

ASTRON

Radboud University



**university of
 groningen**



**Vrije
 Universiteit
 Brussel**

Cosmic-Ray Key Science Project : review on recent results and future perspective

Laura Rossetto

on behalf of the CR-KSP :

**A. Bonardi, S. Buitink, A. Corstanje, H. Falcke, B. Hare, J.R. Hörandel, P. Mitra, K. Mulrey, A. Nelles,
 J.P. Rachen, L. Rossetto, P. Schellart, O. Scholten, S. ter Veen, S. Thoudam, T.N.G. Trinh, T. Winchen**

- **P. Schellart et al., A&A 560, A98 (2013):** Detecting cosmic rays with the LOFAR radio telescope
- **P. Schellart et al., NIMPA 742, 115 (2014):** Recent results from cosmic-ray measurements with LOFAR
- **P. Schellart et al., JCAP 10, 014 (2014):** Polarized radio emission from extensive air showers measured with LOFAR
- **S. Buitink et al., PRD 90, 082003 (2014):** Method for high precision reconstruction of air shower Xmax using two-dimensional radio intensity profiles
- **S. Thoudam et al., NIMPA 767, 339 (2014):** LORA – A scintillator array for LOFAR to measure extensive air showers
- **A. Nelles et al., Aph 60, 13 – 24 (2015):** A parameterization for the radio emission of air showers as predicted by CoREAS simulations and applied to LOFAR measurements
- **A. Corstanje et al., Aph 61, 22 – 31 (2015):** The shape of the radio wavefront of extensive air showers as measured with LOFAR
- **P. Schellart et al., PRL 114, 165001 (2015):** Probing Atmospheric Electric Fields in Thunderstorms through Radio Emission from Cosmic-Ray-Induced Air Showers
- **A. Nelles et al., Aph 65, 11 – 21 (2015):** Measuring a Cherenkov ring in the radio emission from air showers at 110-190 MHz with LOFAR
- **A. Nelles et al., JCAP 05, 018 (2015):** The radio emission pattern of air showers as measured with LOFAR – a tool for the reconstruction of the energy and the shower maximum
- **A. Nelles et al., JInst 10, P11005 (2015):** Calibrated the absolute amplitude scale for air showers measured at LOFAR
- **S. Thoudam et al., Aph 73, 34 – 43 (2016):** Measurement of the cosmic-ray energy spectrum above 10^{16} eV with the LOFAR Radboud Air Shower Array
- **A. Corstanje et al., A&A 590, 41 (2016):** Timing calibration and spectral cleaning of LOFAR time series data
- **S. Buitink et al., Nature 531, 70 (2016):** A large light-mass component of cosmic rays at $10^{17} - 10^{17.5}$ eV from radio observations
- **T.N.G. Trinh et al., PRD 93, 023003 (2016):** Influence of Atmospheric Electric Fields on the Radio Emission from Extensive Air Showers

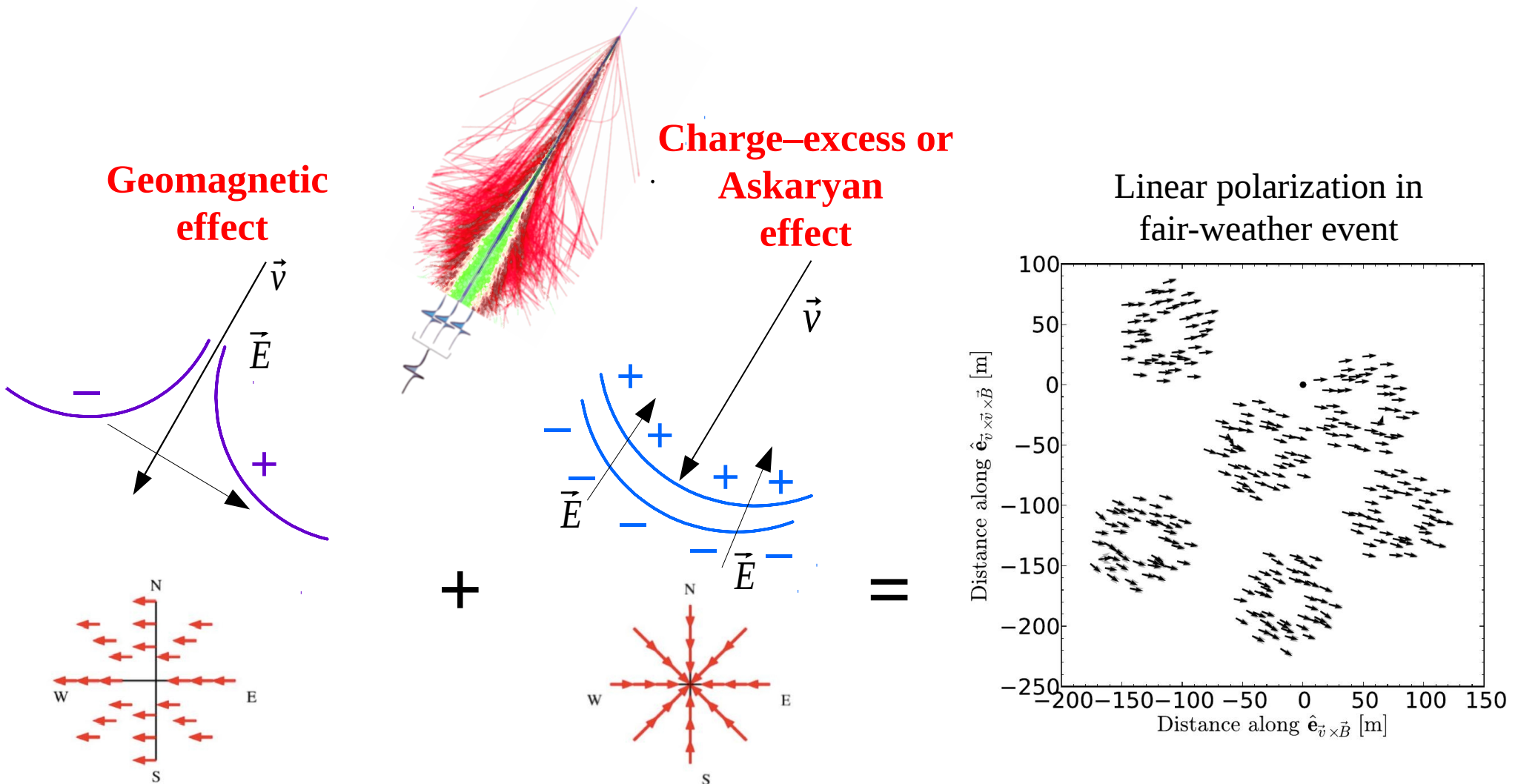
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RECENTLY published papers

