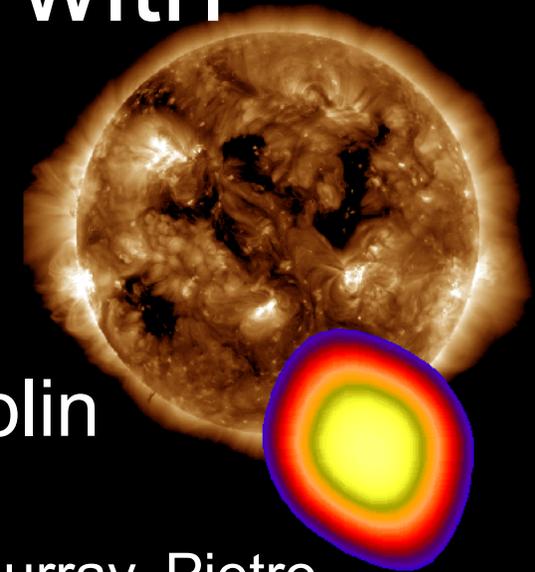


# Imaging Radio Shock Signatures of Solar Coronal Mass Ejections with LOFAR

Diana E. Morosan

University of Helsinki, Trinity College Dublin



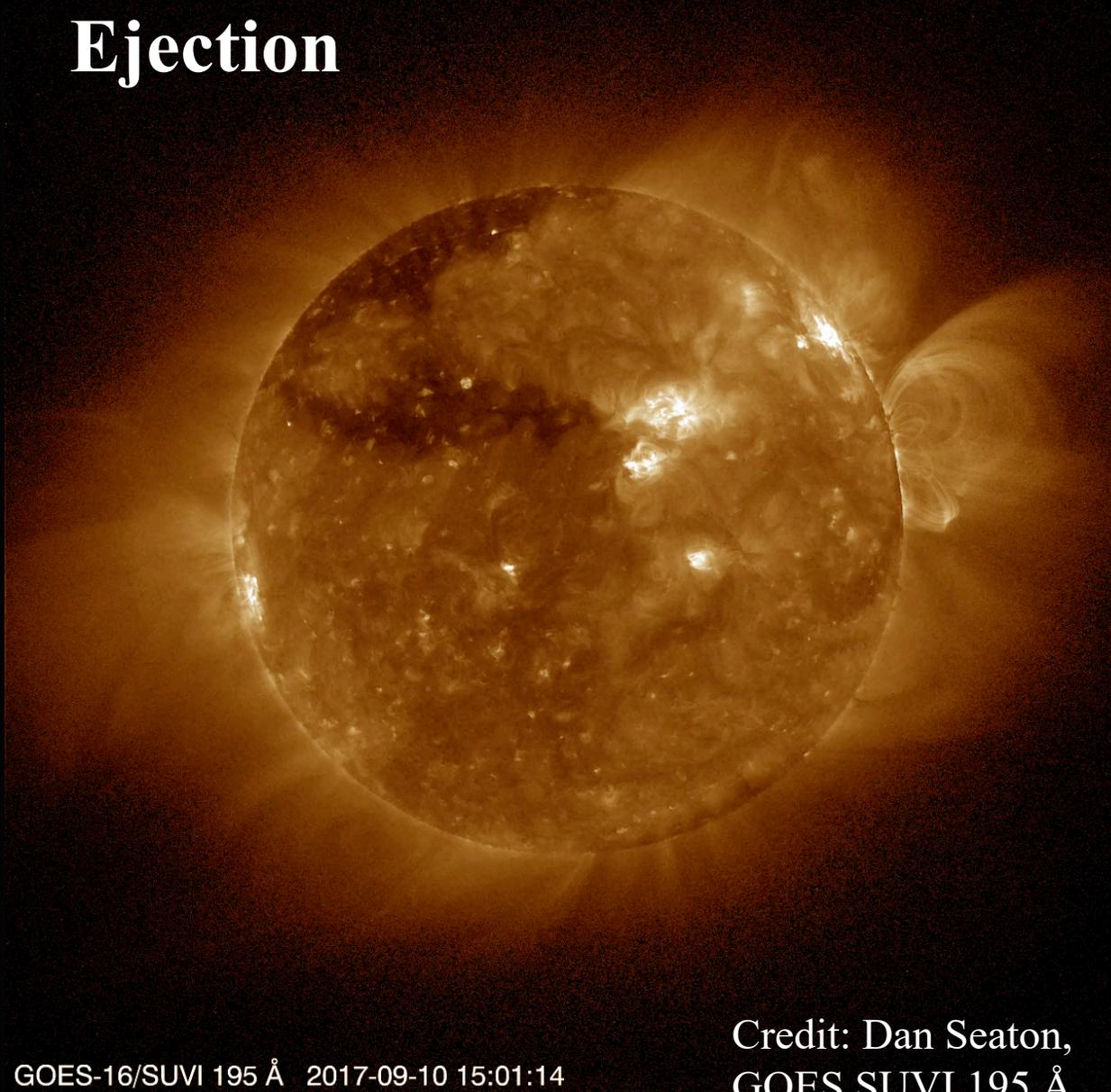
Peter T. Gallagher, Eoin P. Carley, Laura A. Hayes, Sophie A. Murray, Pietro Zucca, Richard Fallows, Joe McCauley, Emilia Kilpua, Gottfried Mann, Christian Vocks, Erika Palmerio, Jens Pomoell, Rami Vainio

Dublin Institute of Advanced Studies, ASTRON, Leibniz-Institut für Astrophysik  
Potsdam, University of Turku

# Solar Activity in September 2017 during LOFAR Observation Campaigns



# The 10 September 2017 Flare and Coronal Mass Ejection

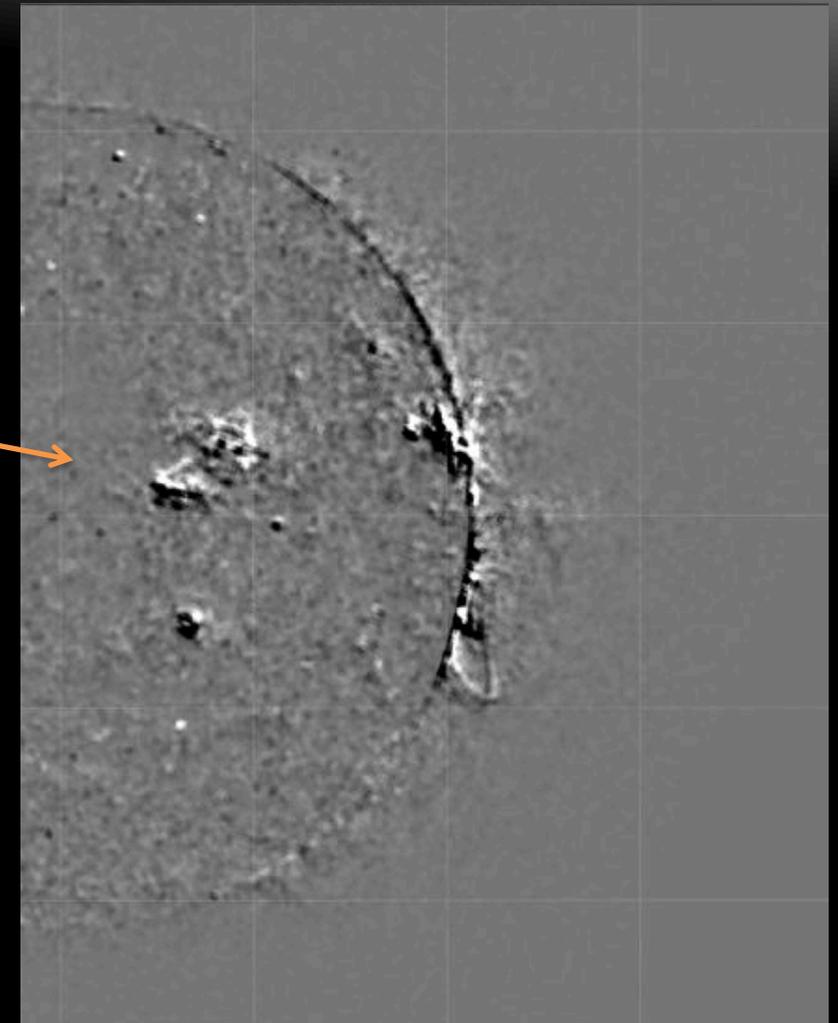
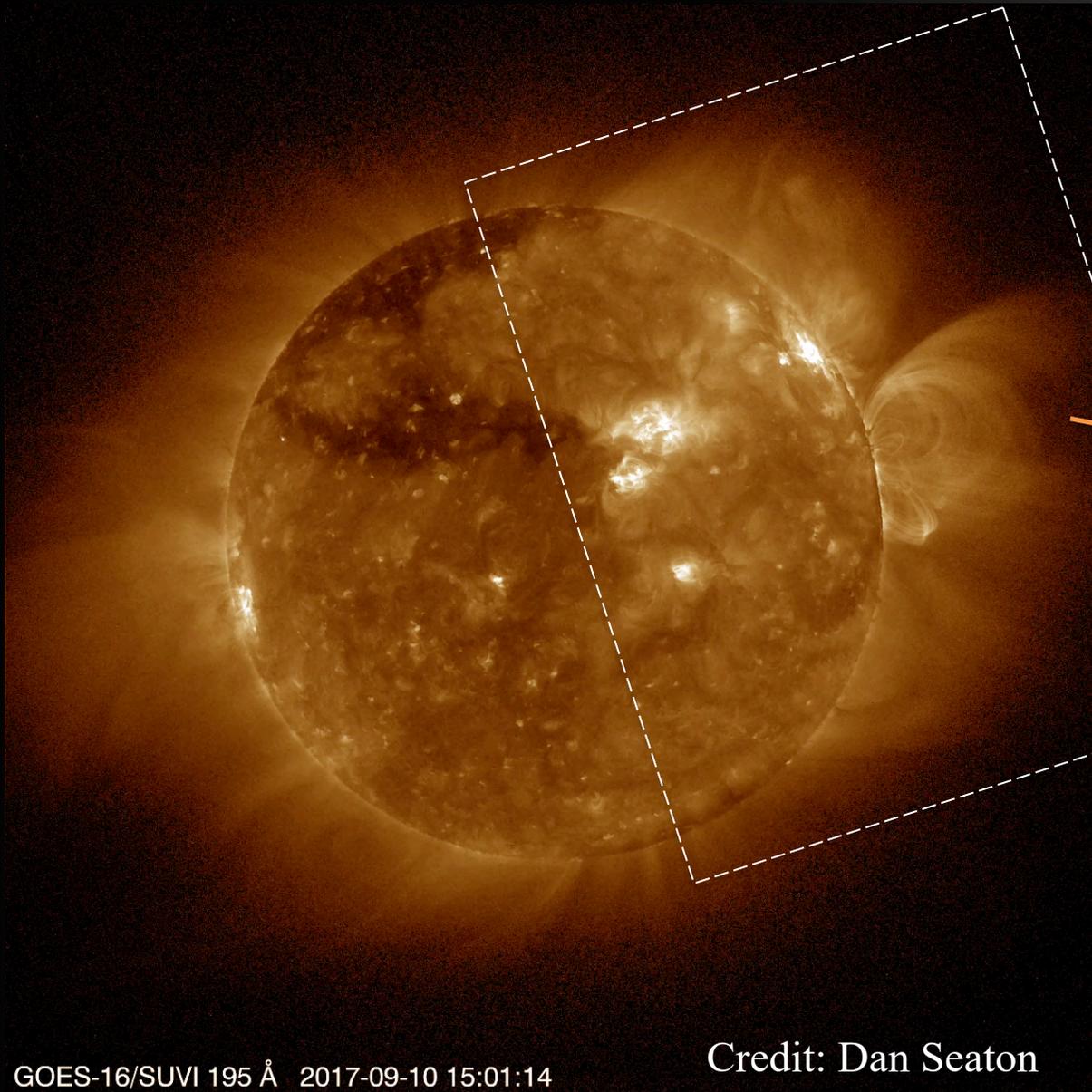


- 10 September 2017**
- **X8.2 Flare on the visible solar limb**
  - **second largest flare of Solar Cycle 24**
  - **accompanied by a very fast CME**

