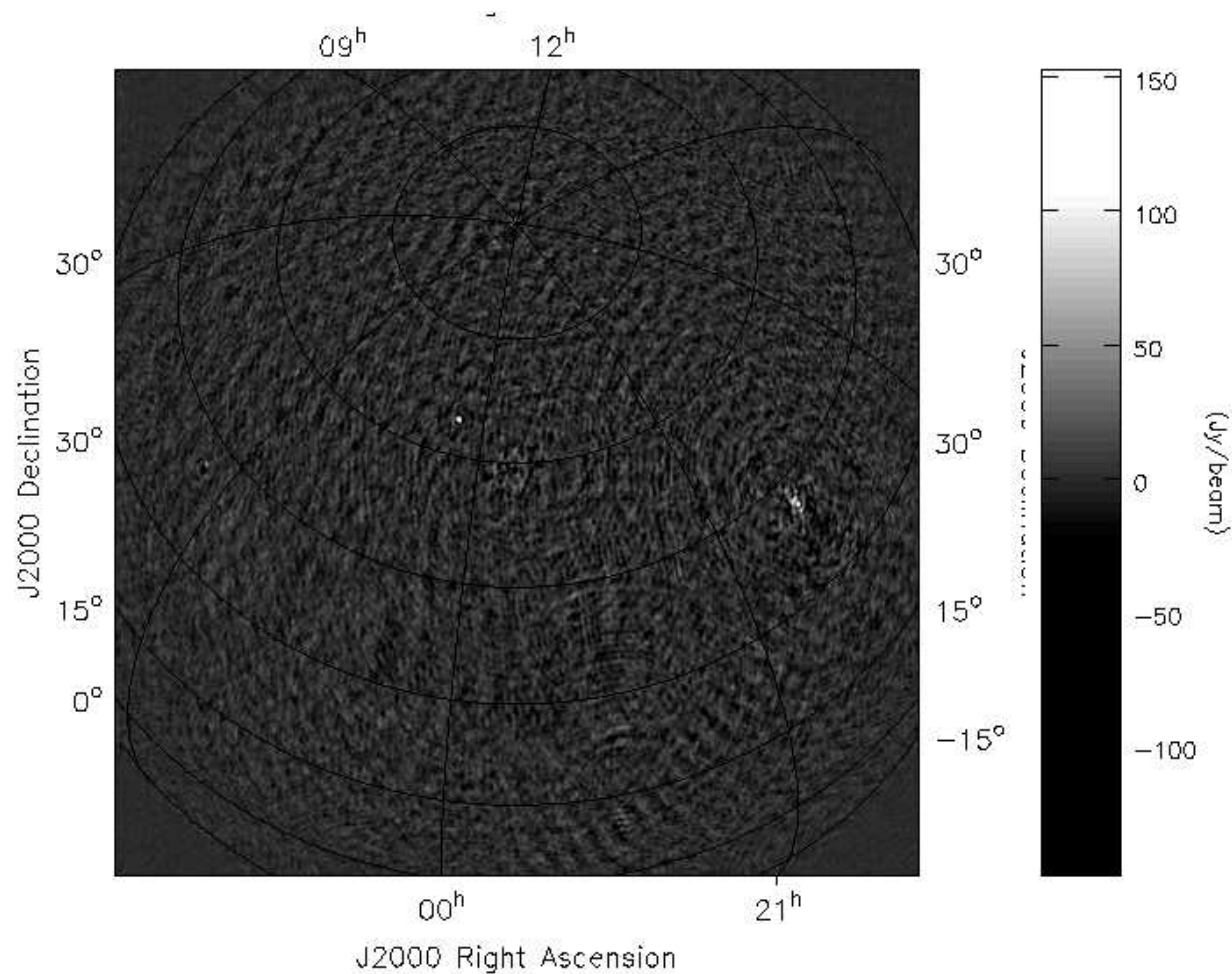


# LOFAR CS1

Calibration and Imaging

(cont.)