- **O. Scholten et al., PRD 94, 103010 (2016):** Measurement of the circular polarization in radio emission from extensive air showers confirms emission mechanisms
- **T.N.G. Trinh et al., PRD 95, 083004 (2017):** Thunderstorm electric fields probed by extensive air showers through their polarized radio emission
- **A. Corstanje et al., Aph 89, 23 – 29 (2017):** The effect of the atmospheric refractive index on the radio signal of extensive air showers

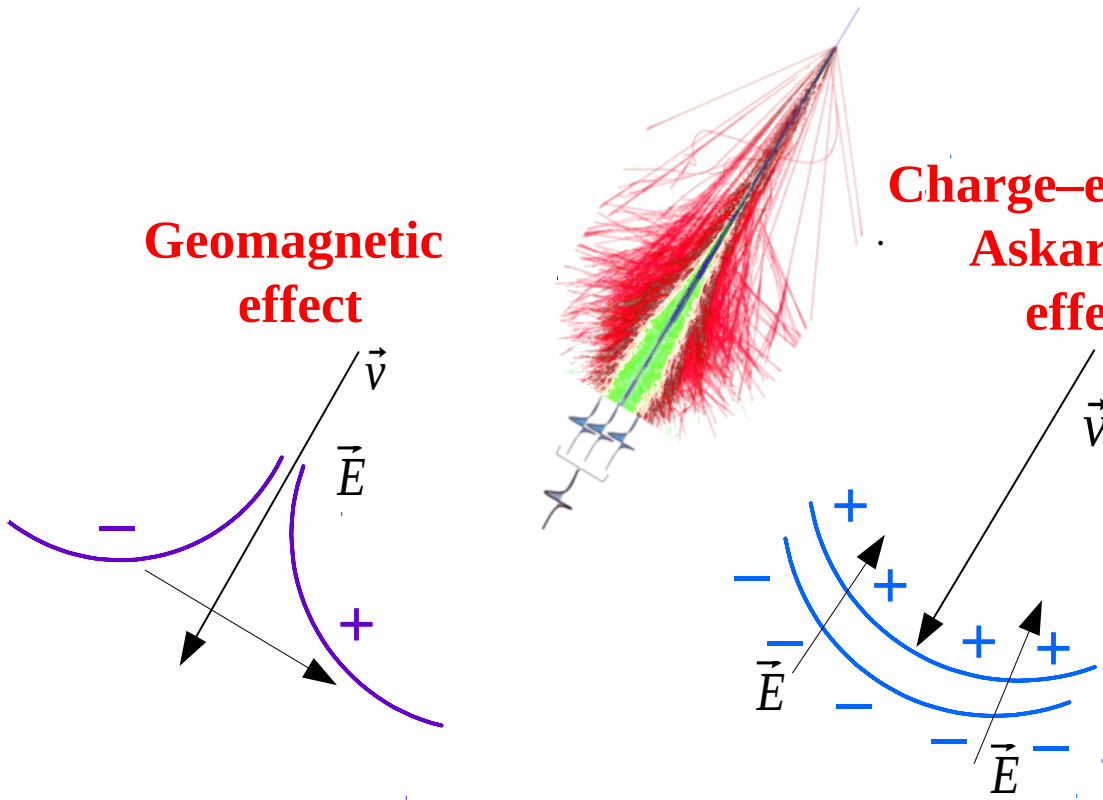
Measurements of circular polarization in fair-weather events

→ **O. Scholten et al., PRD 94, 103010 (2016)**: Measurement of the circular polarization in radio emission from extensive air showers confirms emission mechanisms



Measurements of circular polarization in fair-weather events

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Circular polarization due to a **time-delay** of the **Askaryan effect** of about **1 ns**

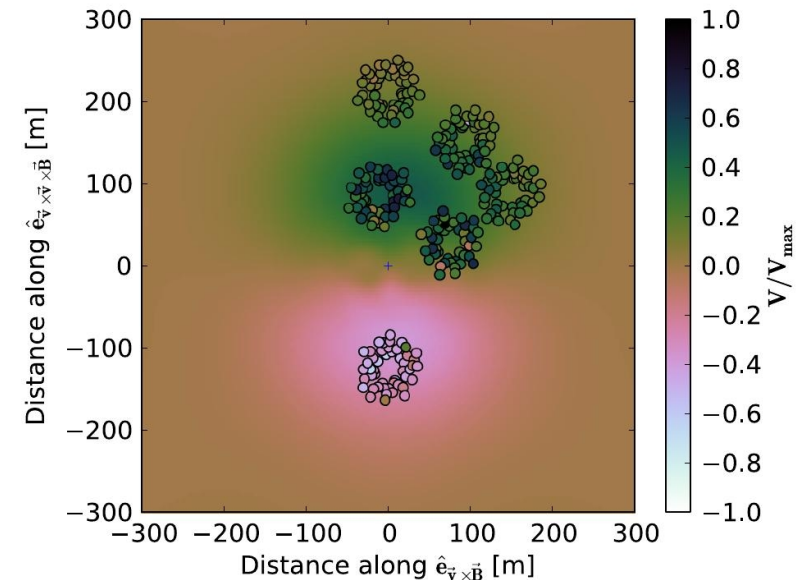
STOKES parameters:

I → intensity

Q, U → linear polarization

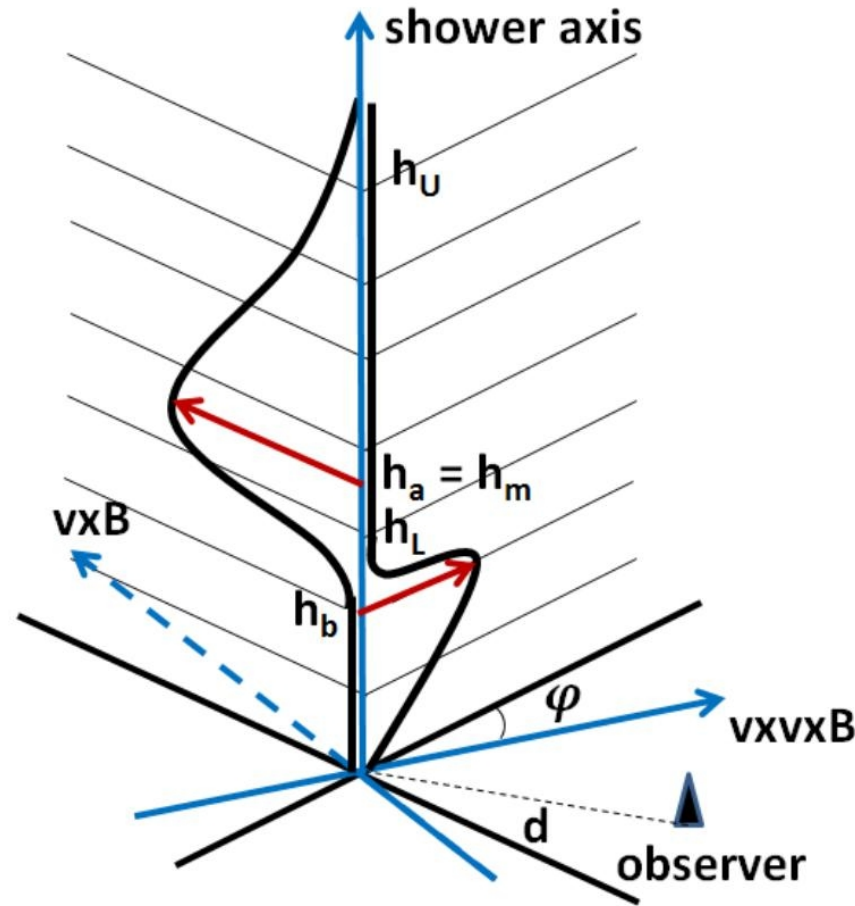
V → circular polarization

Circular polarization in fair-weather event



Measurements of circular polarization in thunderstorm events

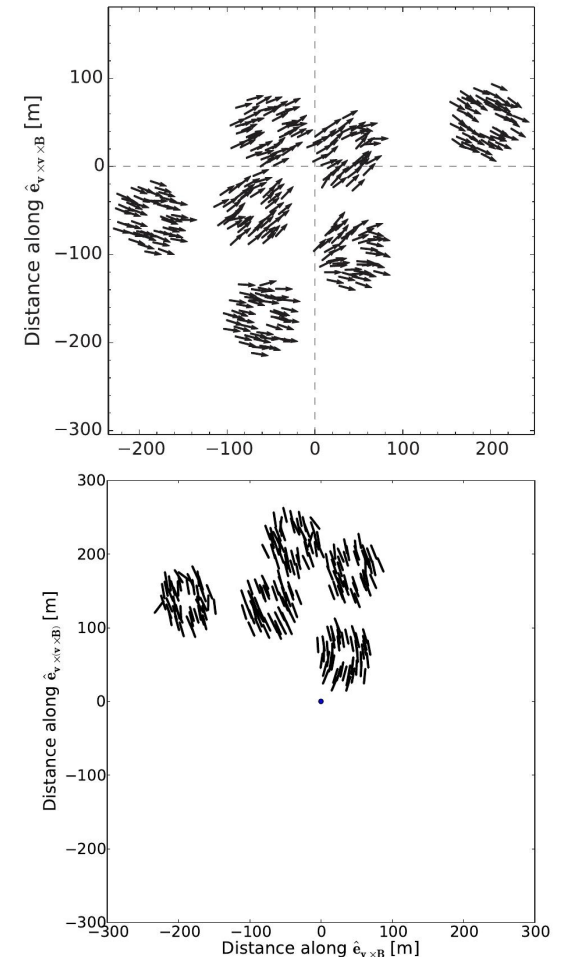
→ T.N.G. Trinh et al., PRD 95, 083004 (2017): Thunderstorm electric fields probed by extensive air showers through their polarized radio emission



The E-field changes direction at different altitudes

→ the **linear polarization** of the signal is not aligned in the $V \times B$ direction

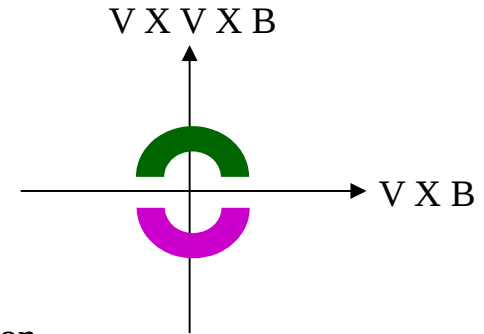
Linear polarization in fair-weather event



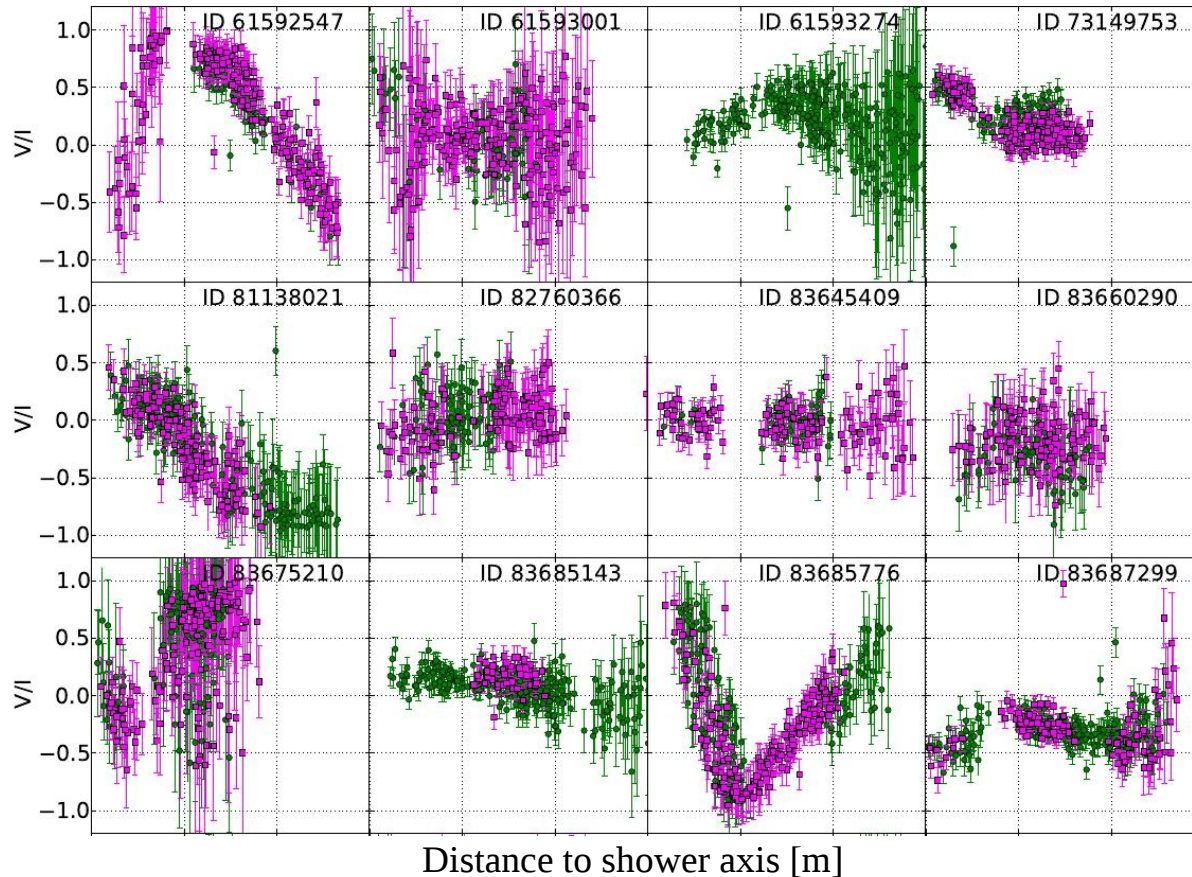
Measurements of circular polarization in thunderstorm events

