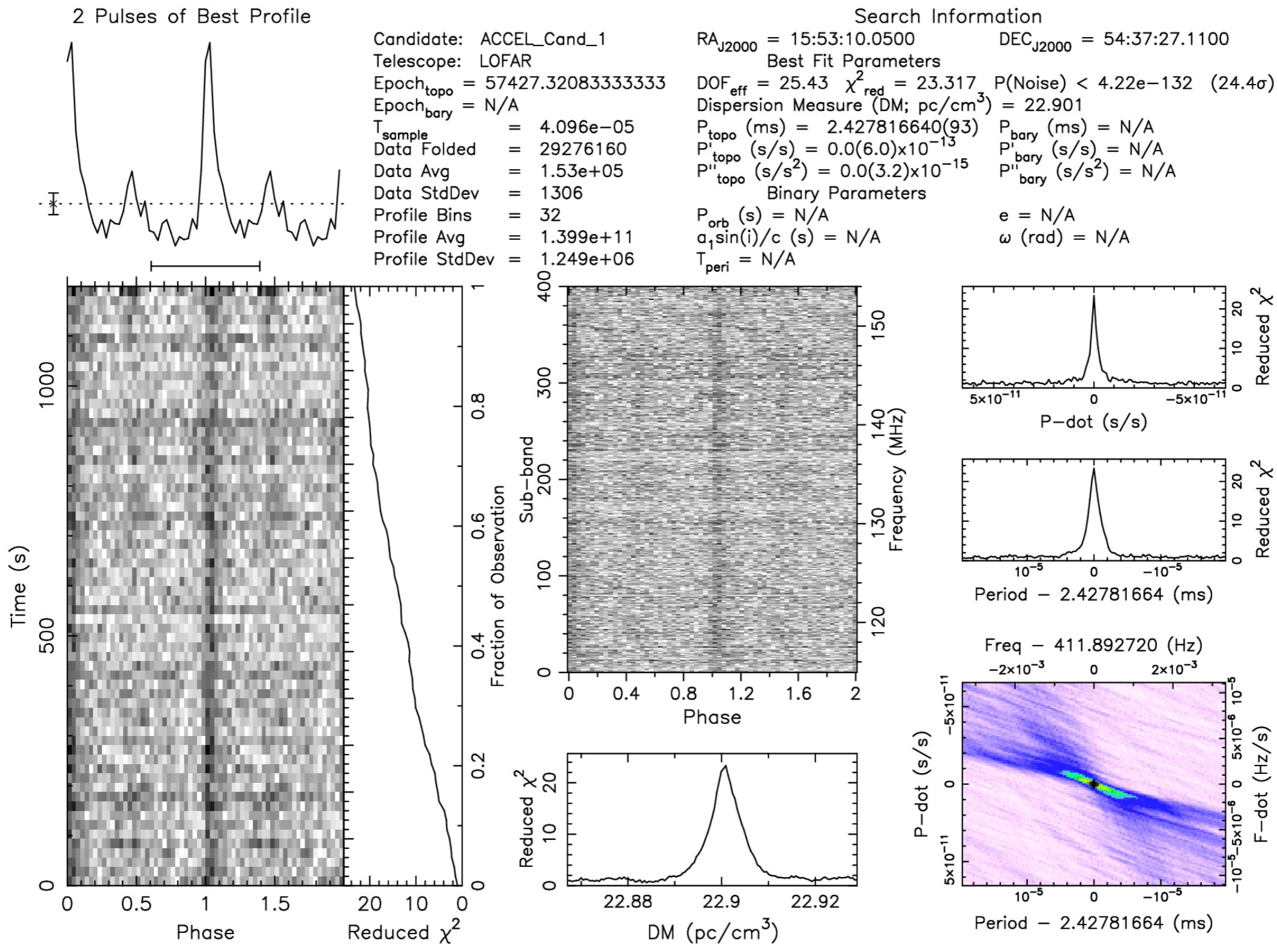
in collaboration with **Jason Hessels, Cees Bassa,  
Vlad Kondratiev, Sotiris Sanidas, Daniele Michilli  
and Alexander van Amesfoort**

LOFAR Science Workshop, April 6, 2016



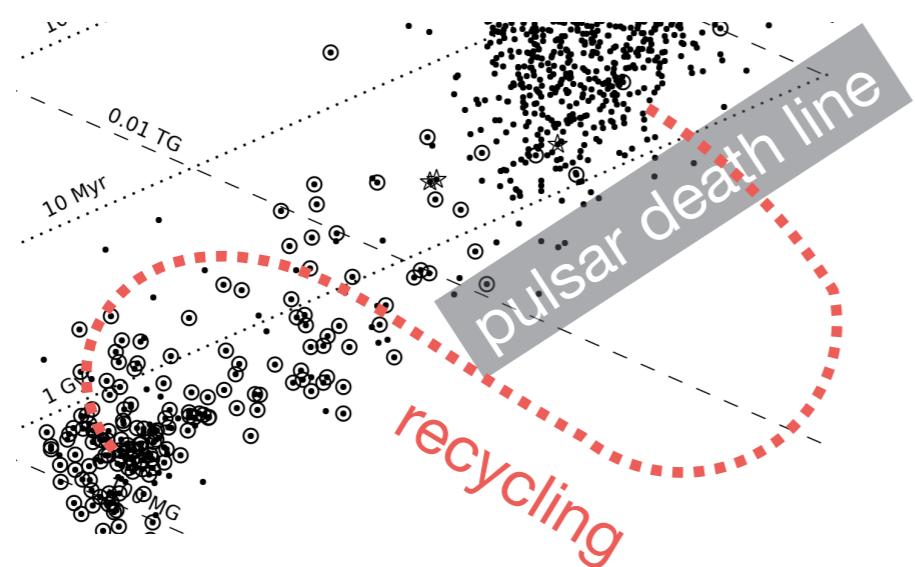
# First LOFAR millisecond pulsar discovery!

Pleunis et al., in prep.

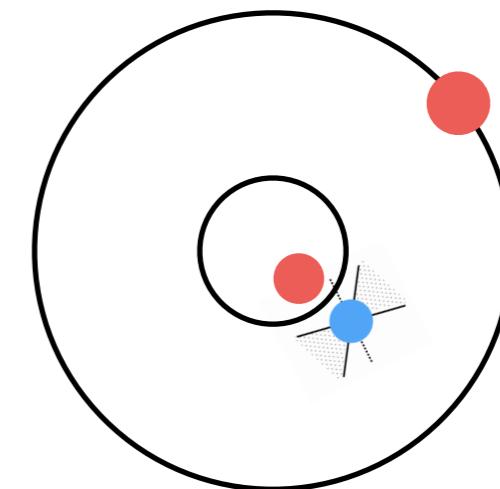


# Survey for millisecond pulsars

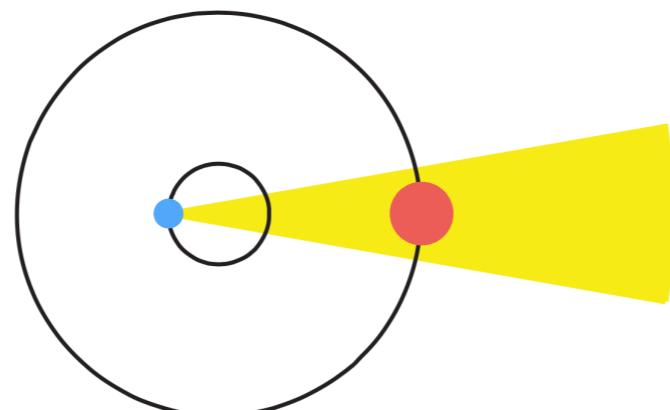
Binary Stellar Evolution  
e.g. Tauris (2011)



