

A first order analysis of Total Power stability using a station beamformer

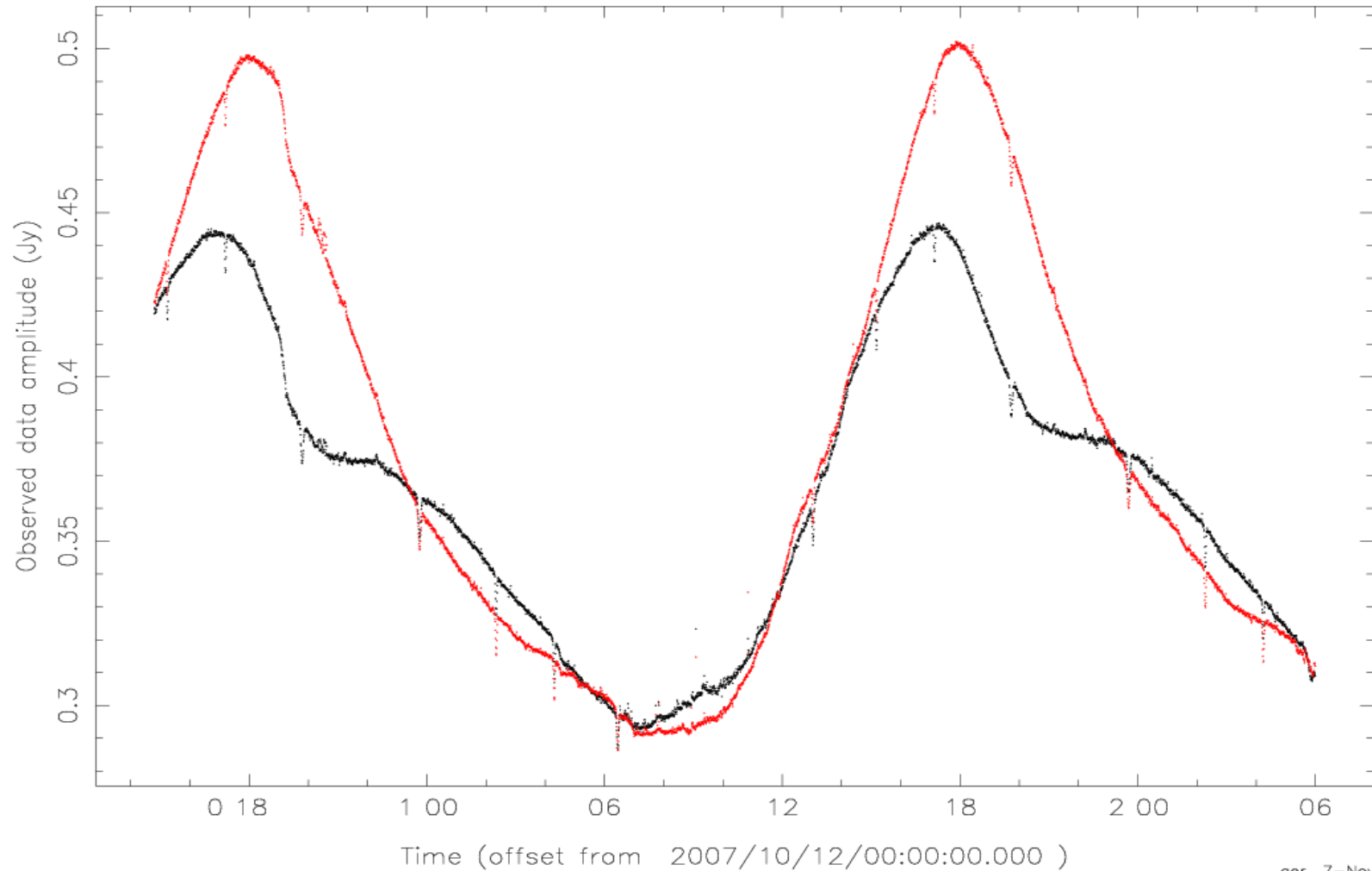
Observation L3980 12-14 October 2007

Ger de Bruyn

Analysis of observation L3980 (36h LBA from 12-14 Oct07)

Total Power on 1 dipole

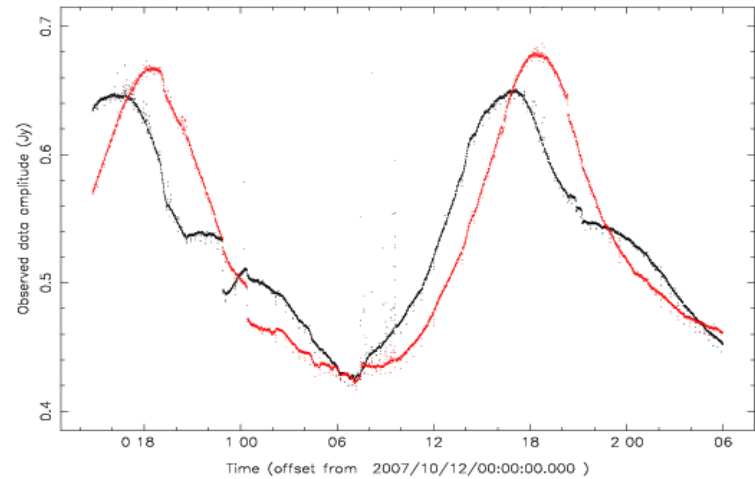
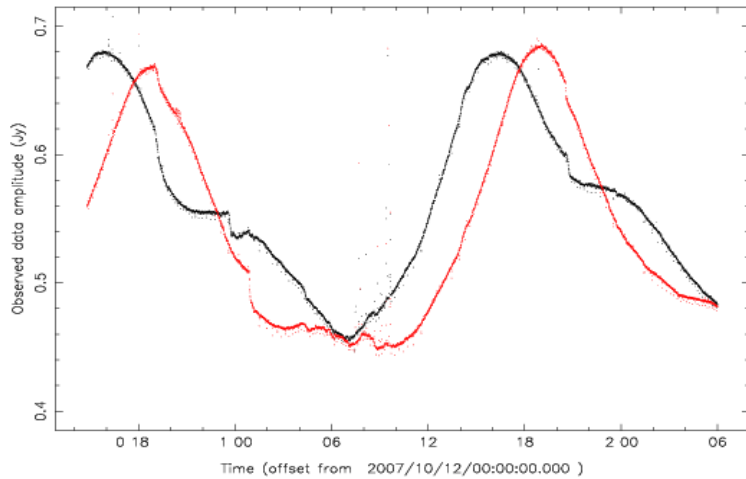