→ **T.N.G. Trinh et al., PRD 95, 083004 (2017)**: Thunderstorm electric fields probed by extensive air showers through their polarized radio emission

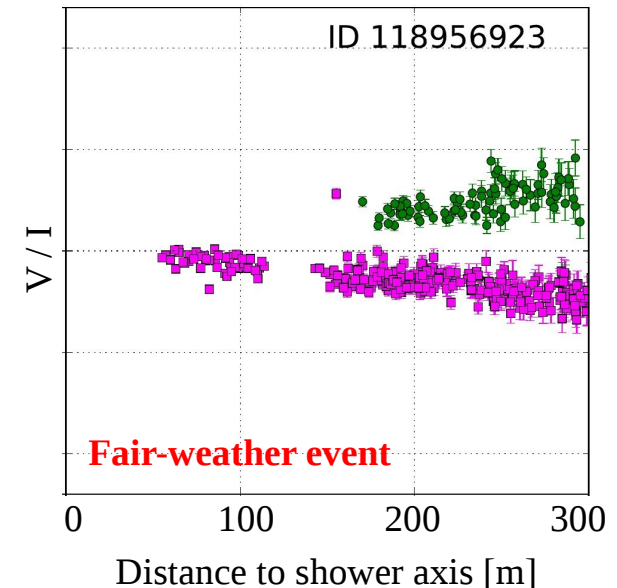
Circular polarization in thunderstorm is enhanced because of the transverse current direction changes at different altitudes



Thunderstorm events



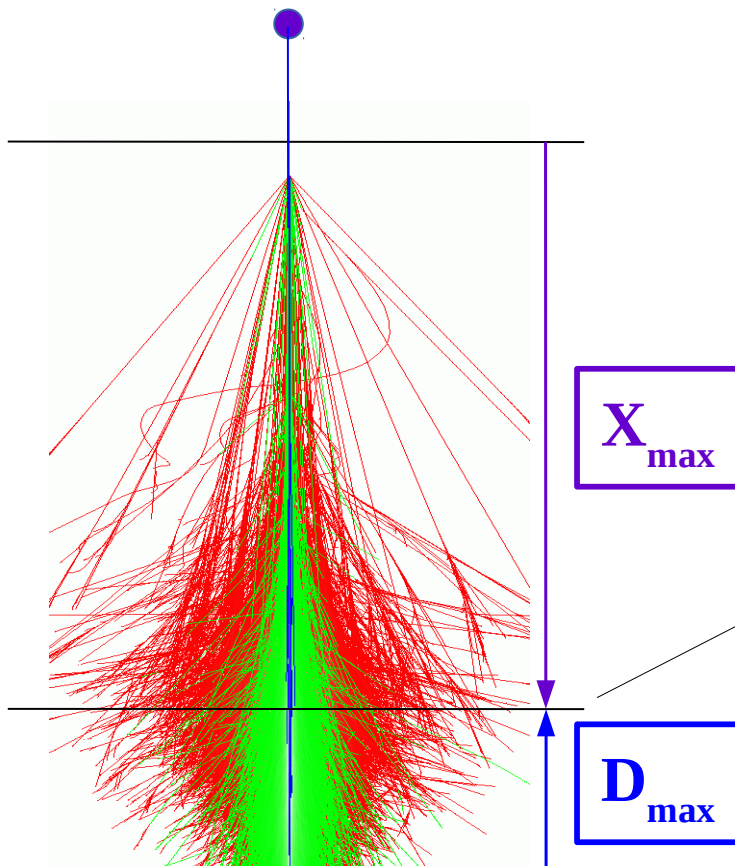
Antenna position
in the shower plane
azimuth = **0 – 180 deg**
azimuth = **180 – 360 deg**



Effect of atmospheric refractive index on mass composition measurements

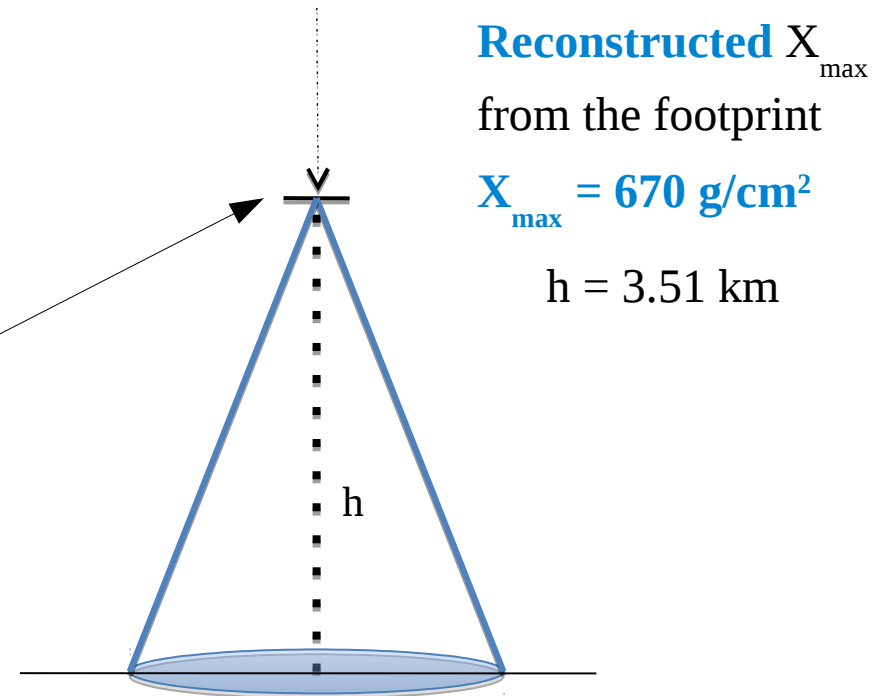
→ **A. Corstanje et al., Aph 89, 23 – 29 (2017)**: The effect of the atmospheric refractive index on the radio signal of extensive air showers