Exotic new  
Millisecond Pulsars  
e.g. Ransom et al. (2014)

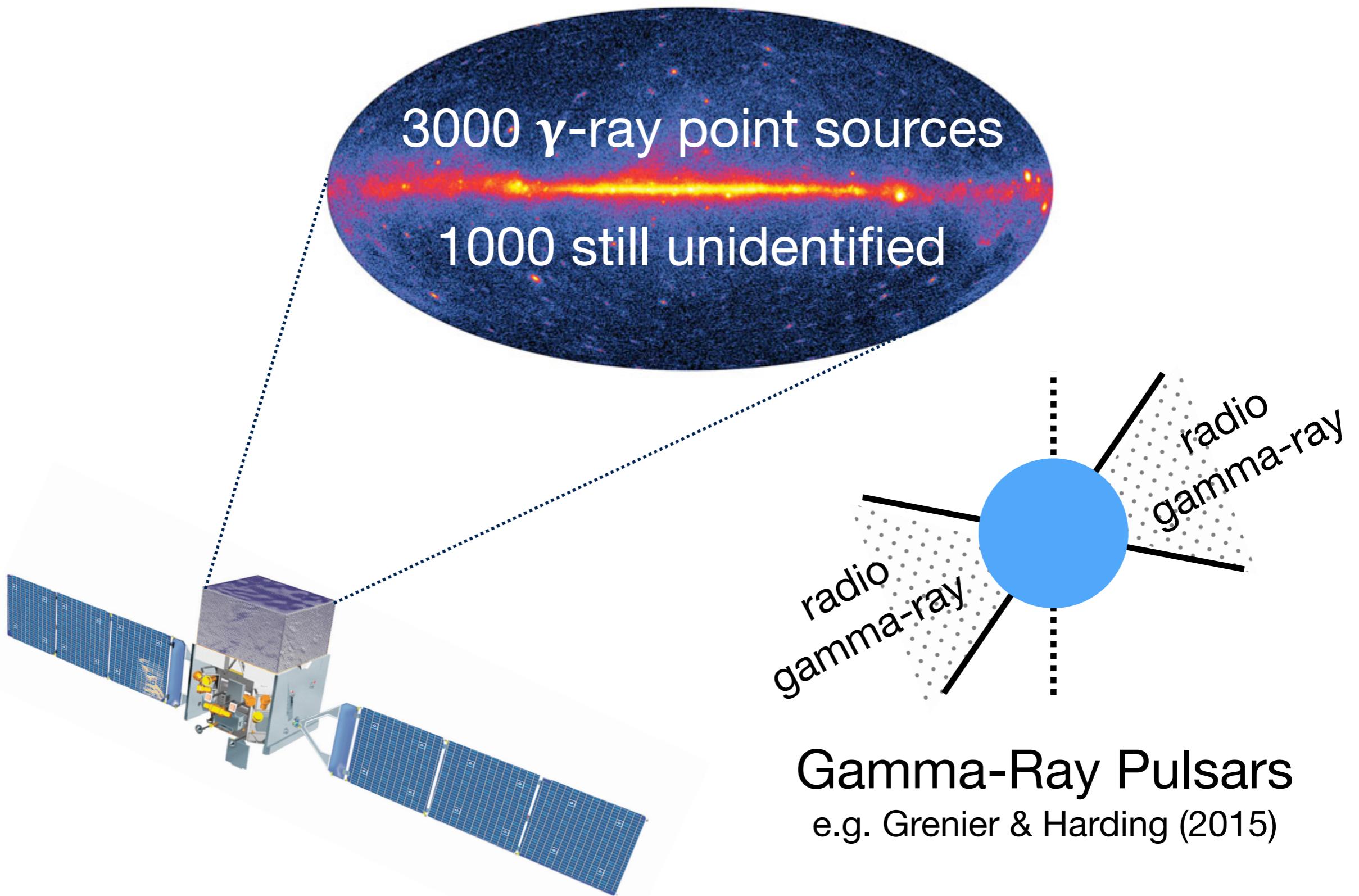


The Neutron Star  
Equation of State  
e.g. Demorest et al. (2010)