name: /dop64_2/ger/LOFAR/CS1/data/13oct07-L3980/SB35.MS Spectral Window: 1 Polarization: 1 Fields: BI
XX YY
Antenna1 = 8 Antenna2 = 8



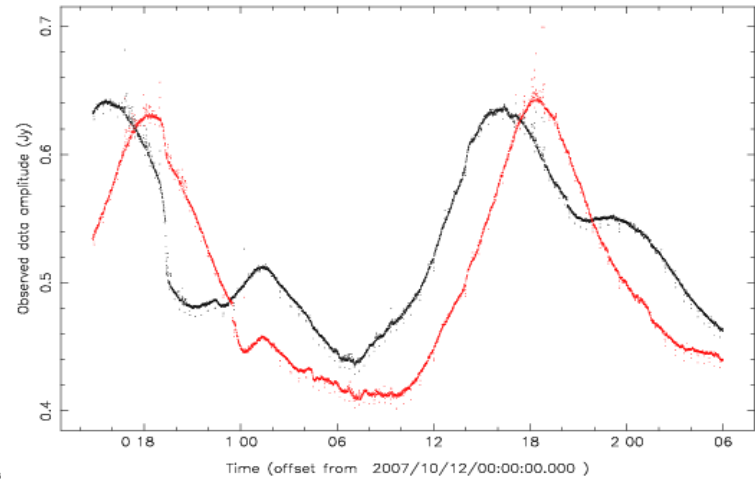
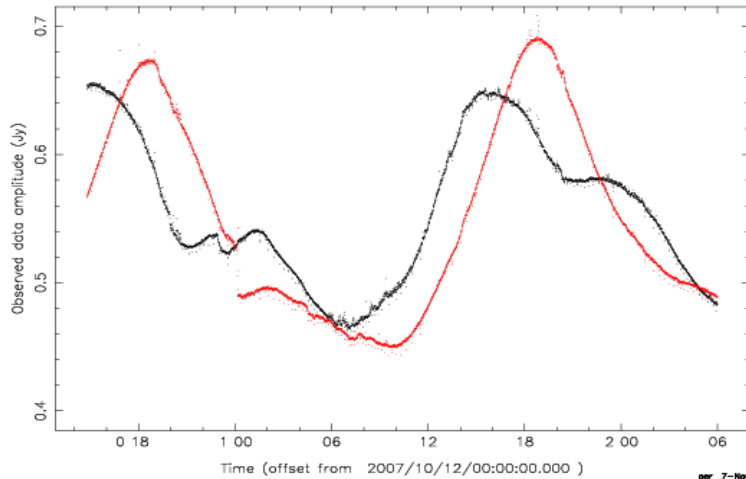
Analysis of observation L3980 (36h LBA from 12-14 Oct07)

Total Power on 4 dipoles in CS008

