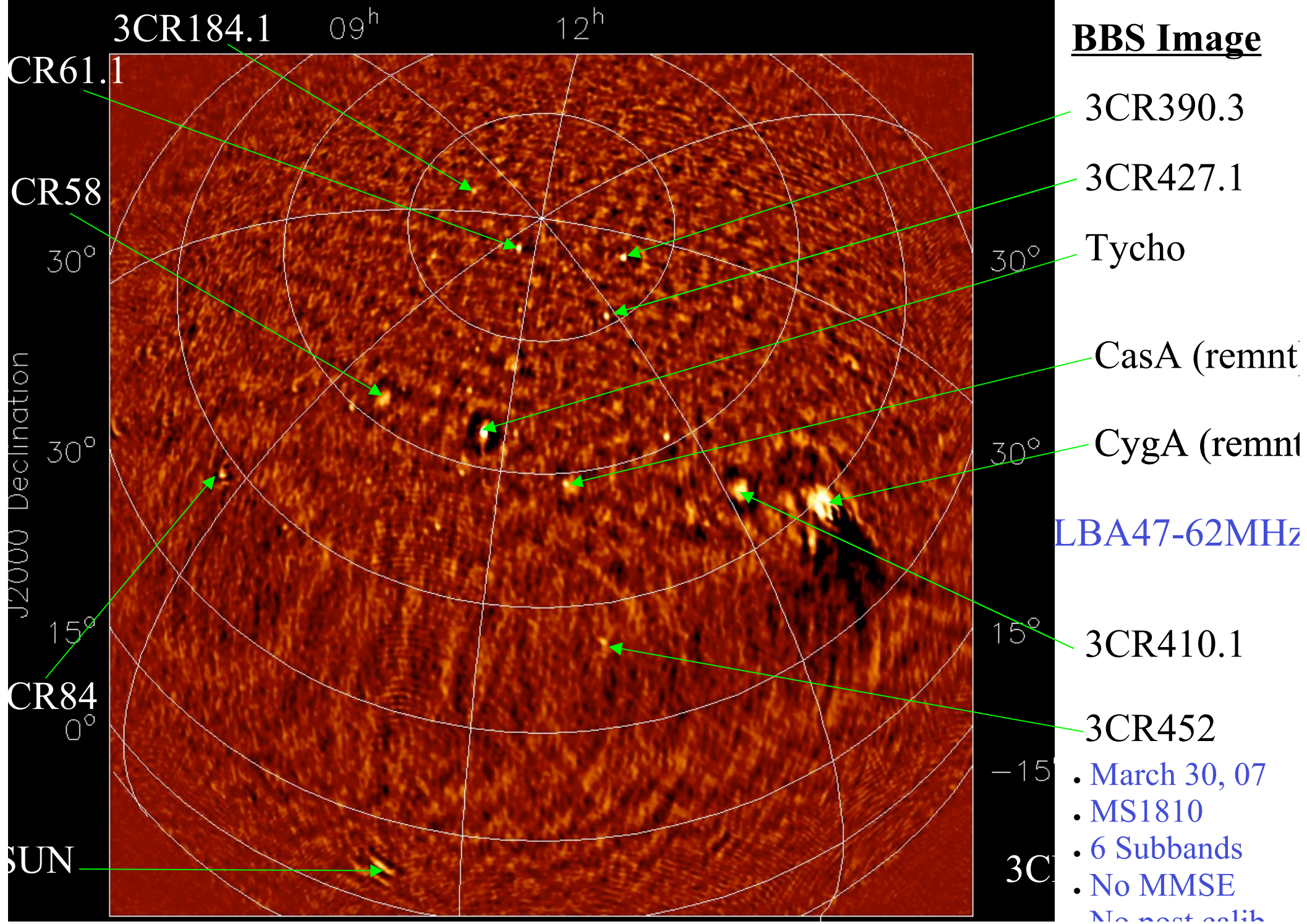


BBS - New Images, Overview and Progress Plan

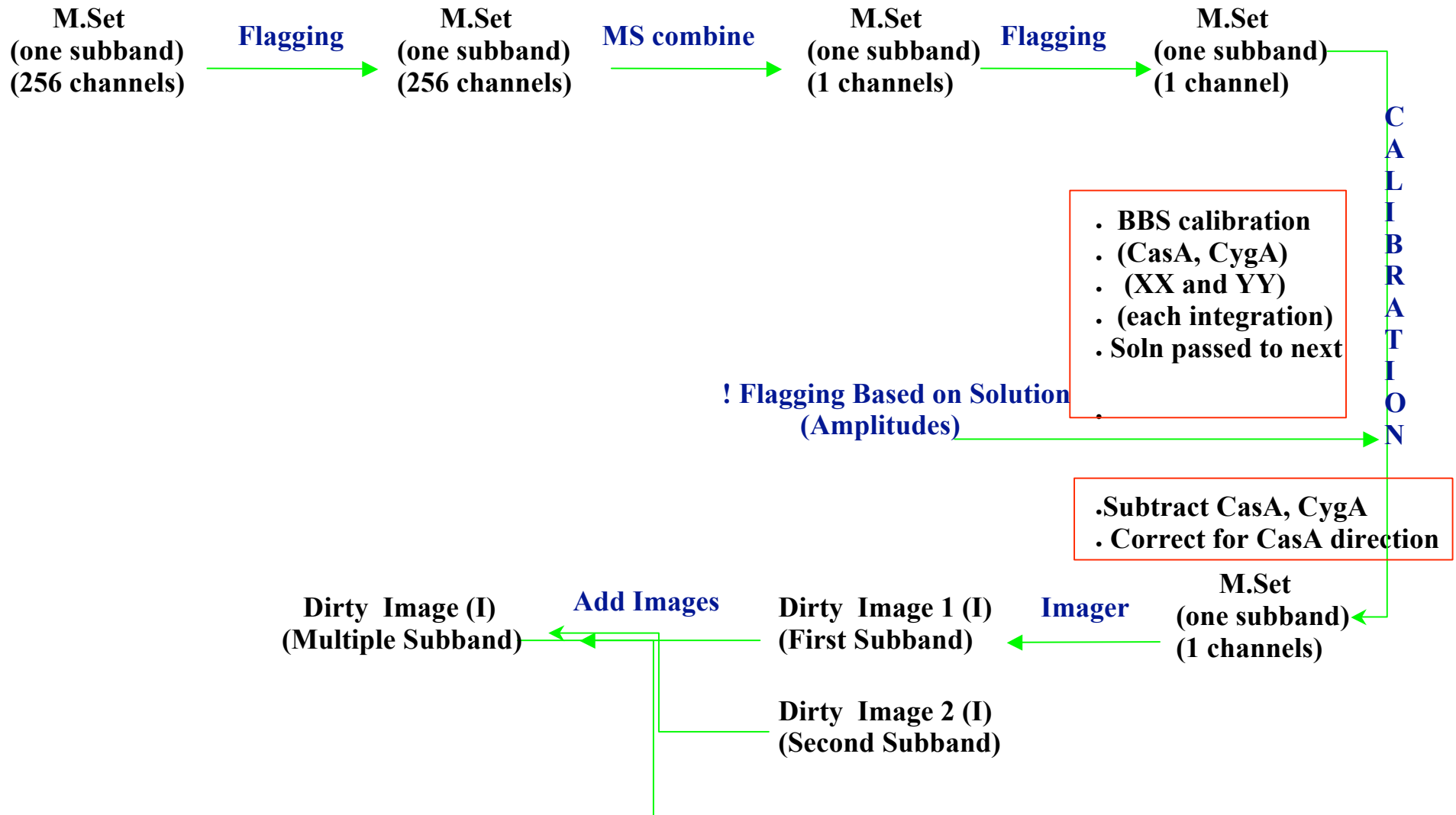
V.N.Pandey, Joris van Zwieten, Ger de Bruyn
CS1 meeting
October 11, 2007

Outline - Recap

- ① Recap - **BBS - First Image** (>40sources) (Mar 30 obsvn) - Sep, 07
 - ② Analysis - Processing Pipeline, Solutions, Limitations
 - ③ **Lessons Learnt**
 - ④ **Processing of Recent LBA observations - New deeper Images**
 - ⑤ **Processing of HBA observations - Initial Images**
 - ⑥ **Conclusions - BBS Status, different versions, Next Course of action**
-



Analysis: Steps-Schematic



Limitations - Bottlenecks

- Out of 15 Sub bands processed (37 to 72MHz) -
>Images for 10 Sub bands seem to be good !!
- Rigorous pre-calib flagging -> 13 Good images.
- Post-calib flagging -> 13 good images.
- Is it necessary to do Post-Calibration Flagging?
- Let us look at one sub band where we failed!!

