# Deep observations with CS1 for a survey of SNRs

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### CSI- LBA Current status

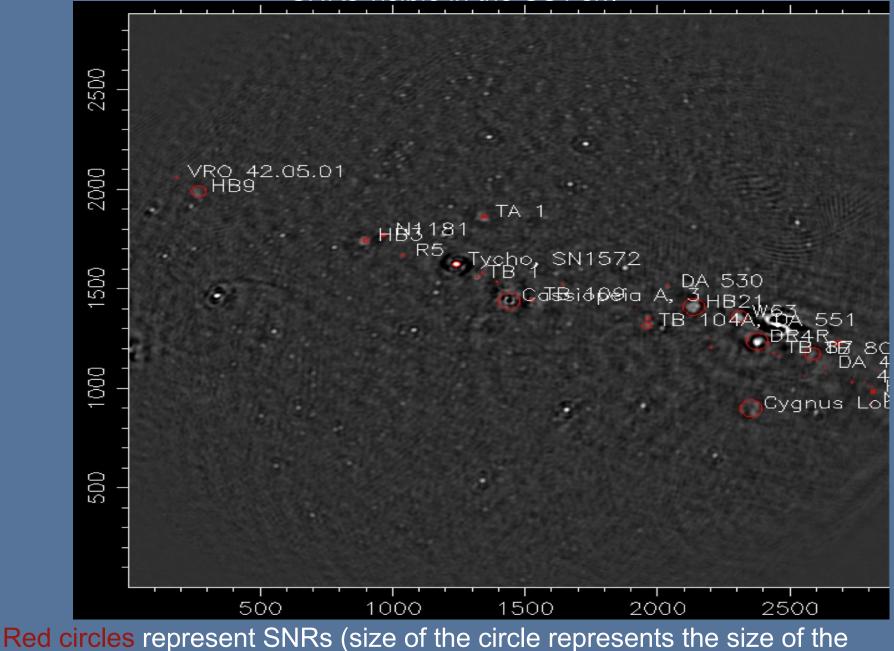
Frequency (MHz)	30	55 - 60	80
Synthesised Beam (arc min)	(~20)	half power (~10)	(~8)
Largest Detected source (degrees)	-	~4	-
Best rms sensitivies achieved (Jy)	-	2 - 3	-

CSI data could be used to study the spectral behavior of bright sources in the Northern Galactic plane.

SNRs are good candidates for this study

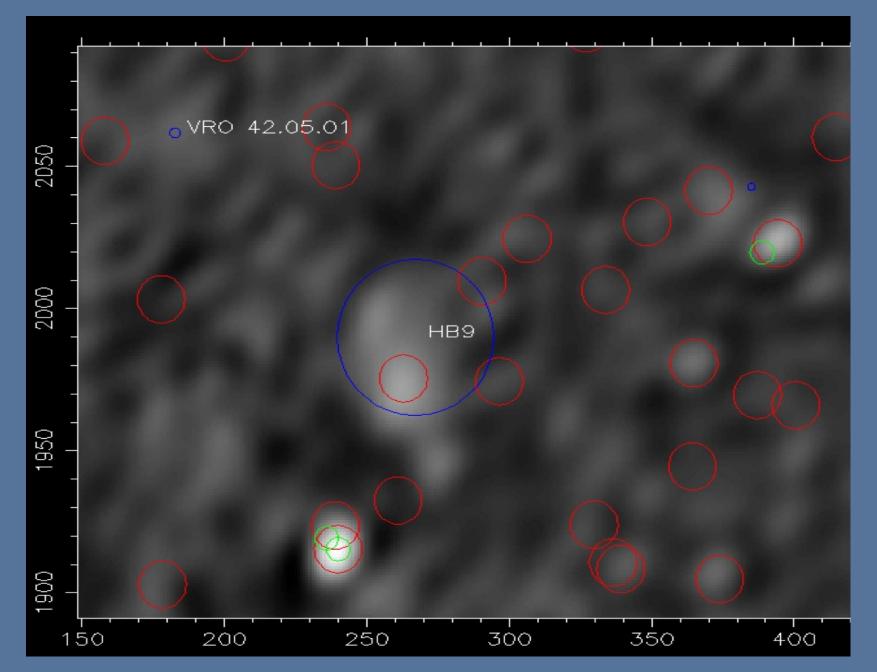
- Niruj has plotted Dave Greens SNR catalogue on 102339 image of Sarod (CS1 data at 60 MHz).
- Resolved shell structure of half a dozen SNRs are seen in the CSI image.
- It will be useful to carry out 30-80 MHz observations on these sources
- Niruj's 'Noise' package + manual is available to carry out image analysis, associating source lists and analyzing source lists against external catalogues.
- A python version of BDSM is available now.