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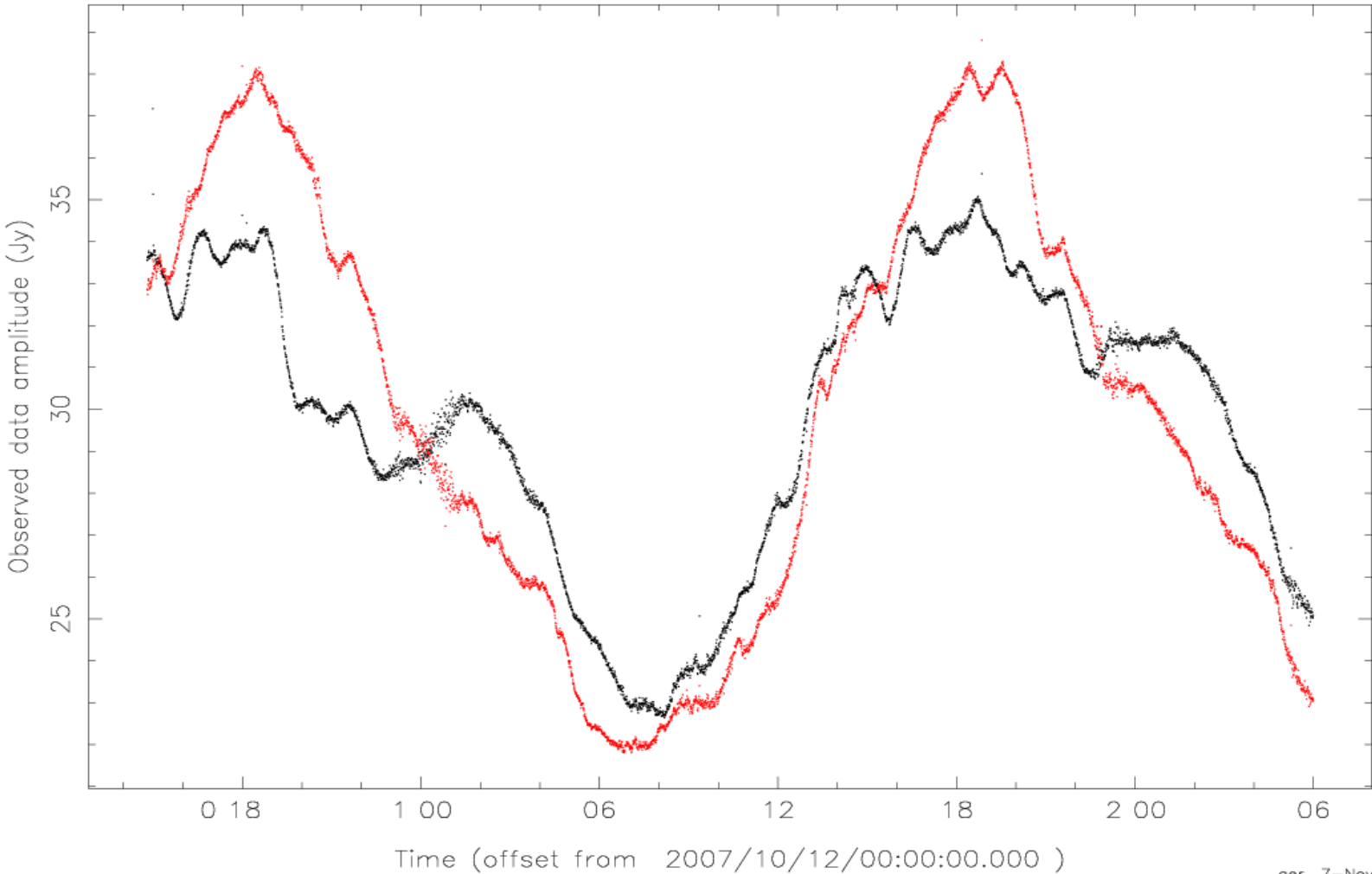


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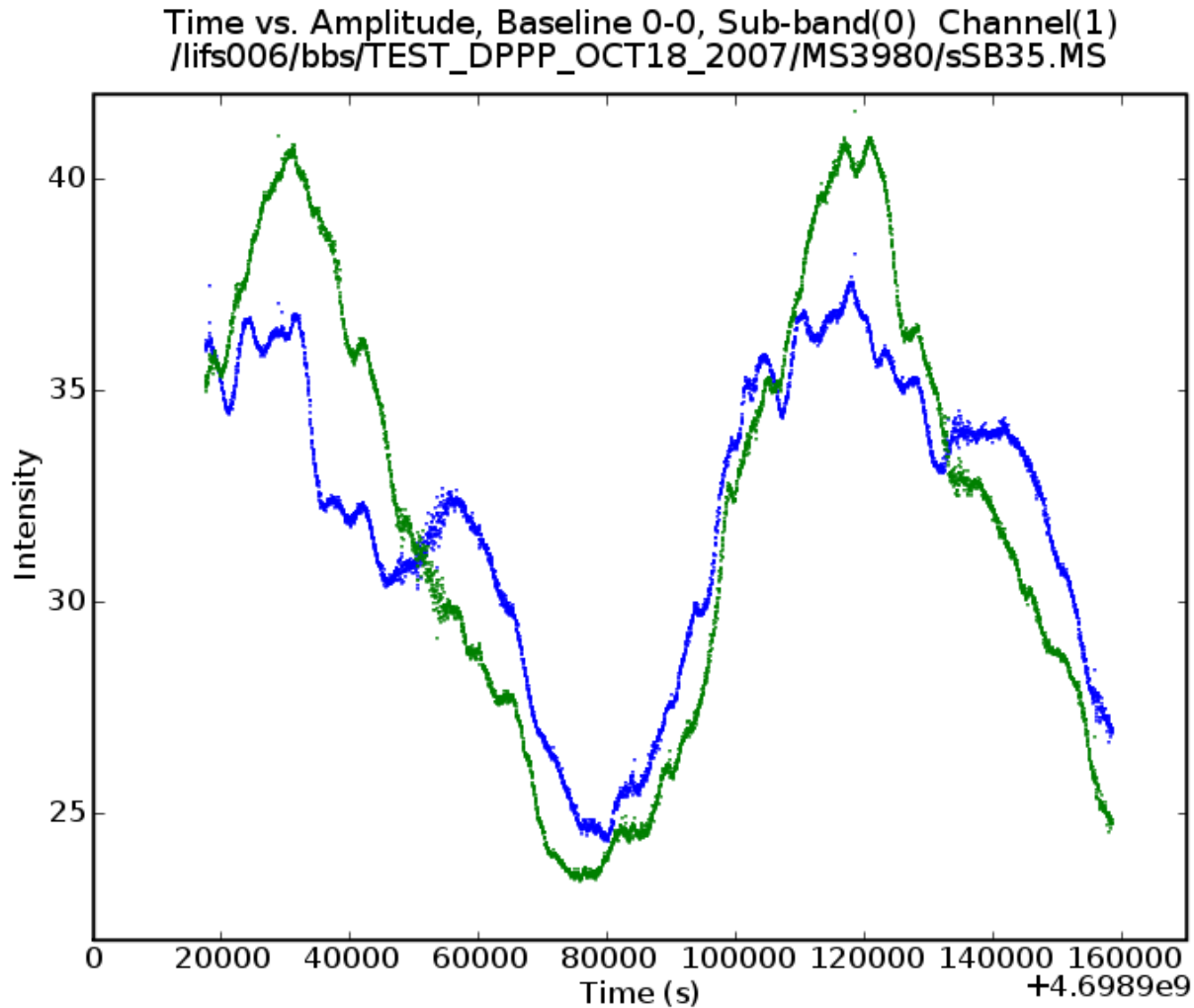


Total power on microstation 0 (Antenna 1, 48 dipole sum)

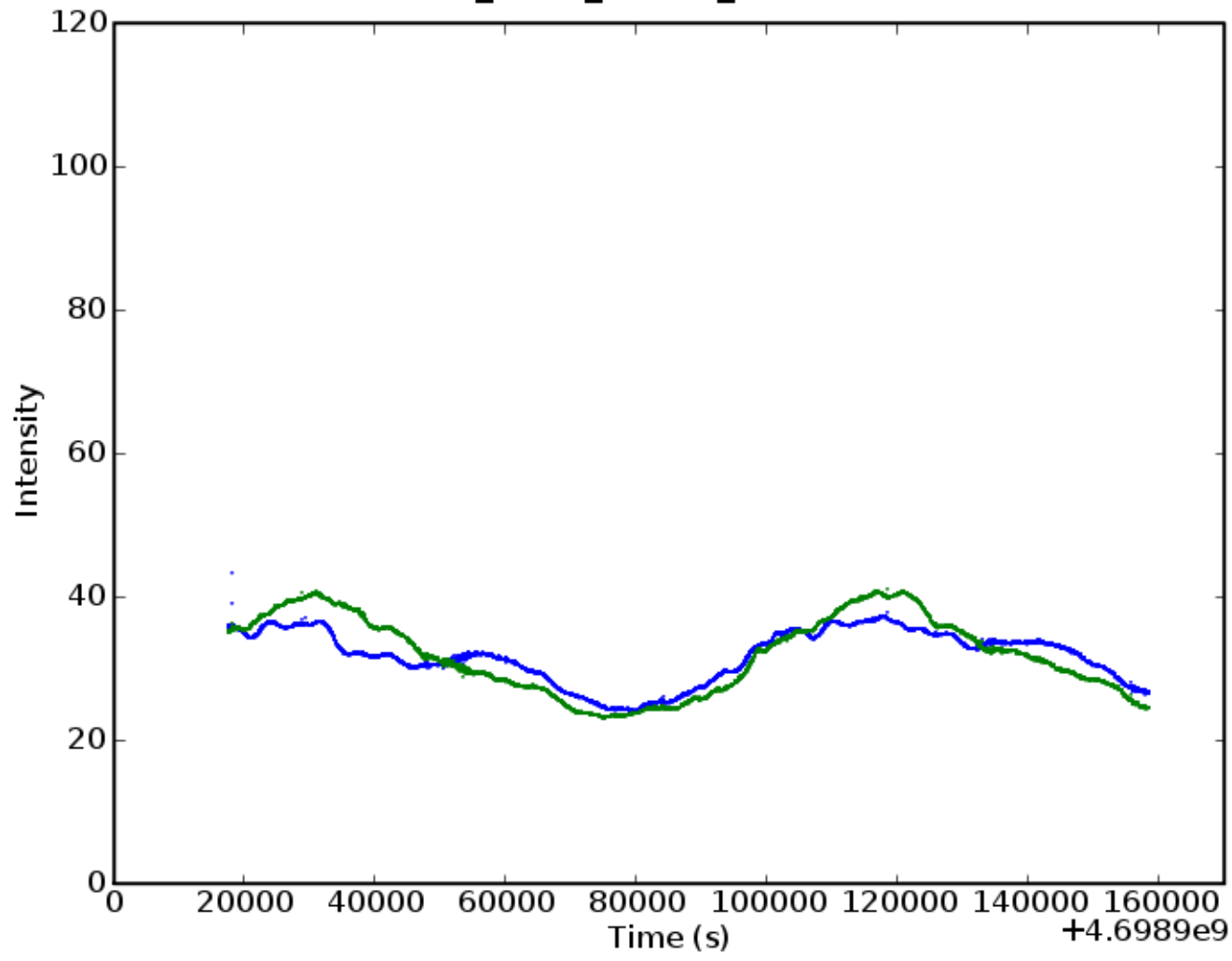