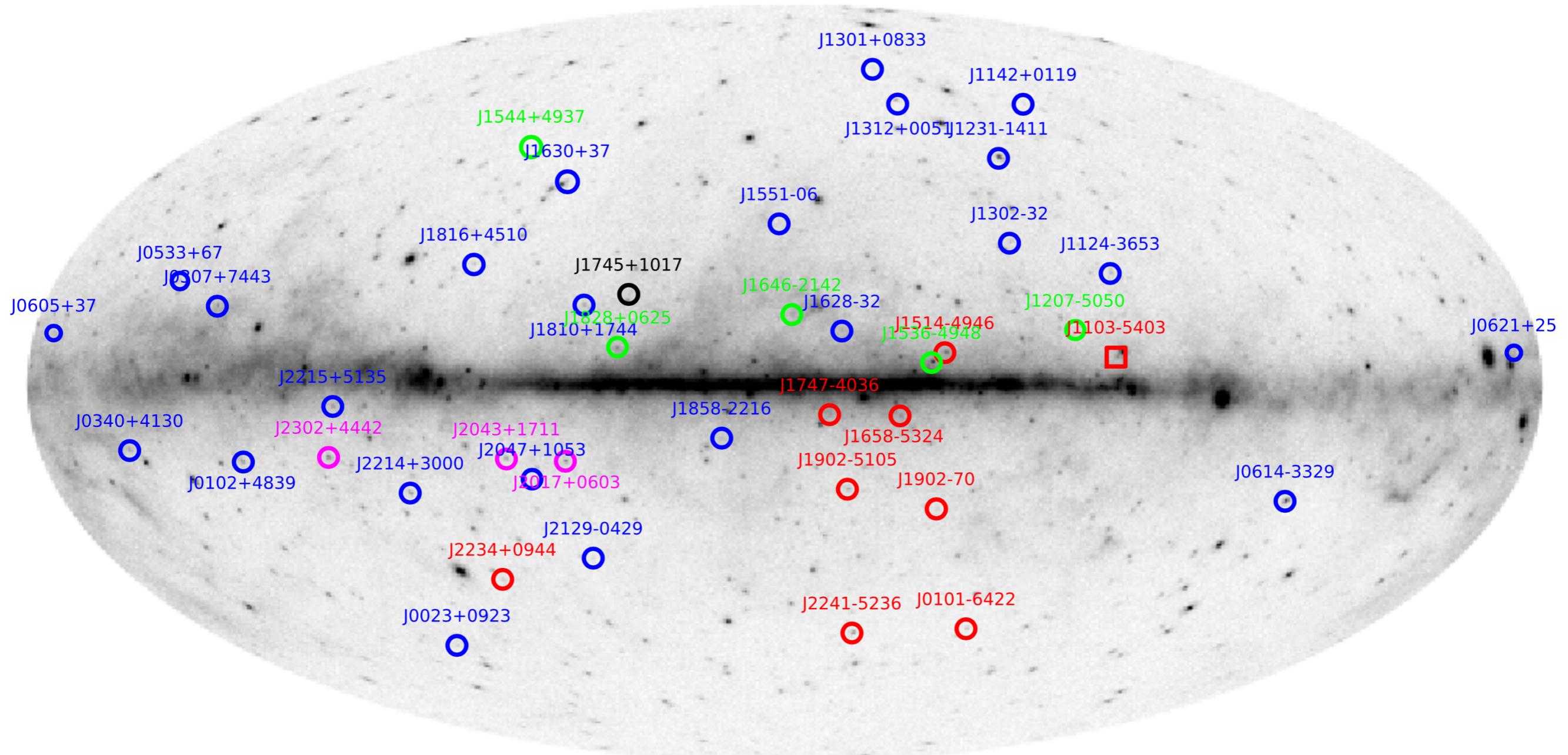


Shapiro  
delay

# *Fermi* unidentified sources



# *Fermi* unidentified sources

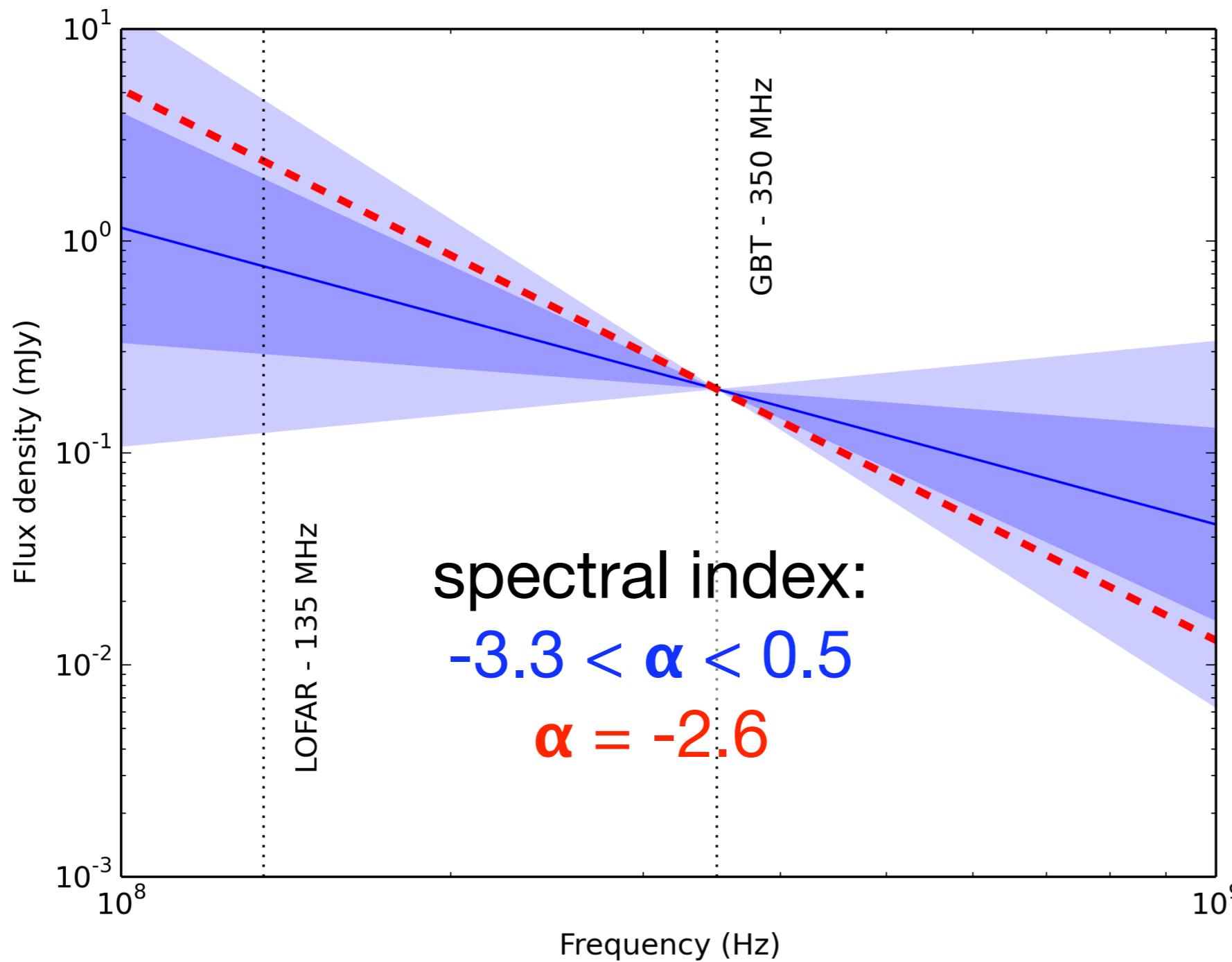


GBT, Parkes, GMRT, Nançay, Effelsberg  
Ray et al. (2012)

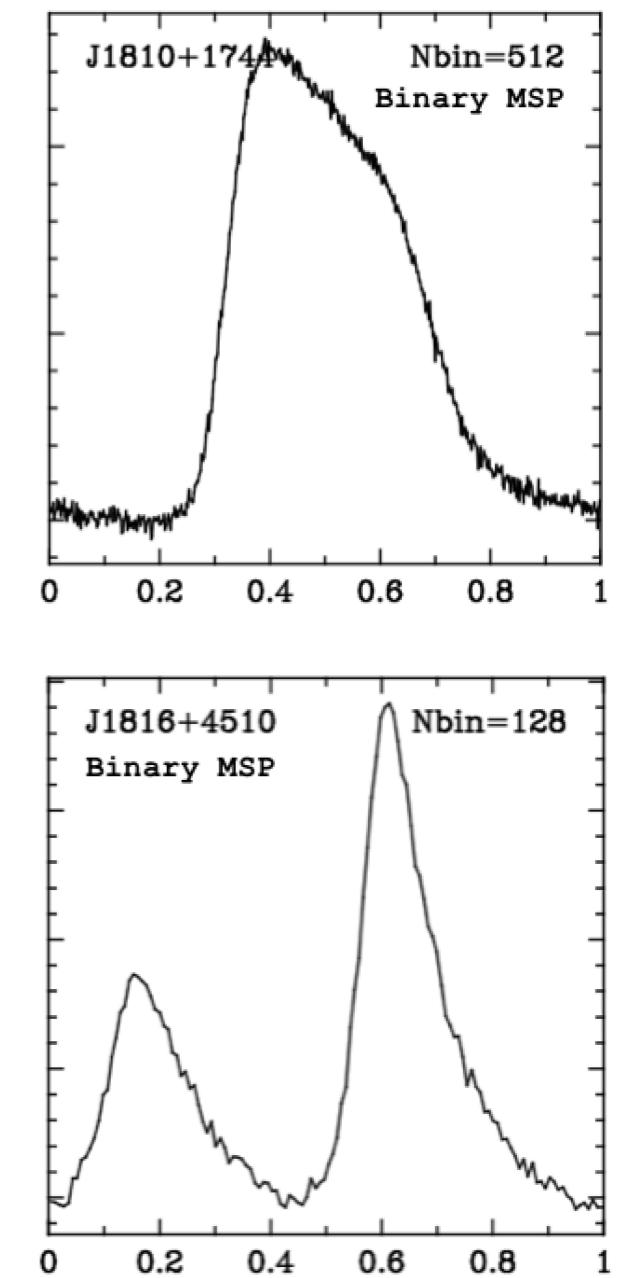
# LOFAR advantages

pulsars can have very steep spectra

Bates et al. (2013)