#### SNRs visible in the CS1 skv

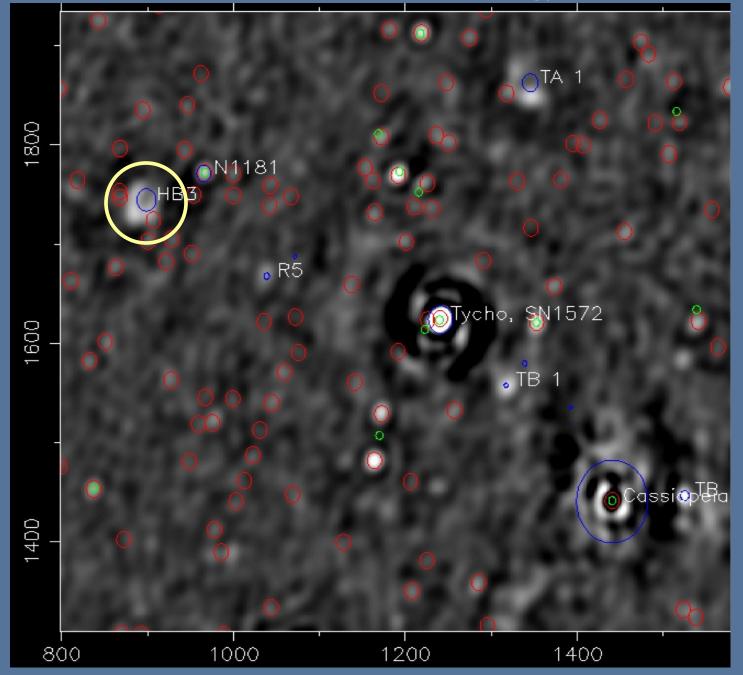


source) from Dave Green's catalog

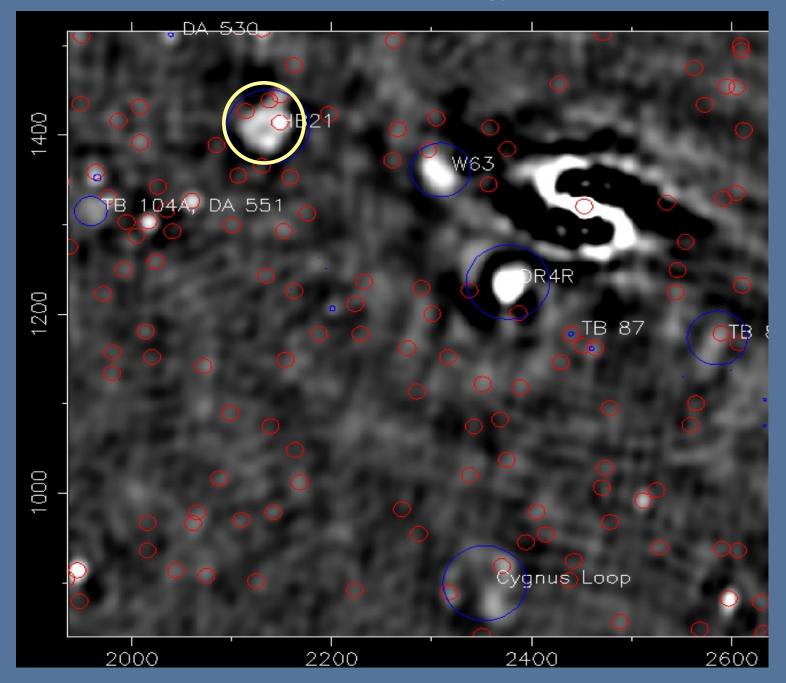
#### Red circles- 4C catalog, Green circles- 3C catalog, Blue circles- DG SNRs



