

LOFAR Long Baseline Busy $2/5$ -Week

George Heald
on behalf of the LLBWG
31 March 2010

Special thanks to Stefan Wijnholds!



Outline

- What's needed for the long baselines [abridged]
- Ongoing work:
 - X,Y order for LBA_OUTER / LBA_INNER
 - Searching for differential Faraday rotation...
 - Cable length possibly cause of 8 MHz ripple?
 - Clock offsets vs. delays

LLBWG:

J. Conway, J. Anderson, V. Tudose, O. Wucknitz, M. Garrett,
J. McKean, N. Jackson, G. Heald, A. Polatidis, N. Pradel

What's needed

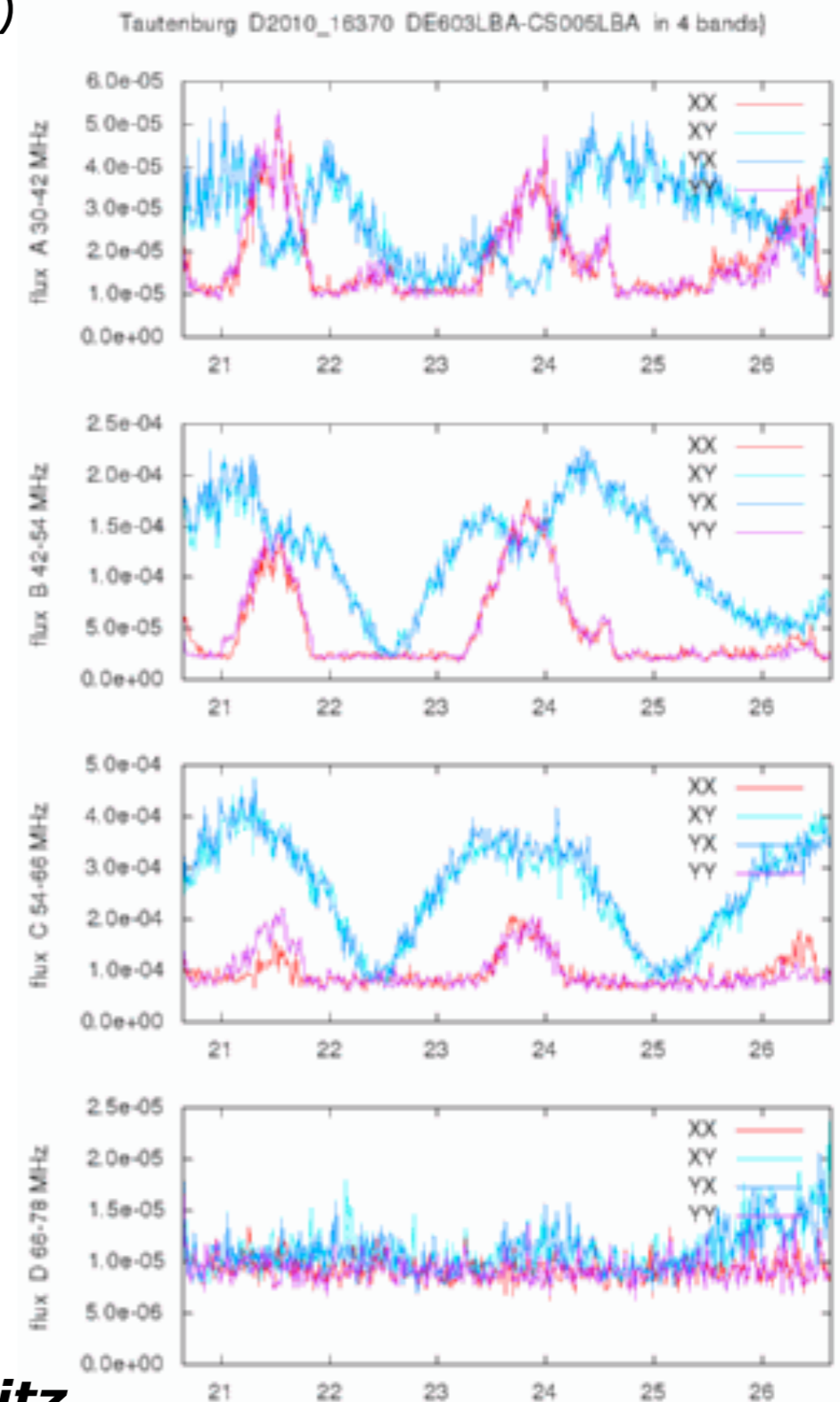
- Prioritization ongoing for several more-or-less long baseline specific items:
 - Fringe fitting in BBS [clock and ionospheric delays over multiple subbands]
 - Faraday rotation [more on this later]
 - Source models for 3C sources [Jackson; using Atlas of DRAGNs]
 - Multiple station beams [one calibration, one science]
 - Calibration transfer between beams [see previous point]
 - Calibrating over multiple beams [later]
 - Imaging with long baselines.....
 - Calibrating on several sources per beam
 - Averaging down uv data after phasing up to source locations
 - Check post-calibration noise levels
 - Speed of BBS
 - Speed of imaging

Cable problem?

- Inspection of Dutch-German baselines led to suspicion that X and Y cables were swapped in Tautenburg (or some similar problem)

(note that $\{XY, YX\} / \{XX, YY\}$ decreases systematically to higher frequencies, implying that at zero wavelength all correlation would be in the crosshands)

- such a cable swap problem seemed to be ruled out by careful inspection at the station



