



LOFAR Sky Models

2nd MSSS Meeting
21 August 2008

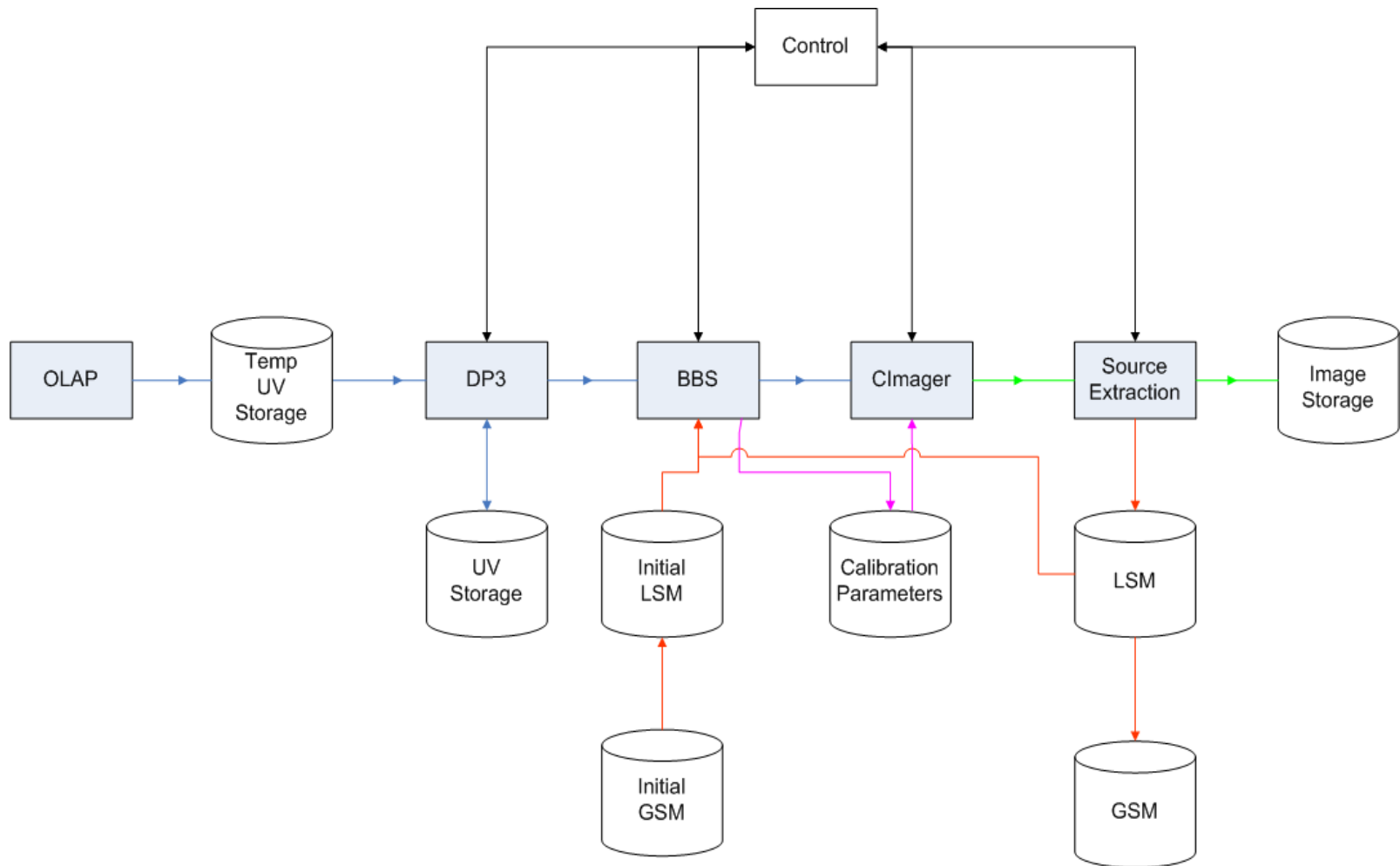
Michael Wise

(Contributions from R. Nijboer, N. Mohan, C. Law, B. Scheers, J. Swinbank, M. Bell)