# More Sources!



L2007\_01576, subband 5 (60 MHz)

# Peeling

## □ Jones Matrices

$$\mathbf{G} = \begin{bmatrix} g_{11} & 0 \\ 0 & g_{22} \end{bmatrix} \quad \mathbf{J} = \begin{bmatrix} J_{11} & J_{12} \\ J_{21} & J_{22} \end{bmatrix}$$

- Peel CasA and CygA. If we solve for  $\mathbf{J}$  Jones what we get is  $\mathbf{JP}$  where  $\mathbf{P}$  is the beam and projection. Since  $\mathbf{P}$  is an image plane effect, we cannot use this to correct the residual data.
- Hence we peel with diagonal  $\mathbf{J}$  Jones and use the solution of CasA to correct the residual data.
- More dipoles: better sensitivity, CasA has high SNR: better solution, 29hr observation, beamforming

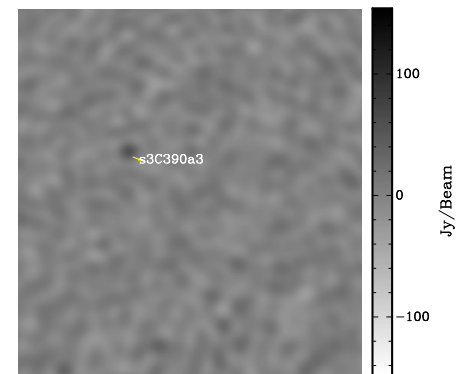
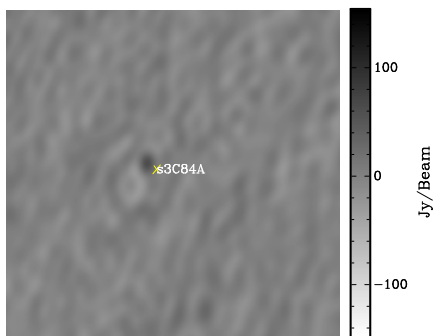
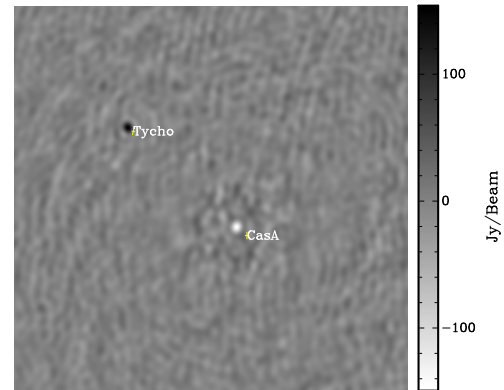
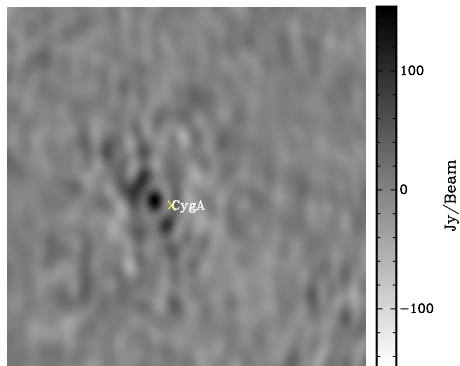
# Peeling...

- Solve for diagonal  $\mathbf{J}$ , real and imaginary parts

$$\mathbf{J} = \begin{bmatrix} J_{11} & 0 \\ 0 & J_{22} \end{bmatrix}$$

- Solution domain: one timeslot (1 min), 10 channels (10 kHz)
- Simultaneous solution in directions of CasA and CygA
- $\mathbf{J}$  include the cumulative effect of electronics ( $uv$  plane), station beam (Az,El), model error in source fluxes