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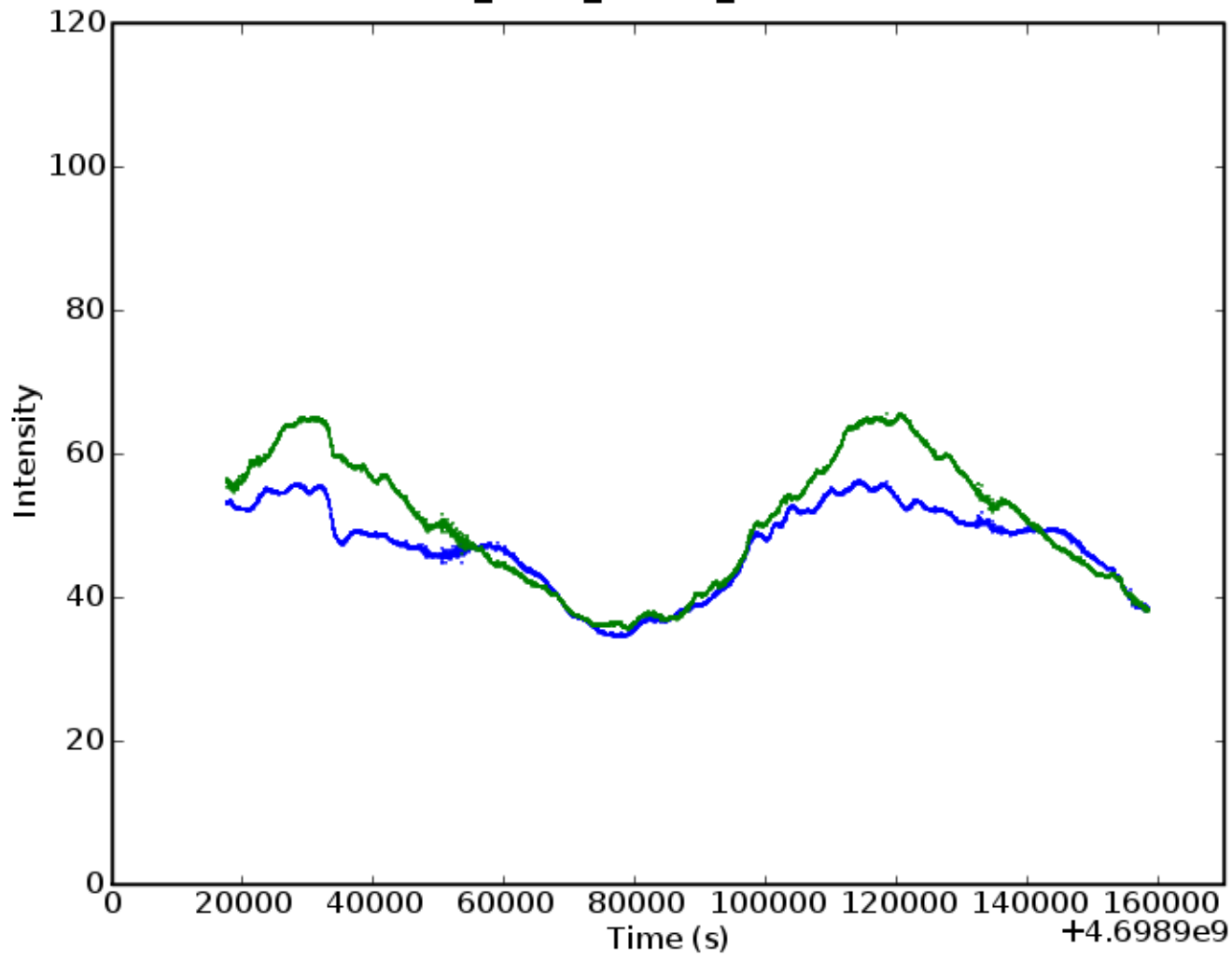
Data agree with those gone through Pandey's DP³
and plotted with Casey and Joe's pyDAL



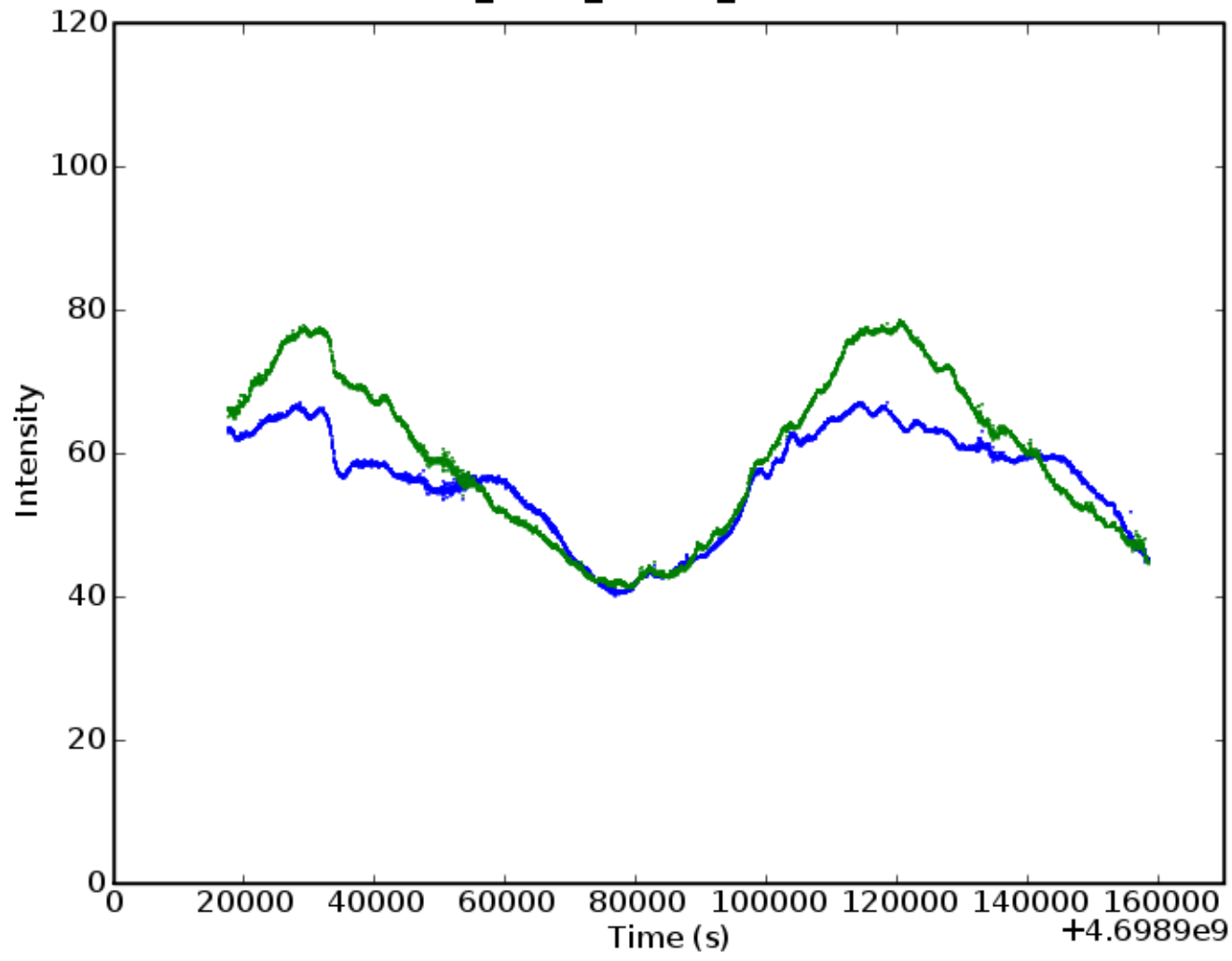
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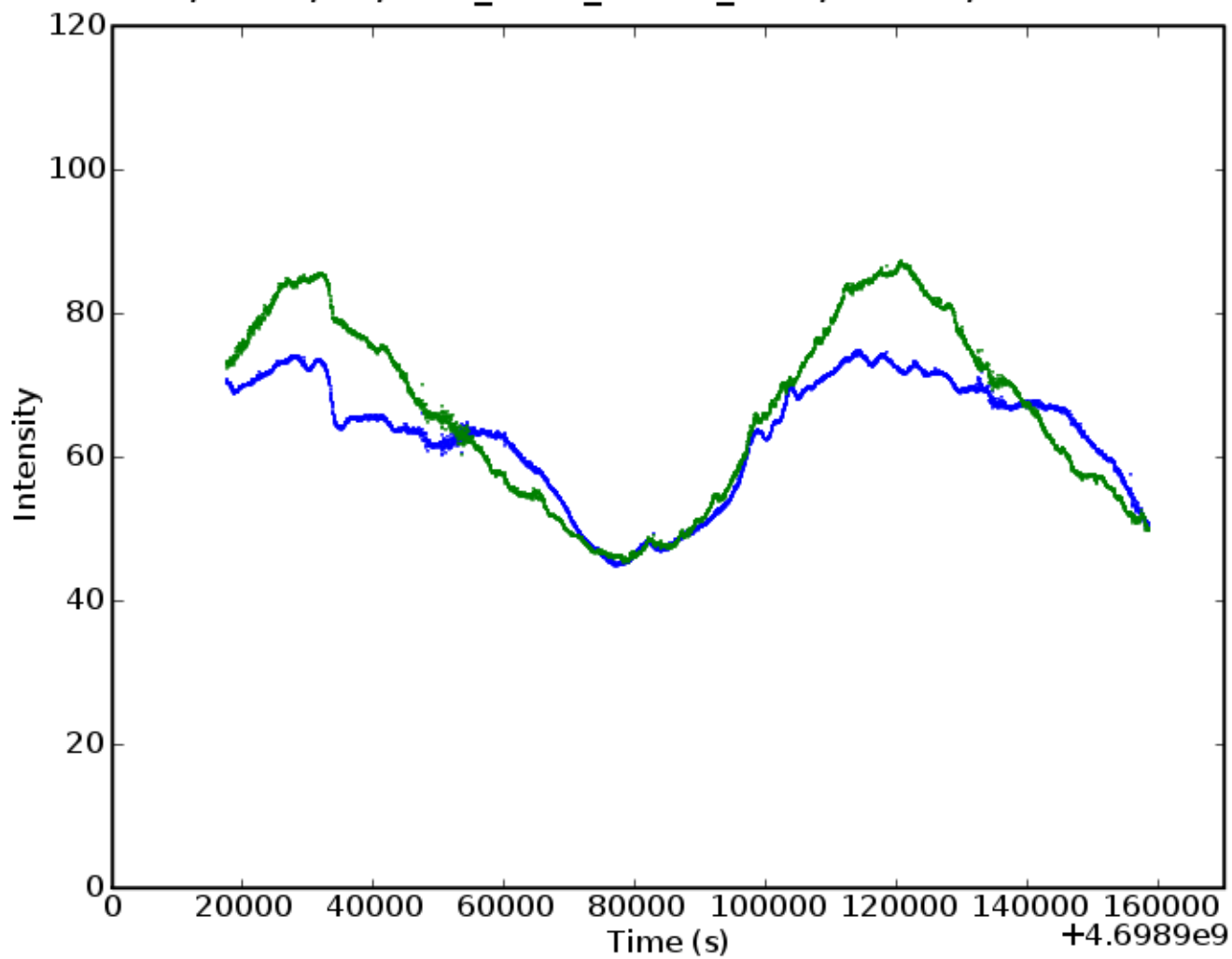