LOFAR Sky Models

Outline

- **Integration**
- **Initialization**
- **Implementation**



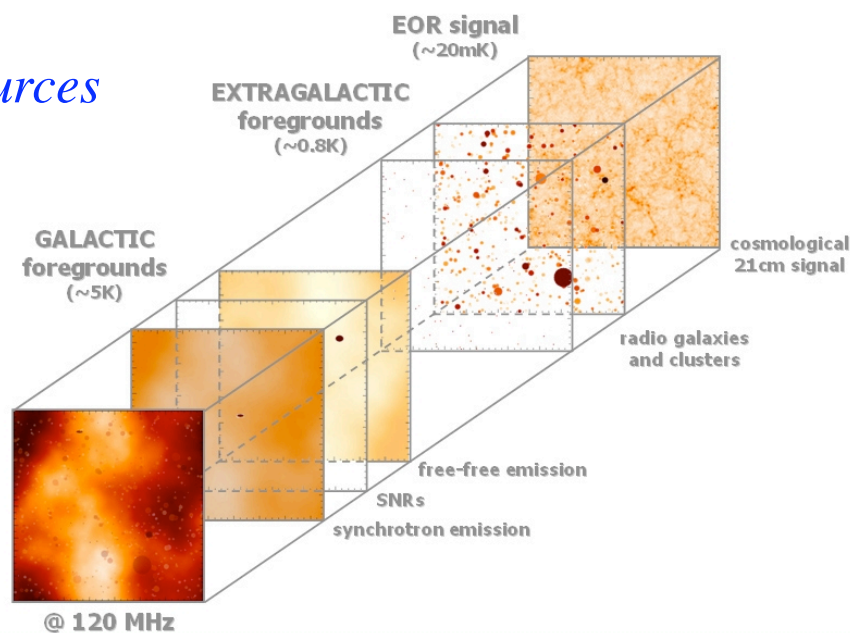
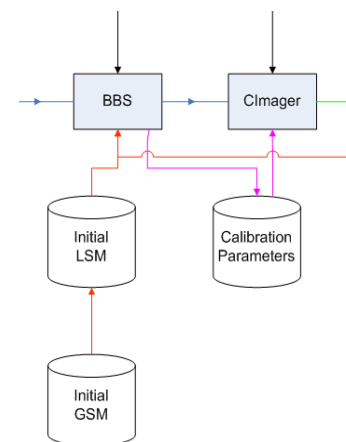
Global Sky Model (GSM)

- All-sky source database
- Spatial, spectral, and temporal(?) information
- Provides initial values for pipeline LSM(s)
- Ingests refined LSM(s) from pipelines
- Improves over time

Local Sky Model (LSM)

- Subset of GSM used by a particular pipeline
- Both input and output for pipeline components
- Multiple interfaces required (*GSM, BBS, DP³, ...*)
- May need pipeline-specific implementations
- Potential data product

- Calibration
 - *Used as initial inputs to BBS*
 - *Final LSM useful for diagnosis and reprocessing*
- Surveys
 - *Used to construct master source lists*
 - *Estimate spectral indices for sources*
- Transients
 - *Identify and monitor known sources*
 - *Detect new transient sources*
- EoR
 - *Subtract foregrounds*