GOES-16/SUVI 195 Å 2017-09-10 15:01:14

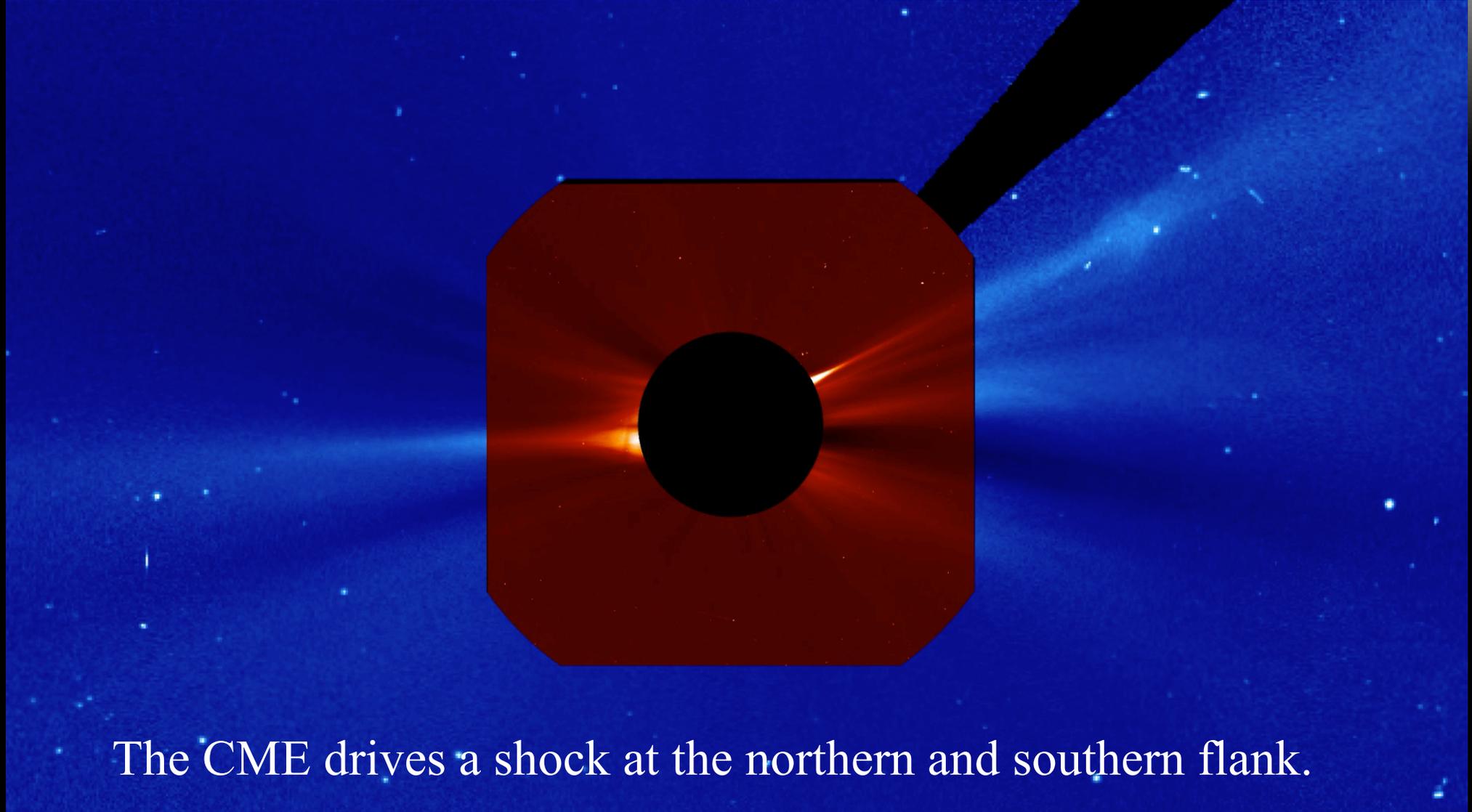
Credit: Dan Seaton,  
GOES SUVI 195 Å

# Observations of a Fast Coronal Mass Ejection (CME)



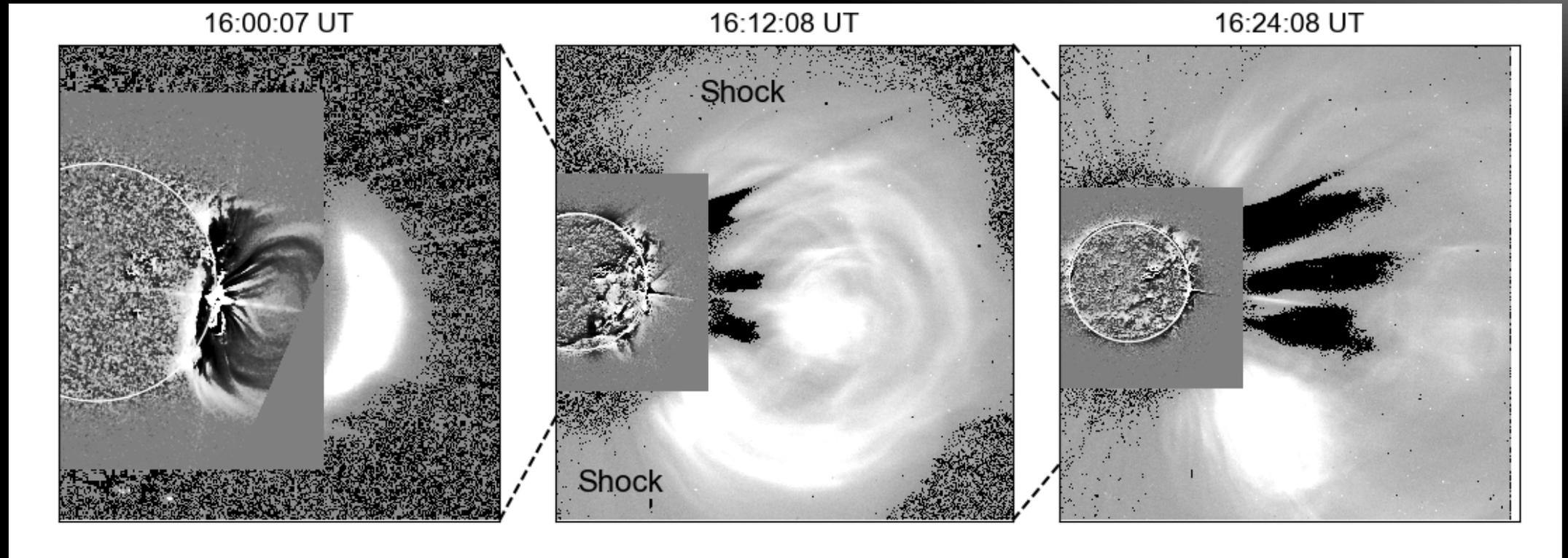
**Very fast CME – 3000 km/s**

# Evolution of the 10 September 2017 CME



The CME drives a shock at the northern and southern flank.

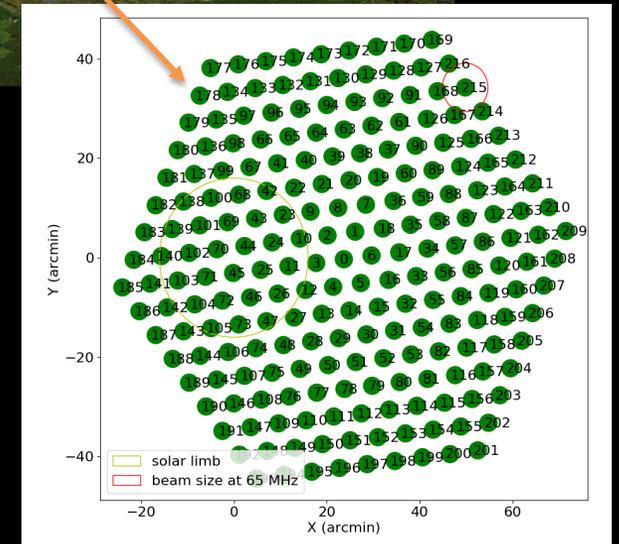
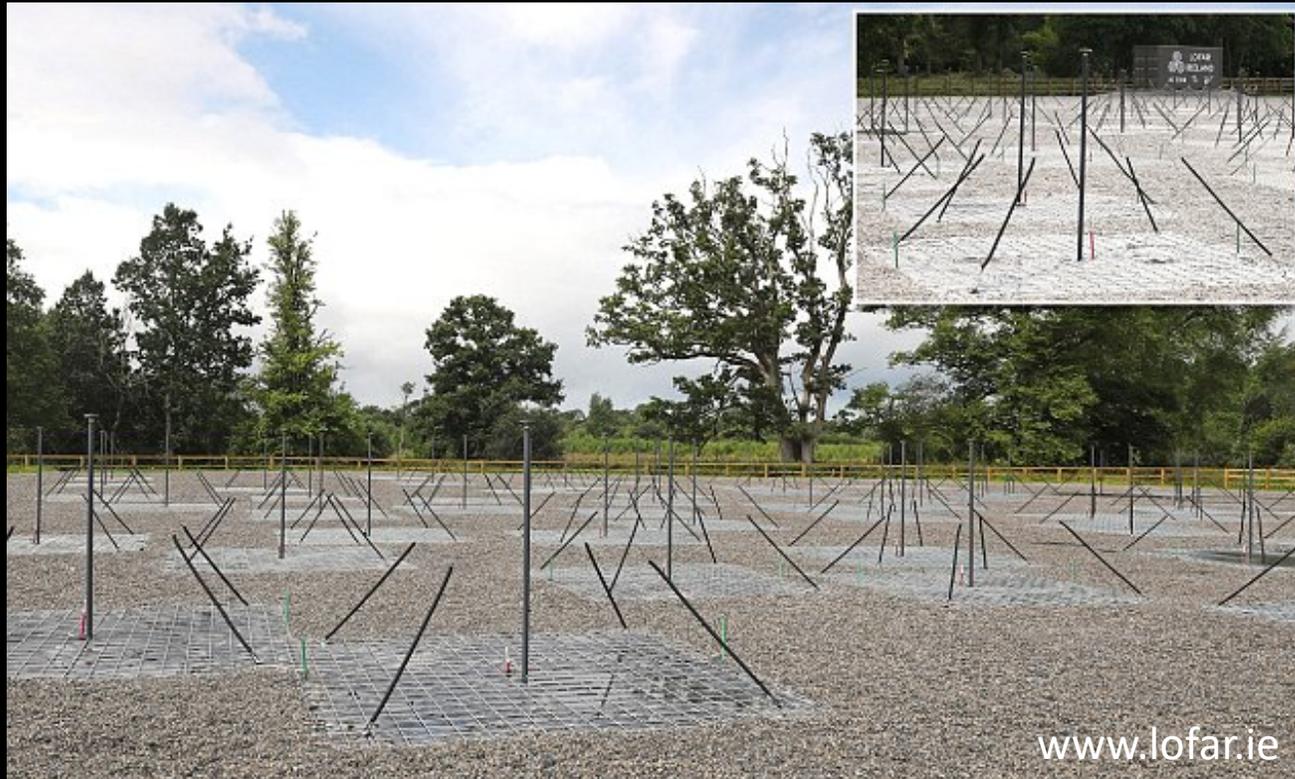
# Evolution of the 10 September 2017 CME



GOES-R SUVI + SOHO/LASCO

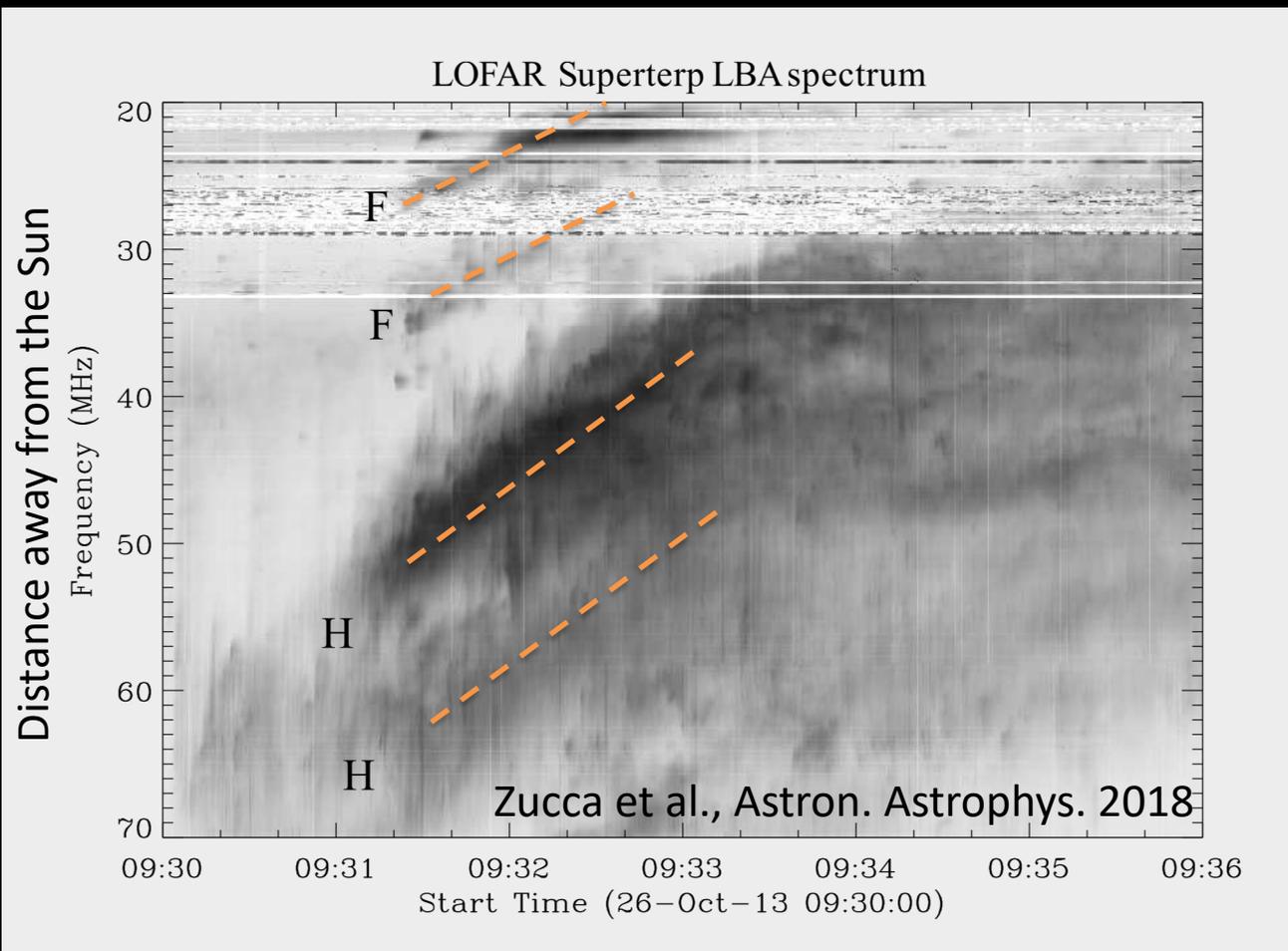
The CME drives a shock at the northern and southern flank.

