

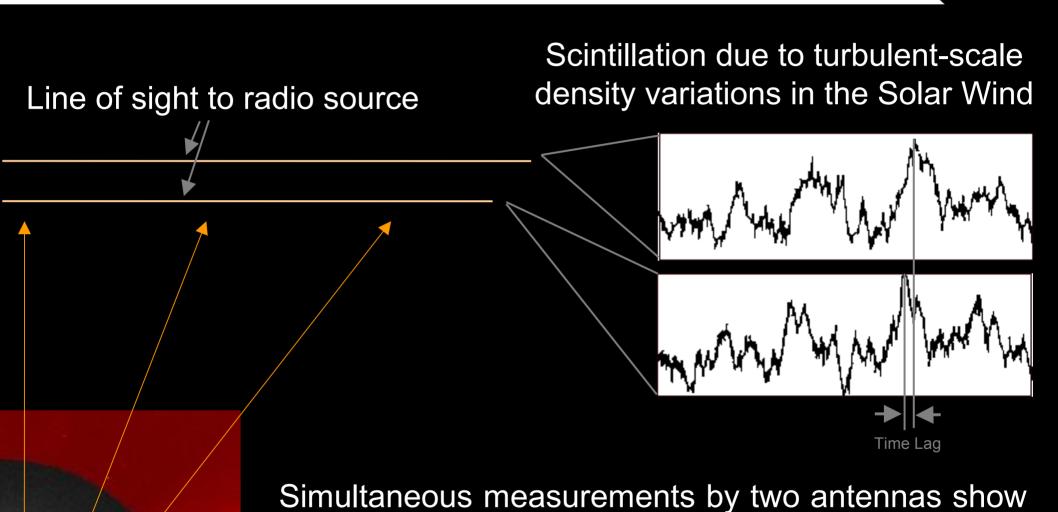


Netherlands Institute for Radio Astronomy



#### Radio Measurements



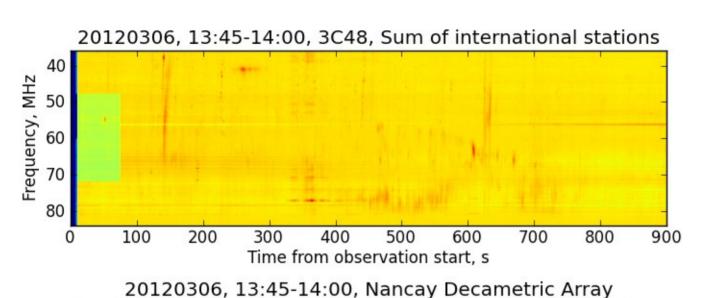


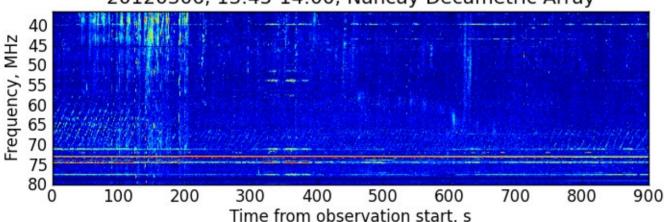
similar patterns of scintillation.

Time-lag for maximum cross-correlation gives estimate of solar wind outflow speed.

# Solar Flare Activity - Comparison with Nançay Decametric Array



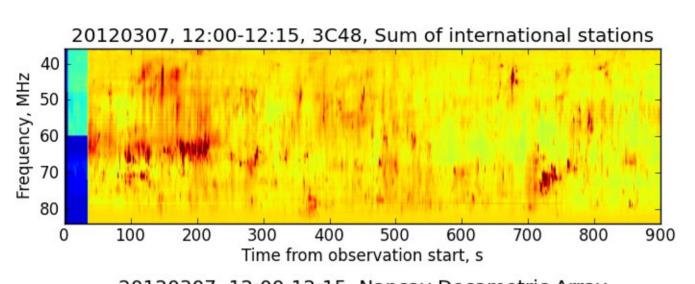


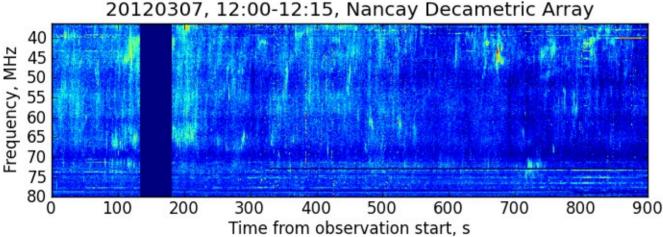


LOFAR IPS data averaged down to match resolution of Nancay (1.24s).