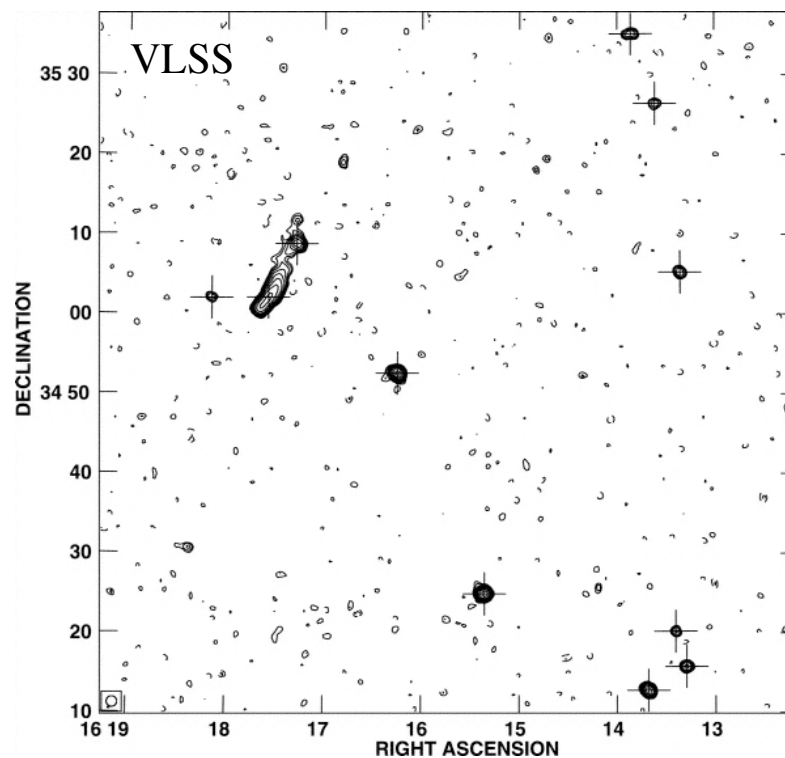
Source parameters

- Positions, fluxes (IQUV), polarization properties
- Size and shape characterization (Gaussians, shapelets, Bessel functions, etc.)
- All as functions of frequency coverage
- Derived quantities (spectral index)
- Source classification information
- Temporal information (time tagging)
- Ephemeris for moving objects(?)
- Errors on everything

Related information

- Ionosphere model parameters (time, frequency)
- RFI sources (time, position, frequency)

- VLSS
 - 74 MHz, 80 arcsec FWHM
 - 70,000 sources
- WENSS
 - 325 MHz, 54 arcsec FWHM
 - ~300,000 sources
- NVSS
 - 1.4 GHz, 45 arcsec FWHM
 - 1.8×10^6 sources
- 8C
 - 38 Mhz, 4.5 arcmin FWHM
 - ~5000 sources



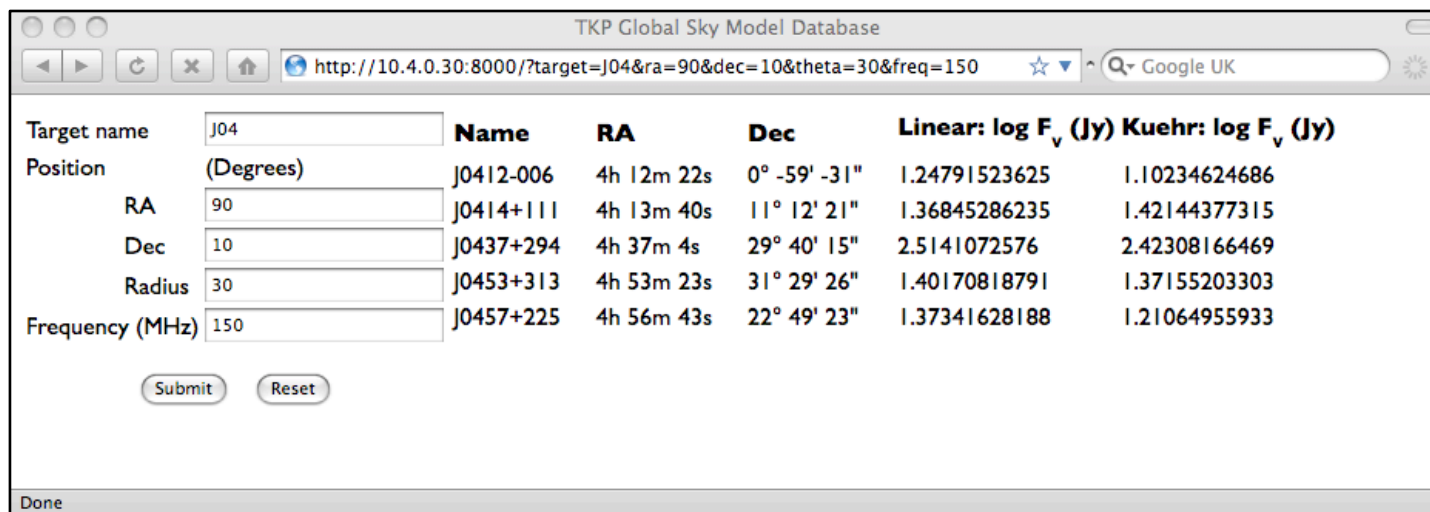
Cohen et al. (2007)