# LOFAR Core and I-LOFAR Observations

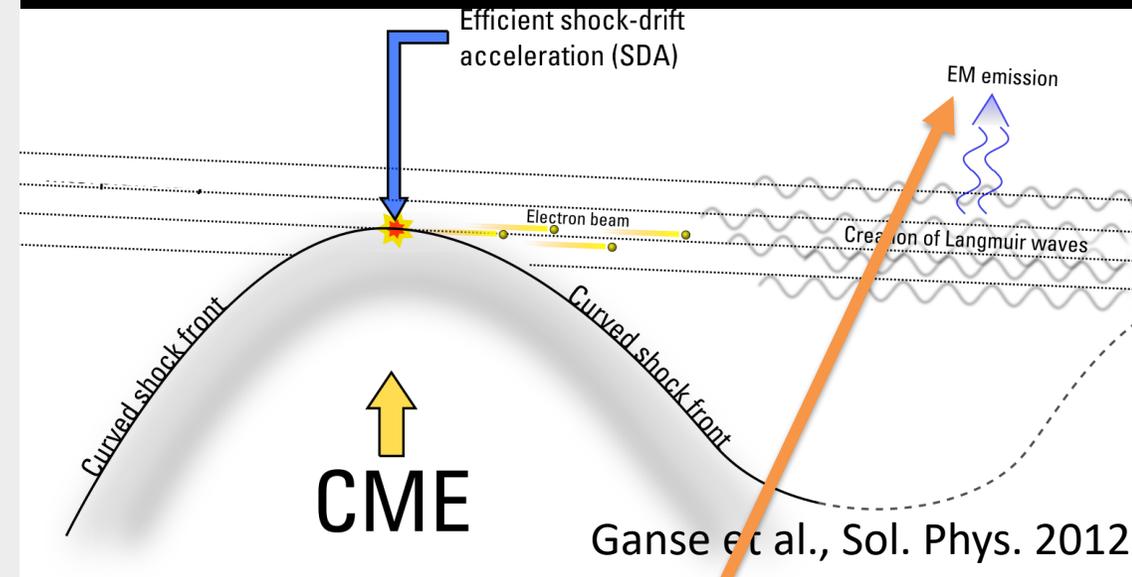


LBA's from the LOFAR core and I-LOFAR station (Ireland) observed the Sun on 10 September 2017

# What did we expect to see with LOFAR? – Type II radio bursts

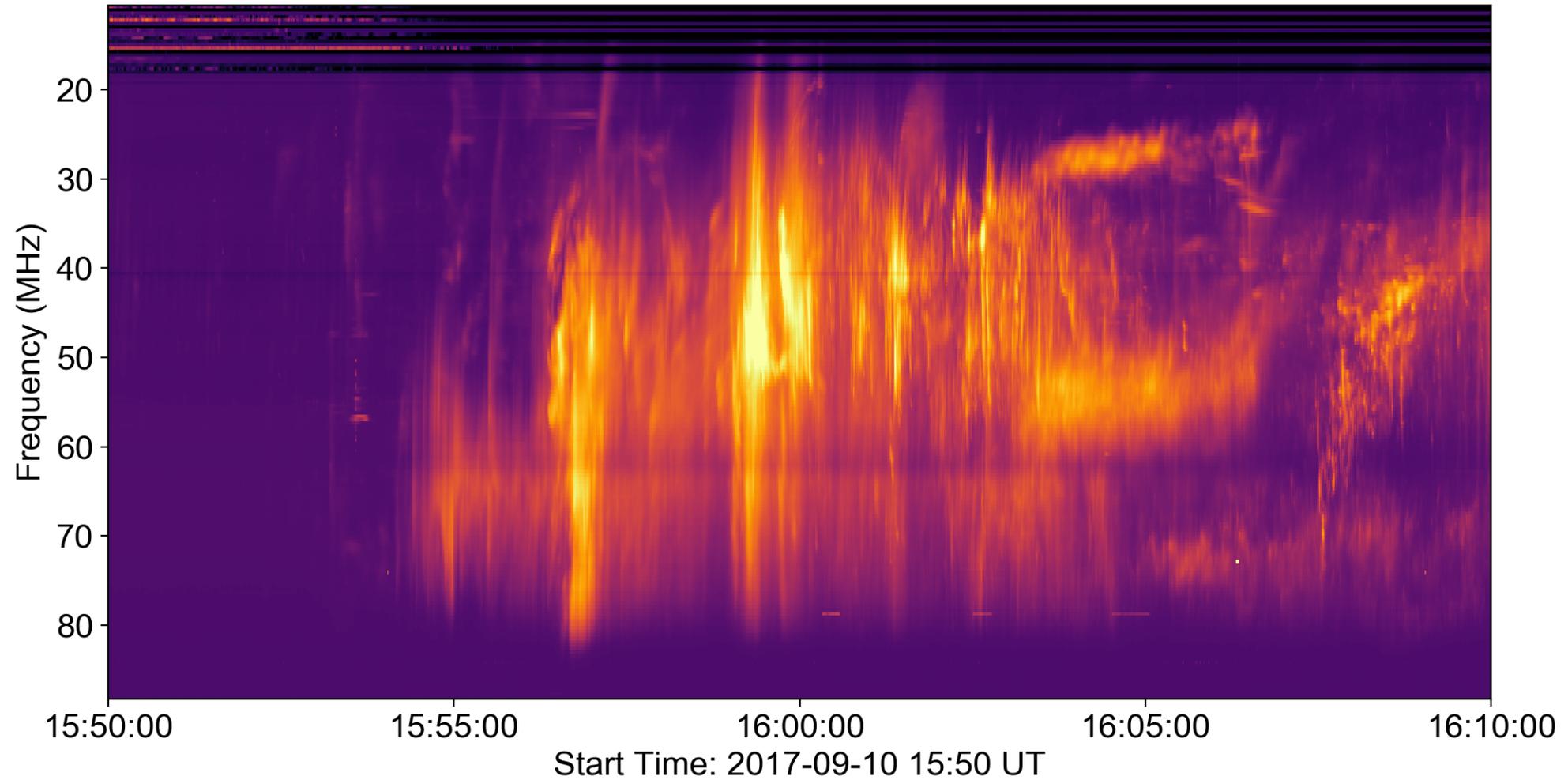


How do shocks accelerate electrons in the corona?

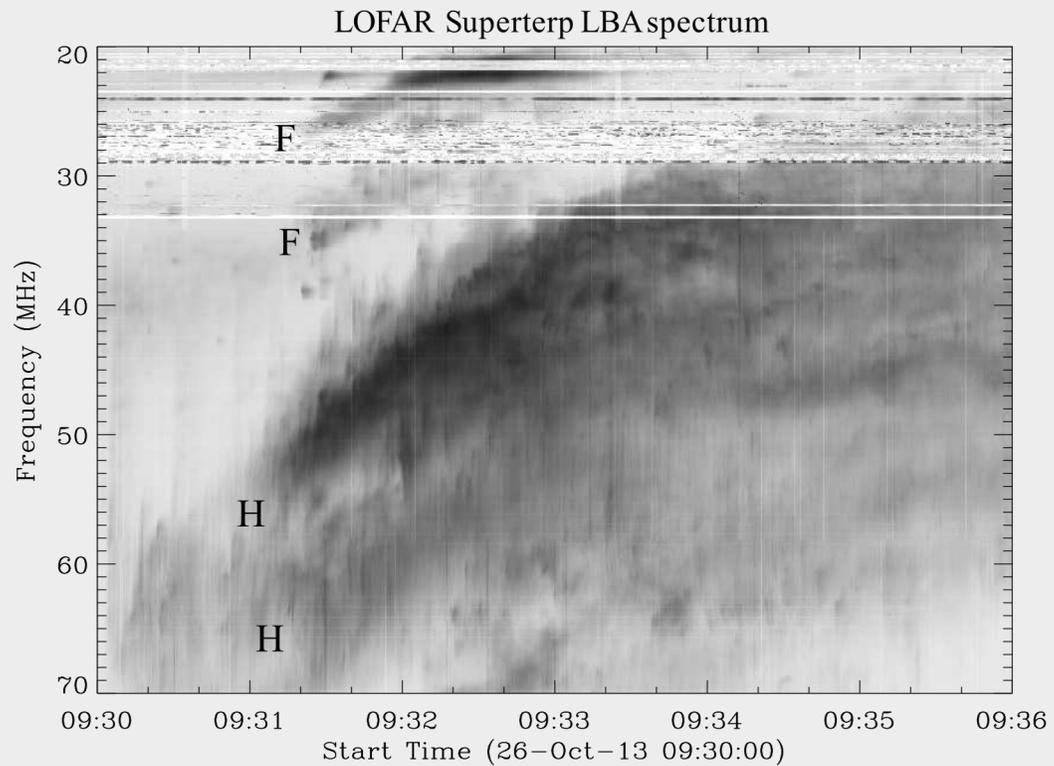
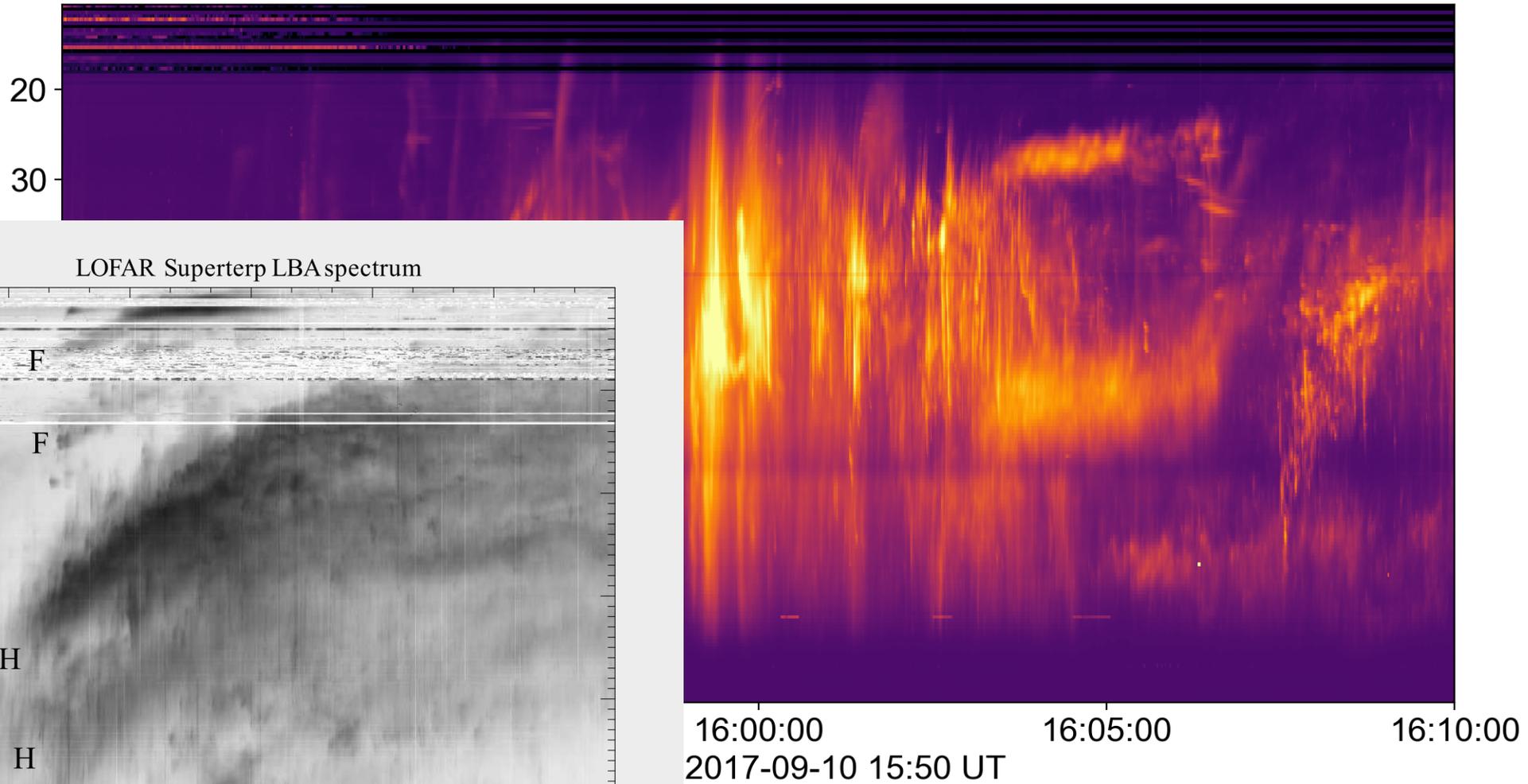