09^h

12^h

16 hours

J2000 Declination

30°

30°

15°

0°

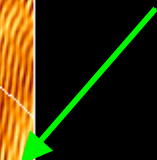
30°

30°

15°

-15°

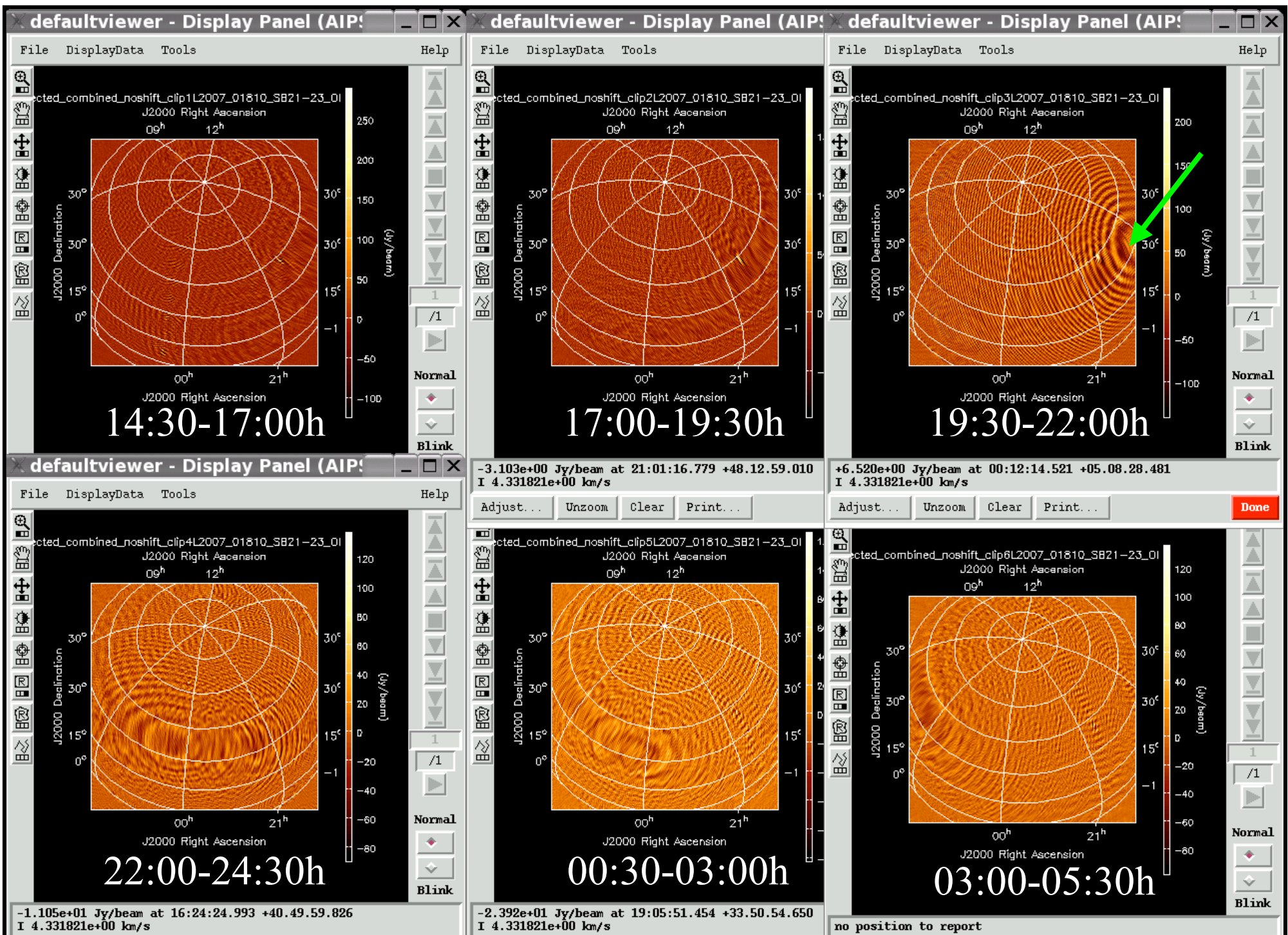
?????

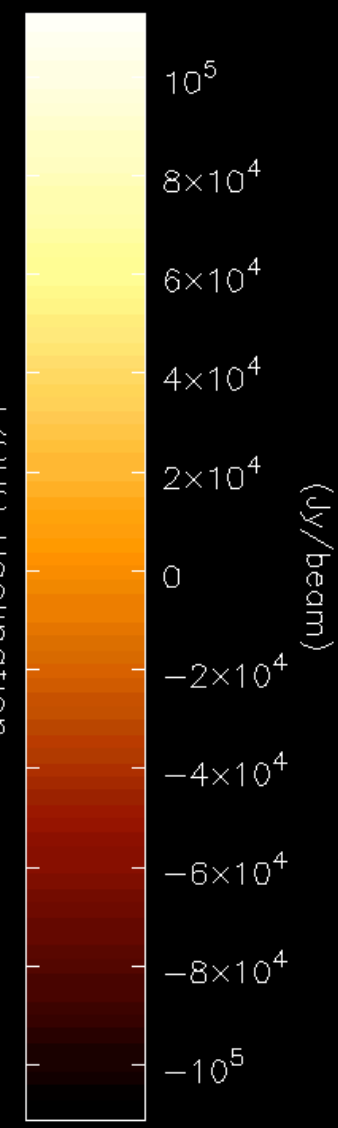
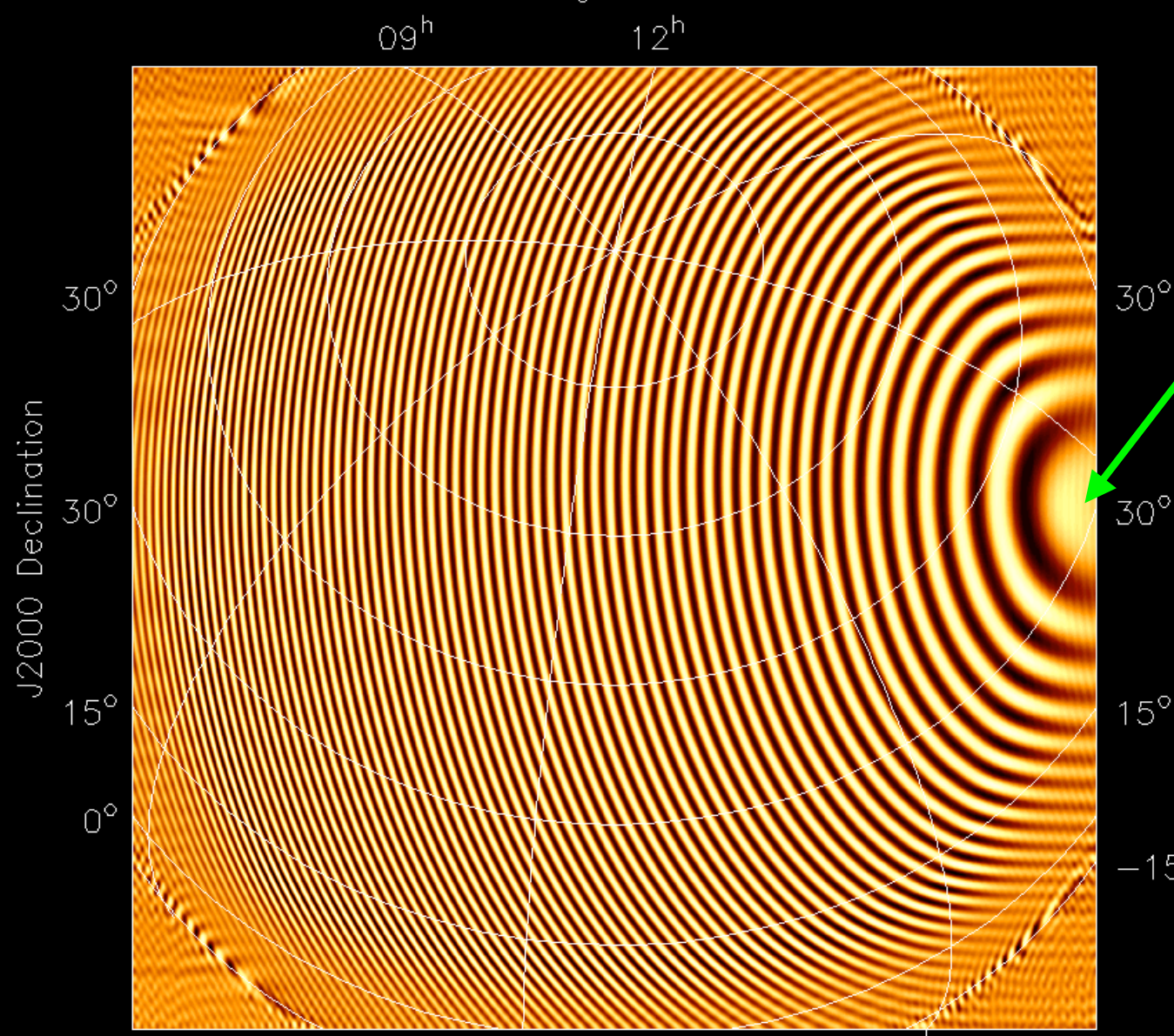