MSSS \Rightarrow 60 MHz, ~82 arcsec FWHM
150 MHz, ~33 arcsec FWHM

- Extract LSM from GSM
 - *Return source lists and parameters for FOV*
 - *Fast cone search on position and radius*
- Predicted source fluxes
 - *Return predicted flux at intermediate frequencies*
 - *Support various standard spectral models*
 - *Helmboldt et al. (2008) spectral fits to 74 MHz sources*
- Predicted sky maps
 - *Convolve source shape models to desired resolution*
 - *How to handle poorly resolved sources?*
- Associate catalog source lists

Sky Models

- GSM is a database (*MySQL, PostgreSQL?, MonetDB?, Oracle?*)
- Many predictive functions implemented in database
- Python interfaces provided for database access
- Prototype already exists
- Simulated maps created using external tools (*ORGSM, N. Mohan*)



Target name	Name	RA	Dec	Linear: log F _v (Jy)	Kuehr: log F _v (Jy)
J04	J0412-006	4h 12m 22s	0° -59' -31"	1.24791523625	1.10234624686
RA: 90	J0414+111	4h 13m 40s	11° 12' 21"	1.36845286235	1.42144377315
Dec: 10	J0437+294	4h 37m 4s	29° 40' 15"	2.5141072576	2.42308166469
Radius: 30	J0453+313	4h 53m 23s	31° 29' 26"	1.40170818791	1.37155203303
Frequency (MHz): 150	J0457+225	4h 56m 43s	22° 49' 23"	1.37341628188	1.21064955933