# Solar Flare Activity - Comparison with Nançay Decametric Array





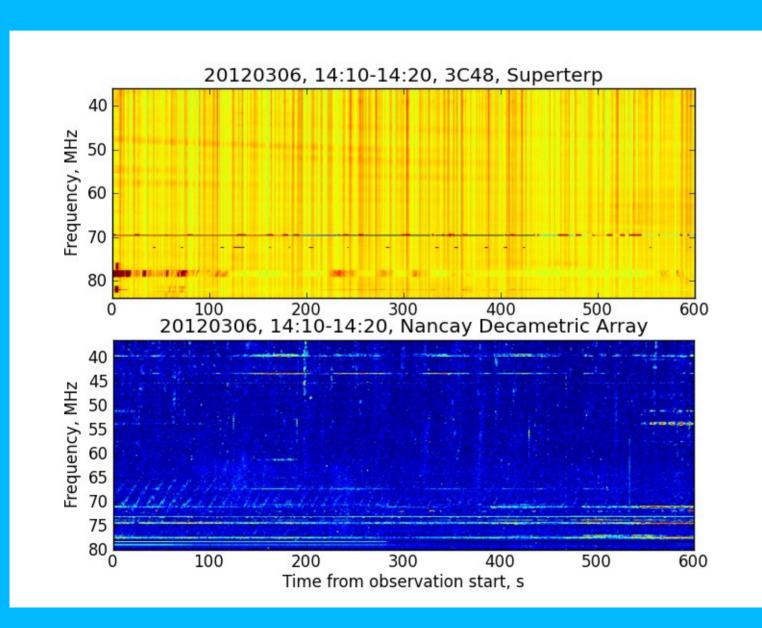


LOFAR IPS data averaged down to match resolution of Nancay (1.24s).

A lot of radio activity seen by both instruments.

# Solar Flare Activity – Comparison with Nancay Decametric Array





LOFAR IPS data averaged down to match resolution of Nancay (1.24s).

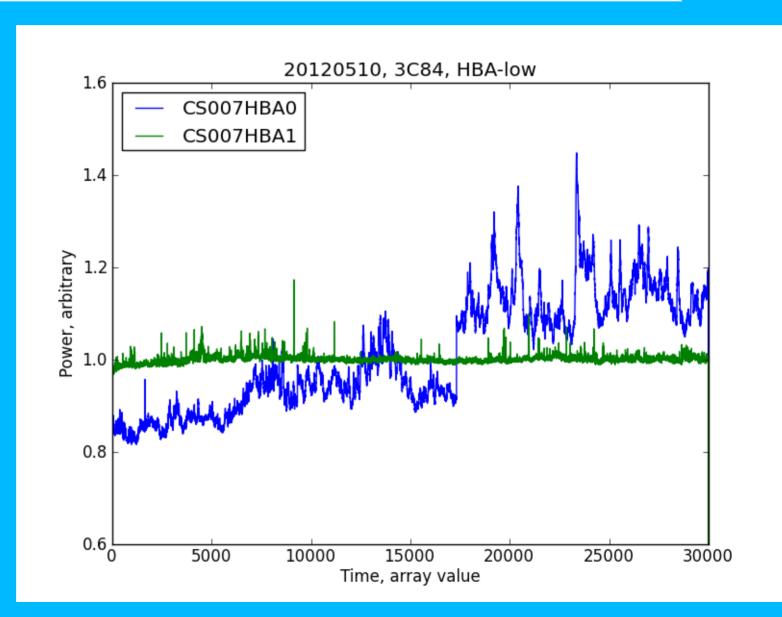
IPS signal (vertical stripes) seen very clearly in Superterp, as it should be, and no sign of the Sun!

## Solar Flare Activity – CS007



Time series' of observation of 3C84 taken with Superterp in Fly's Eye mode show evidence of some solar activity in CS007HBA0.

Little sign of activity in other stations, including CS007HBA1.