LOFAR profiles  
Kondratiev et al. (2015)

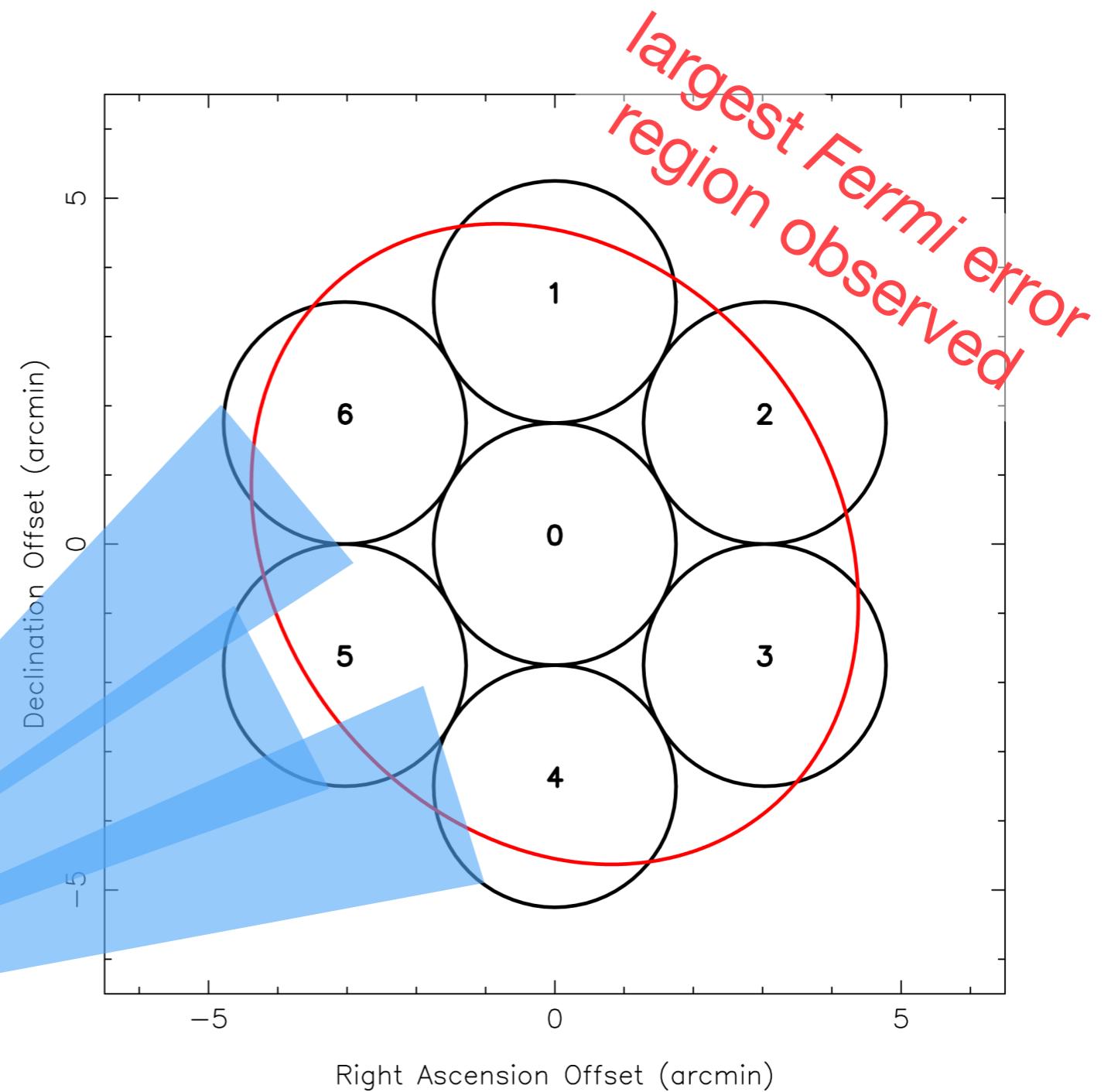
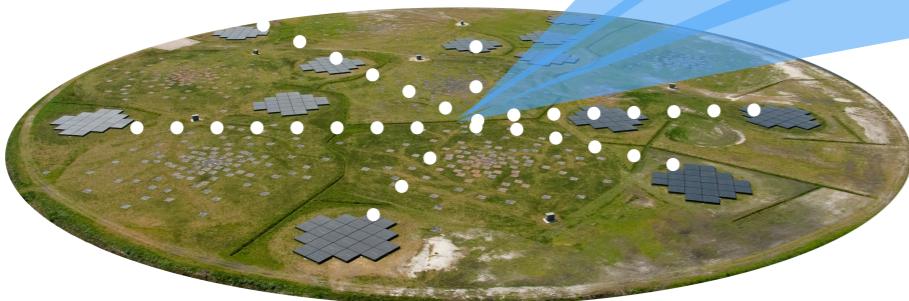