$X_{\max}(\text{protons}) > X_{\max}(\text{Fe})$
because $\sigma_{\text{proton}} < \sigma_{\text{Fe}}$



Simple simulation model

- assuming all radiation comes from X_{\max} level
- assuming the footprint size scales with the Cherenkov angle (Cherenkov angle model)

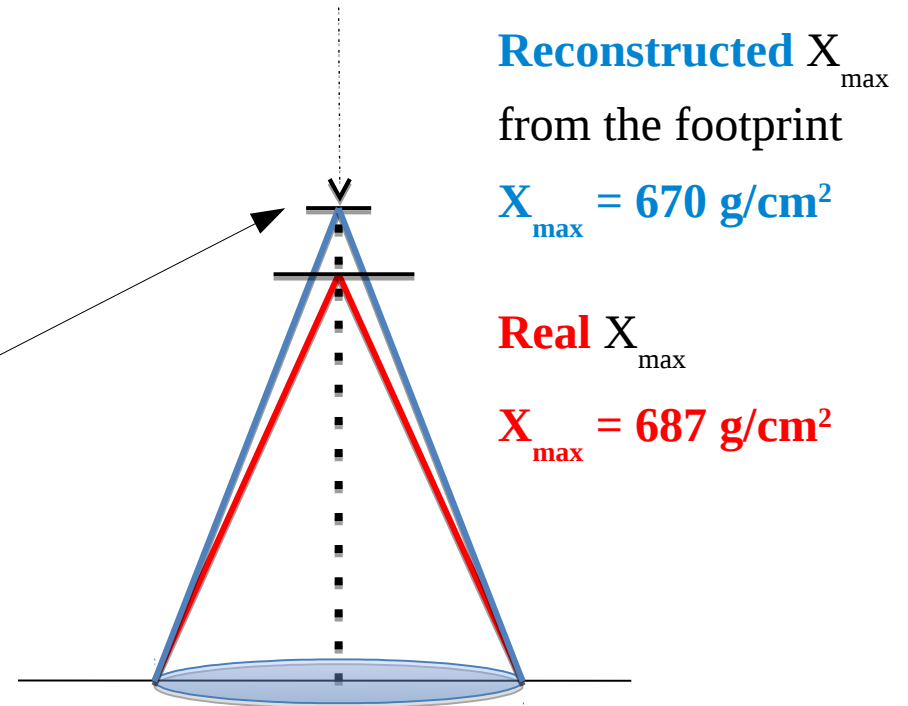
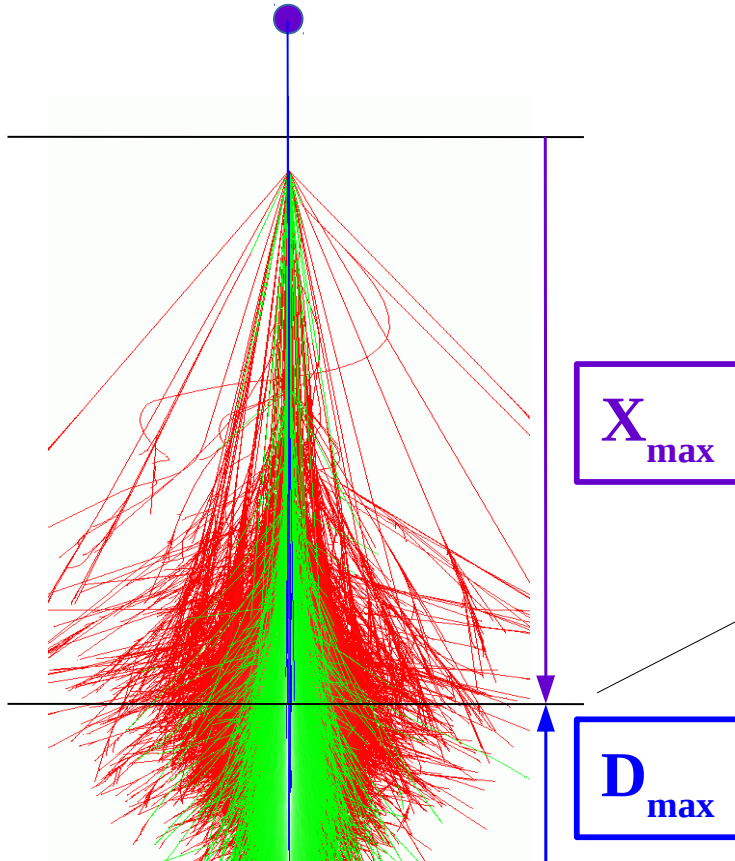