$$f = n f_p = 9000 \sqrt{n_e} \text{ (Hz)}$$

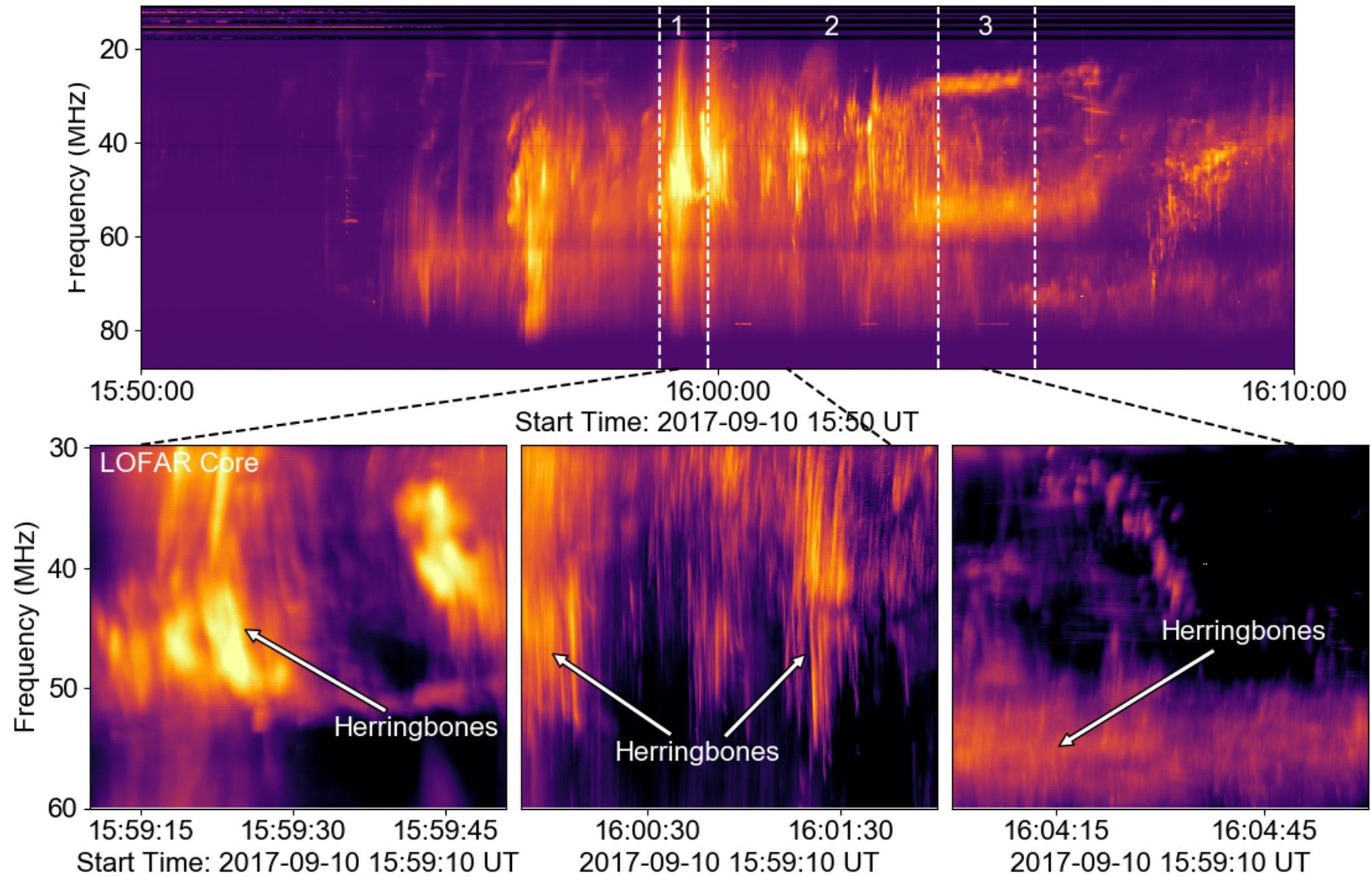
# What we actually saw with LOFAR:



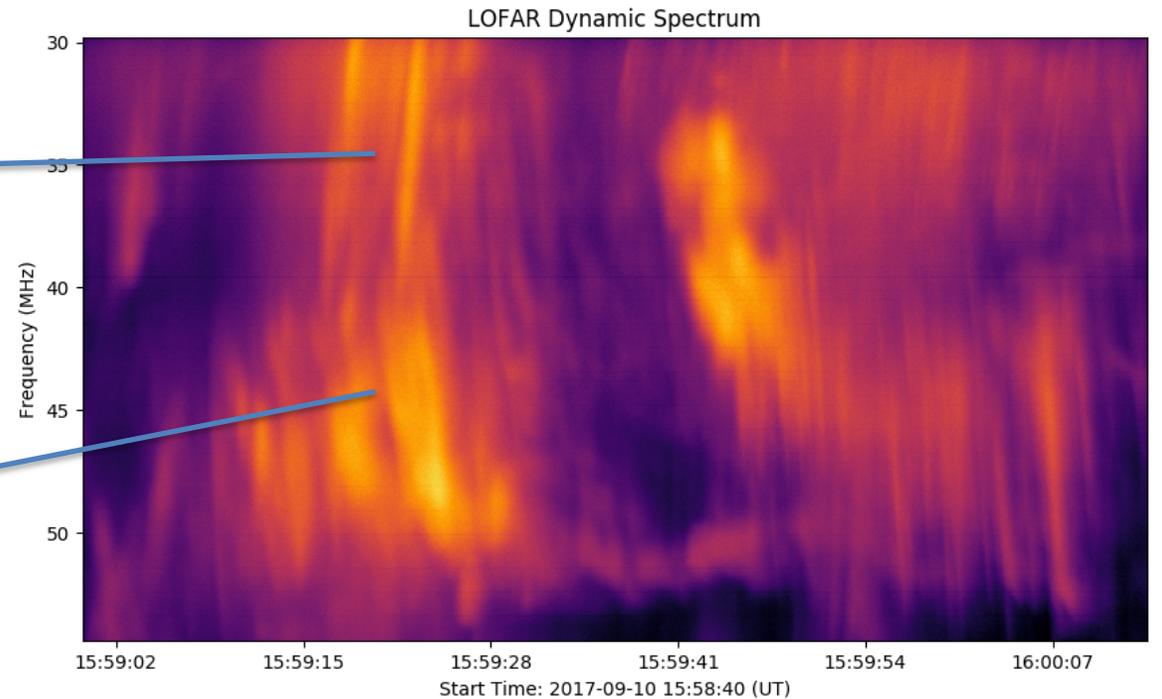
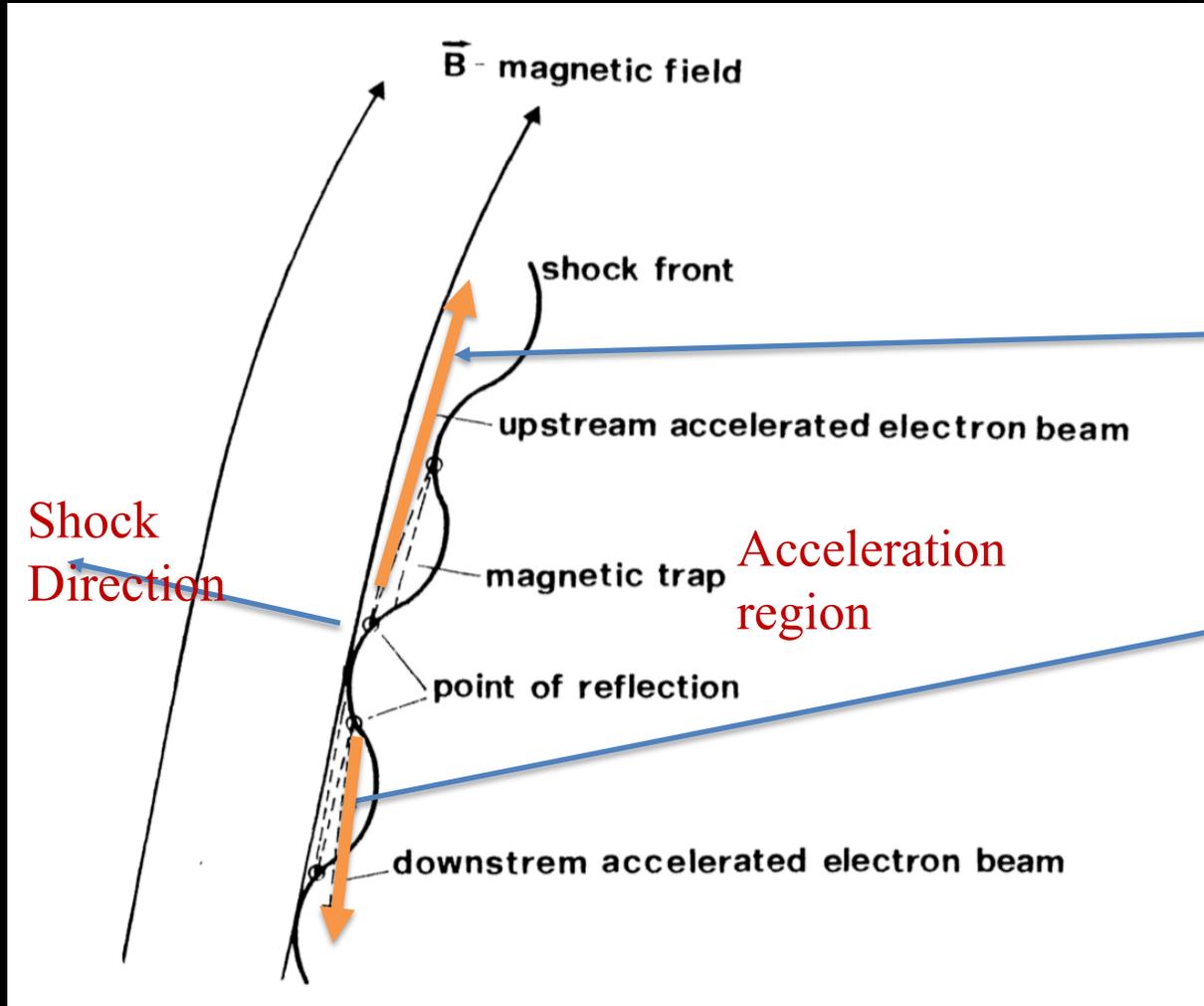
# I-LOFAR Observations of the Event