# Survey setup

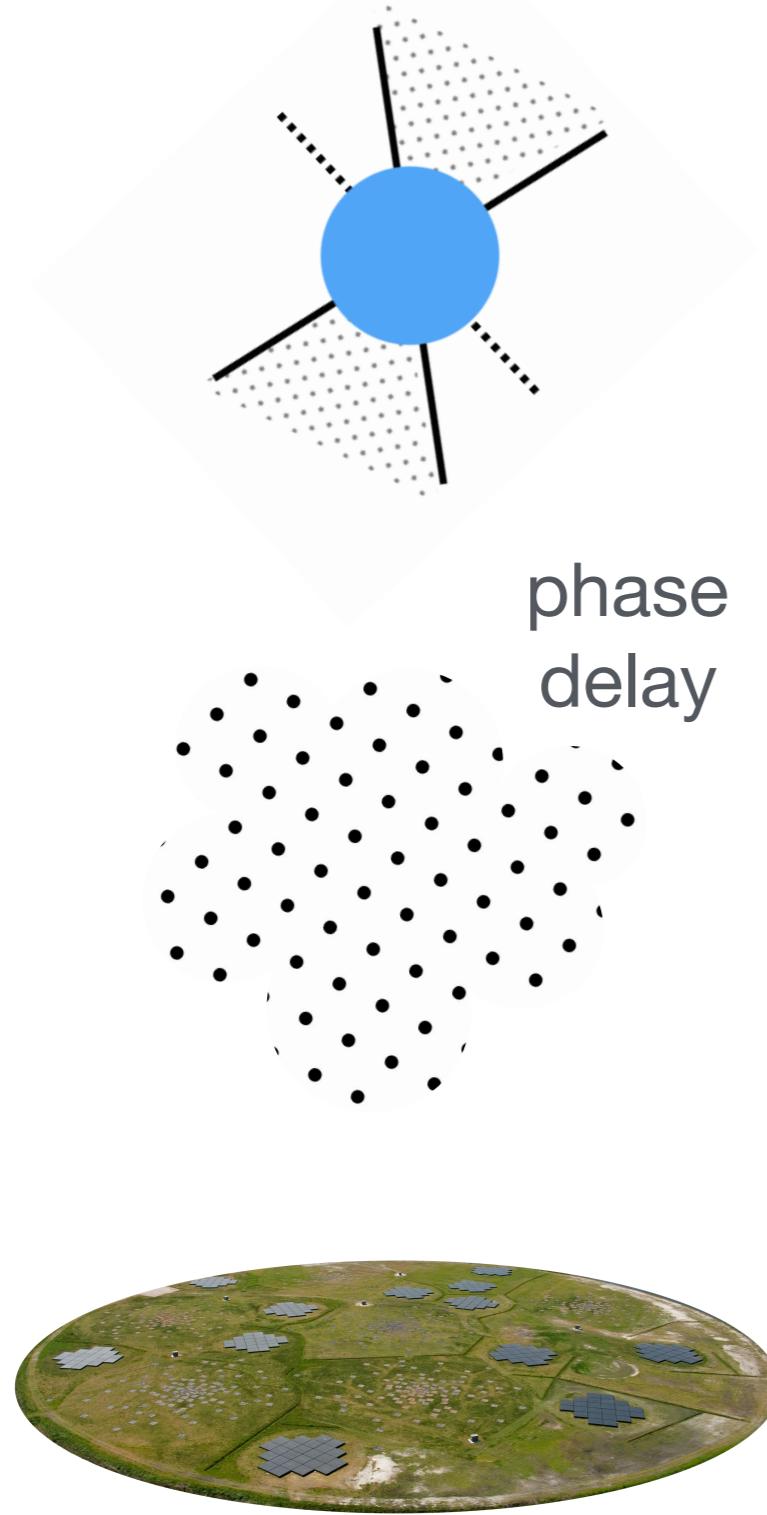
need complex voltage data;  
want maximum data rate

48 sources  
(galactic longitude  $> 10^\circ$ )  
+ 2 test pulsars

21 LOFAR  
Core stations



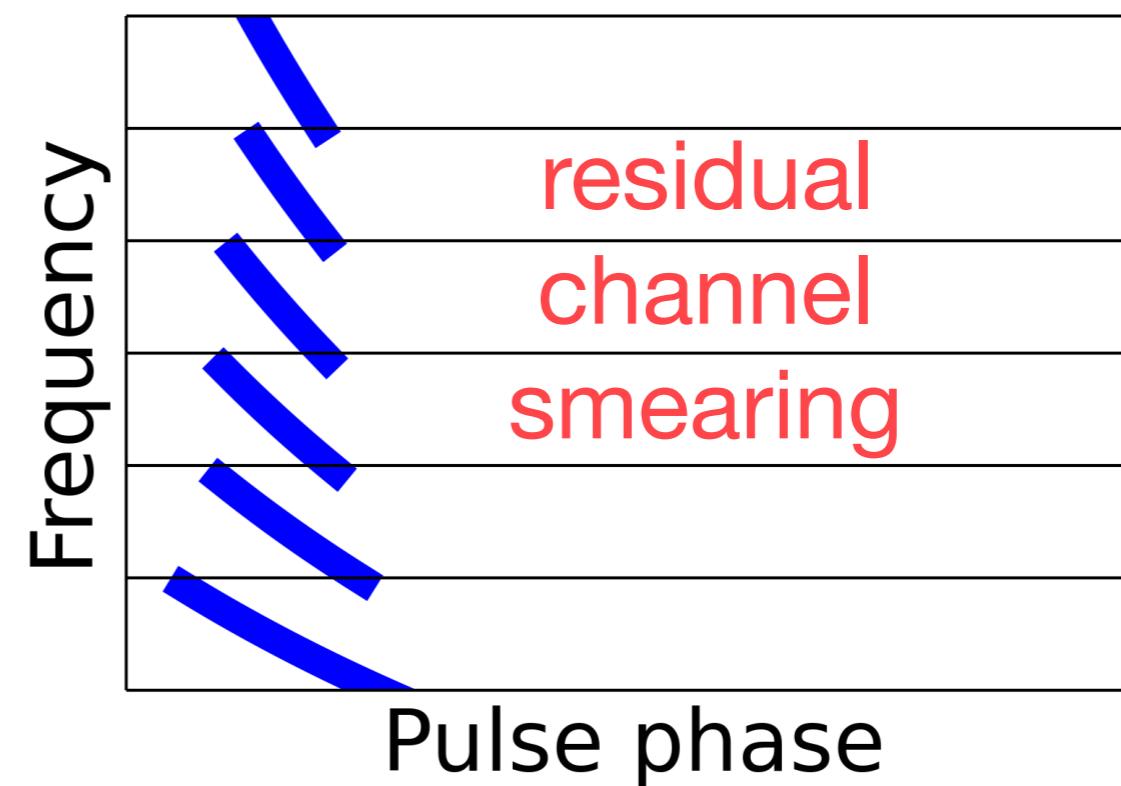
# LOFAR challenge: correcting for the ISM