Effect of atmospheric refractive index on mass composition measurements

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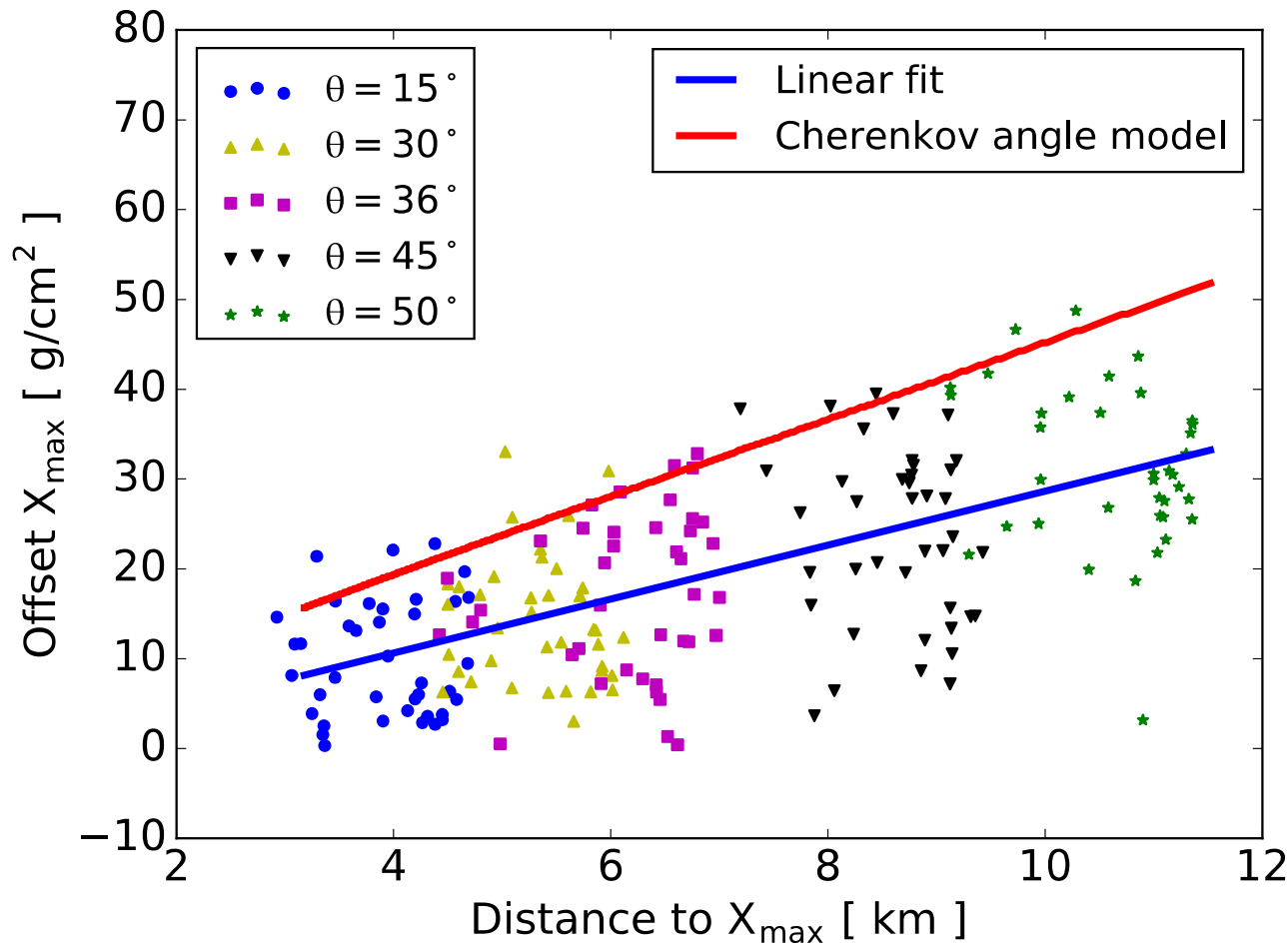
$X_{\max}(\text{protons}) > X_{\max}(\text{Fe})$
 because $\sigma_{\text{proton}} < \sigma_{\text{Fe}}$

- the reconstructed X_{\max} depends on the refractive index n (**higher n → smaller X_{\max}**)
- the refractive index depends on temperature, humidity and air pressure



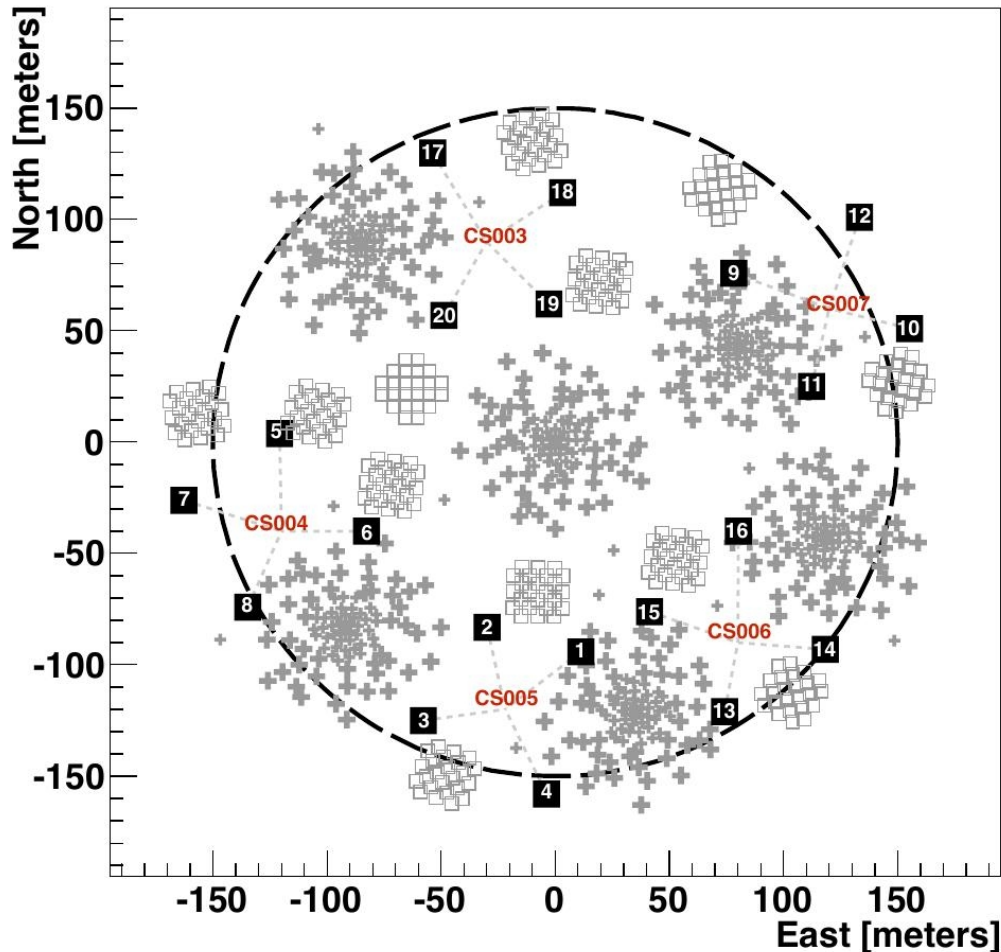
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Simulation results