(Jy/beam)

00^h

21^h

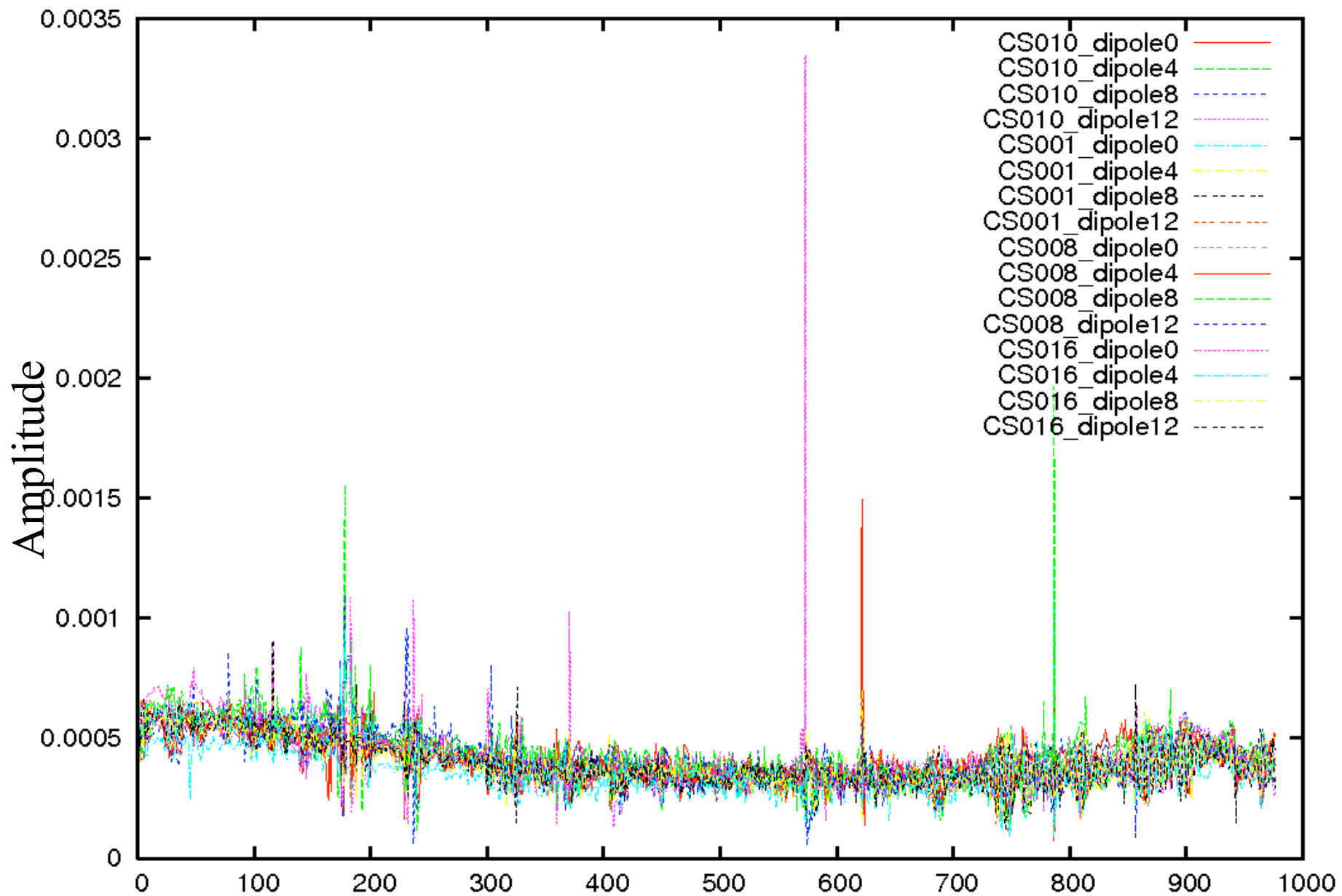




19:31:15-19:32:30

J2000 Right Ascension

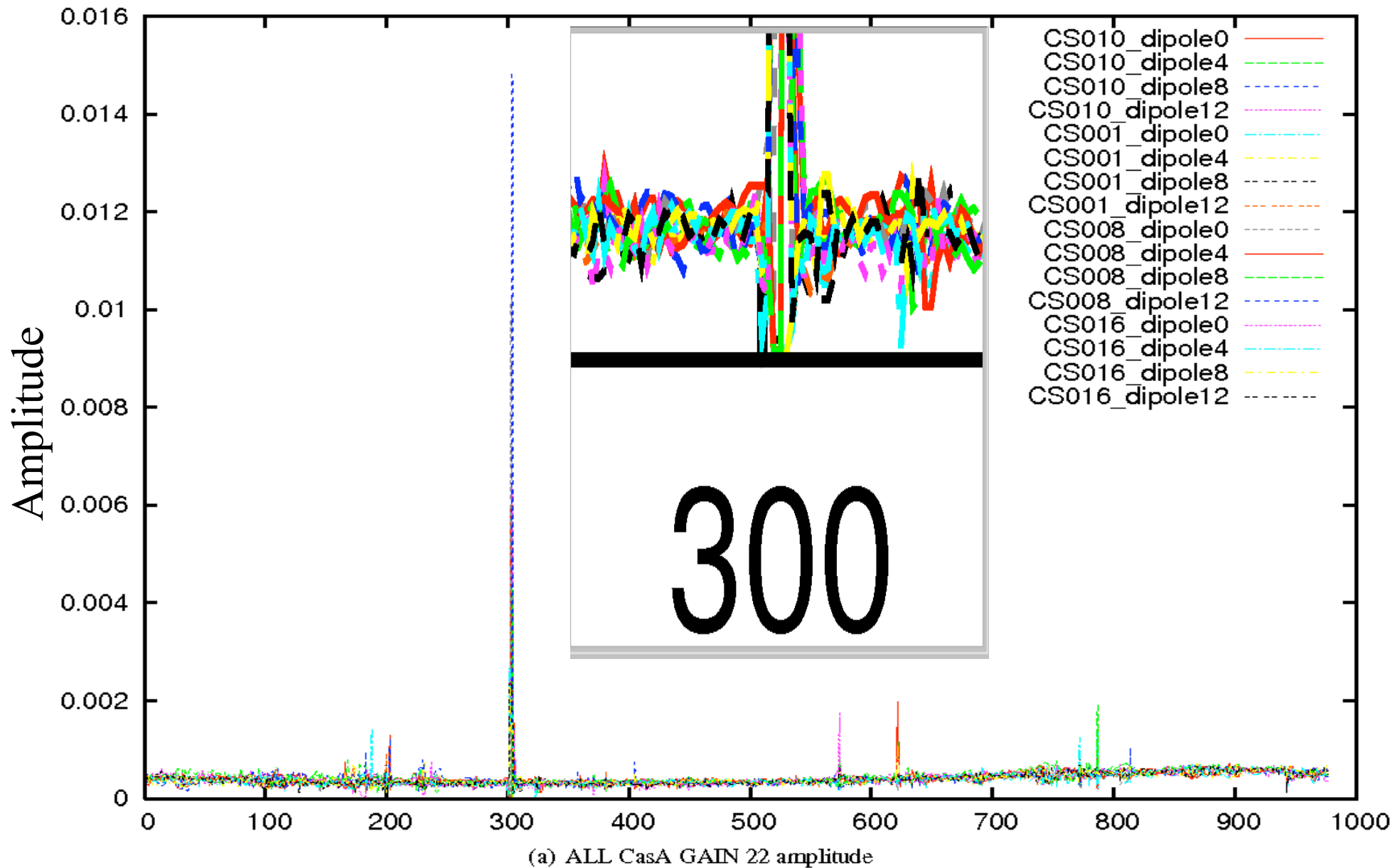
Only one bad visibility



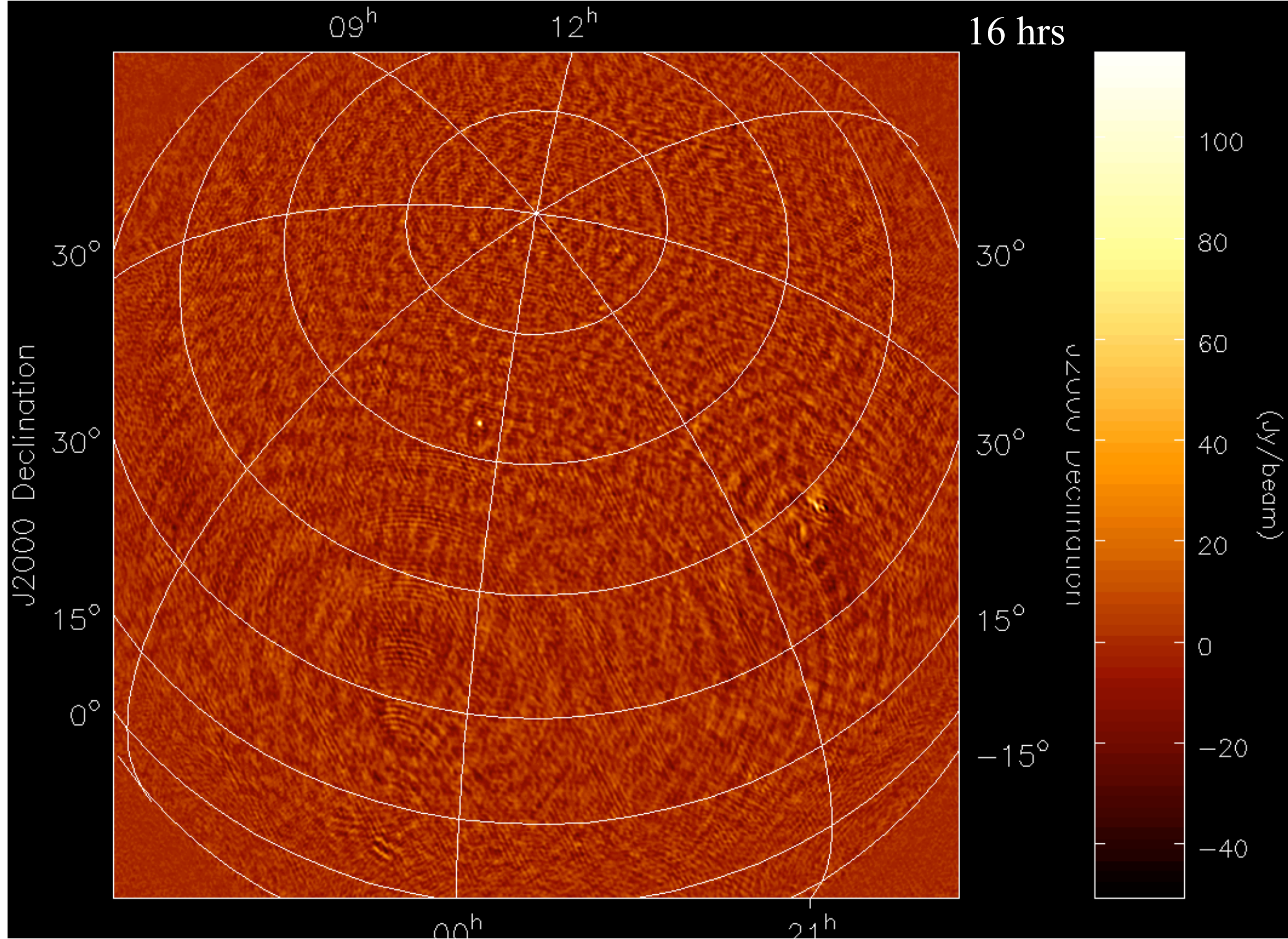
(a) ALL CasA GAIN 11 amplitude

Time (Integration number~minute)

CasA 11, Amplitude



Time (Integration number~minute) CasA, (YY) Amplitude



Lessons Learnt / Open Issues

- Bad solutions -> may introduce distortions -> a unphysical image!
- Extra intelligence/flagging needed during the solving process! -> careful while passing the solutions to next time slot!
- χ^2 based ideas, error bar on each solved parameter, condition number