# LOFAR Core and I-LOFAR Observations of the Event

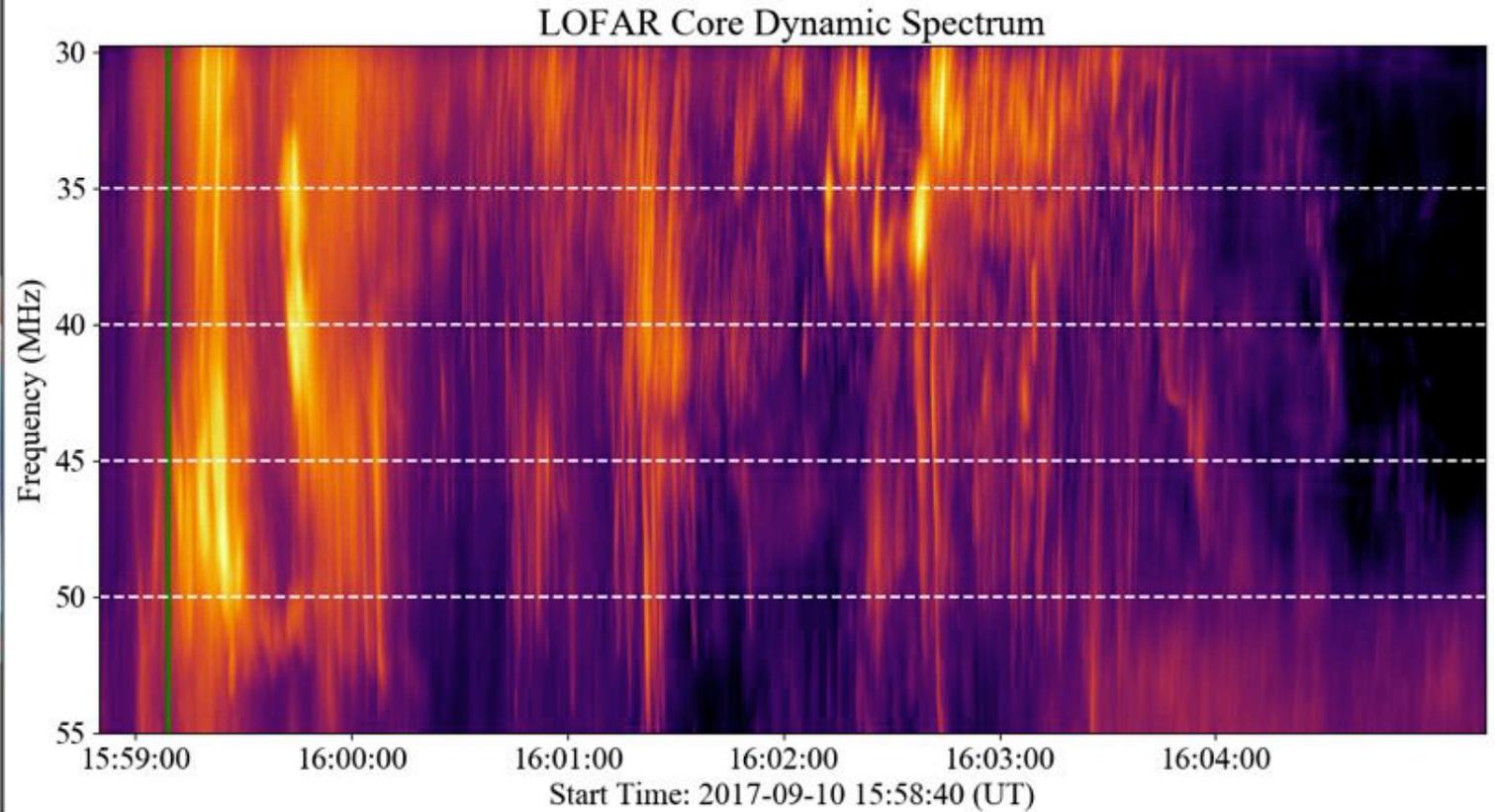
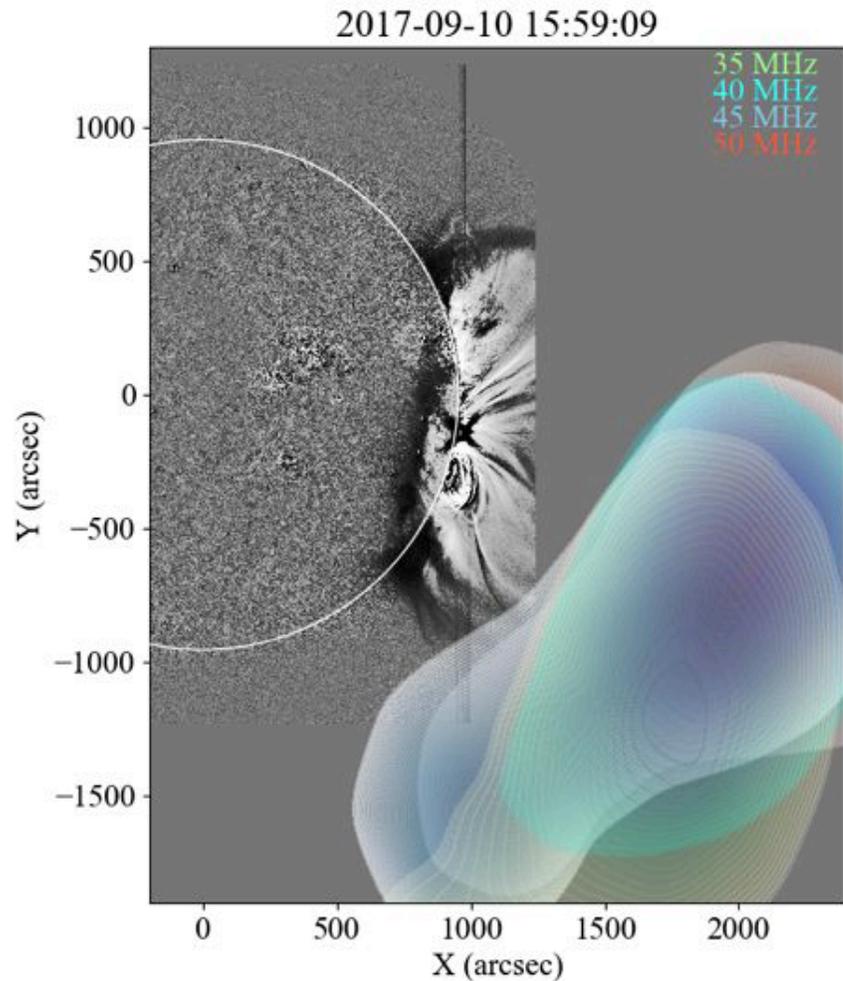


# Theoretical Model of Herringbones

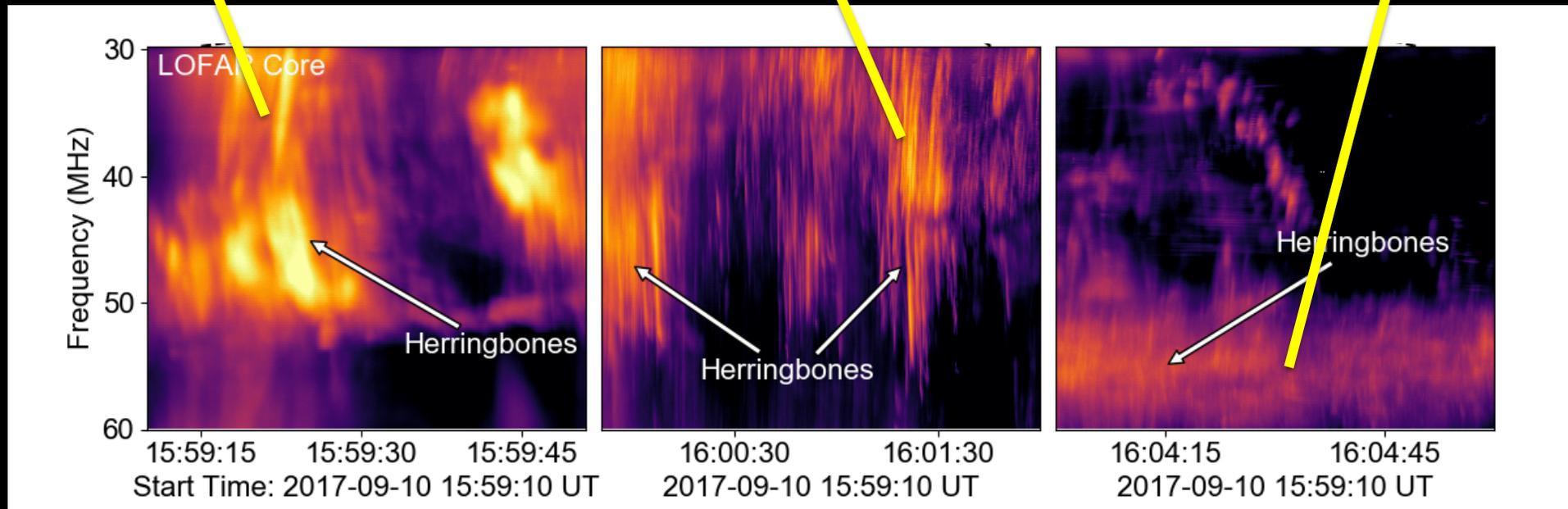
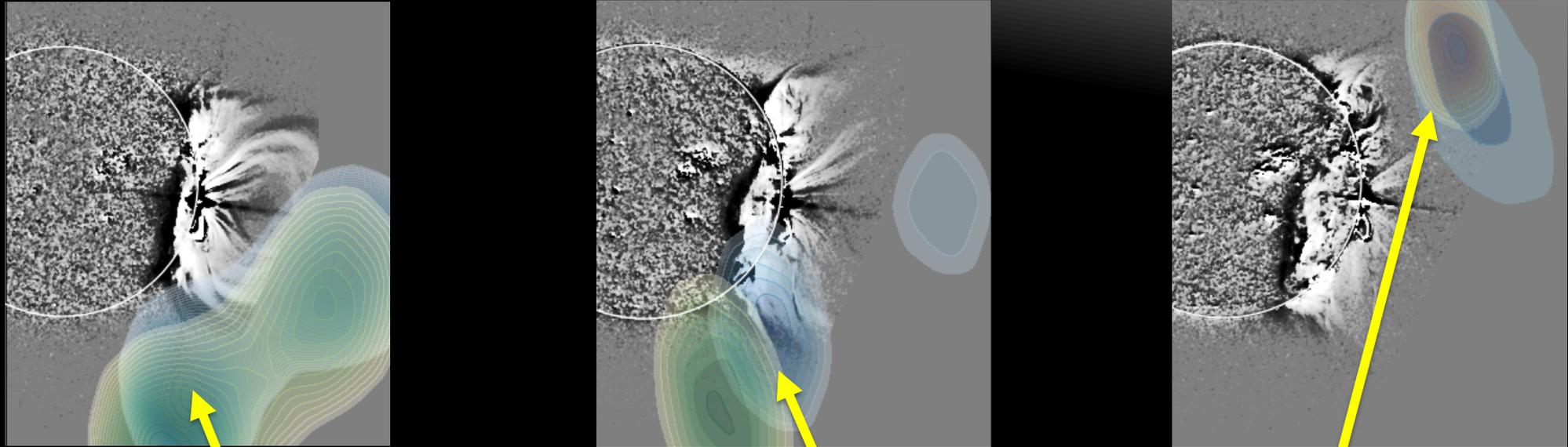