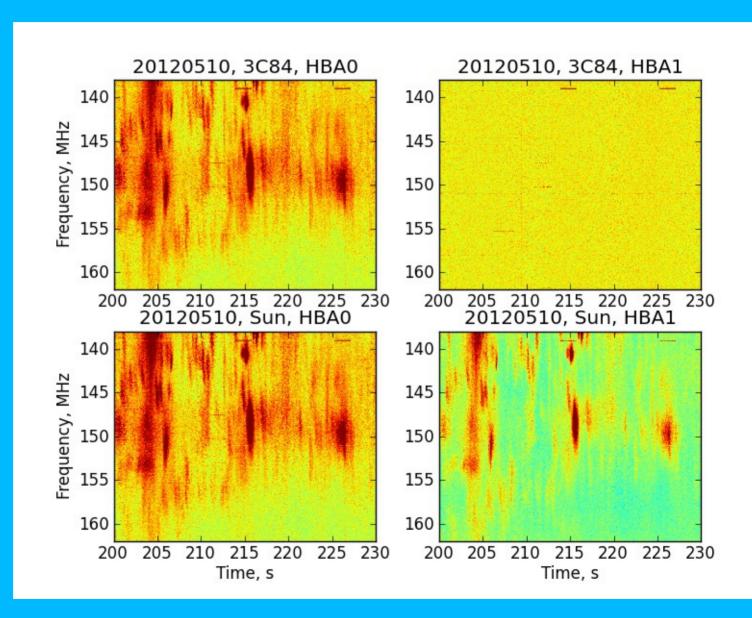


### Solar Flare Activity - CS007



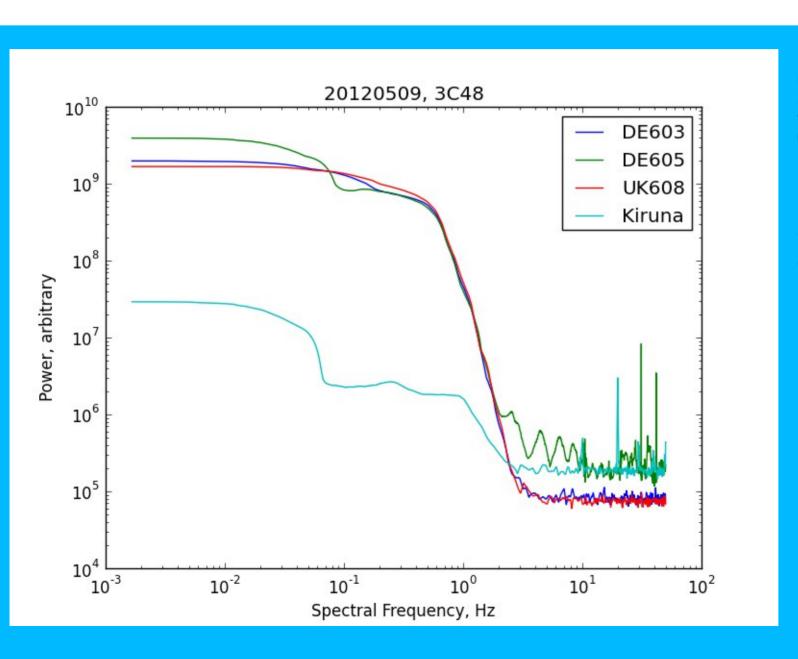
Second beam pointed at the Sun for all early May IPS observations:

Dynamic spectra confirm that Sun is seen in CS007HBA0.



### **EISCAT Comparison**



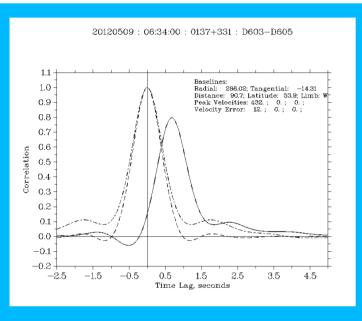


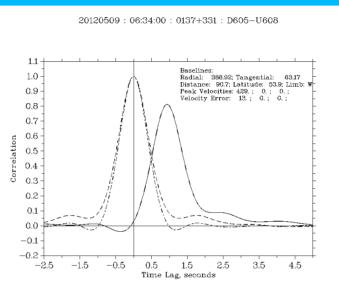
Observation of 3C48: 20120509 06:35 to 06:45 UT

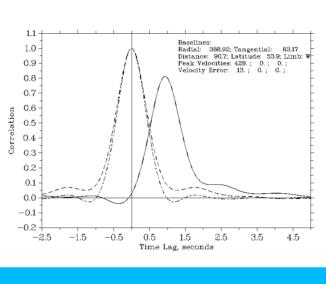
Plot of power spectra from different stations

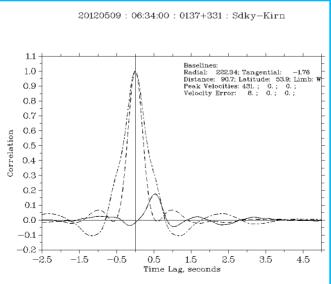
#### **EISCAT Comparison**











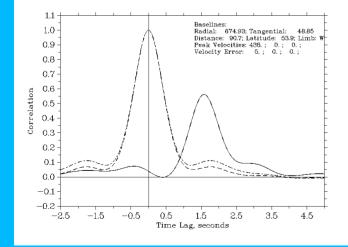
Observation of 3C48: 20120509 06:35 to 06:45 UT

Top left: DE603-5 Speed 432 km/s

Top right: DE605-UK608 Speed 429 km/s

Lower left: DE603-UK608 Speed 436 km/s

Lower right: Kir-Sod Speed 431 km/s



20120509: 06:34:00: 0137+331: D603-U608

## Summary



- Comparisons with full-resolution data from the Nancay
  Decametric Array confirm solar flare activity seen in March
  IPS observations.
- In observations in May, a second beam was placed on the Sun:
  - Approximately half the observations affected by flare activity.
  - CS007HBA0 strongly affected in one observation, other
    Superterp stations not affected as strongly.
- First comparisons of cross-correlation functions with EISCAT:
  - Estimated solar wind speeds match exactly!