- % of data flagged in a solve domain
- pre-post calibration (including based on solutions, residuals, corrected data) flagging
- MMSE !!
- More Ideas!!?

Processing recent LBA Observations

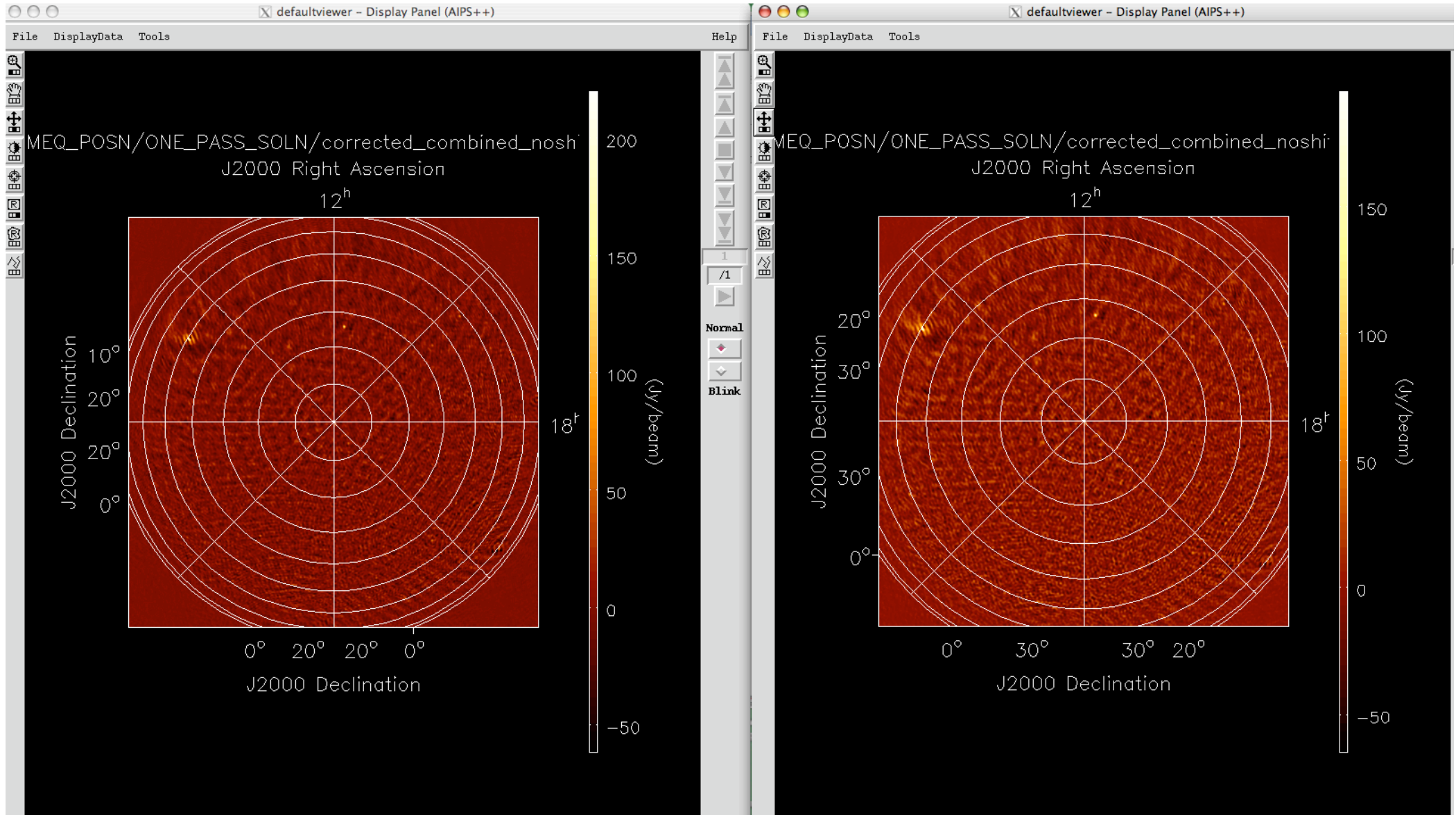
- L3463 (July 27, 07), L3464 (July 28, 07), L3565 (Aug3, 07) , ...
- 36 subbands, from 38-62MHz. Phased at NCP.
- **Same scheme of processing.**
- **Initial flagging -> solving complex gains for CasA and CygnusA**
- **Subtract CasA and CygnusA**
- **Correct for the direction of CasA**
- **Image using ALPS++ imager**
- **Individual Subband images, and then combined image**