LBA_INNER vs. LBA_OUTER

- X,Y inputs swapped (*on purpose*) between LBA-high and LBA-low on RCU boards

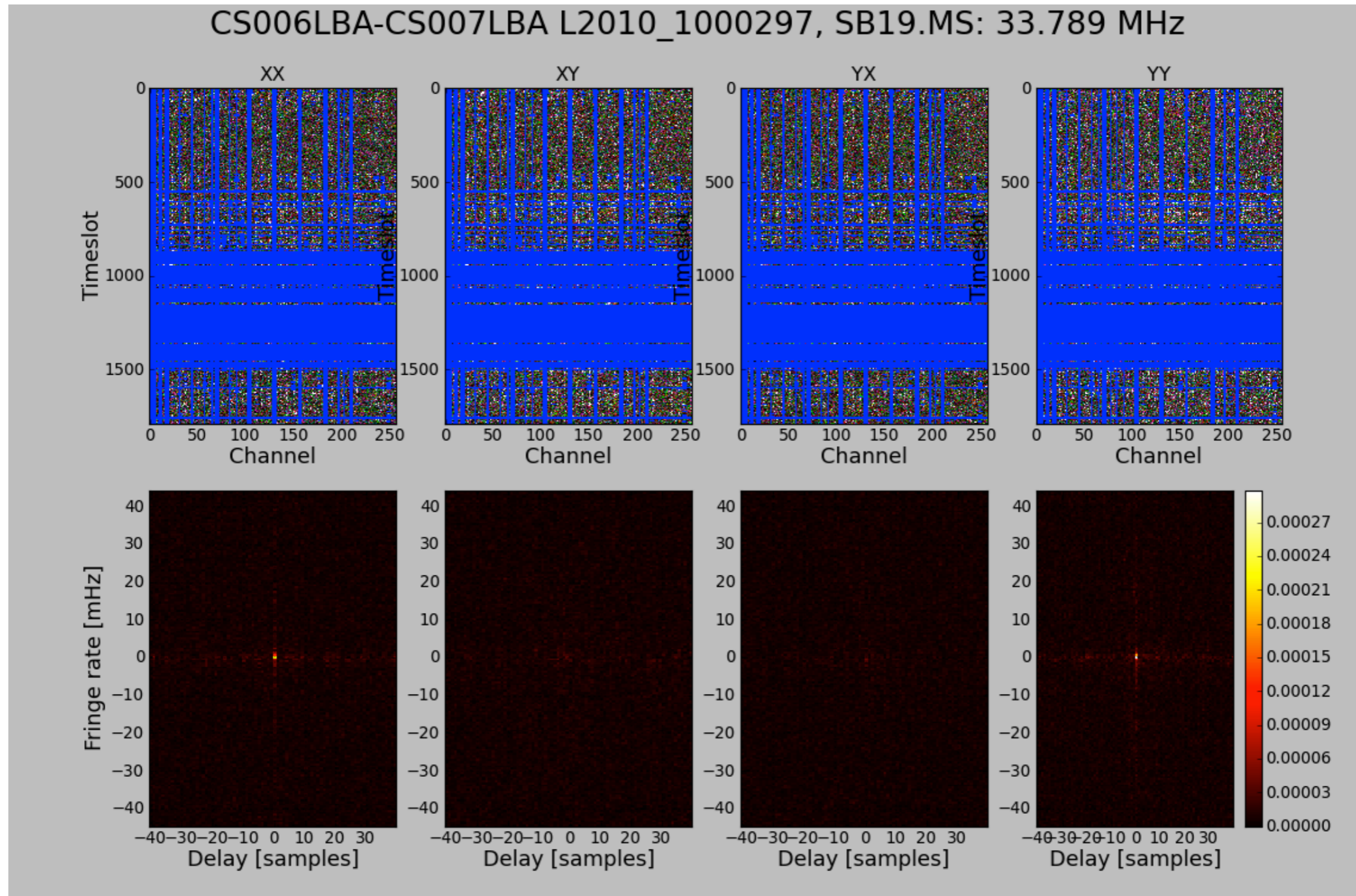
RCU-LBA connection scheme SUBRACK 0 P 2/4

| | | | | | | | | | | | |
|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|--|
| RCU7 | RCU6 | RCU5 | RCU4 | RCU15 | RCU14 | RCU23 | RCU22 | RCU31 | RCU30 | RCU29 | |
| H3_Y | H3_X | H2_Y | H2_X | H7_Y | H7_X | H11_Y | H11_X | H15_Y | H15_X | H14_Y | |
| 3_Y | 3_X | 2_Y | 2_X | 7_Y | 7_X | 11_Y | 11_X | 15_Y | 15_X | 14_Y | |
| L51_X | L51_Y | L50_X | L50_Y | L55_X | L55_Y | L59_X | L59_Y | L63_X | L63_Y | L62_X | |
| RCU3 | RCU2 | RCU11 | RCU10 | RCU13 | RCU12 | RCU21 | RCU20 | RCU27 | RCU26 | RCU28 | |
| H1_Y | H1_X | H5_Y | H5_X | H6_Y | H6_X | H10_Y | H10_X | H13_Y | H13_X | H14_X | |
| 1_Y | 1_X | 5_Y | 5_X | 6_Y | 6_X | 10_Y | 10_X | 13_Y | 13_X | 14_X | |
| L49_X | L49_Y | L53_X | L53_Y | L54_X | L54_Y | L58_X | L58_Y | L61_X | L61_Y | L62_Y | |
| RCU1 | RCU0 | RCU9 | RCU8 | RCU17 | RCU16 | RCU19 | RCU18 | RCU25 | RCU24 | | |
| H0_Y | H0_X | H4_Y | H4_X | H8_Y | H8_X | H9_Y | H9_X | H12_Y | H12_X | | |
| 0_Y | 0_X | 4_Y | 4_X | 8_Y | 8_X | 9_Y | 9_X | 12_Y | 12_X | | |
| L48_X | L48_Y | L52_X | L52_Y | L56_X | L56_Y | L57_X | L57_Y | L60_X | L60_Y | | |

- RSP0
- RSP1
- RSP2
- RSP3

LBA_INNER vs. LBA_OUTER

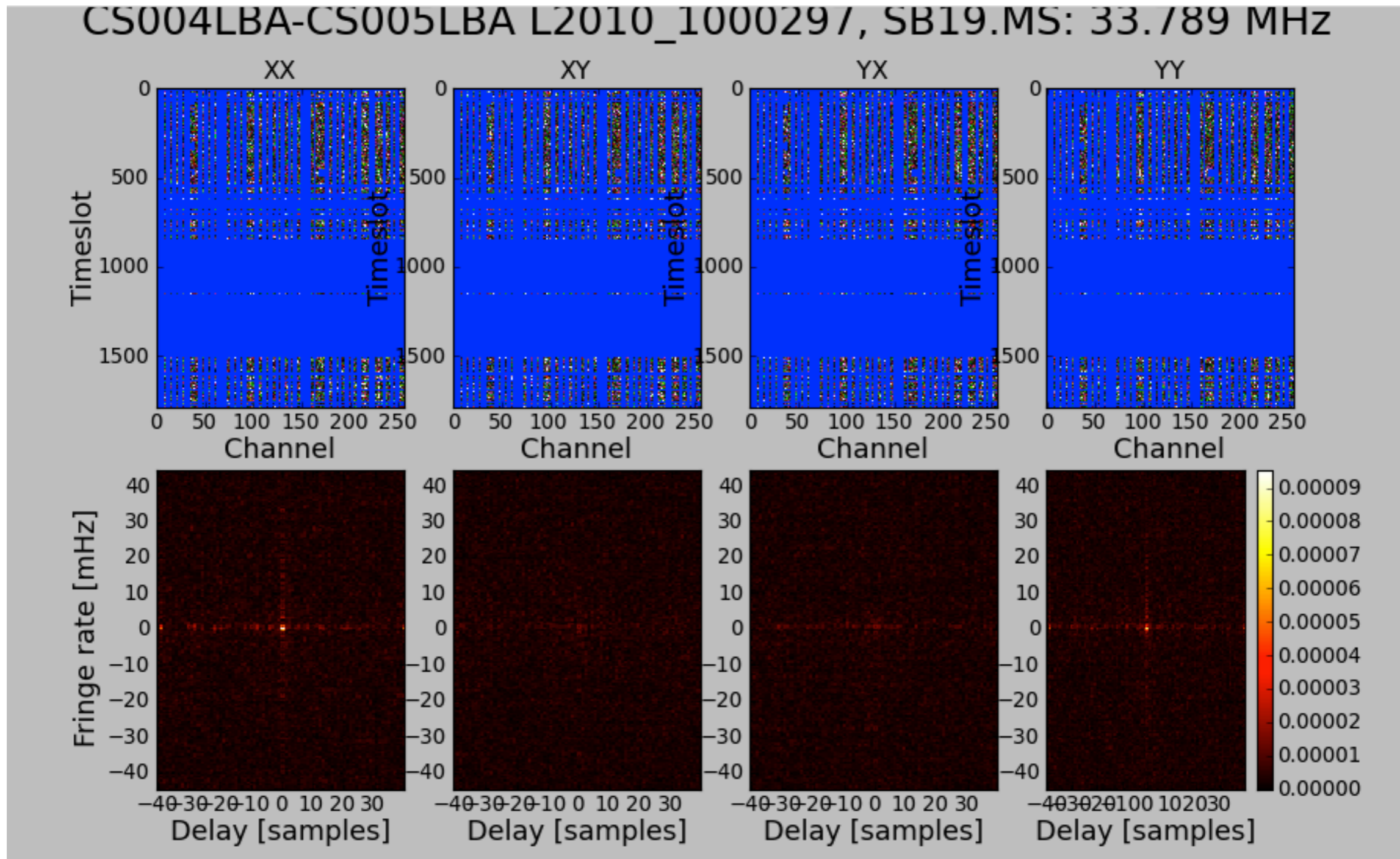
- Test experiment: several superterp stations, some in LBA_INNER and some in LBA_OUTER, correlated together in same experiment.