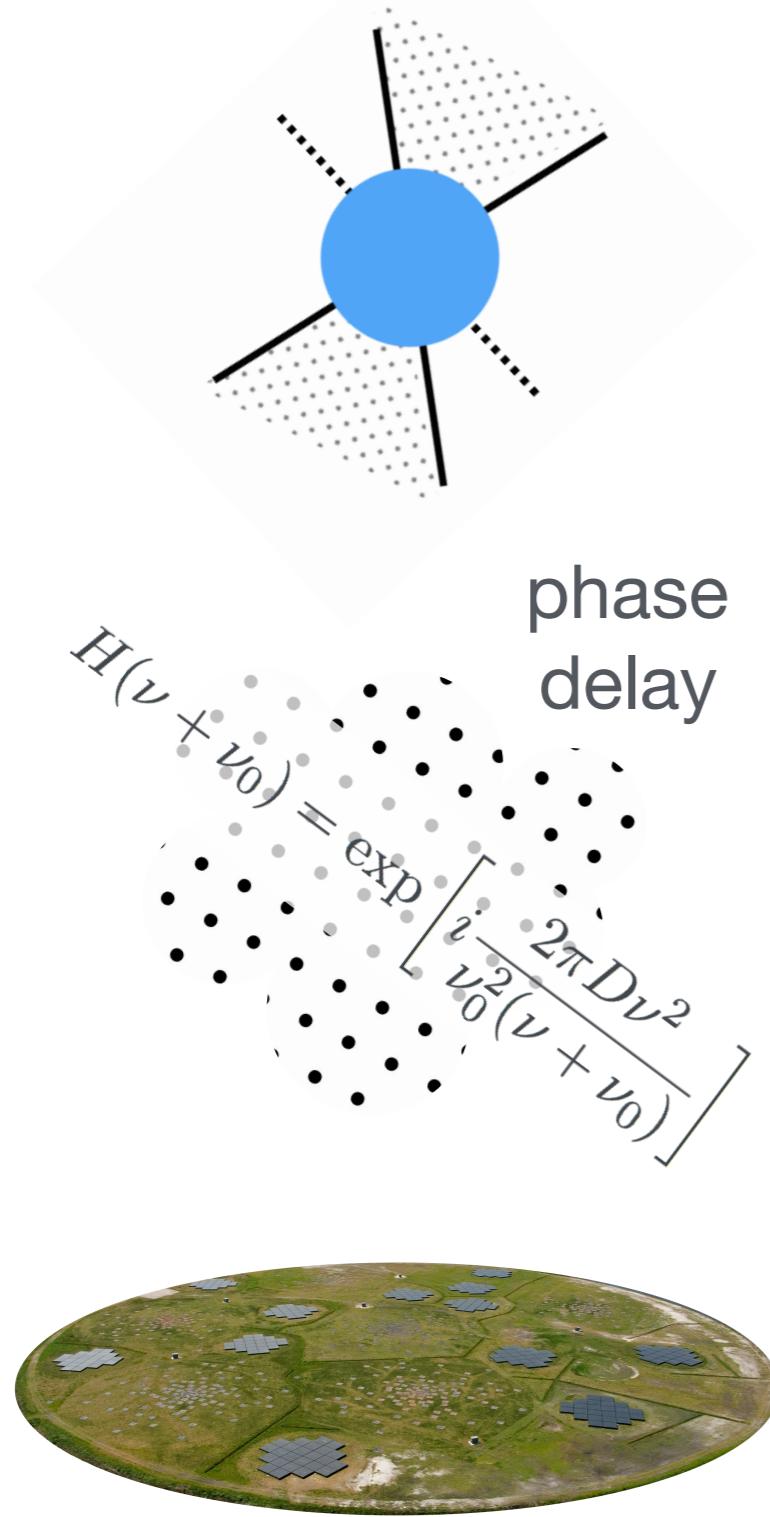
dispersive delay  
*a priori* unknown

$$\tau_{\text{DM}} \propto \text{DM} \nu^{-2}$$

Incoherent dedispersion:



# LOFAR challenge: correcting for the ISM



dispersive delay  
*a priori* unknown

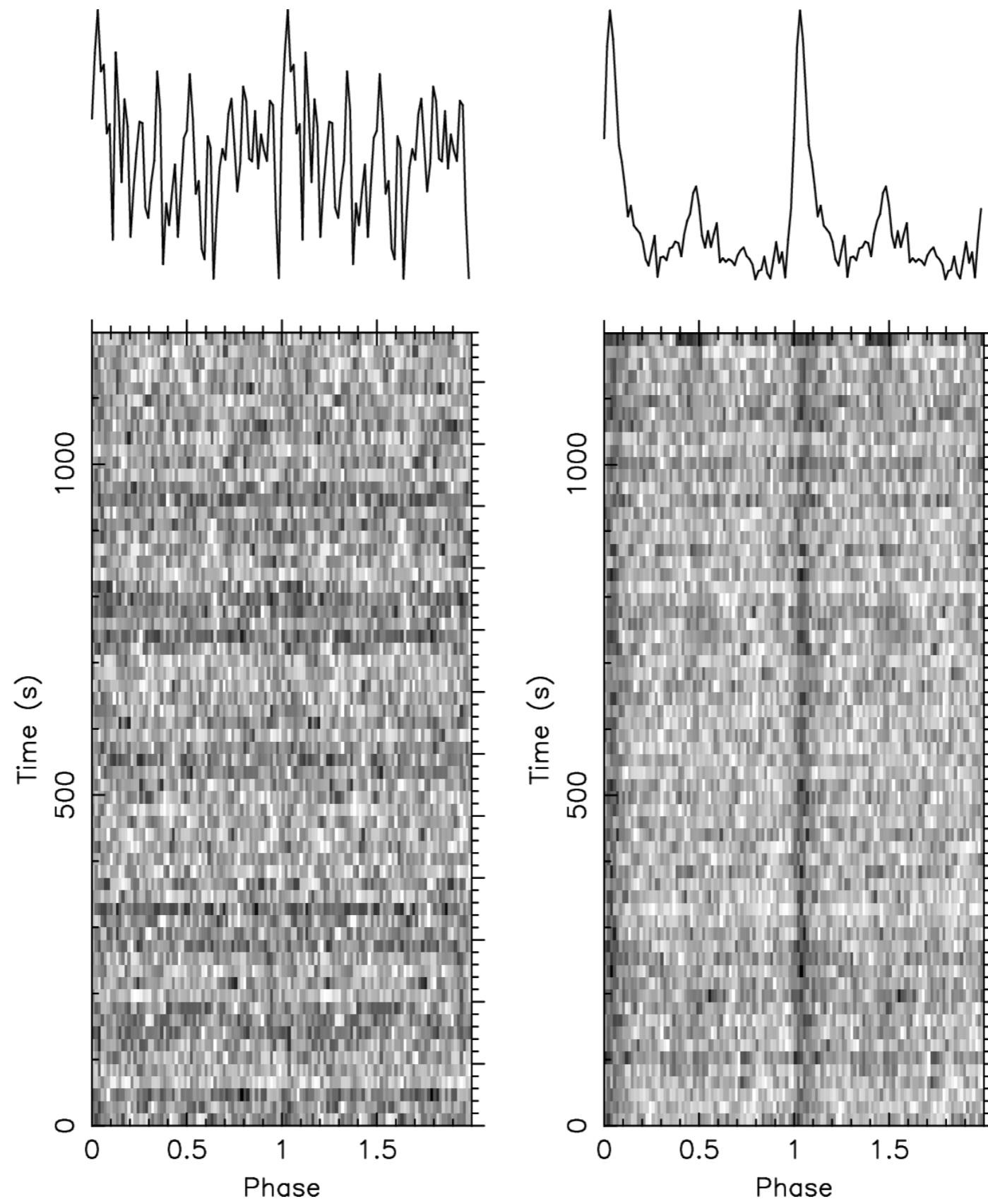
$$\tau_{\text{DM}} \propto \text{DM} \nu^{-2}$$

Coherent dedispersion:

e.g. Hankins (1971), Hankins & Rickett (1975),  
van Straten (2003)

“the convolution of raw signal voltages with the inverse of the ISM transfer function”