MS3463; Sub Band: 18

Small improvement

After Solution Flags

Before Solution Flags

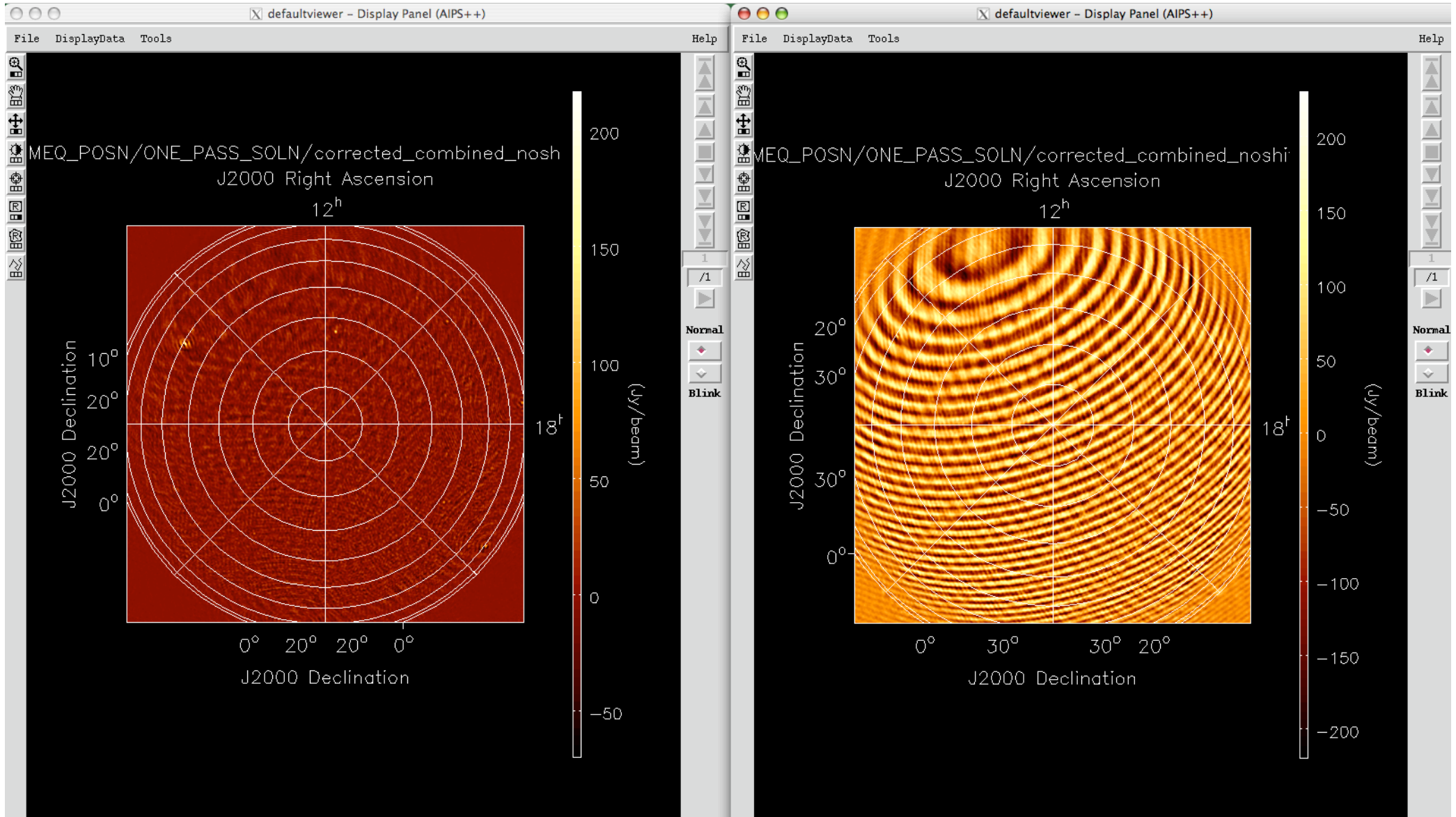


MS3463; Sub Band: 19

Improvement !!!