Two LBA_INNER stations...

LBA_INNER vs. LBA_OUTER

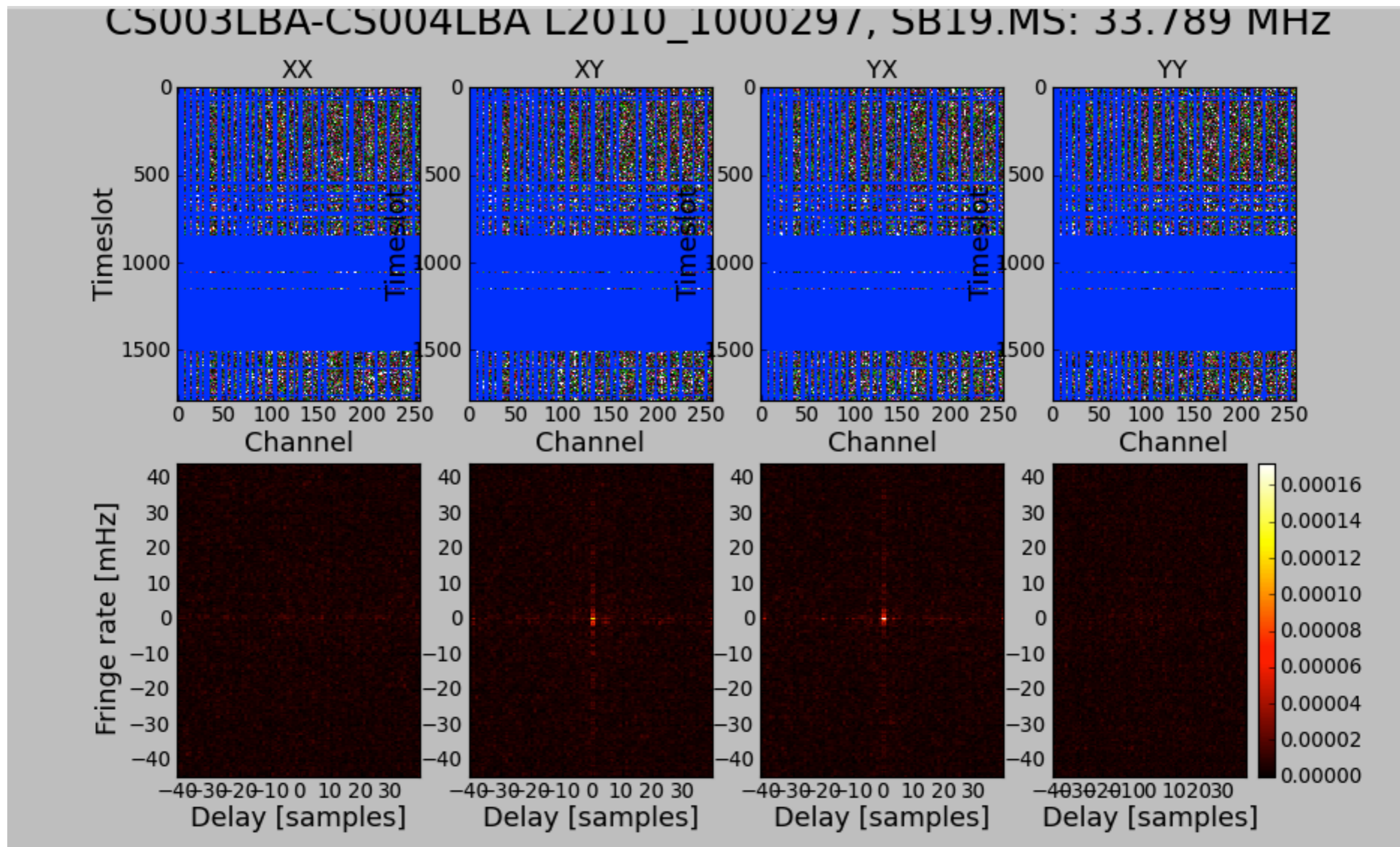
- Test experiment: several superterp stations, some in LBA_INNER and some in LBA_OUTER, correlated together in same experiment.



Two LBA_OUTER stations...

LBA_INNER vs. LBA_OUTER

- Test experiment: several superterp stations, some in LBA_INNER and some in LBA_OUTER, correlated together in same experiment.



One LBA_INNER and one LBA_OUTER station!

Therefore ...

**(0 1
1 0)**

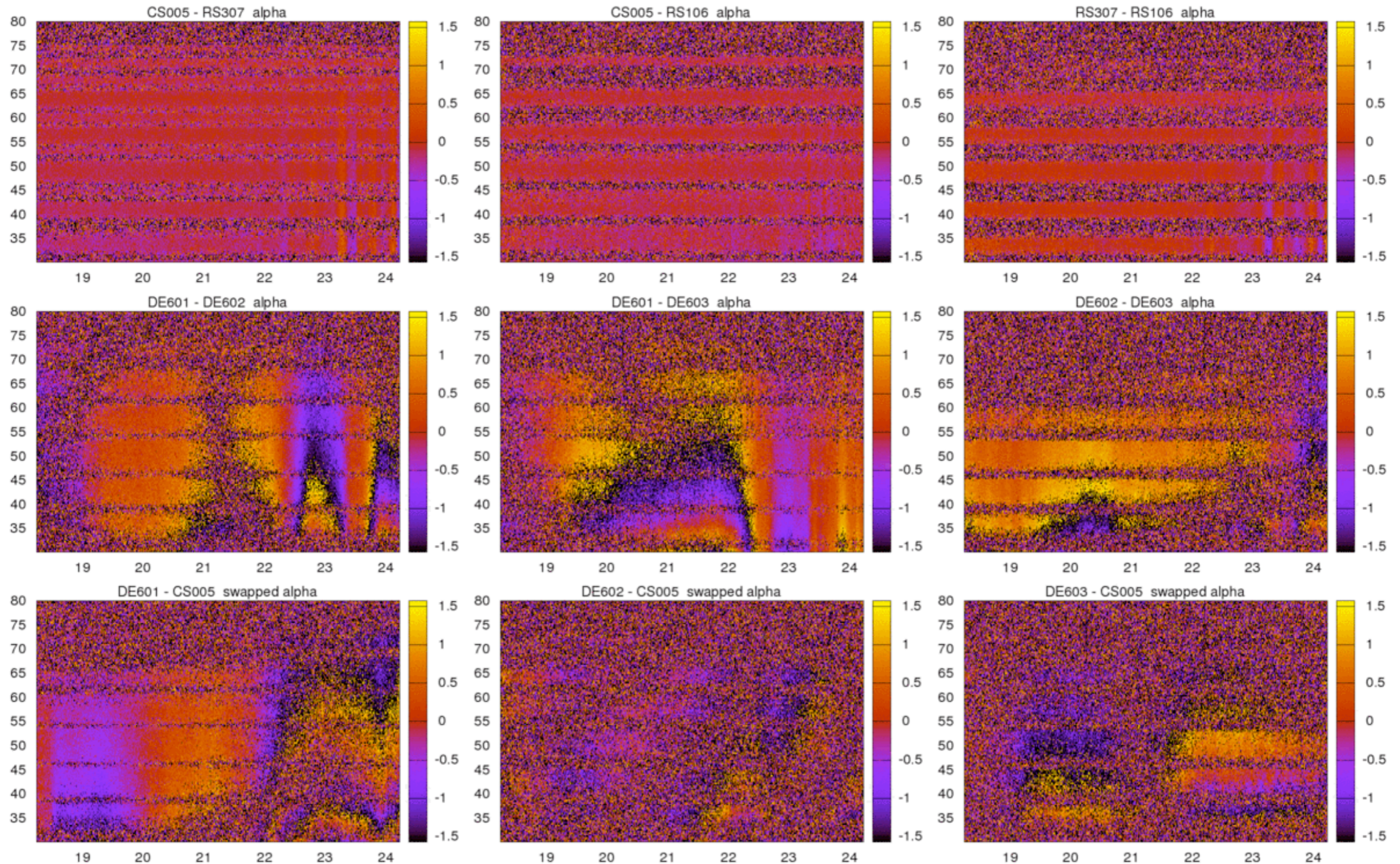
Therefore ...