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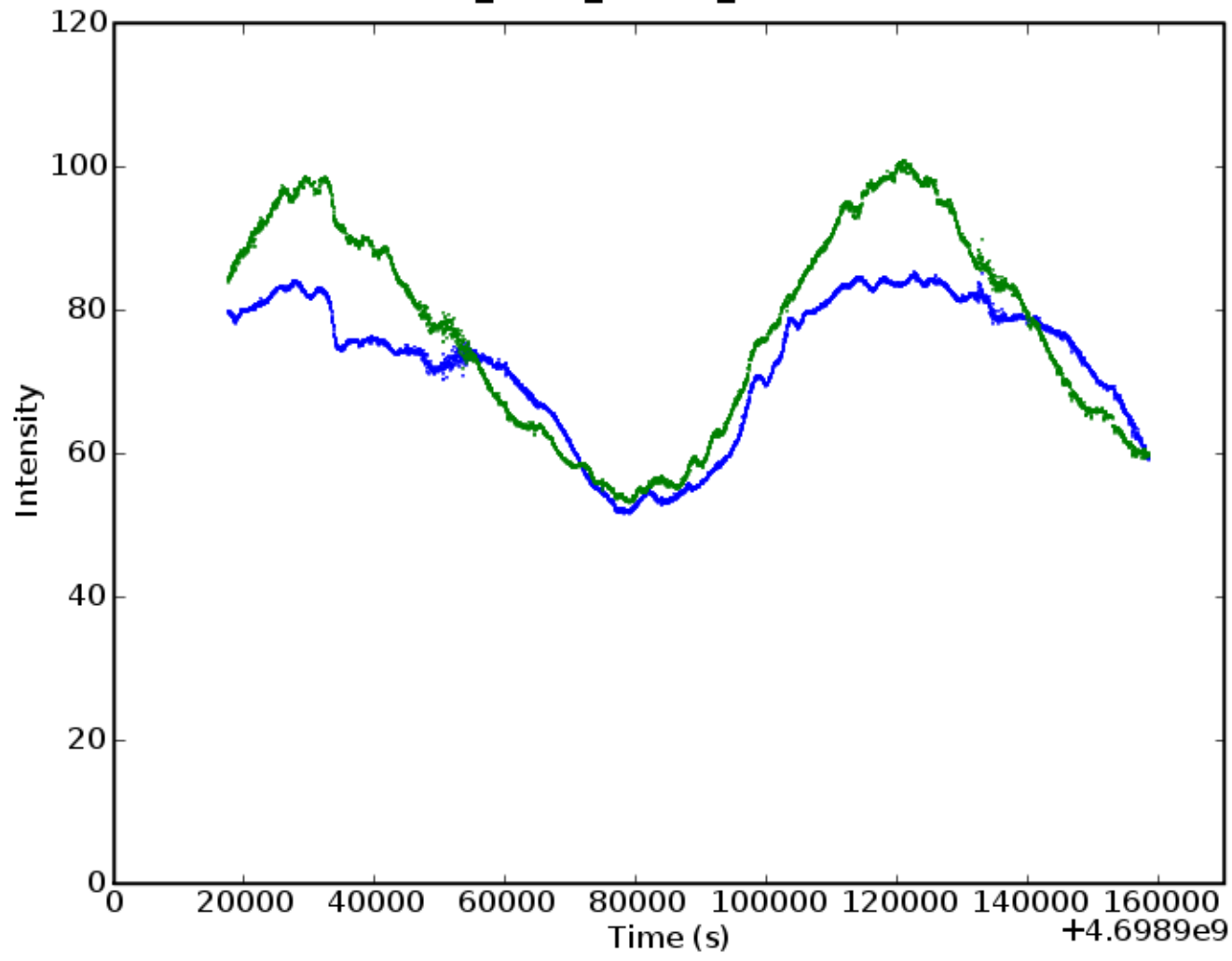
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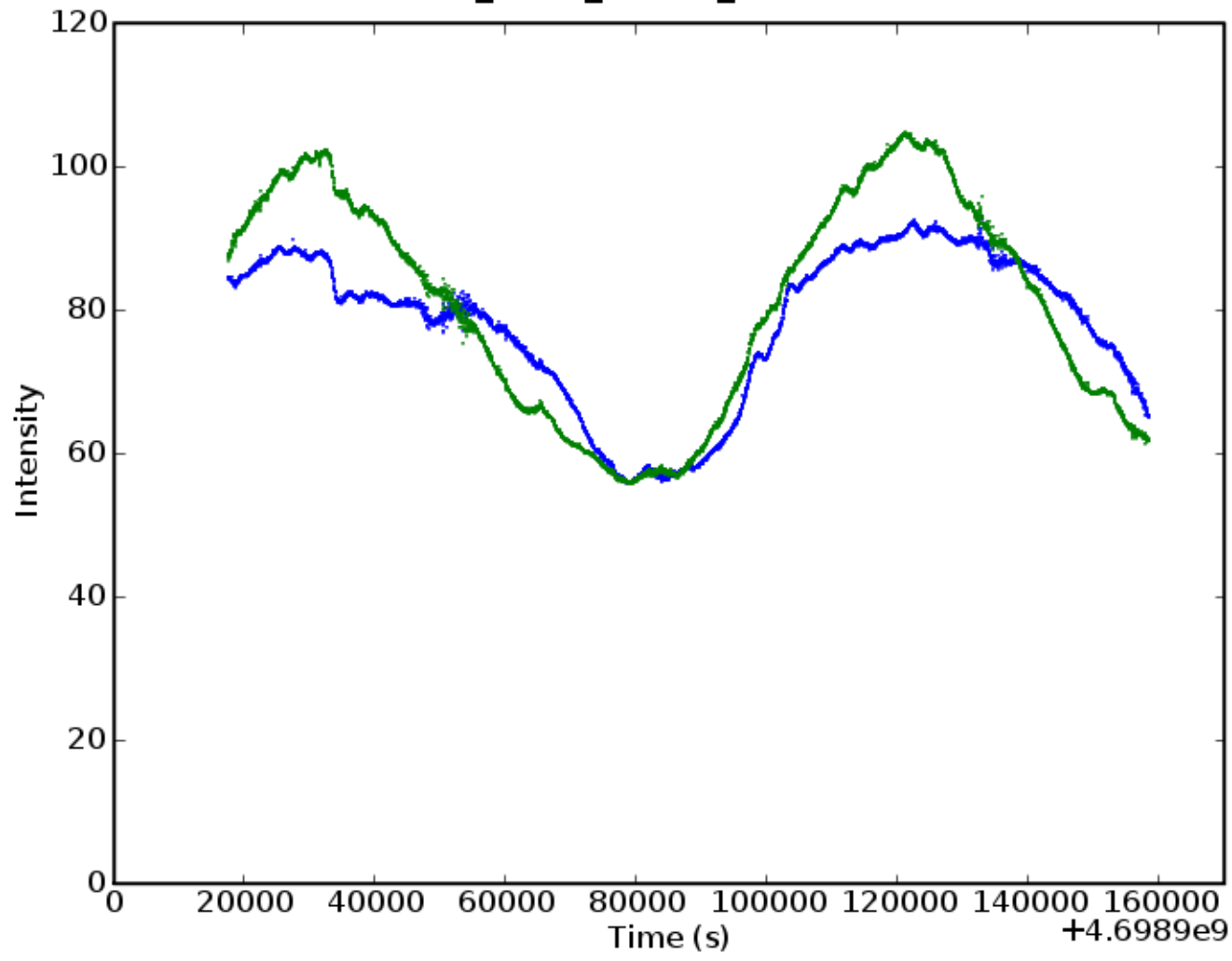
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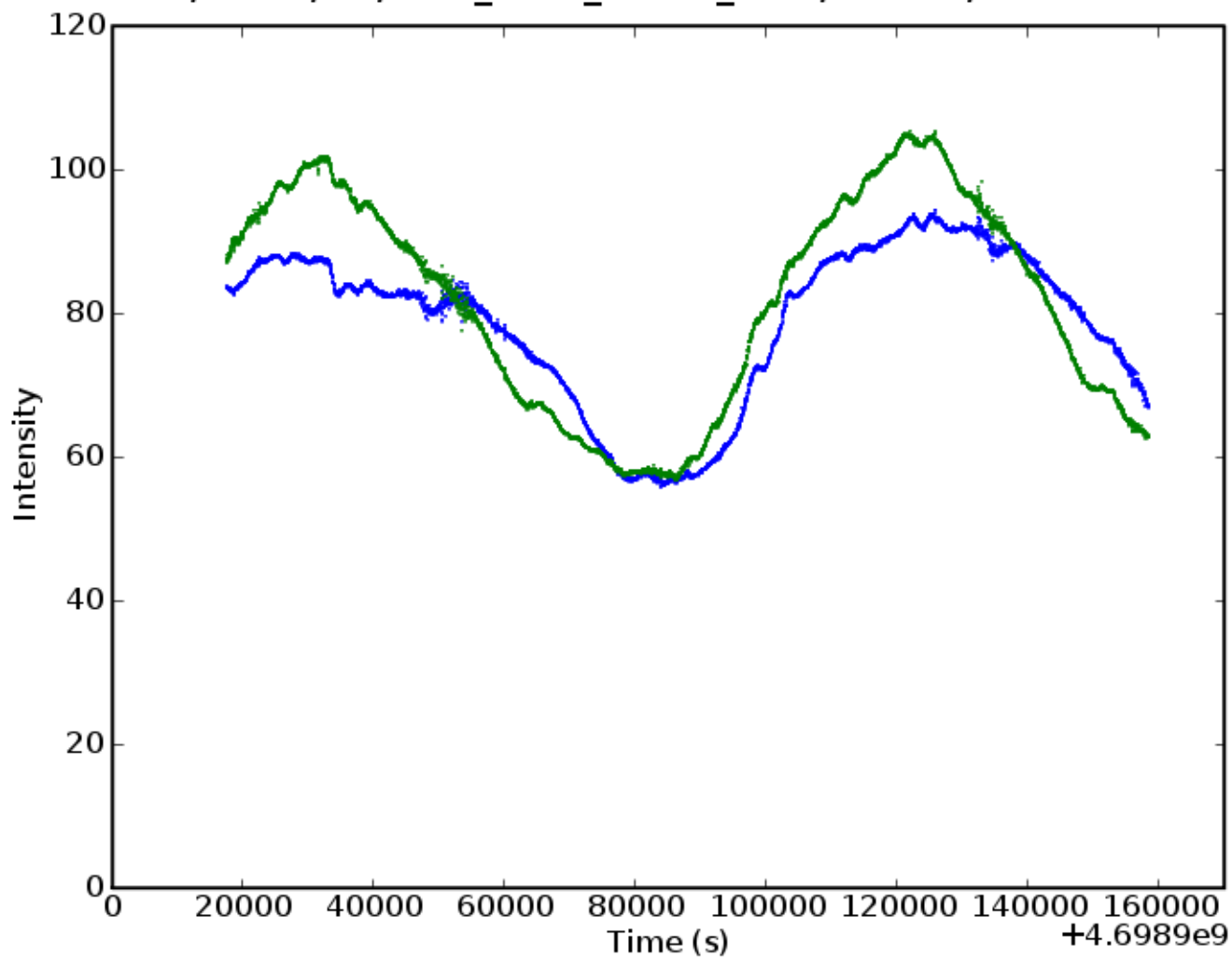