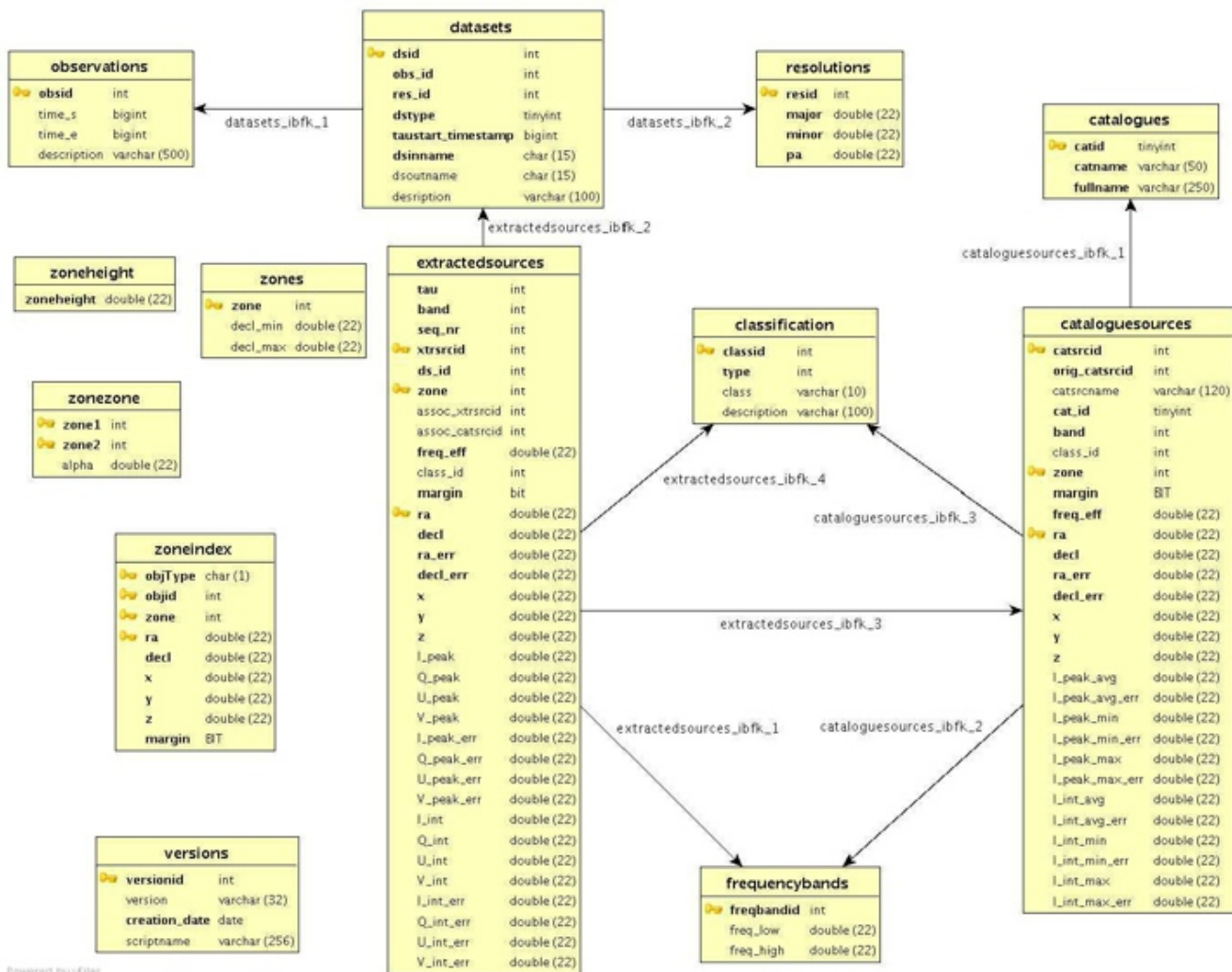
Example query on prototype GSM database

Sky Models

- GSM is a database (*MySQL, PostgreSQL?, MonetDB?, Oracle?*)
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```
SELECT alpha_gt_300 * LOG10(@nu/74) + LOG10(f_ext) AS 'linear: log F_{nu} (Jy)'
  ,CASE WHEN param_a IS NULL
    THEN NULL
    ELSE param_a +
      IFNULL(param_b, 0) * LOG10(@nu/74) +
      IFNULL(param_c, 0) * EXP(IFNULL(param_d, 0) * LOG10(@nu/74))
    END AS 'Kuehr: log F_{nu} (Jy)'
FROM sources src
  ,spectralparameters sp
WHERE spectral_params_id = spectral_paramsid
AND src_name = @source_name
```

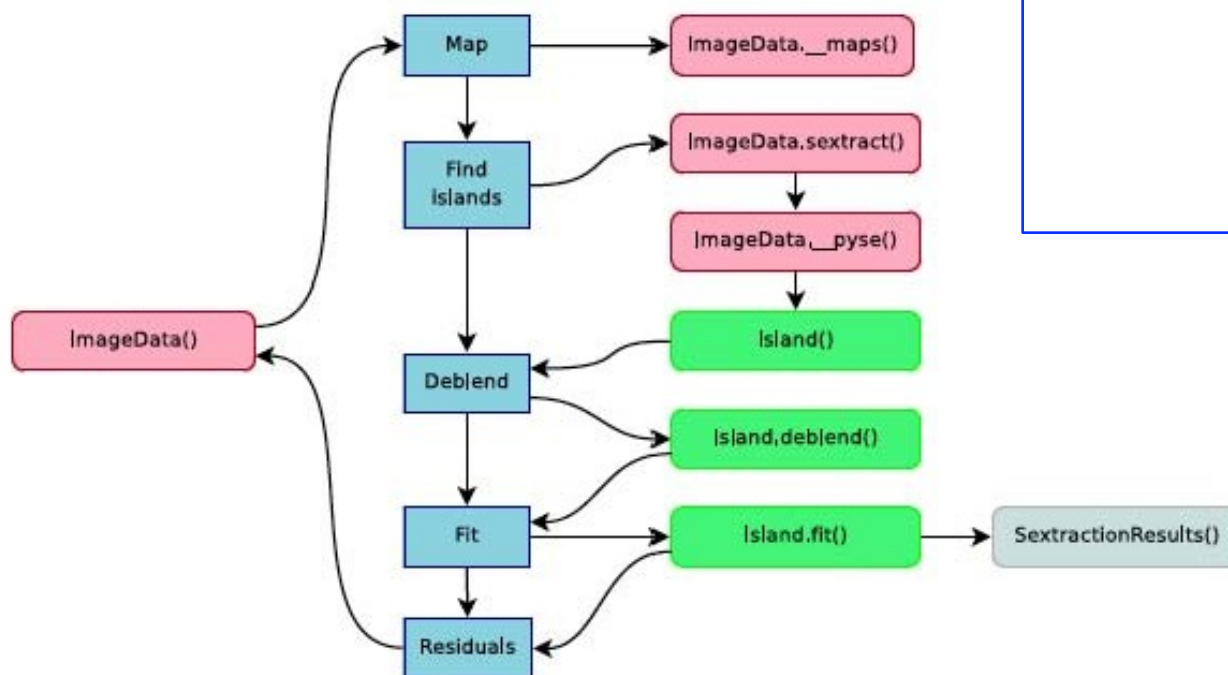
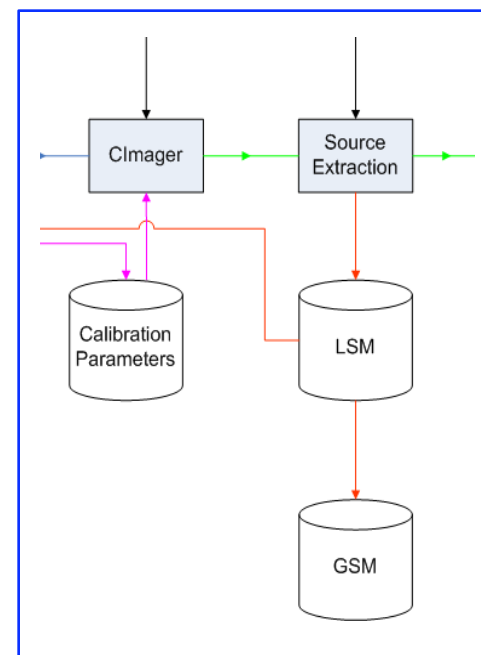
Example of MySQL stored procedure



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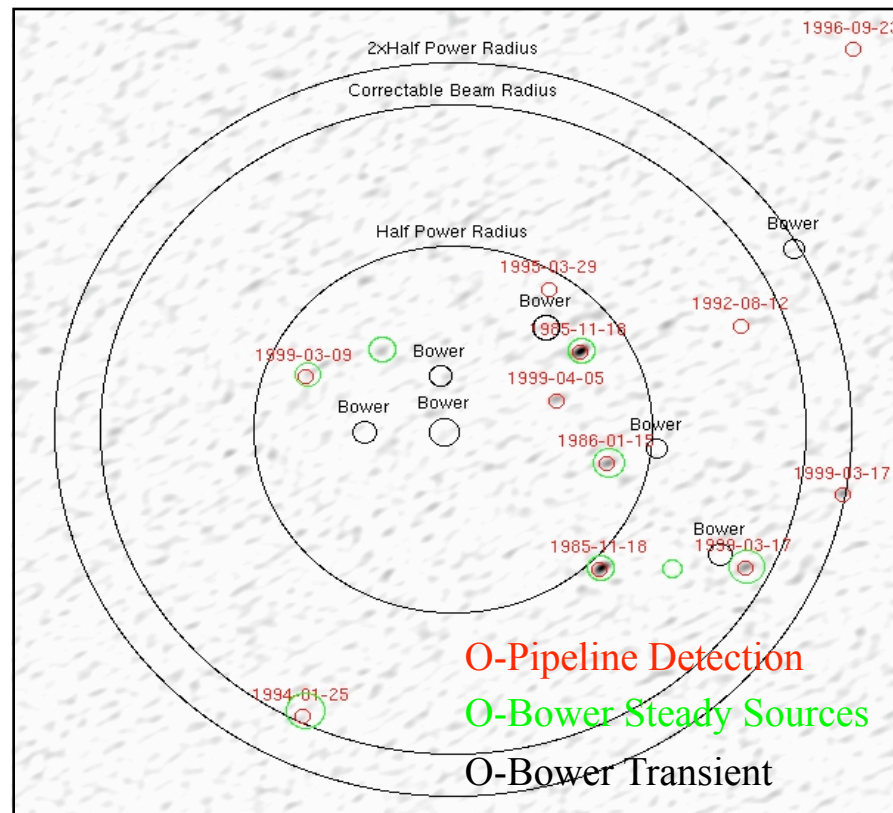
Database scheme for the Transients pipeline area (B. Scheers 2008)

- Detect sources in image cubes
- Determine source parameters
- Update LSM for next major cycle loop
- Merge final results into GSM database



- Two implementations currently available
- Transient pipeline version optimized for speed
- PyBDSM version has more functions and visualization
- Both python-based
- Both being tested
- Both easily connected to GSM database

Bower VLA deep field



M. Bell (2008)

Next steps for MSSS

- Set-up prototype on output cluster
- Integrate into MSSS pipeline structure
- Connect source-finding package to database
- Create LSM for BBS using database

Open questions

- How do we handle sources which are multiple sources at different frequencies?
- GSM on european baselines?
- How represent extended sources?
- What do we do about polarized sources?
- Need to specify LSM format(s)