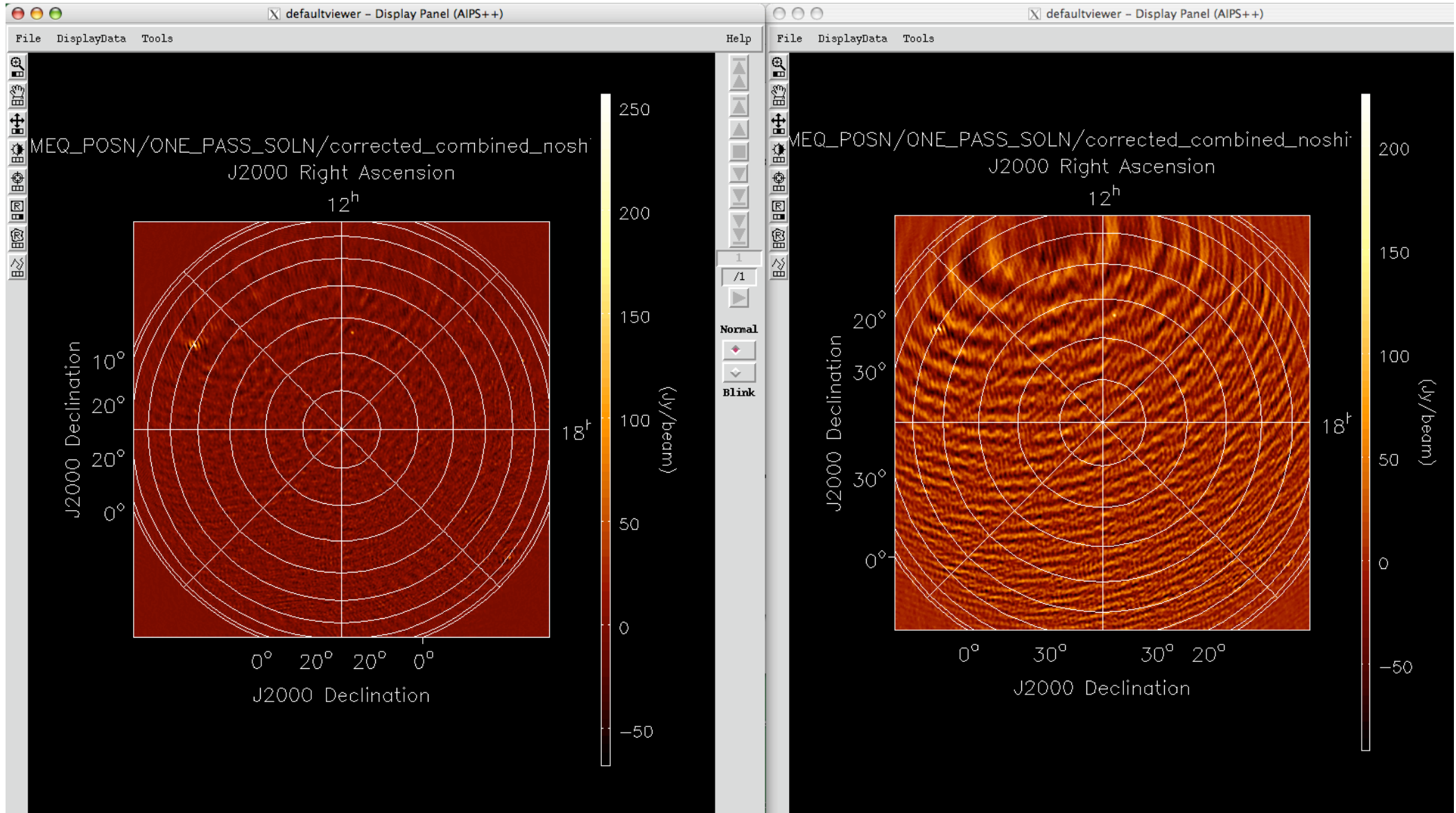
After Solution Flags

Before Solution Flags

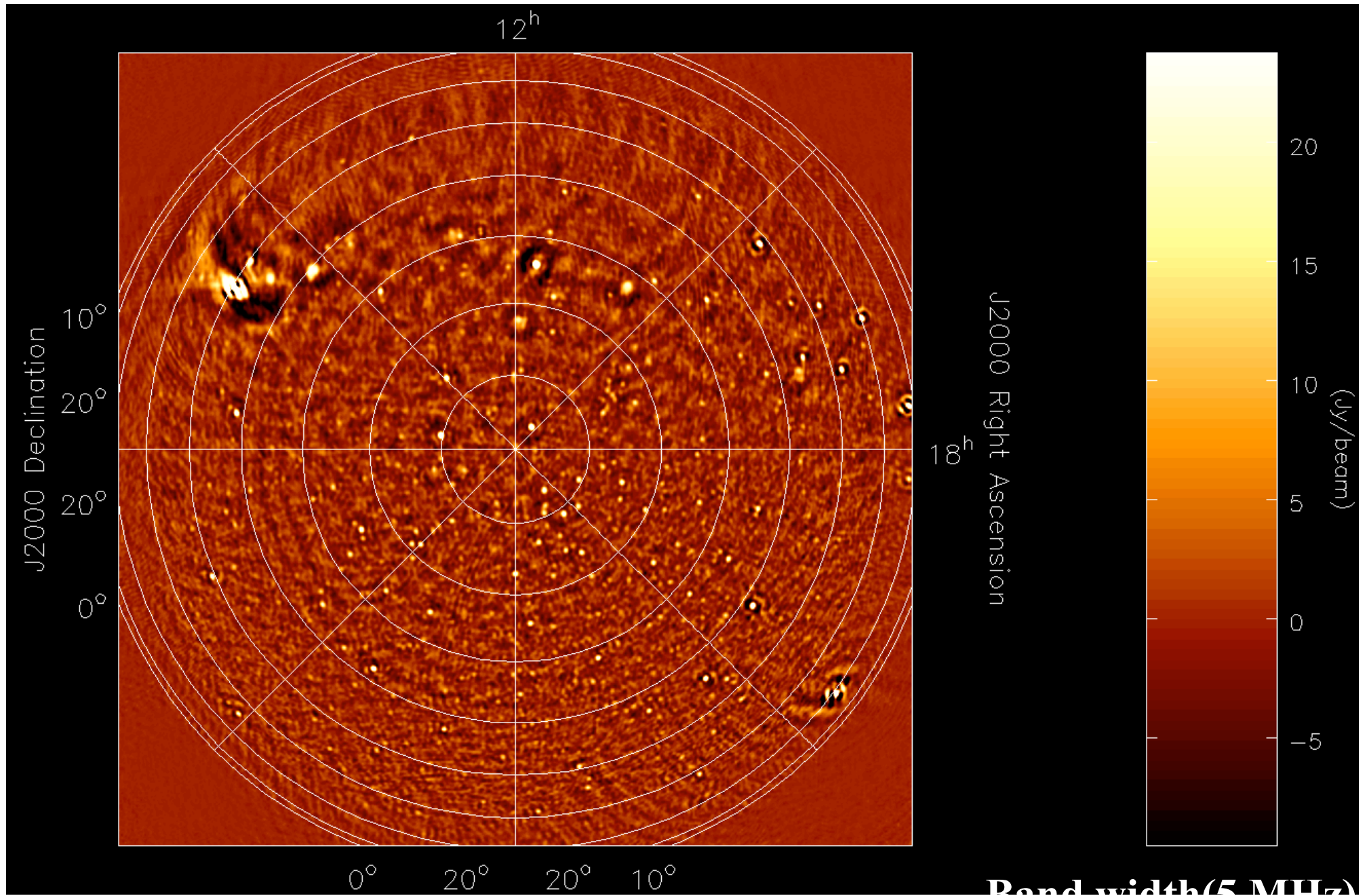


MS3463; Sub Band: 21 After Solution Flags

Improvement
25 before 36 after
Before Solution Flags



MS3463; Average {Sub Band: 0-35} 99.8%. **>400 sources**



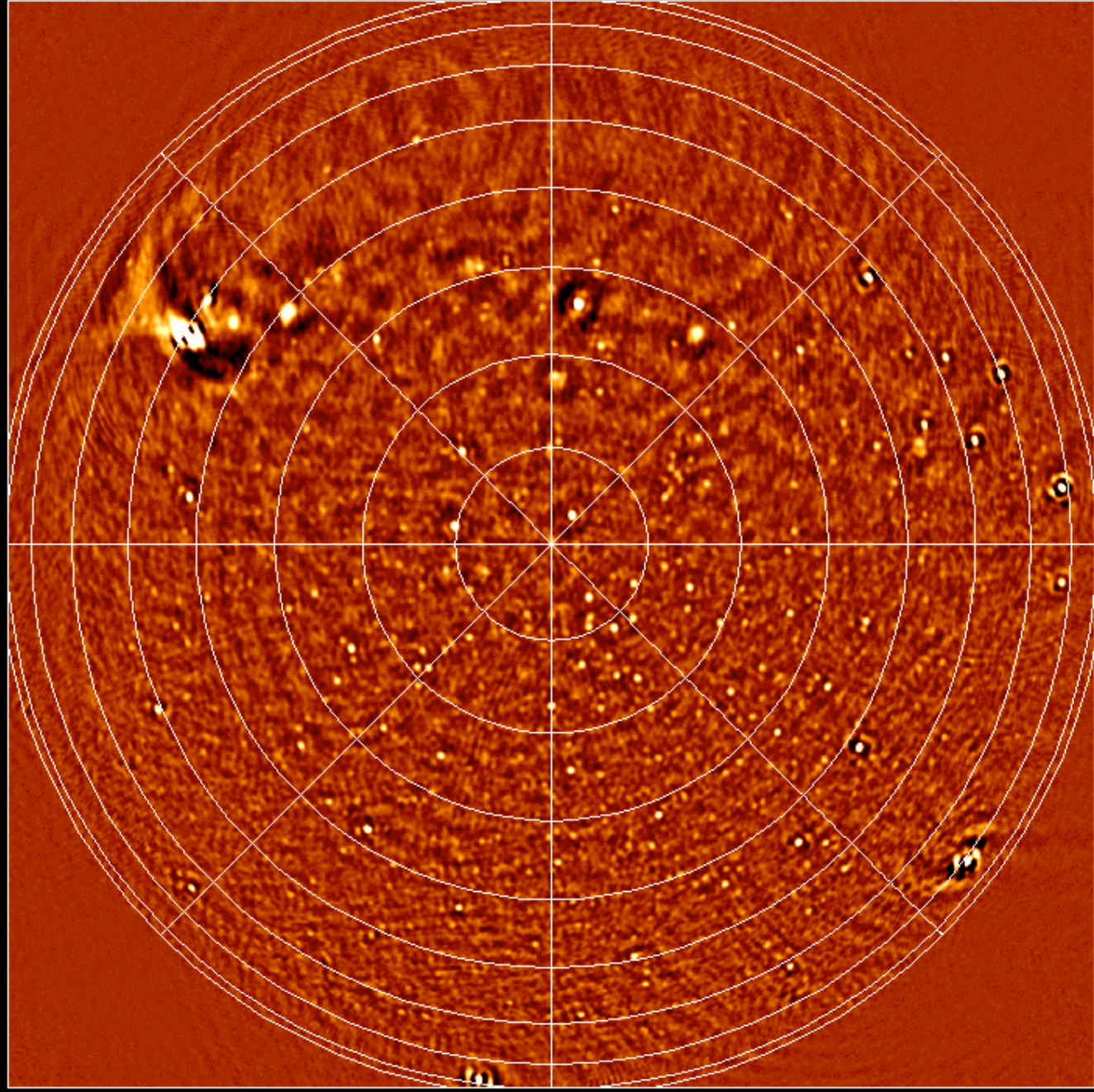
Movie

J2000 Right Ascension

18^{h}

J2000 Declination

0° 10° 10° 0°



J2000 Declination

0°

10°

10°

0°

(Jy/beam)