... we apologize for blaming Tautenburg

(0 1
1 0)

... but wait, there's more!

- The X,Y flip is not the only effect. There is still differential Faraday rotation:



Differential Faraday rotation

- Joris has implemented a Faraday rotation Jones matrix in BBS
- To verify the need for solving F-Jones in this case, look to the data!
- Solve for all 4 elements of Gain (G) in BBS

$$V = G_i C G_j^\dagger$$

- Ansatz: Faraday rotation part is separable, and remainder is diagonal:

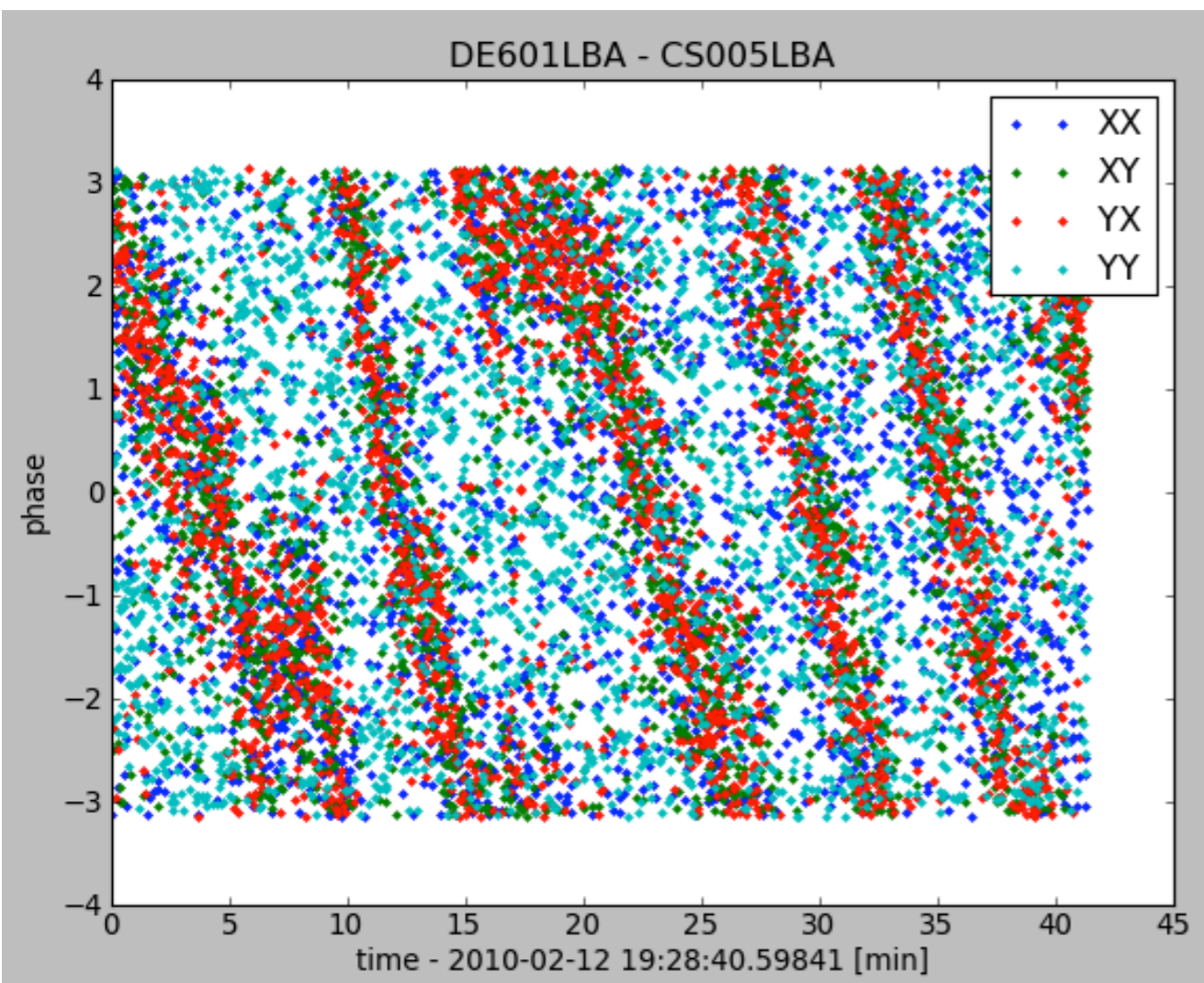
$$G = \begin{pmatrix} G_{00} & G_{01} \\ G_{10} & G_{11} \end{pmatrix} = \begin{pmatrix} G'_{00} & 0 \\ 0 & G'_{11} \end{pmatrix} \begin{pmatrix} \cos(\phi\lambda^2) & -\sin(\phi\lambda^2) \\ \sin(\phi\lambda^2) & \cos(\phi\lambda^2) \end{pmatrix}$$

- In this case, combinations of the parameters that BBS solves for should give an estimate of the rotation measure:

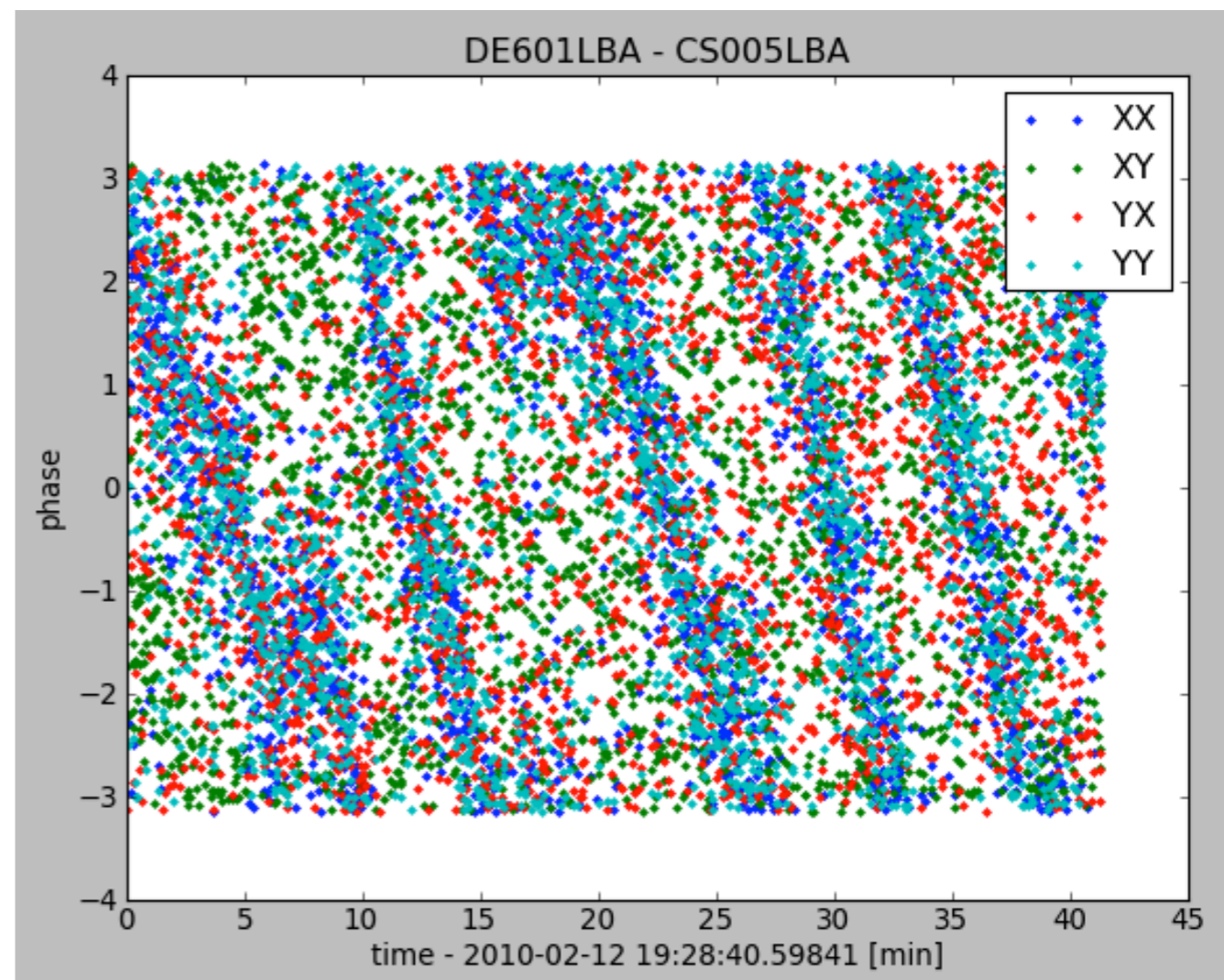
$$\frac{G_{01} - G_{10}}{G_{00} + G_{11}} = \tan(\phi\lambda^2)$$

Faraday rotation?

- Calibrated using BBS, after:
 - running NDPPP to flag, and then compress in frequency
 - splitting out a small (~ 40 min) time range
 - eliminating DE603LBA and RS503LBA
 - flipping X,Y for DE601 - (CS*/RS*): \sim seven lines of python, using pyrap!



BEFORE



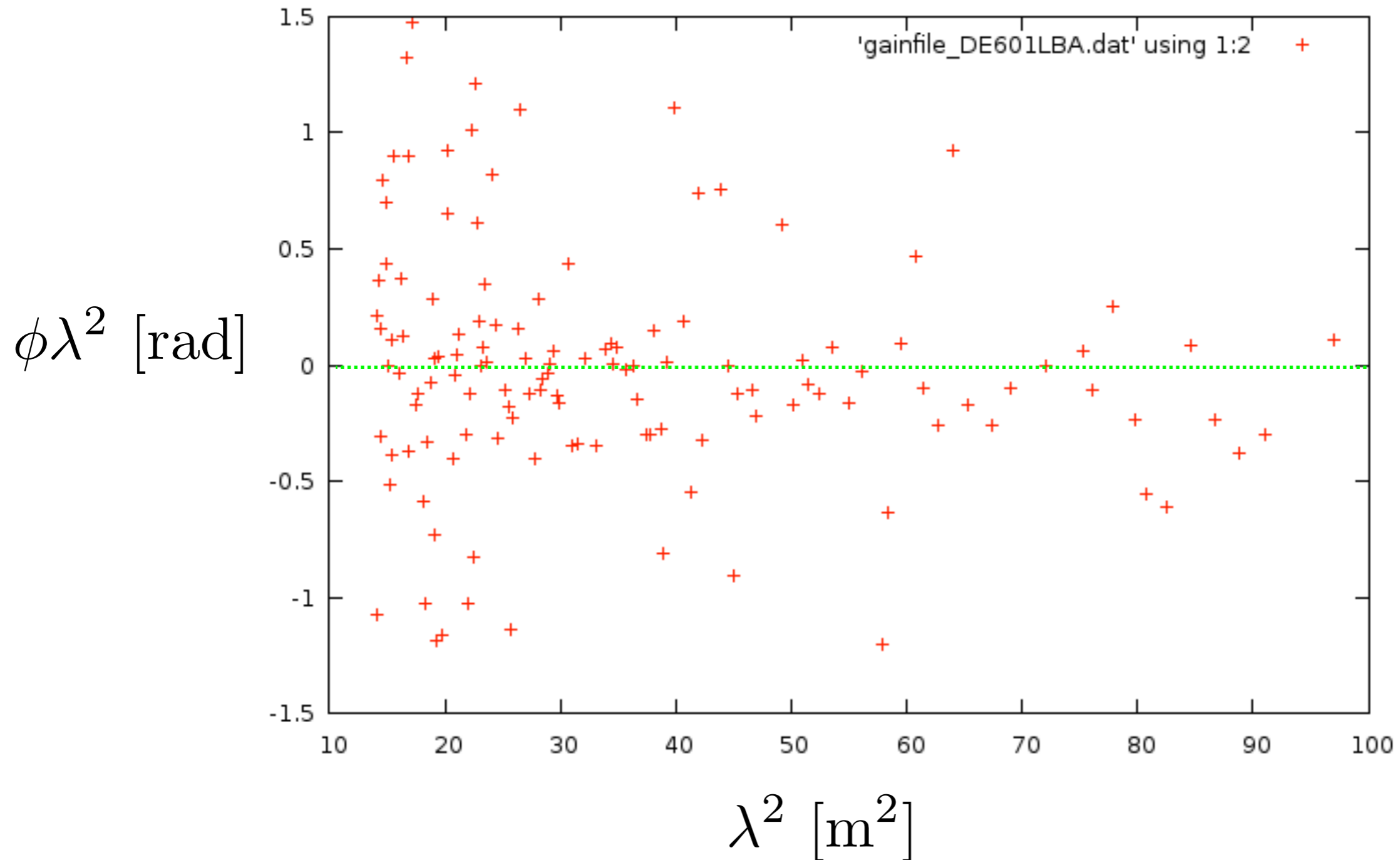
AFTER

Faraday rotation?