Herringbones Model – Zlobec et al., 1993

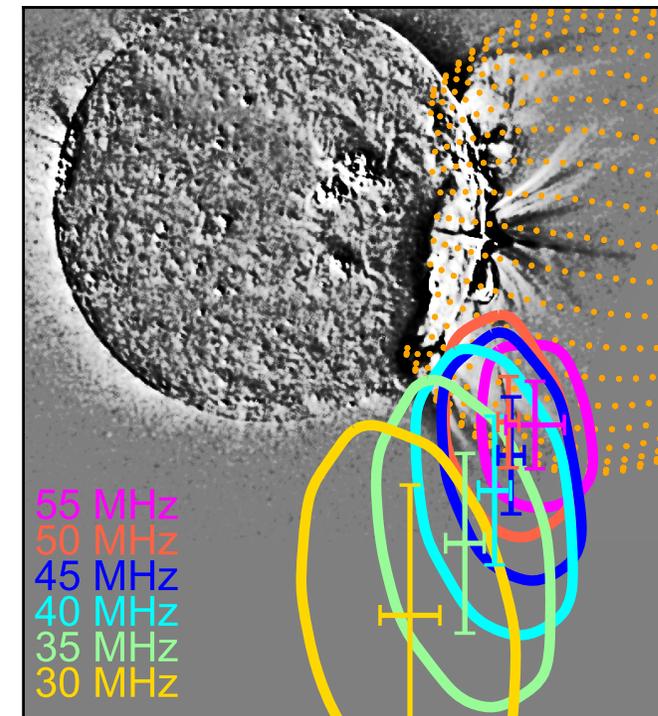
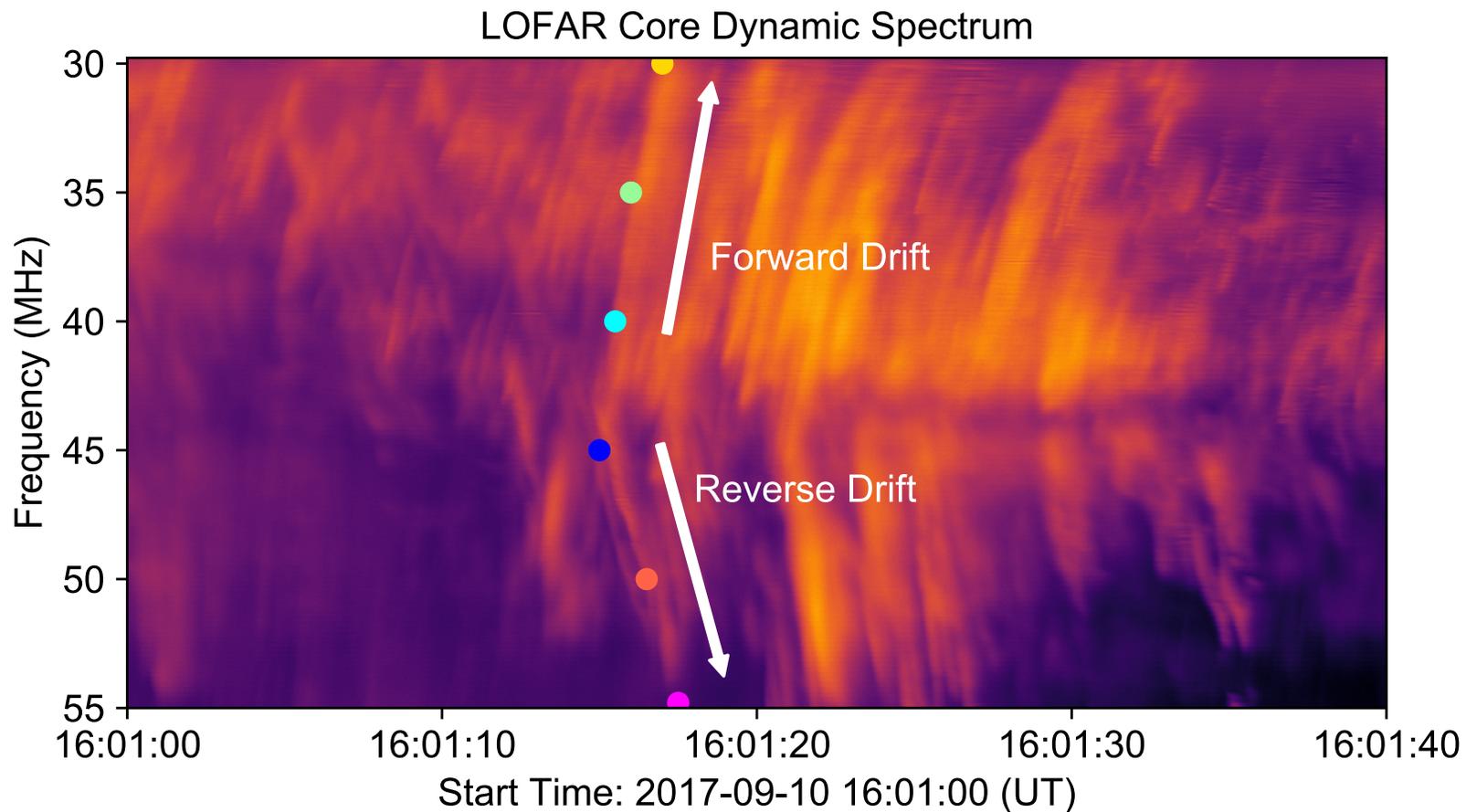
# LOFAR Tied-array Imaging of Herringbones



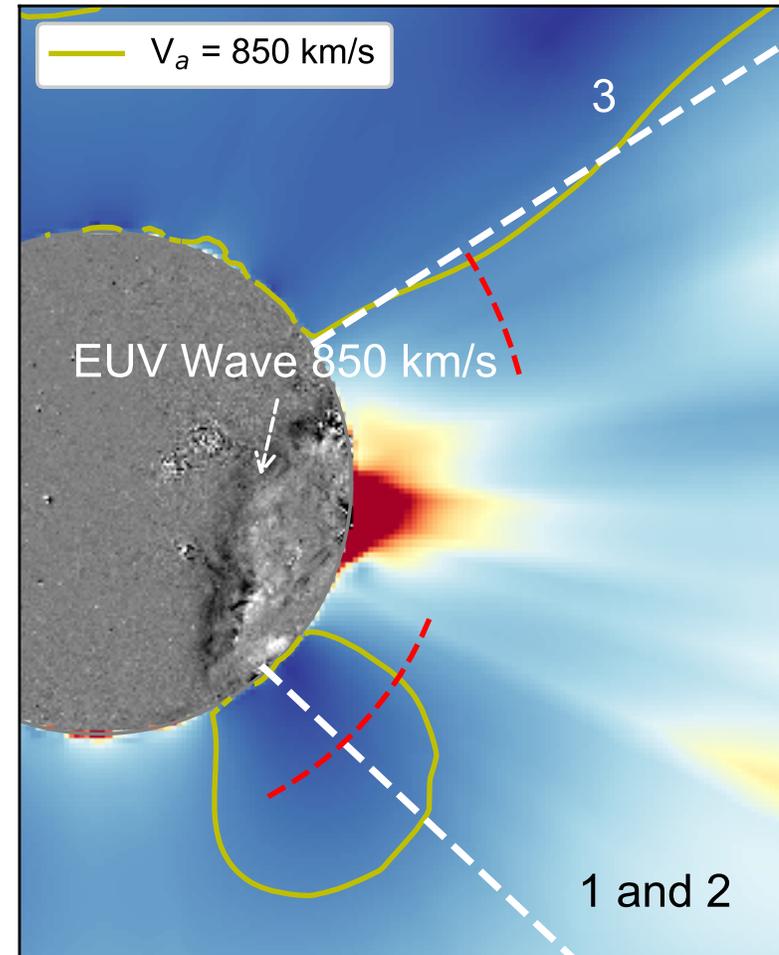
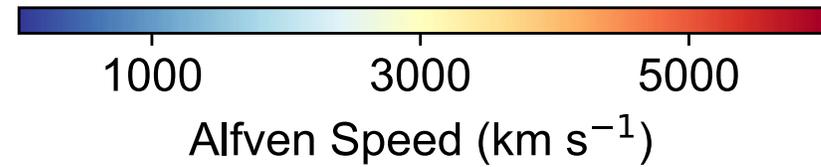
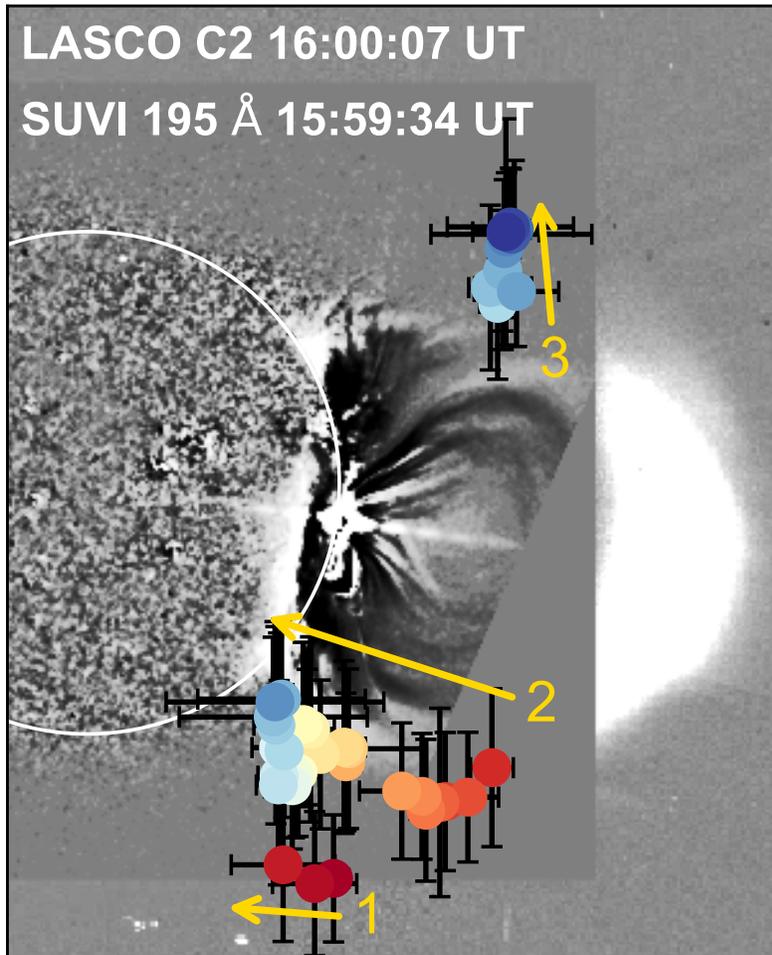
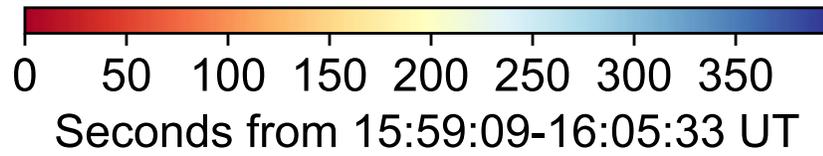
# Three separate region where particles are accelerated:



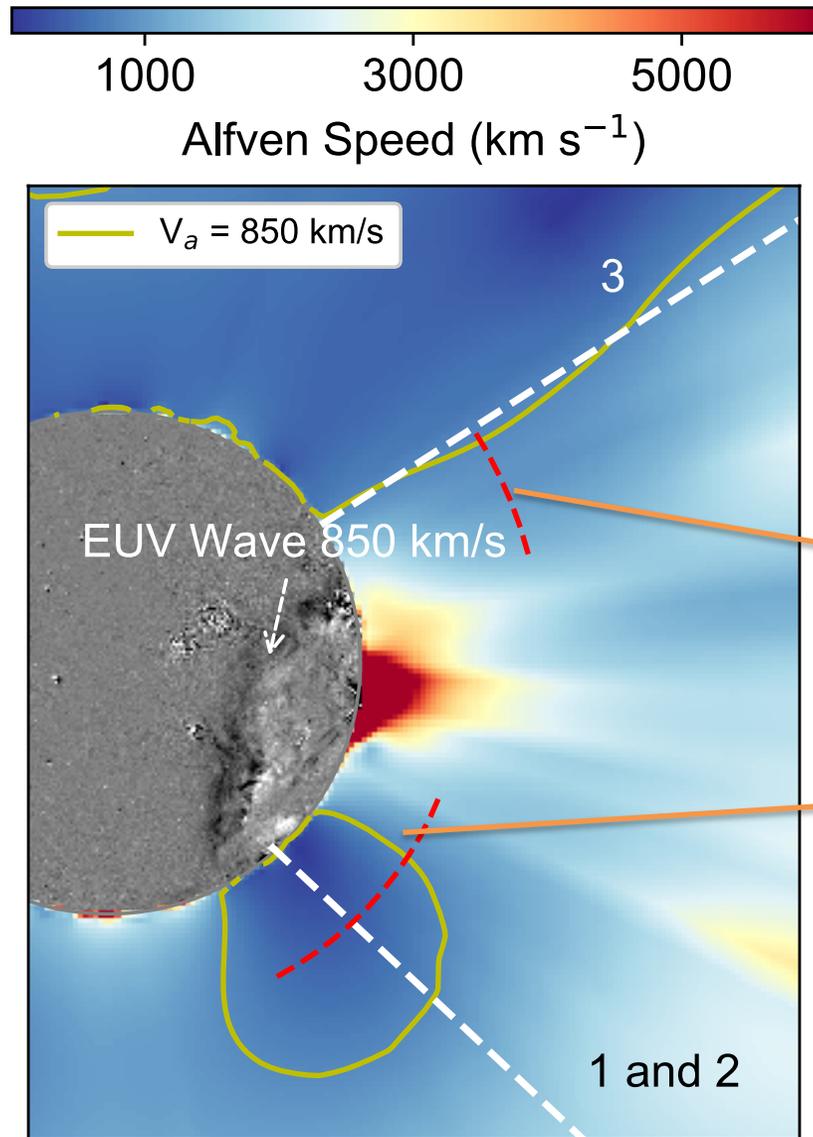
# Finding the location of an individual herringbone



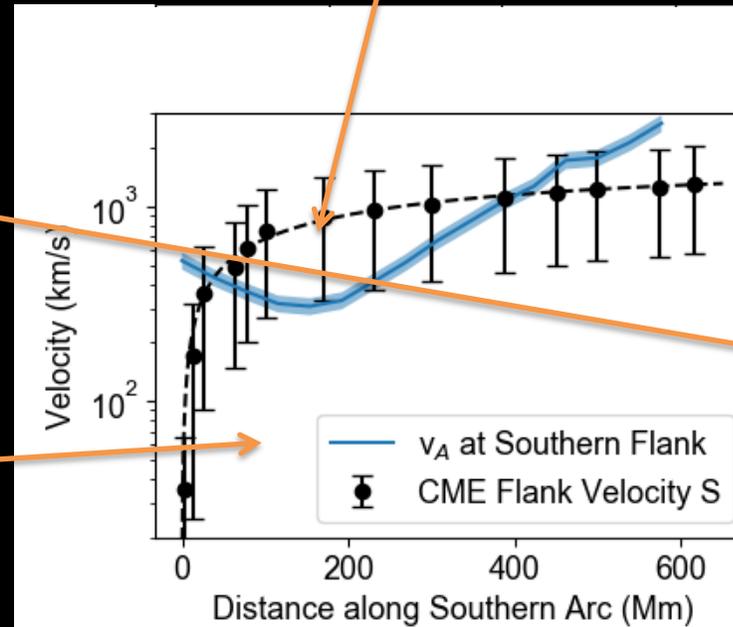
# Is there a shock at the CME flank in the pane-of-sky?