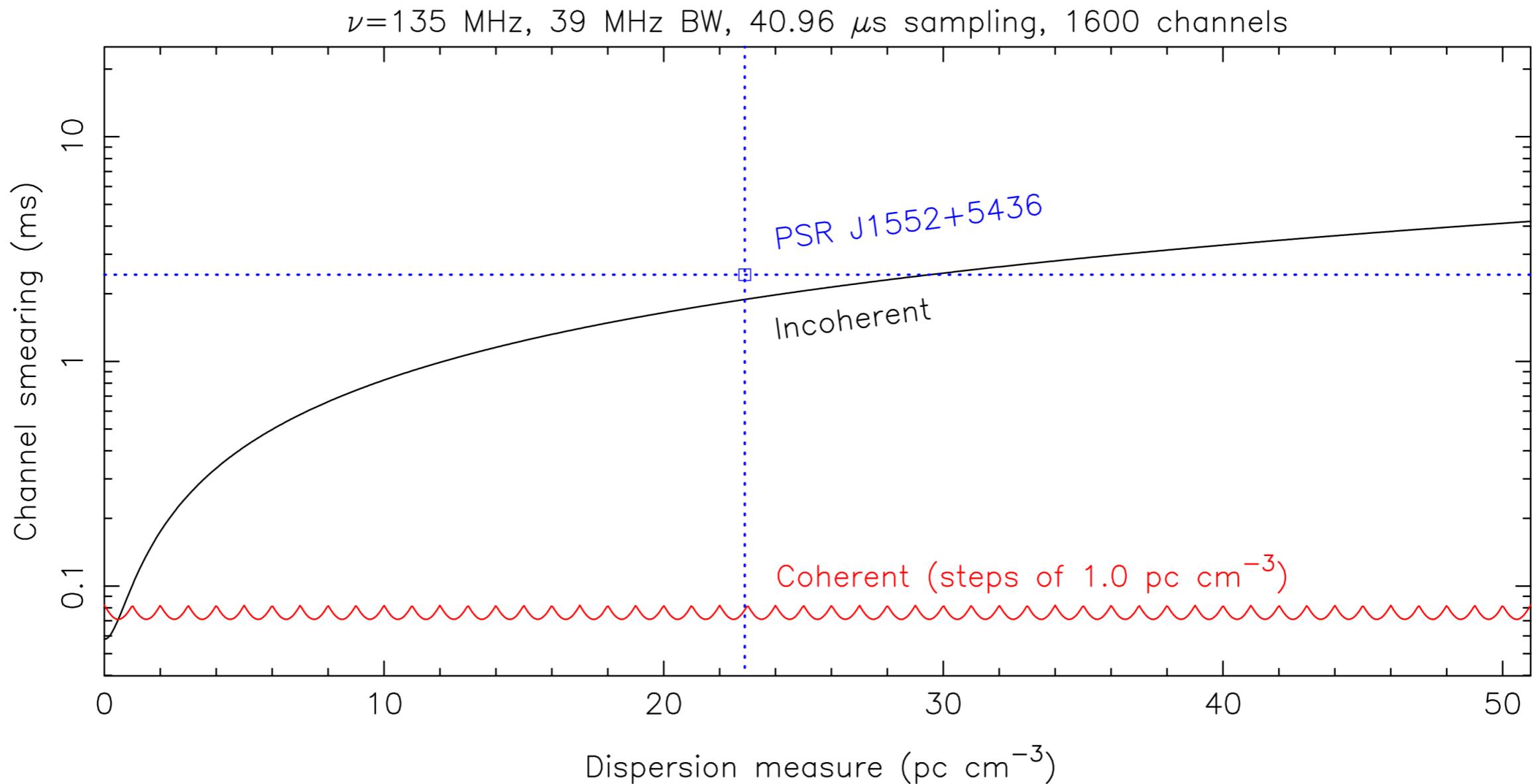
# Incoherent vs. Coherent dedispersion



# LOFAR (semi-)coherent dedispersion search

Bassa et al., in prep.

coherent DM trial step size =  $1 \text{ pc cm}^{-3}$   
incoherent DM trial step size =  $0.002 \text{ pc cm}^{-3}$

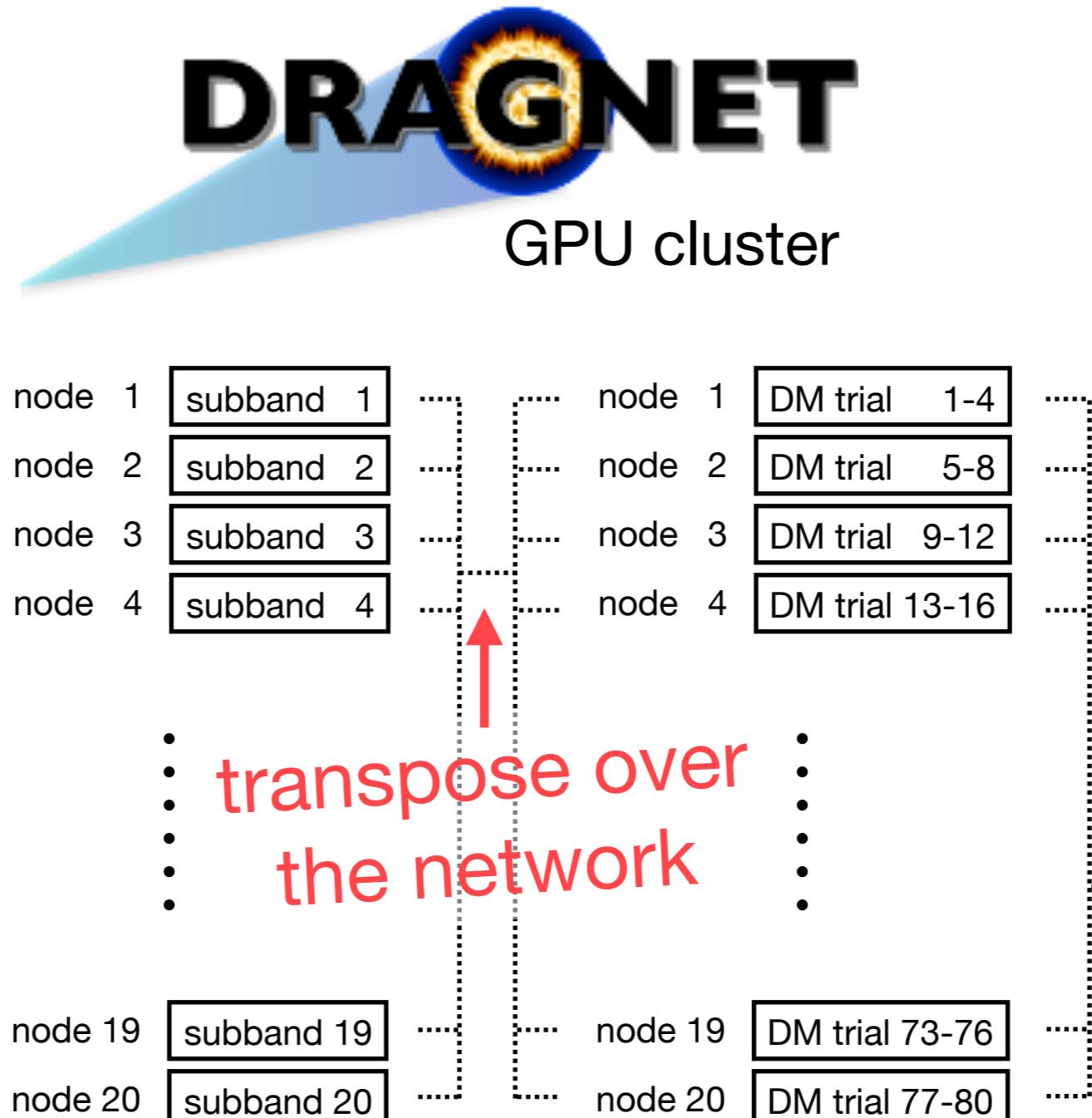


# LOFAR (semi-)coherent dedispersion search

Bassa et al., in prep.