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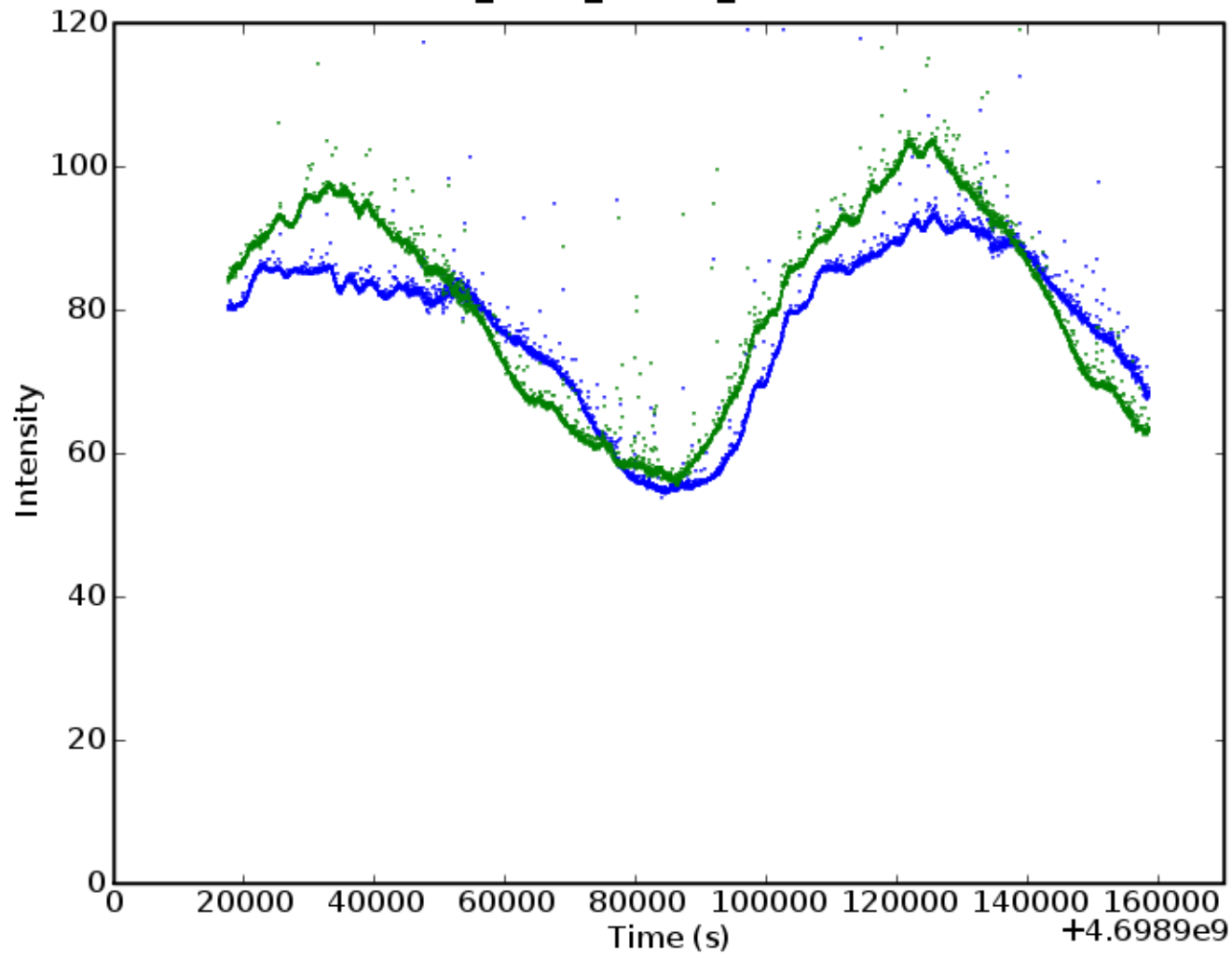
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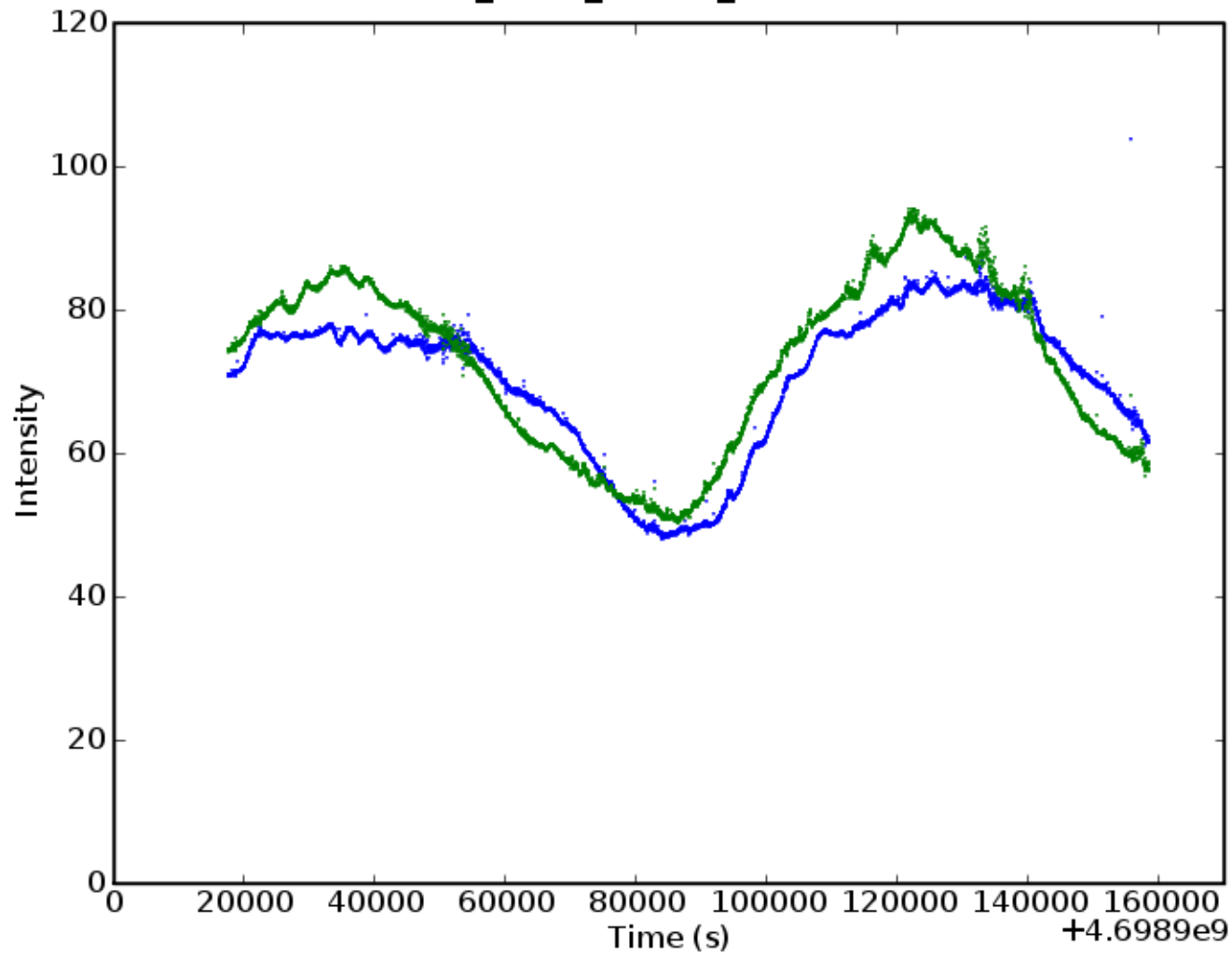
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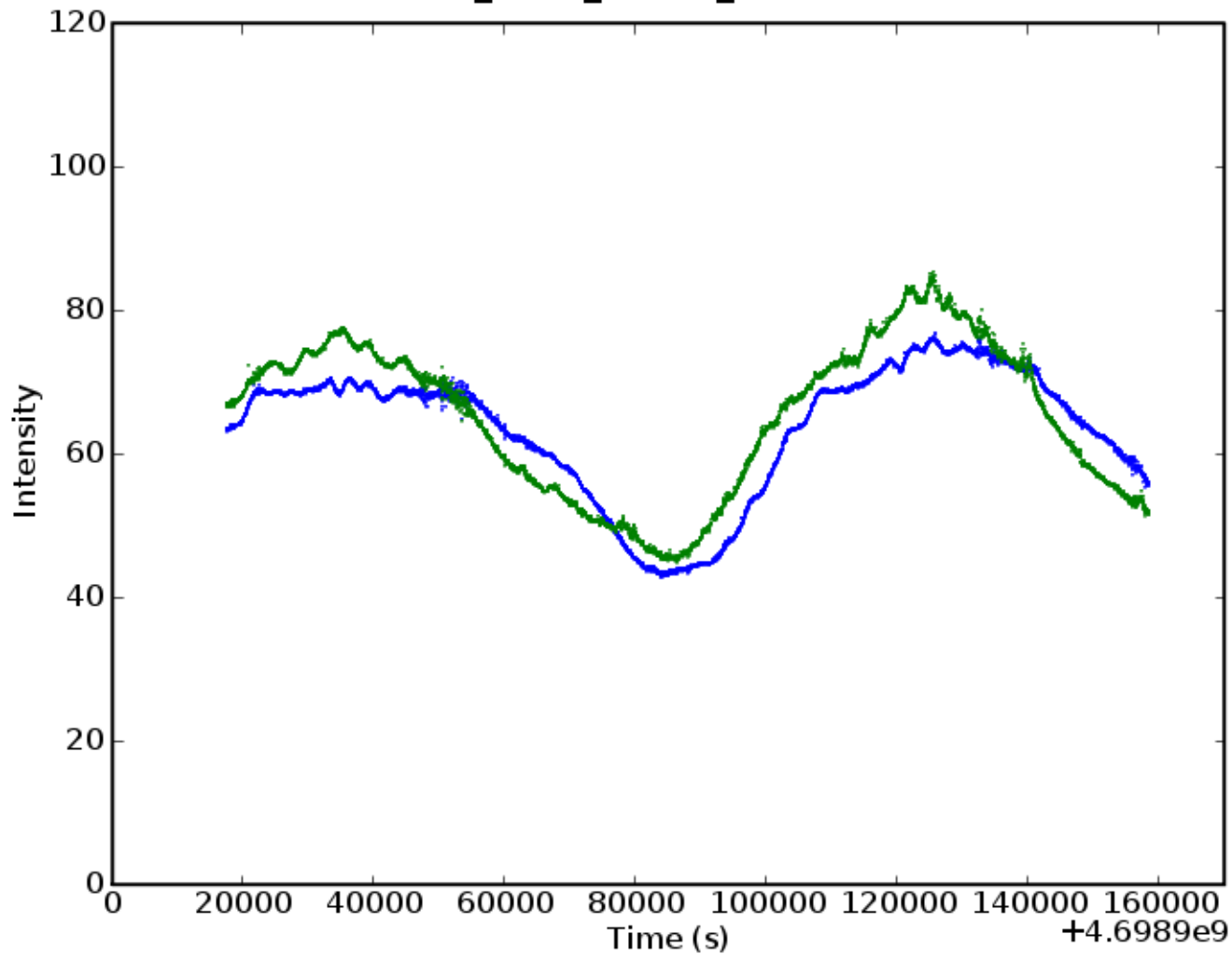
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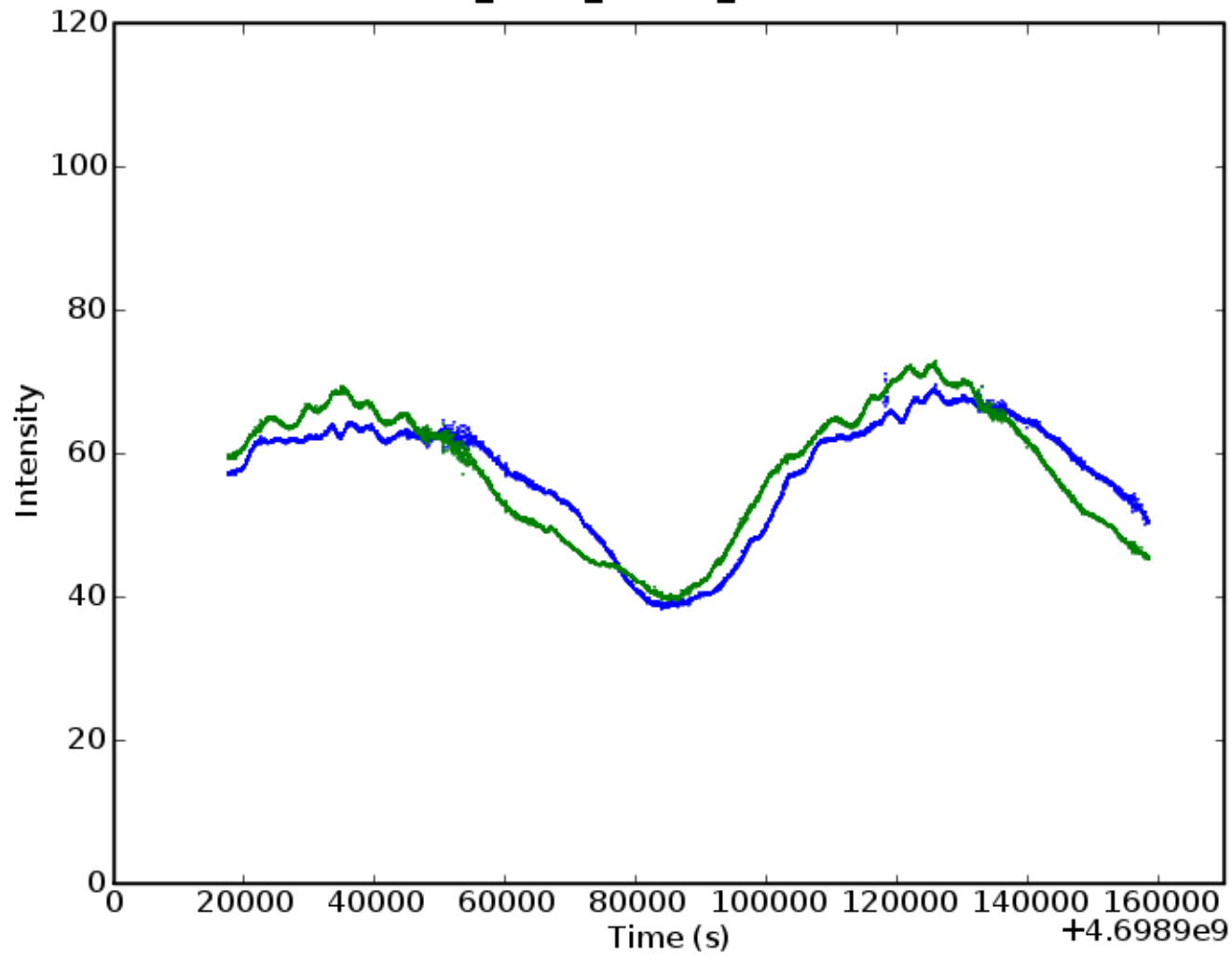