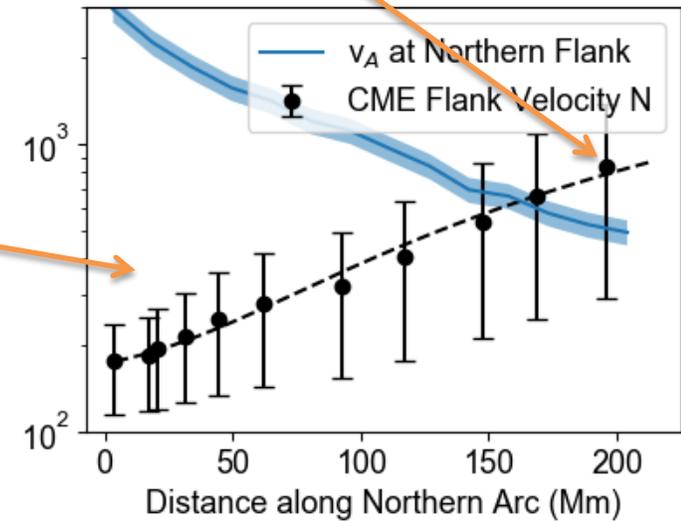
# Plane-of-sky Shock Locations



Shock -  
Mach 3

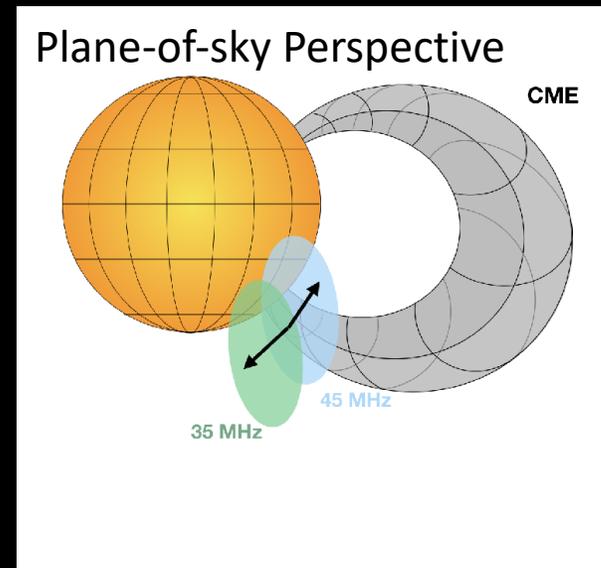
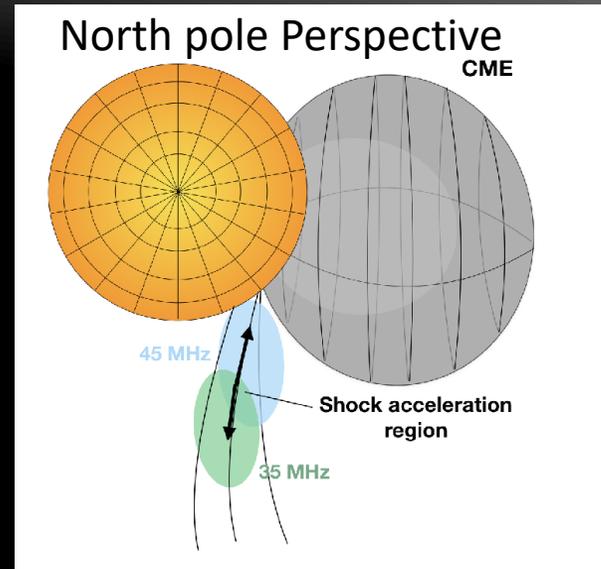
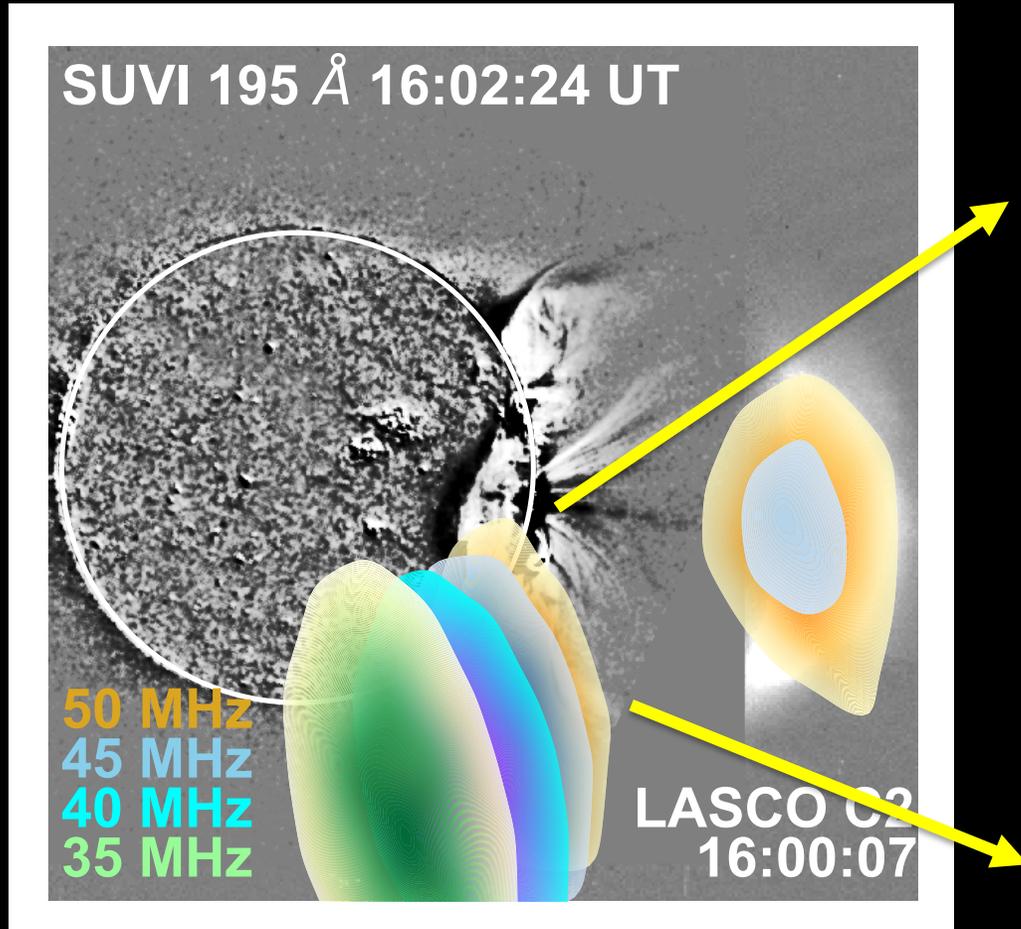


Shock -  
Mach 1.7

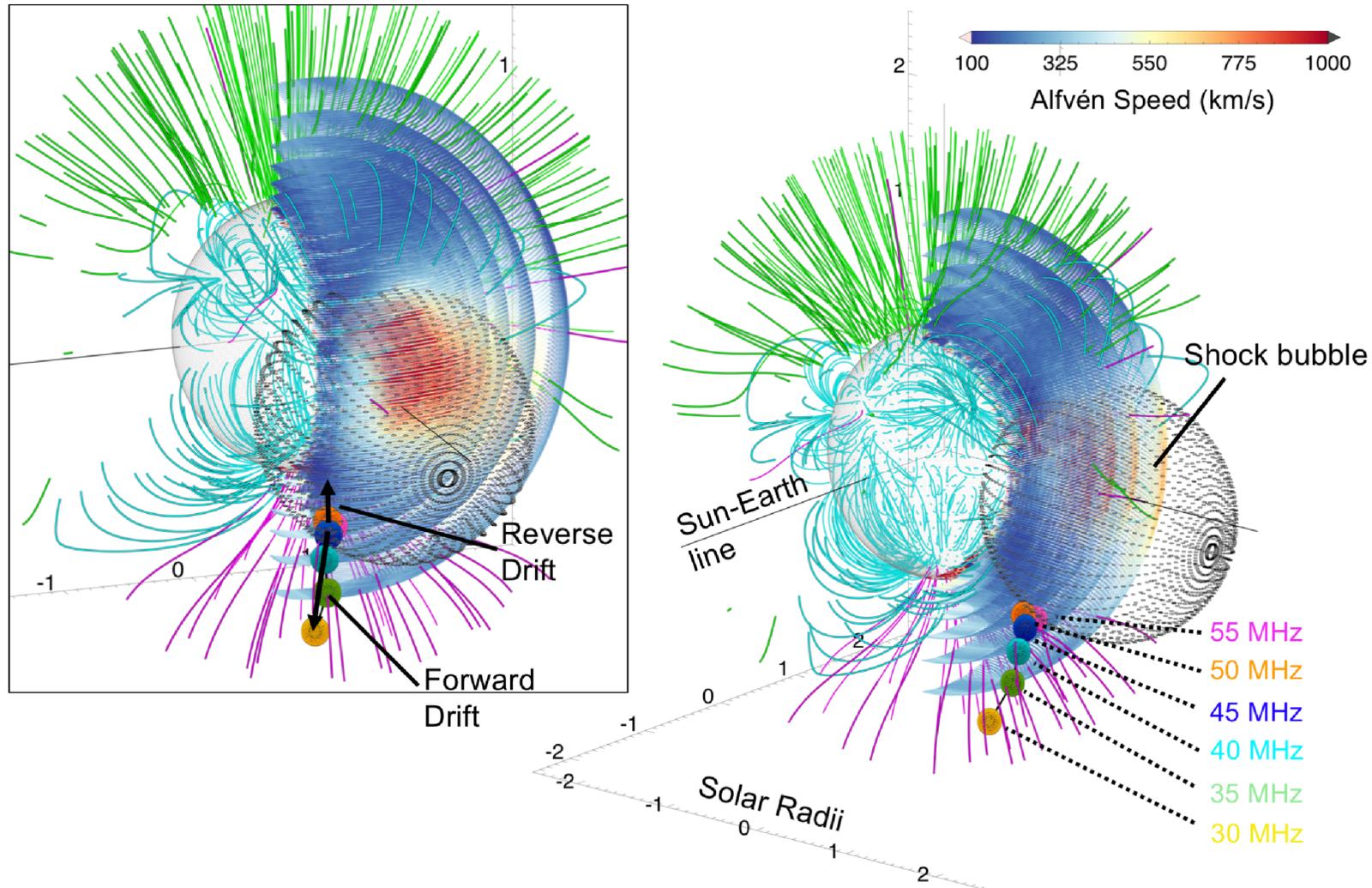


Morosan et al., in review

# Plane-of-sky to 3D Herringbone Locations



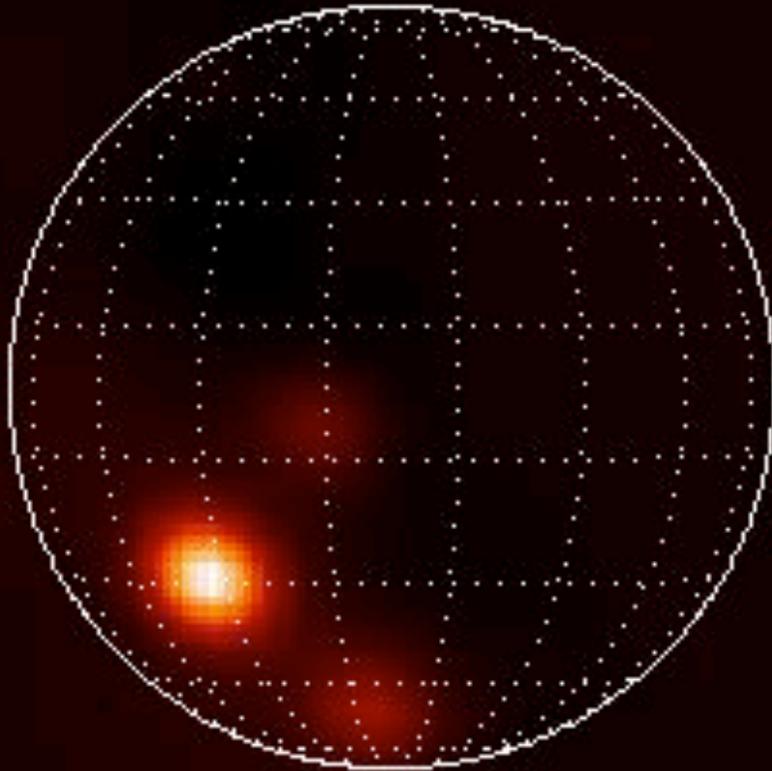
# Reconstruction of 3D Herringbone and Shock Locations