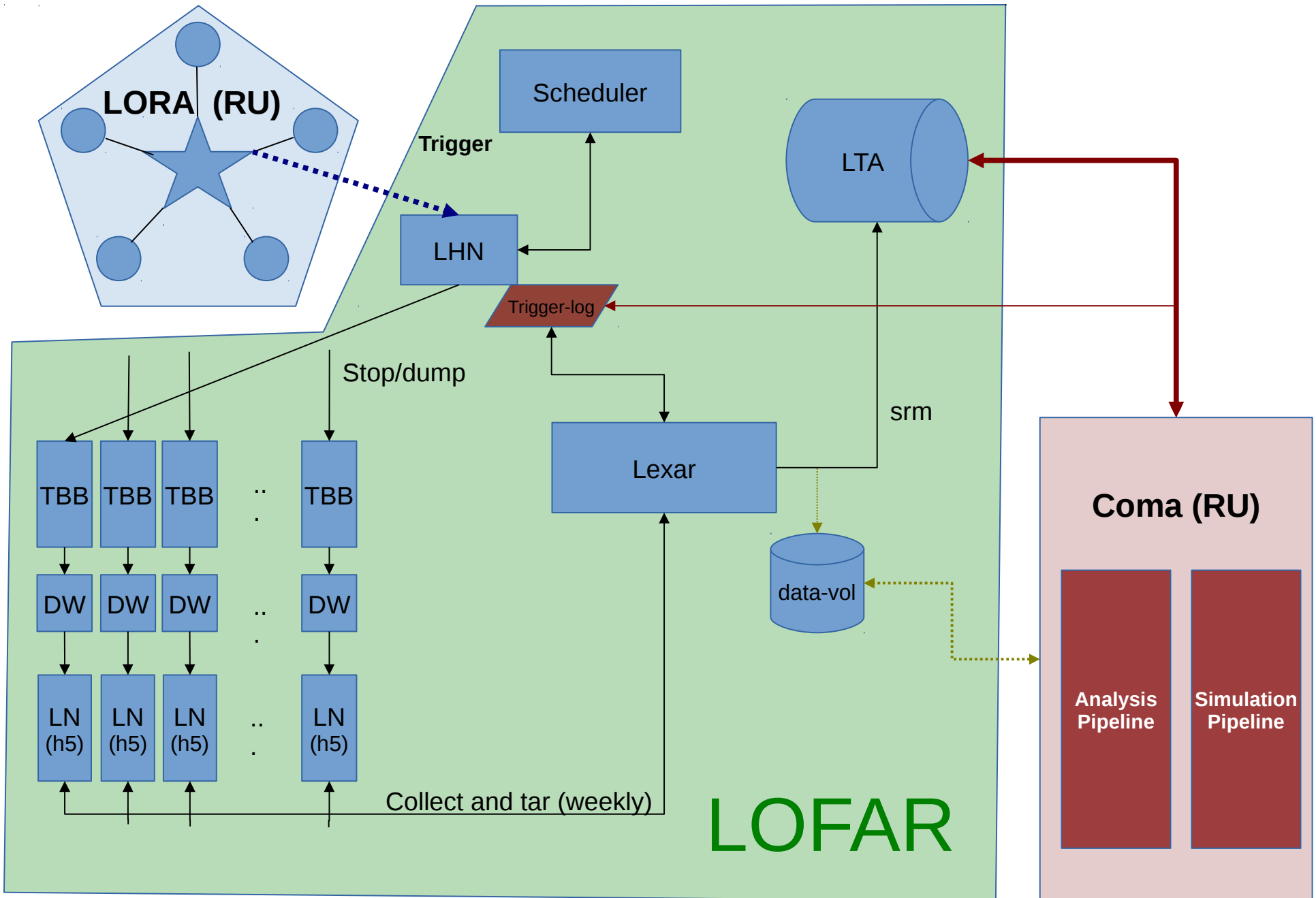
- a **10 % variation** of n gives a systematic error on the reconstructed X_{\max} of about **5 – 20 g/cm²**
- the systematic error on X_{\max} depends on the **zenith angle** of the arrival direction (up to 30 g/cm² for very inclined showers)
- **this affects composition analysis**