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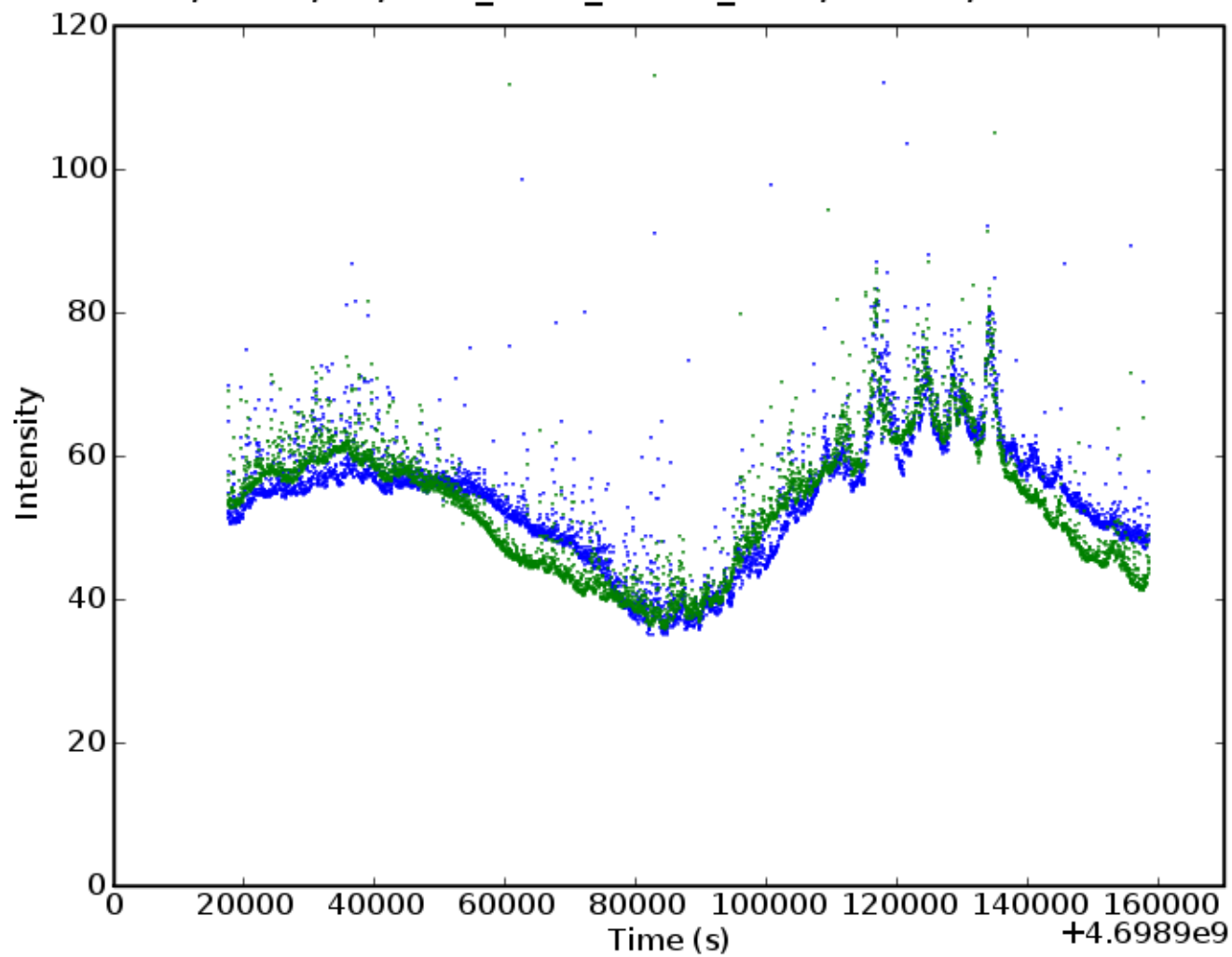
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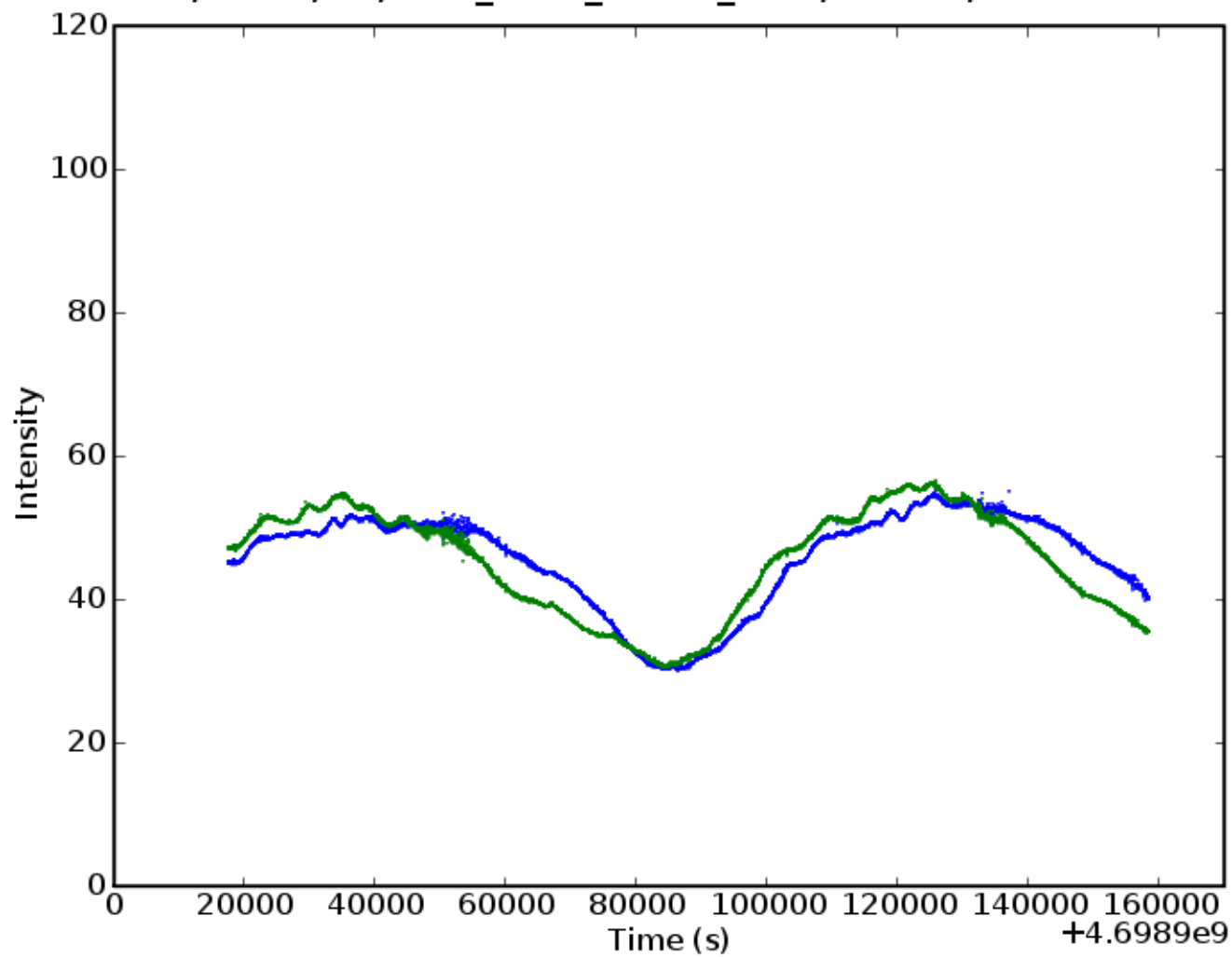
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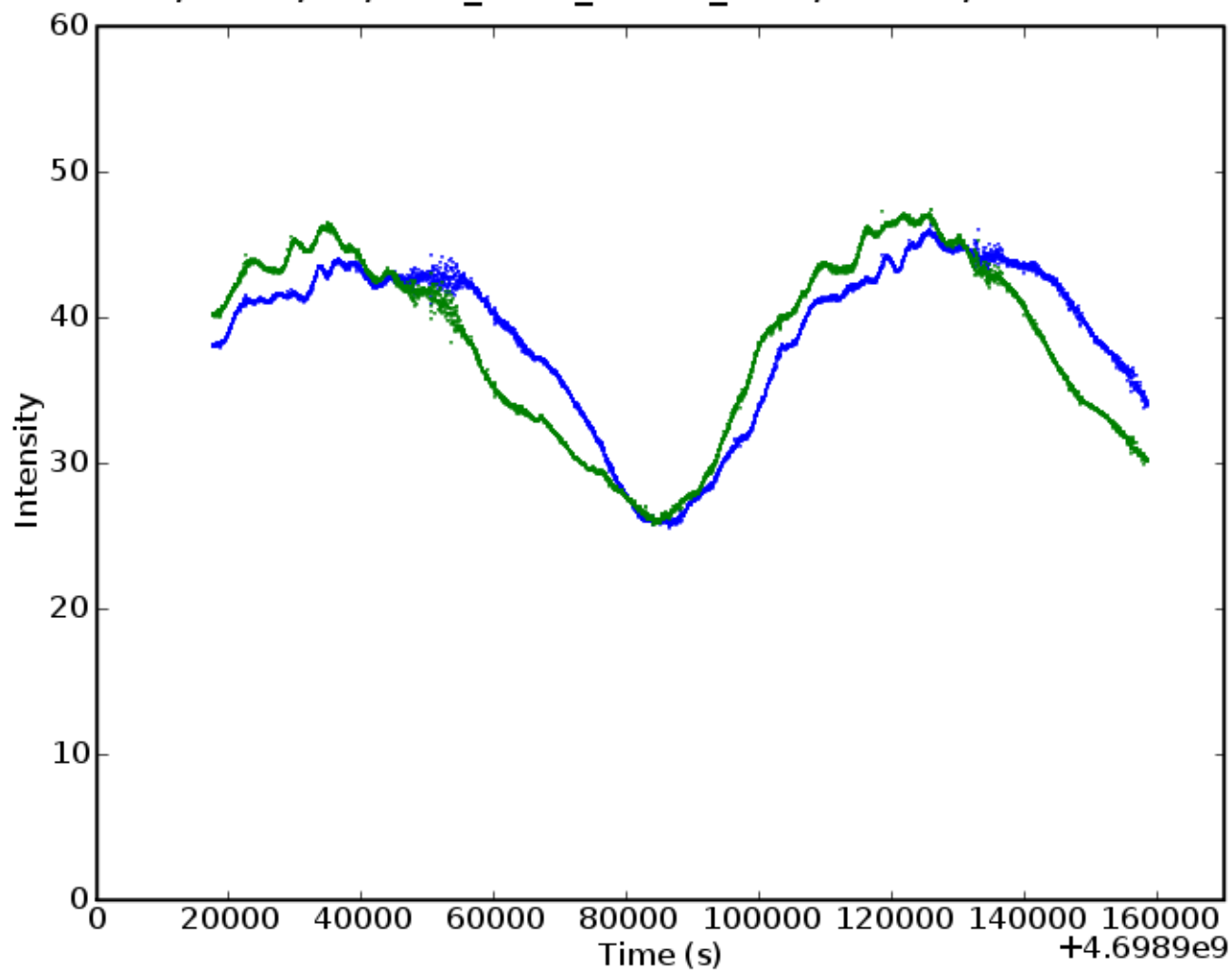