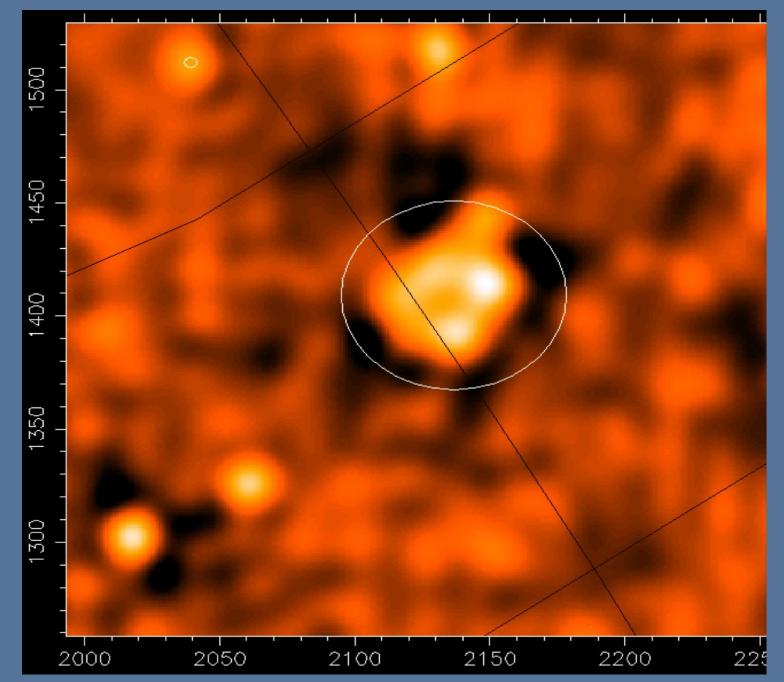
#### Note HB3 has a resolved shell morphology



Note HB21 shows a well resolved shell morphology with multiple components



#### • Blown up image of HB 21showing multiple components



# Multiple components of HB 21 are consistent with this image at 232 MHz with the WSRT (Zhang et al. 2001 ChJAA, 1, 443)

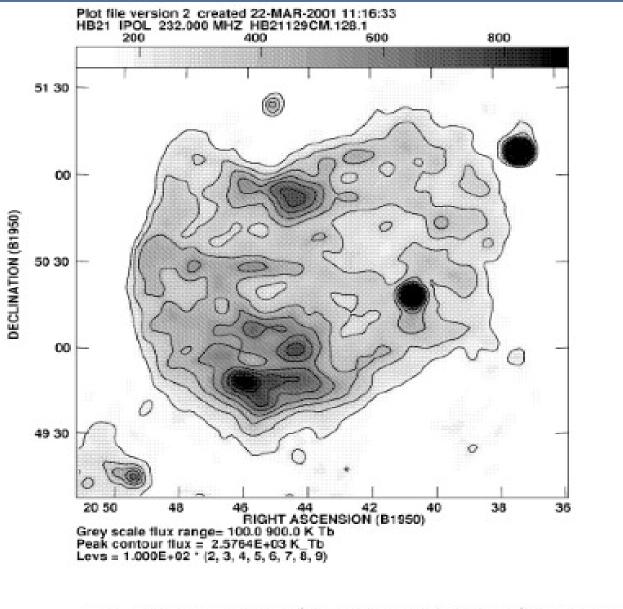


Fig. 1 The intensity map of HB21 observed with the MSRT at 232 MHz.

We would like to carry out a similar study for the following bright spiral galaxies and halos:

Source	Flux density (57.5 MHz) (Jy)	
1- NGC 253	48 ± 9	
2- NGC 1068 (M77)	39 ± 8	
3- NGC 3034 (M82)	29 ± 6	
4- NGC 5236 (M83)	29 ± 5	
5- Abell 2256	25	