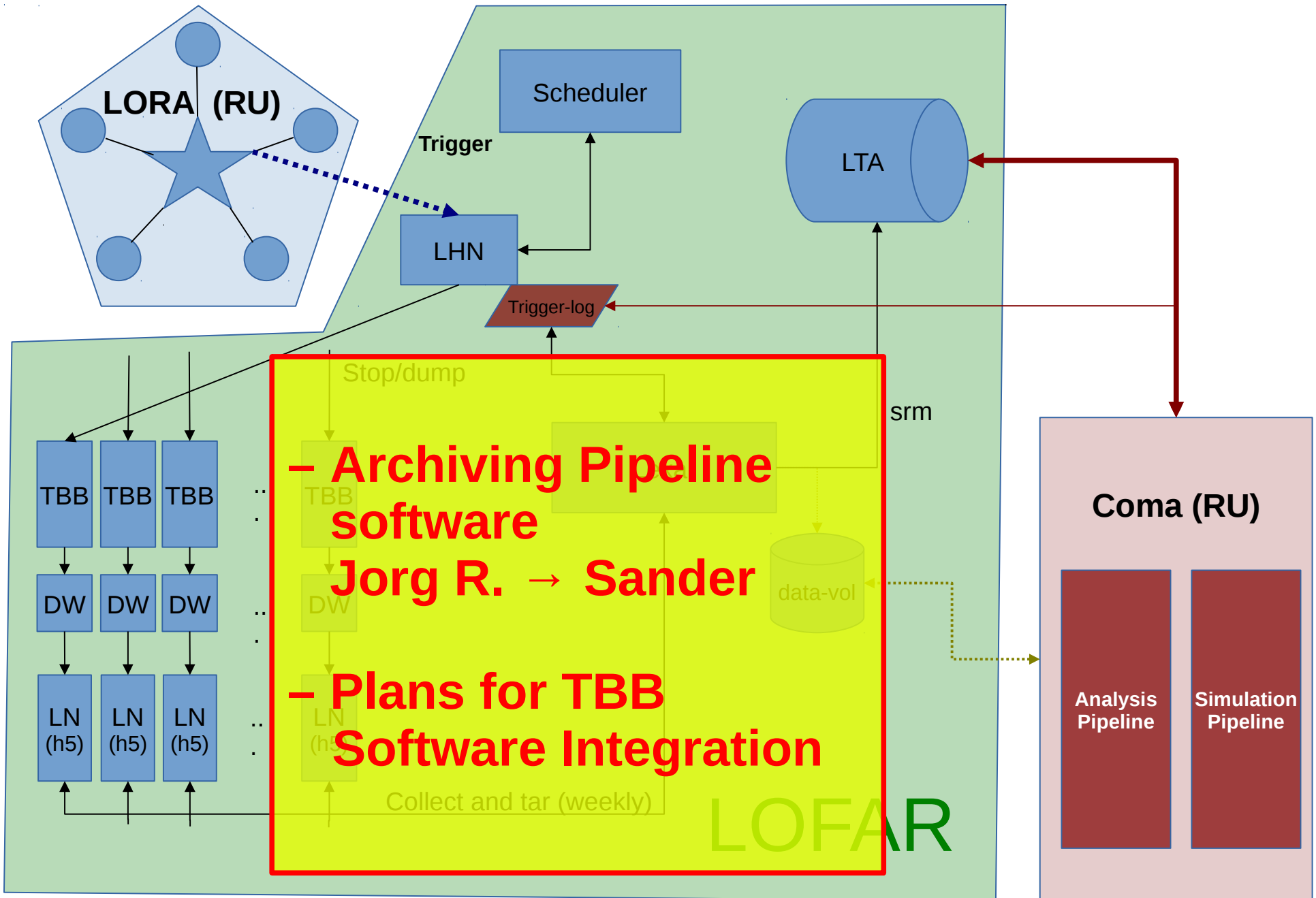


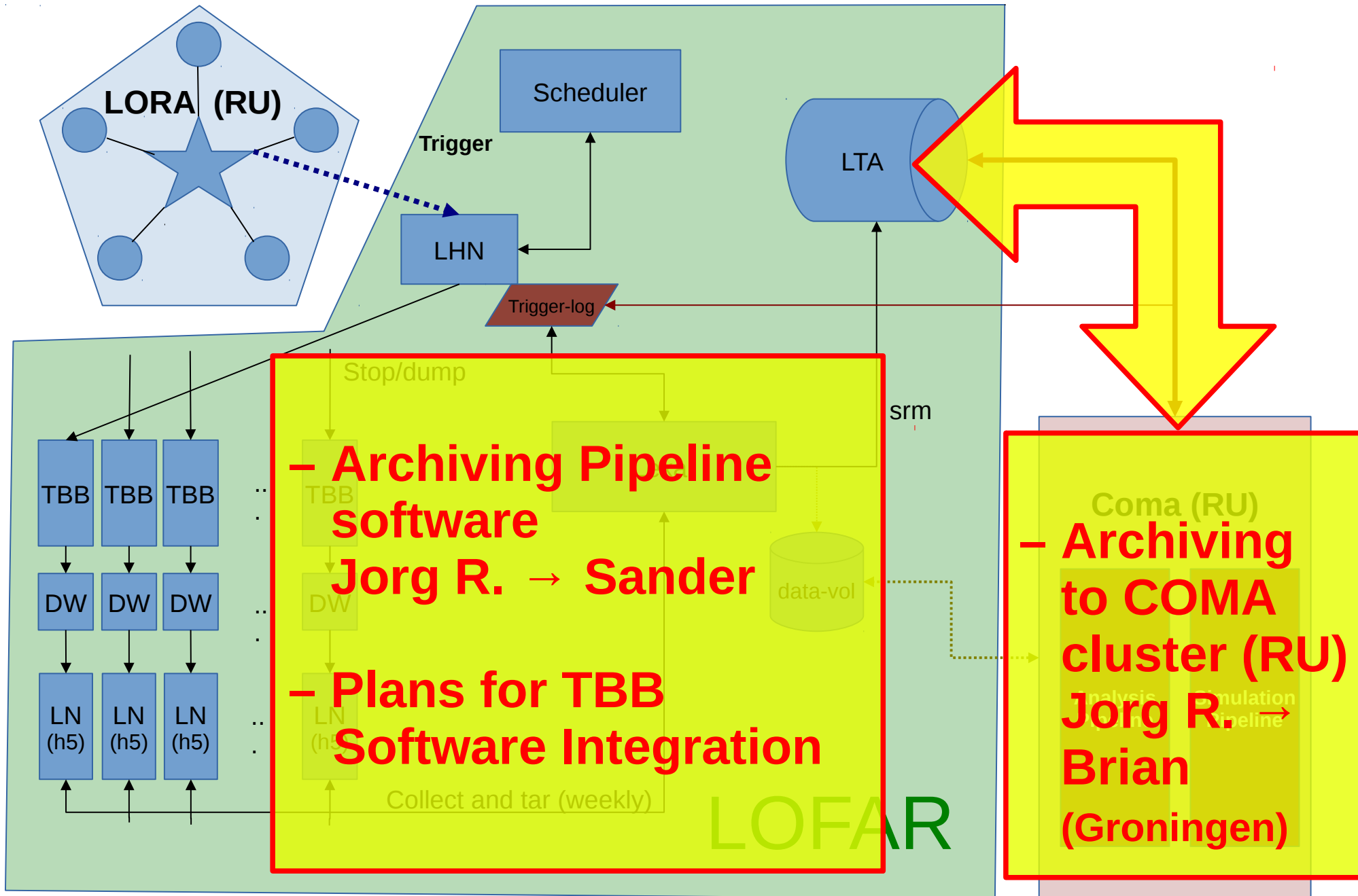
- **Campaign in June 2016:**
 - most **connectors** of the HV and signal cables have been replaced
 - HV and signal cables of detector 5 have been replaced

- **since June 2016** we are back to approximately **1 trigger / hour**

- **FUTURE plans for LORA:**
 - to upgrade/expand LORA with other 20 detectors
 - to implement an hybrid trigger (scintillator detectors + radio antennas)







Nijmegen group:

- **Frequency spectrum analysis** (includes repeating calibration analysis of the LBA antennas) – in preparation
- **Extended analysis on mass composition** (includes refractive index changes, atmospheric model, etc.) – in preparation
- study of a radio self-trigger for cosmic rays detection – includes data analysis of observation with the current self-trigger (a new data acquisition should happen soon) and development of a new ROUTINE

Groningen group:

- study of **circular polarization in fair-weather and thunderstorm events** and comparison with simulations – in preparation
- study of lightnings and air-showers propagation in thunderclouds

Brussel group:

- **simulation study of refractive index** using models for the whole atmosphere
- LORA extension with other 20 detectors and implementation of an hybrid trigger (scintillators + radio)
- NuMoon analysis – searching for radio pulses emitted by particles with energy $> 10^{21}$ eV hitting the Moon