- Calibrated using BBS, after:
 - running NDPPP to flag, and then compress in frequency
 - splitting out a small (~ 40 min) time range
 - eliminating DE603LBA and RS503LBA
 - flipping X,Y for DE601 - (CS*/RS*): \sim seven lines of python, using pyrap
- Used following strategy:
 - use CS005LBA as the reference station
 - hold reference phase fixed to 0 for diagonal G terms
 - hold offdiagonal G terms fixed to 0
 - fit all four terms of G for every other station, using a 20sec time interval, and an independent solution for each (compressed) subband
 - inspect fitted gains as function of frequency

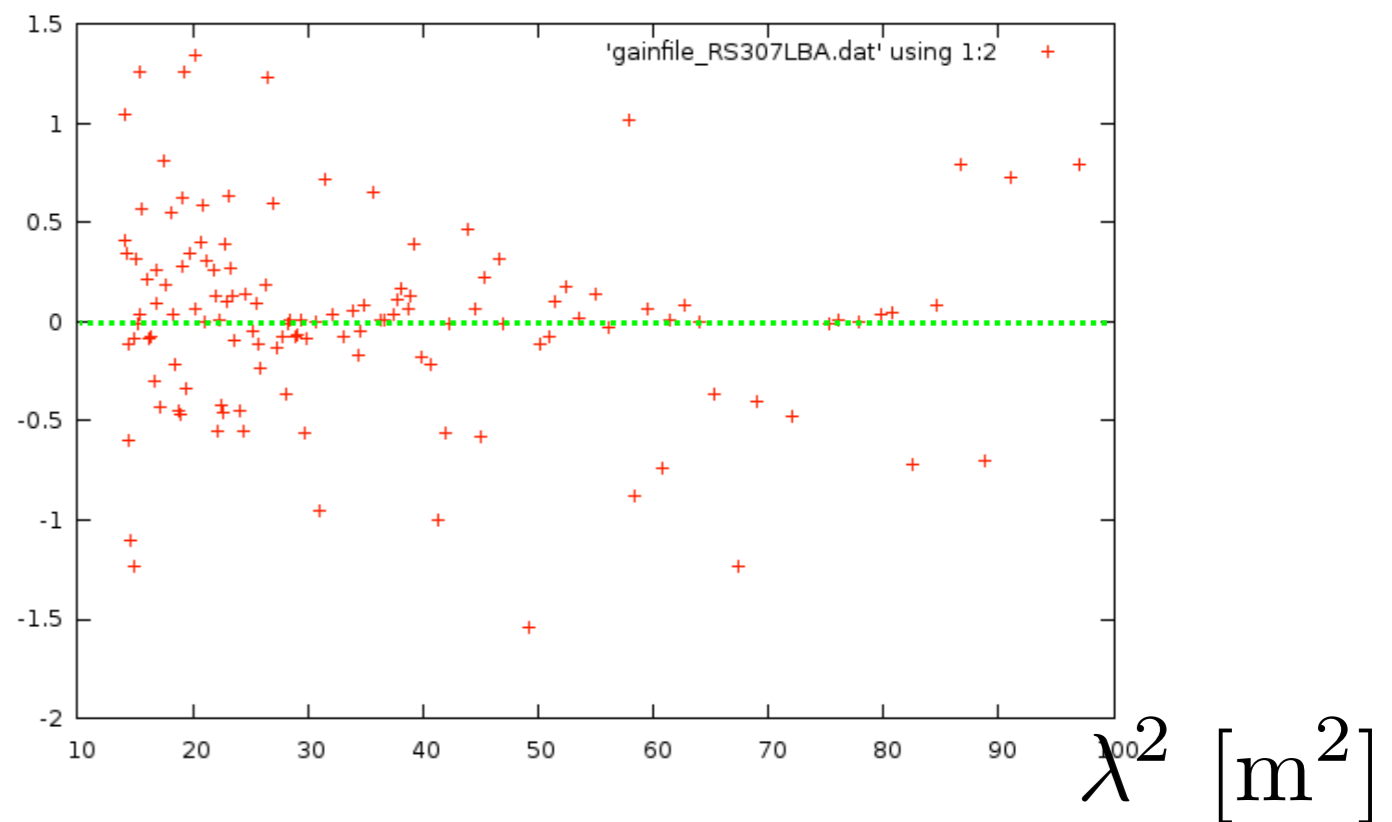
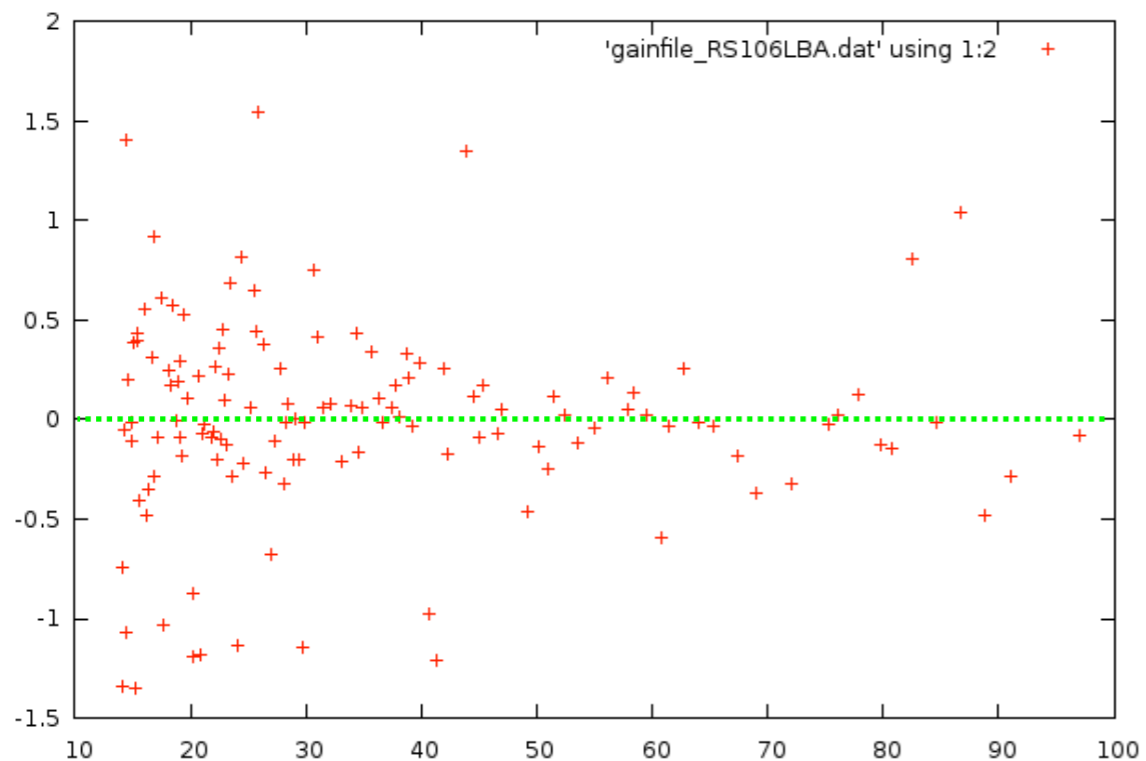
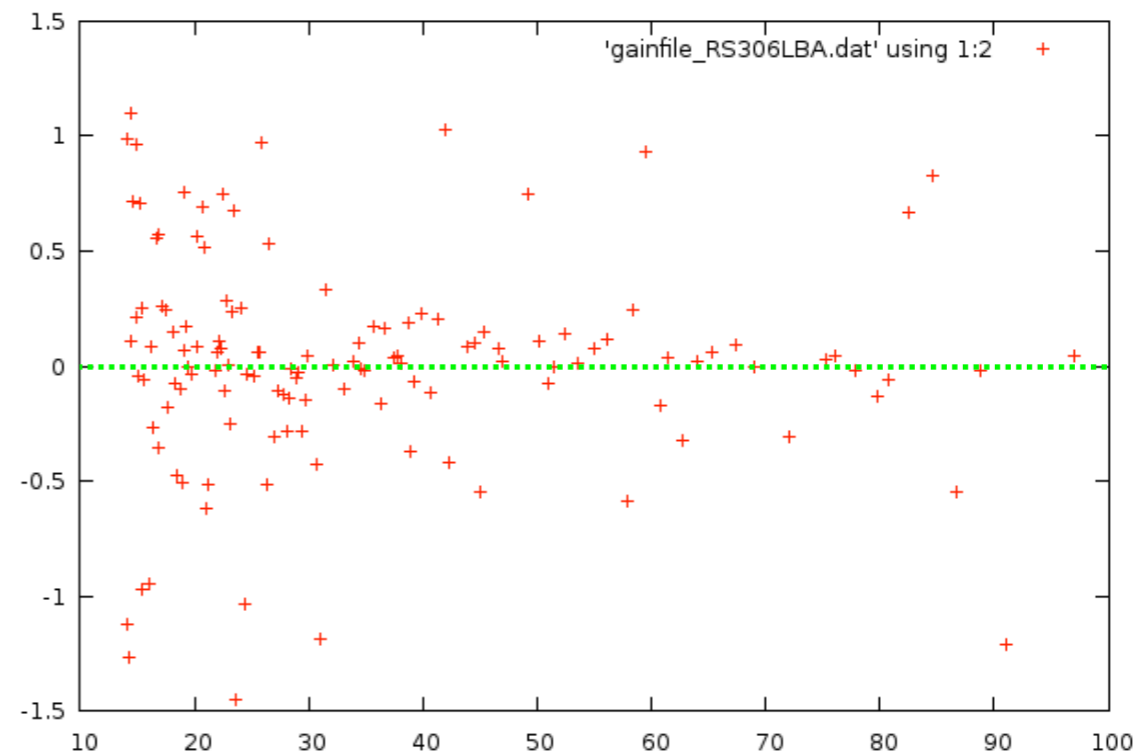
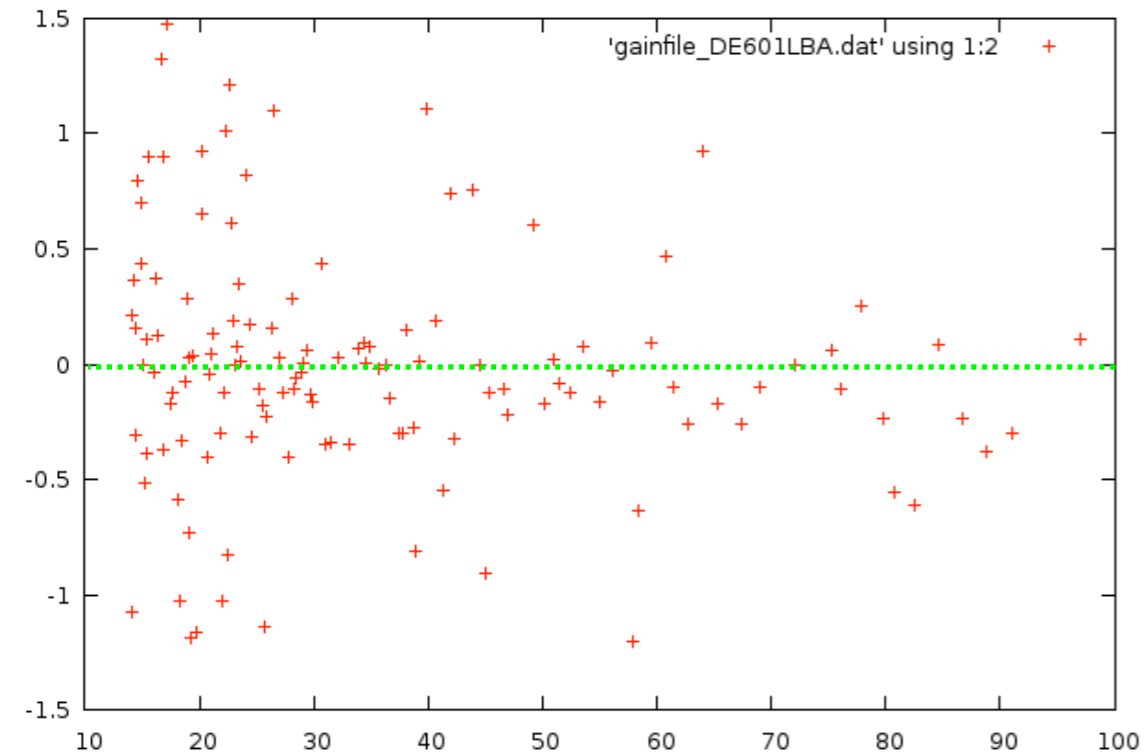
Faraday rotation?