4 GPUs +  
16 CPUs  
per node

5.5 TB  
raw data per  
observation

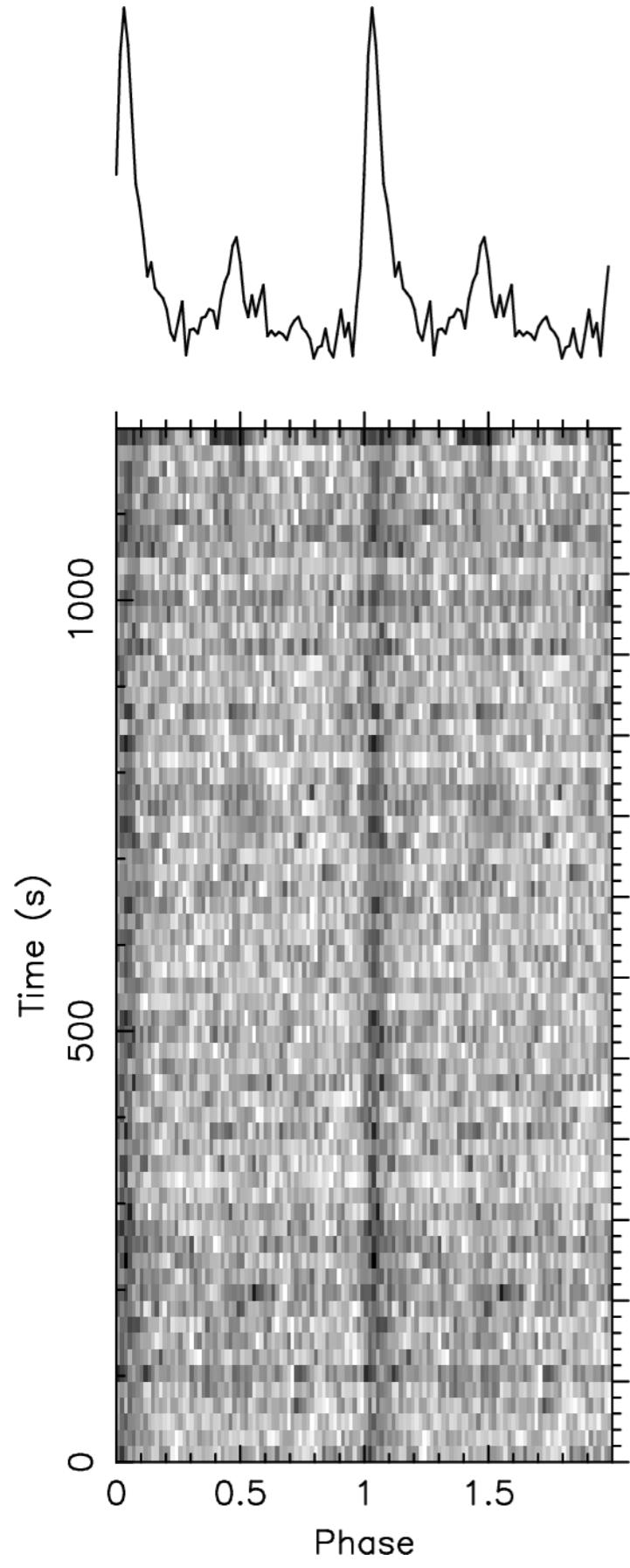


after 24  
hours of  
processing  
with  
up to 0.5  
petaFLOPS

pulsar  
candidates

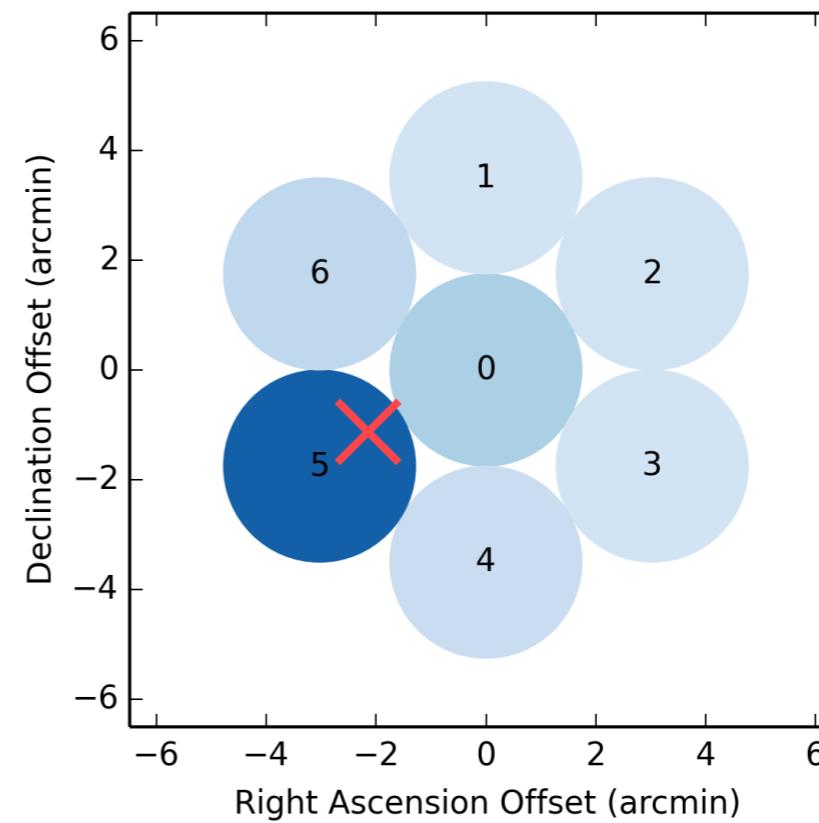
# PSR J1552+5436

Pleunis et al., in prep.



*first* LOFAR millisecond pulsar  
*first* aperture array millisecond pulsar  
lowest observing frequency  
at which an MSP is discovered

$p_s = 2.43$  ms;  $DM = 22.9$  pc cm $^{-3}$ ;  $d = 1.225$  kpc  
isolated or in a long period binary



# Future prospects

PSR J1552+5436

non detections using Lovell and  
Nançay at L-band

LOFAR DDT proposal approved:  
start timing and measure spectrum

millisecond pulsar discovered at the  
lowest observing frequency

LOFAR survey for  
*Fermi* gamma-ray  
sources



GPU cluster



Pulsar surveys with SKA-low for ultra-steep spectrum pulsars