20

15

10

5

0

-5

-10

Processing HBA Observations

- L3743 (Aug 28, 07), ...
- 36 subbands, from 209 MHz -243 MHz, phased at NCP.
- Same scheme of processing. CS10 Flagged completely
- Number of baselines ~40.
- Initial flagging -> solving complex gains for CasA and CygnusA
- Subtract CasA and CygnusA
- Correct for the direction of CasA
- Image using AIPS++ imager
- Individual Subband images, and then combined image

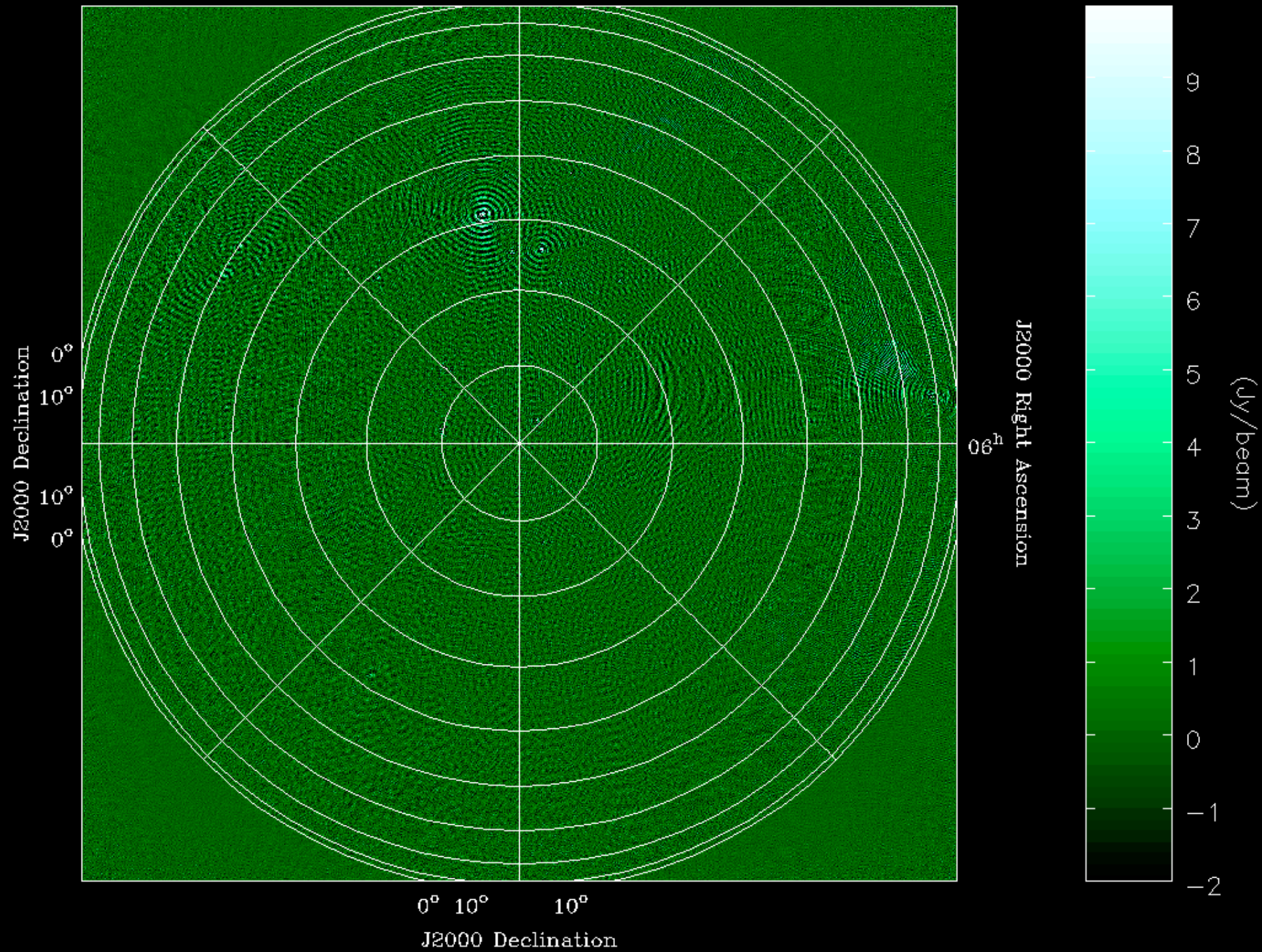
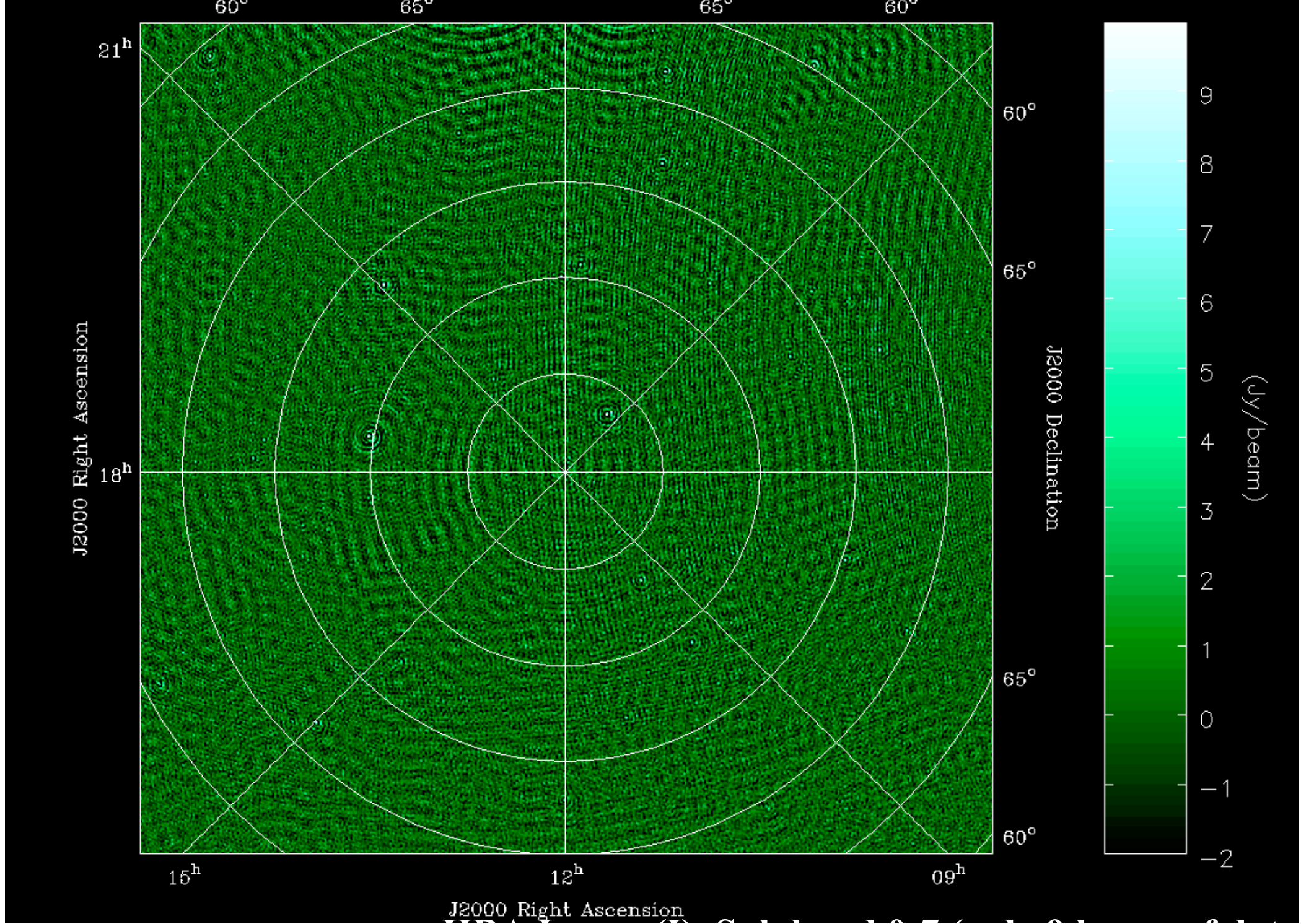
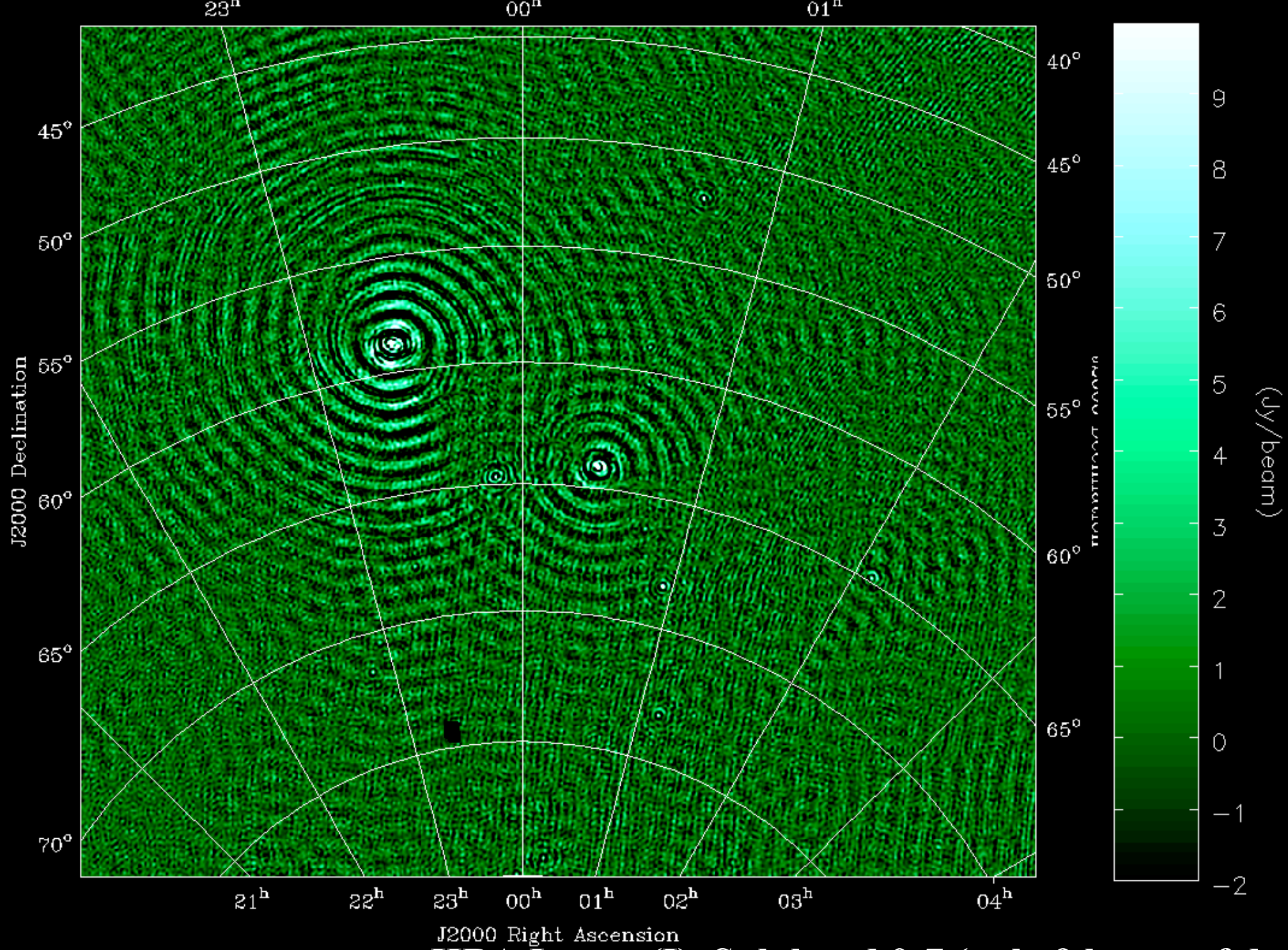