- The time range that I picked seems to have little to no differential Faraday rotation between Effelsberg and Exloo:



Faraday rotation?

- and also none between Dutch stations

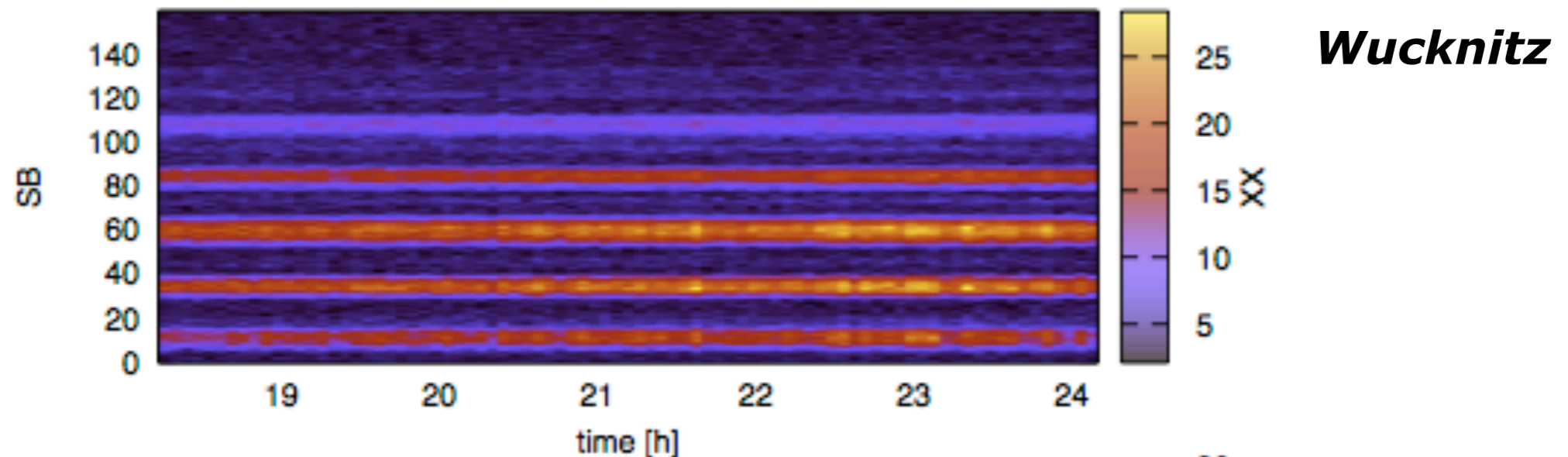
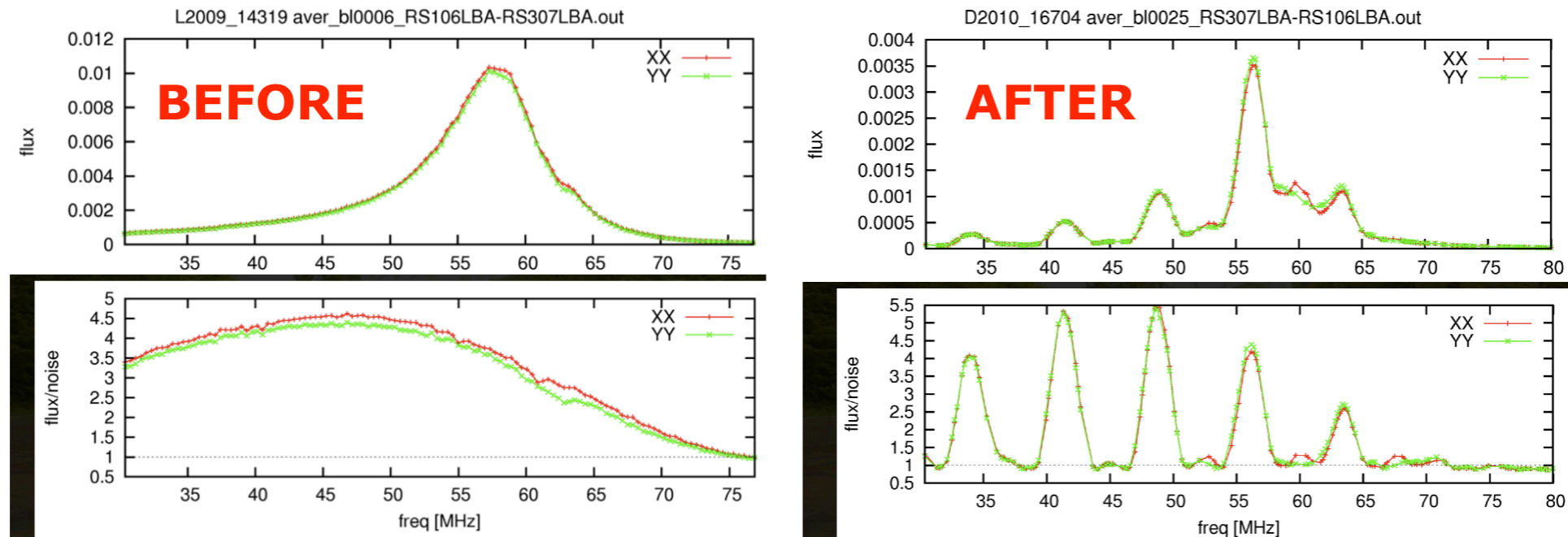


Faraday rotation?

- To do:
 - Find a more appropriate dataset / time range. Which data set did Sarod use to extract differential Faraday rotation in MeqTrees?
 - Use a strong polarized source

8 MHz "ripple"

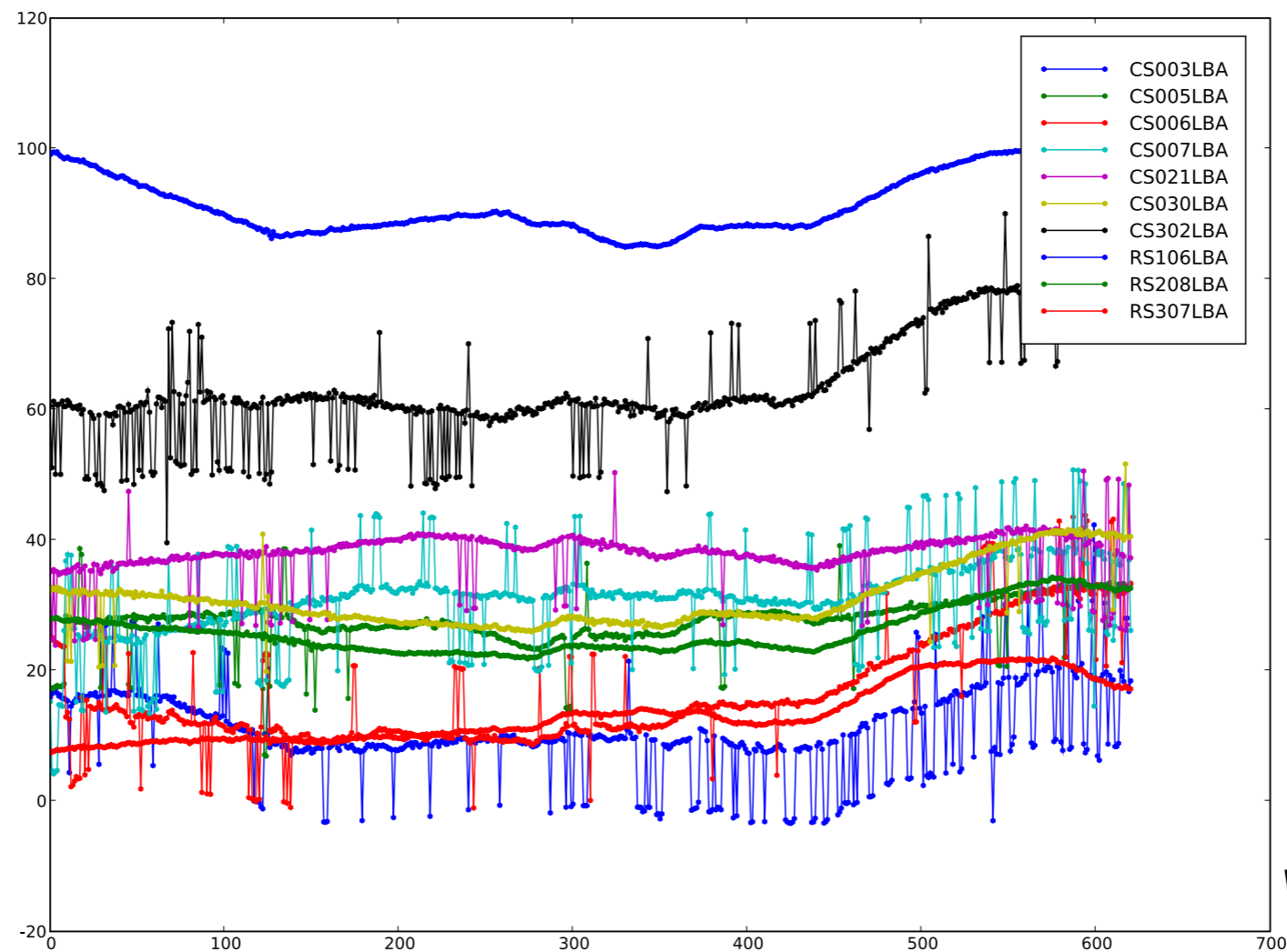
- Olaf detected a strong sensitivity variation with frequency, period ~ 8 MHz.



- Problem began mid-January, and is roughly consistent with different cable lengths in the stations. **Did we stop correcting for cable length?**

Clock offsets

- Are observed $\sim 100\text{ns}$ - $1\mu\text{s}$ "clock errors" really due to clock problems, or are they delays?



- Future observation strategy under development to
 - find out if "clock errors" are constant with time, or variable
 - compare their values with properties of individual stations e.g. cable lengths