# Closeups

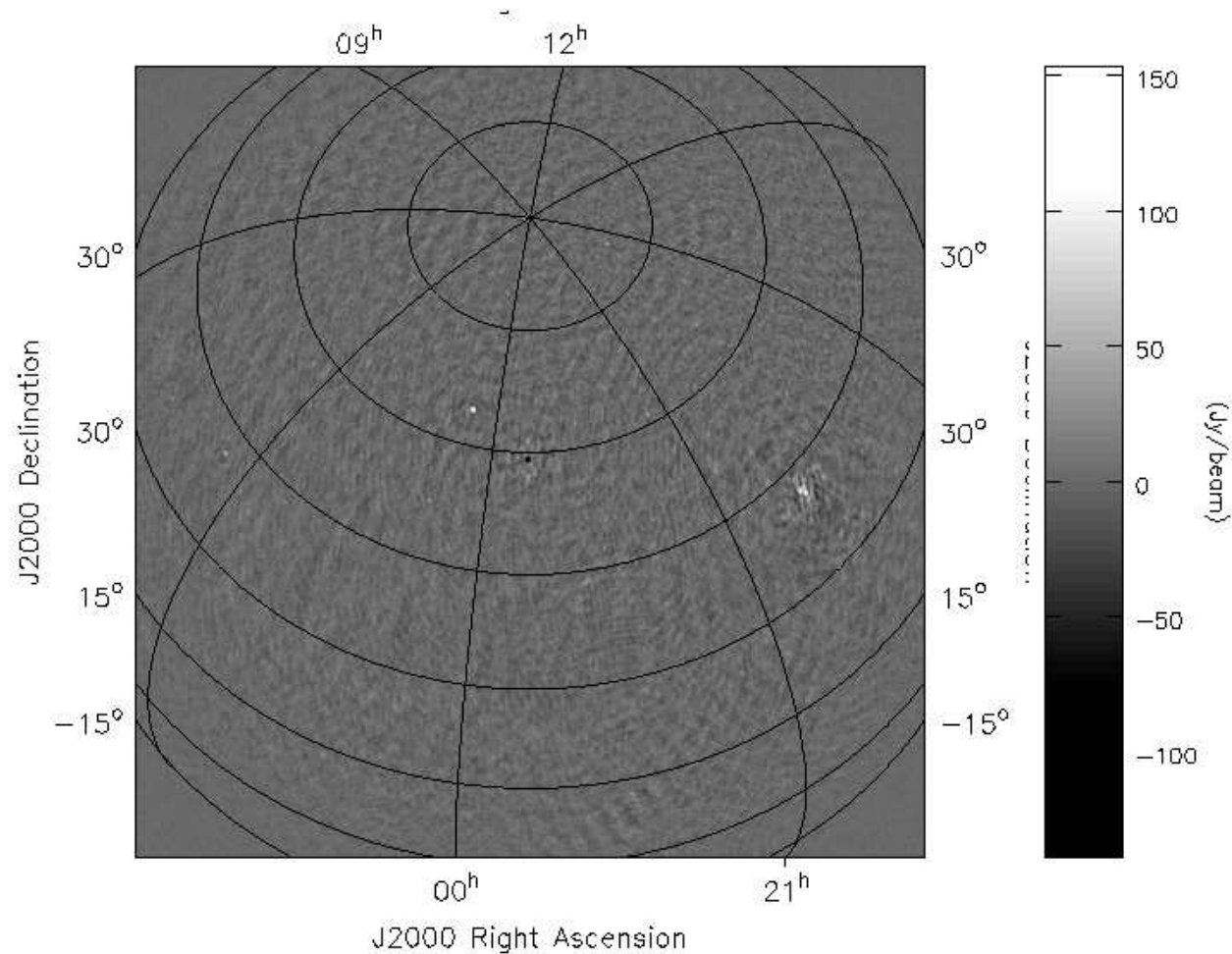


□ Dynamic range close to CasA  $\approx 1:2000$

# Improvements in Speed

- ☐ From 4 hr 30 min to 10 min
- ☐ Do not save solutions:  $\times 7$
- ☐ Subtiling solutions: parallel solvers  $\times 4$
- ☐ Next: Solve for an entire subband at a time, using 2nd order freq polynomials for all parameters.
- ☐ Next: Resampling when shifting the phase-centre.

# Again!



L2007\_01576, subband 4 (58 MHz) + subband 5 (60 MHz)