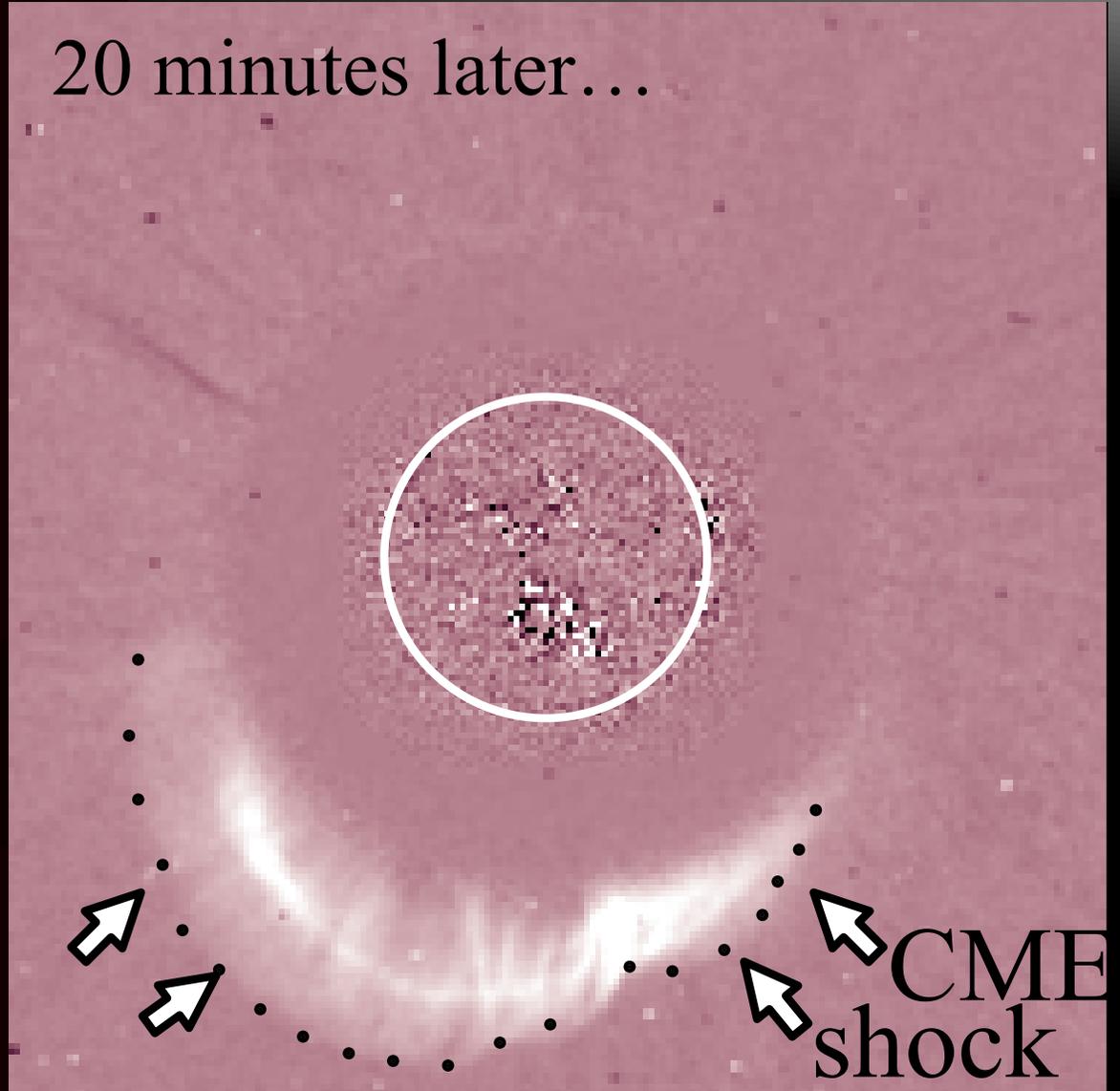


# More Shock Accelerated Electrons

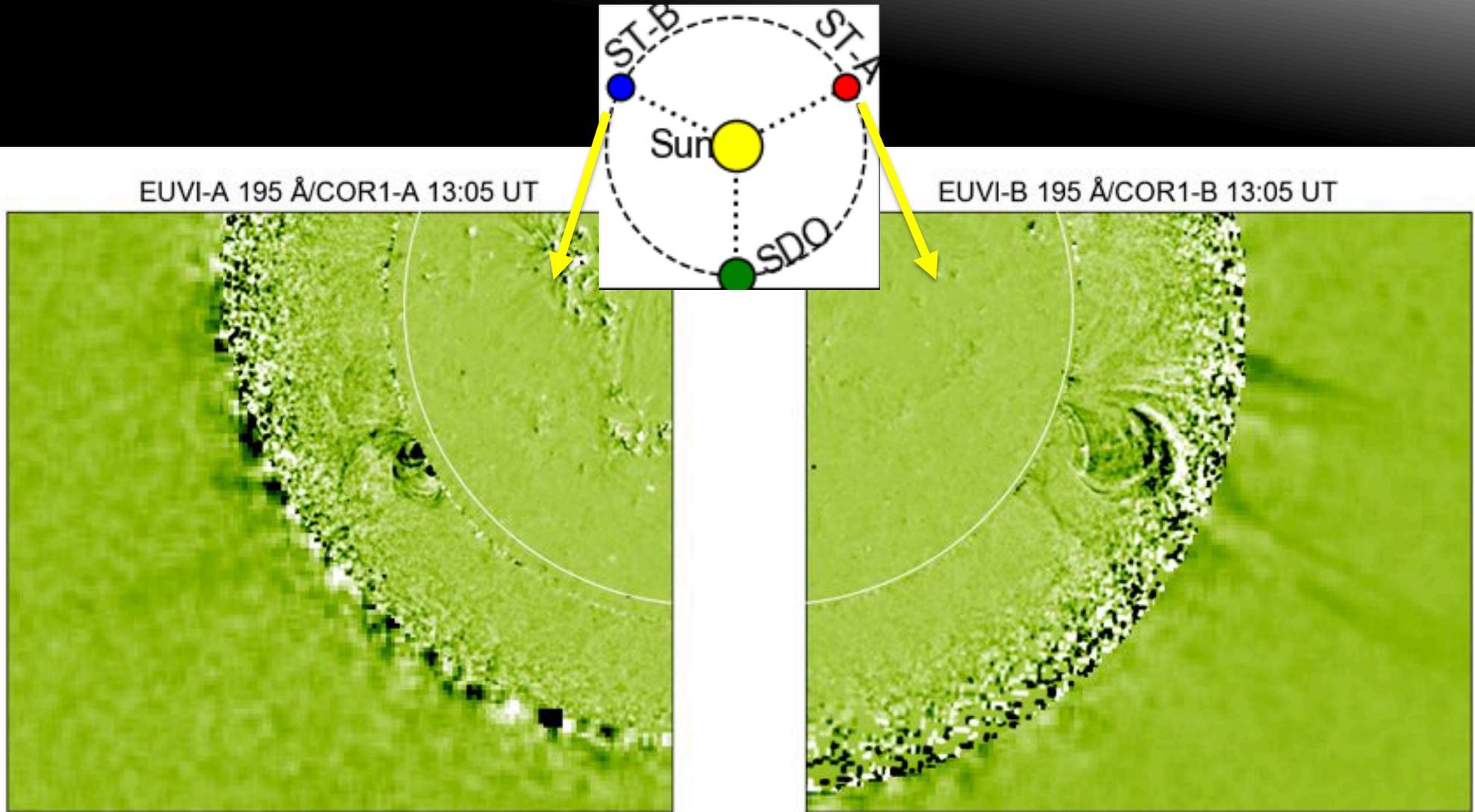
Nançay Radioheliograph



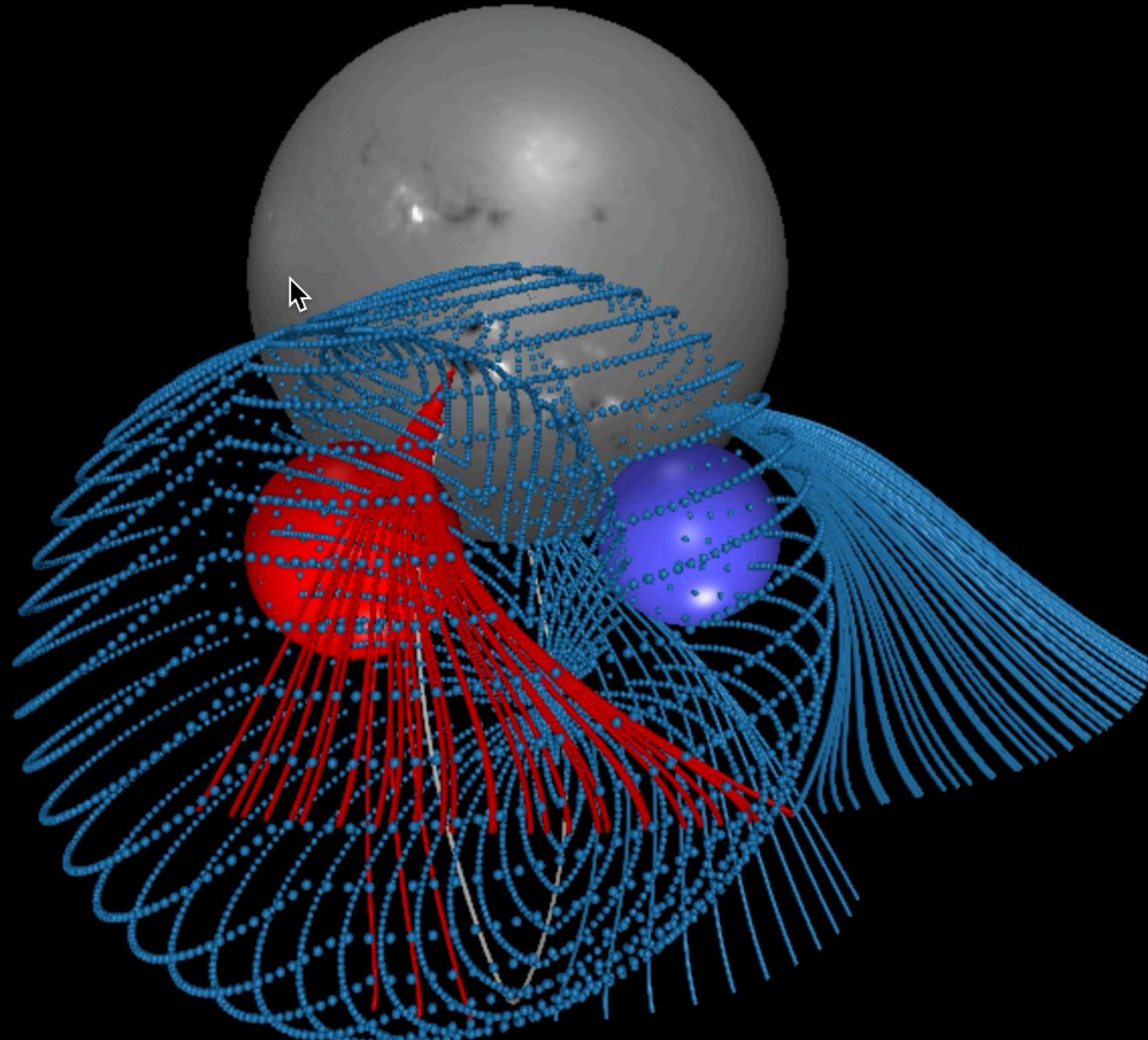
20 minutes later...



# More Shock Accelerated Electrons



# More Shock Accelerated Electrons



# LOFAR Breakthroughs and Future Work

- LOFAR made it possible to track an individual electron beam and to find evidence for electrons accelerated at multiple locations on the expanding CME flank.
- But, there are still many unanswered questions and more ToO LOFAR observations of CMEs can help:
  - why are electrons accelerated only at specific locations since CME shocks are large scale structures?
  - can we get any information on shock morphology in the corona from high resolution interferometric observations?