Figure 1: A circular astronomical plot showing the distribution of flux density in the J2000 equatorial coordinate system. The plot is centered on a specific point, with concentric circles representing constant angular distances from the center. The axes are labeled as follows:



J2000 Right Ascension

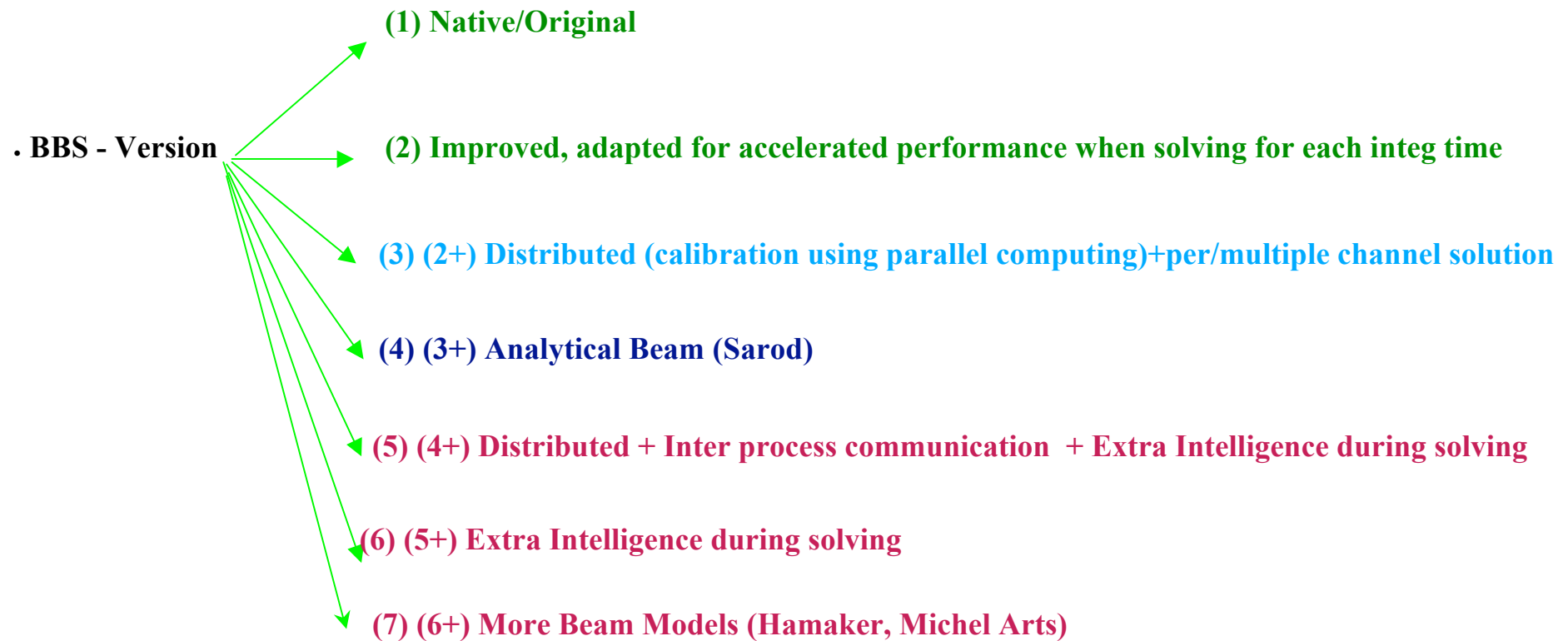
15h 12h 9h 60° 65° 60°



Next to Do.....

- Verify Positions
- Frequency dependent solutions (incl. sky model)
- Including analytical beam in processing
- Flux comparison
- Pipeline...
- and many many more!!

Progress -> Directions



. BBS – Automatic validation of versions!!