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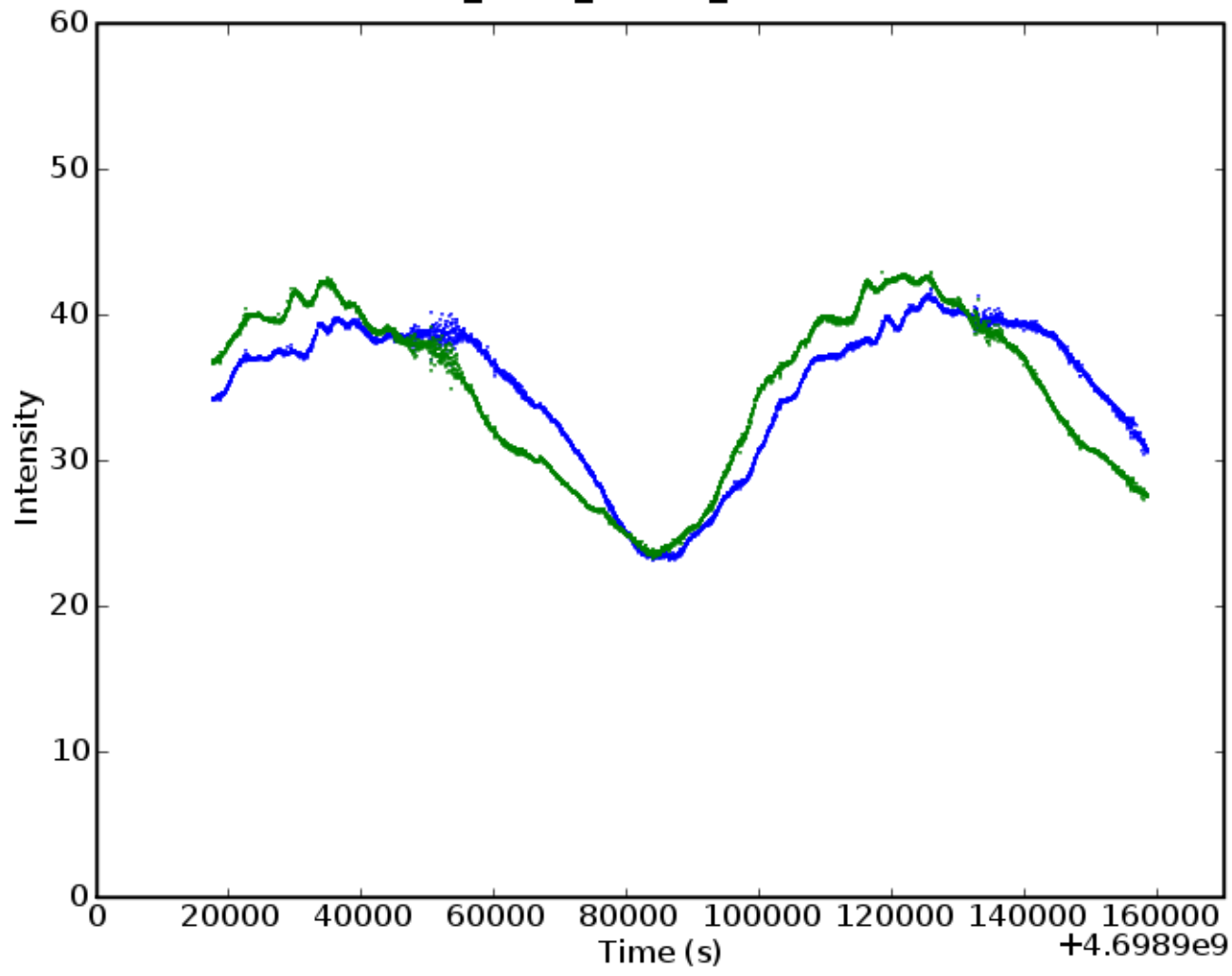
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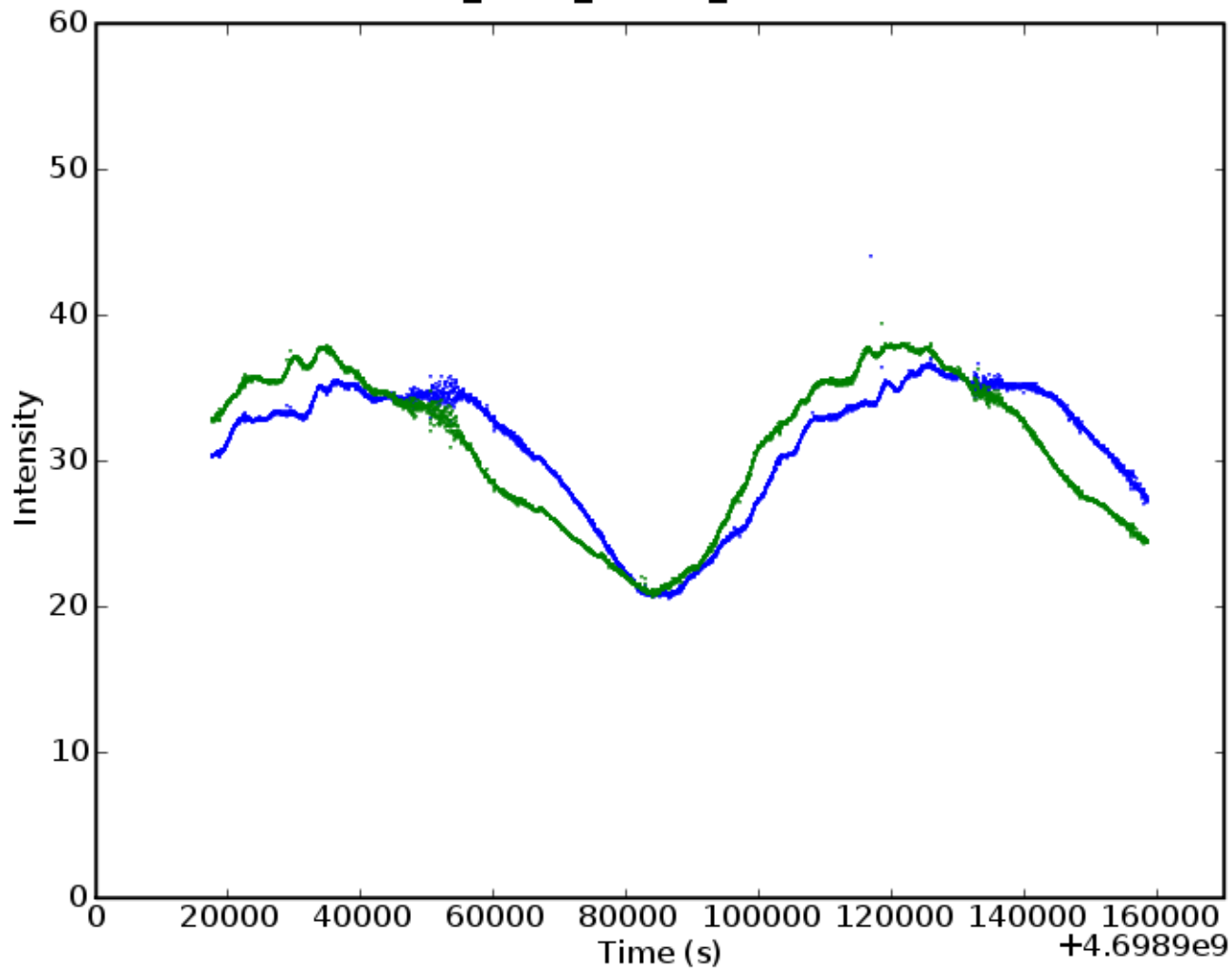
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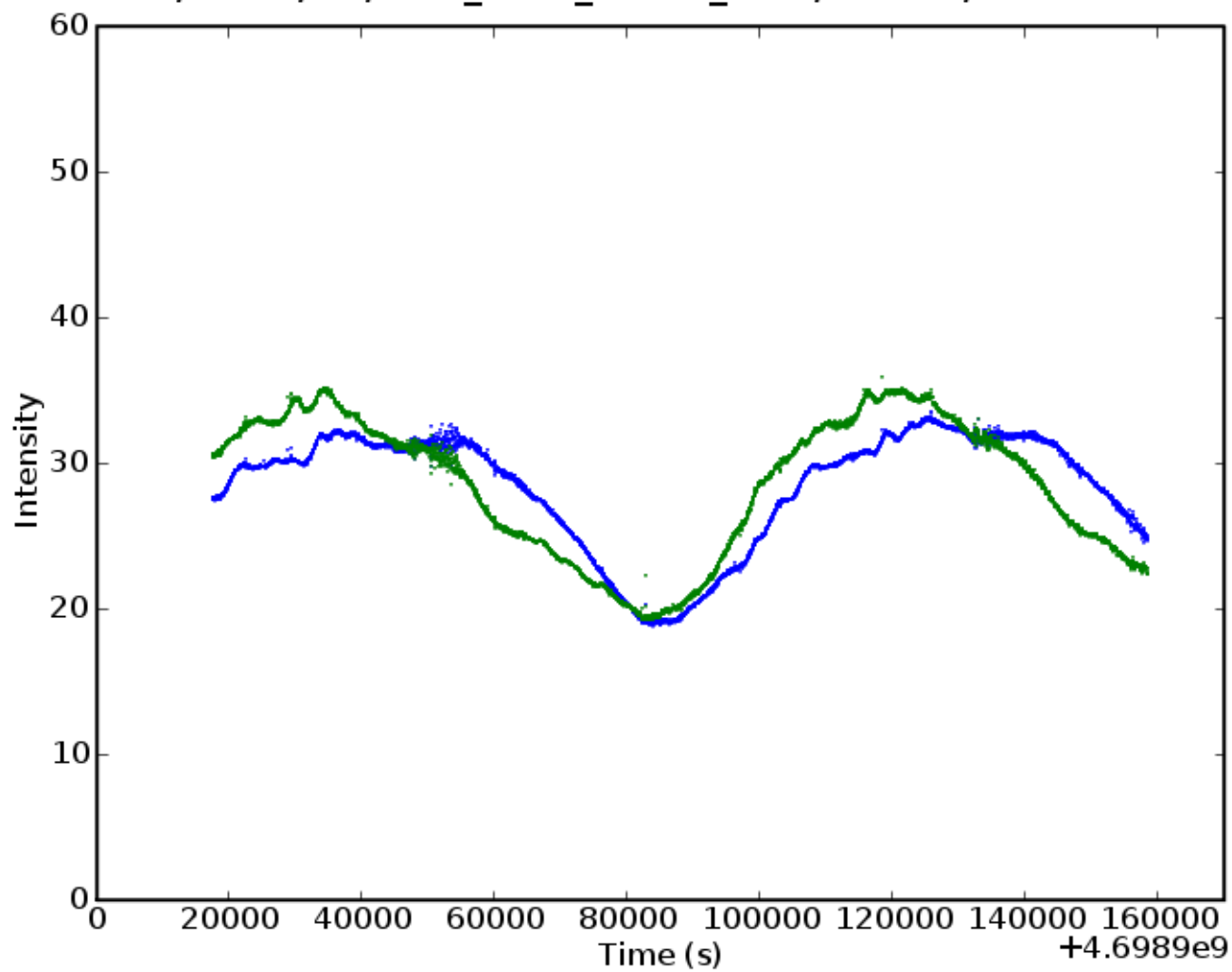
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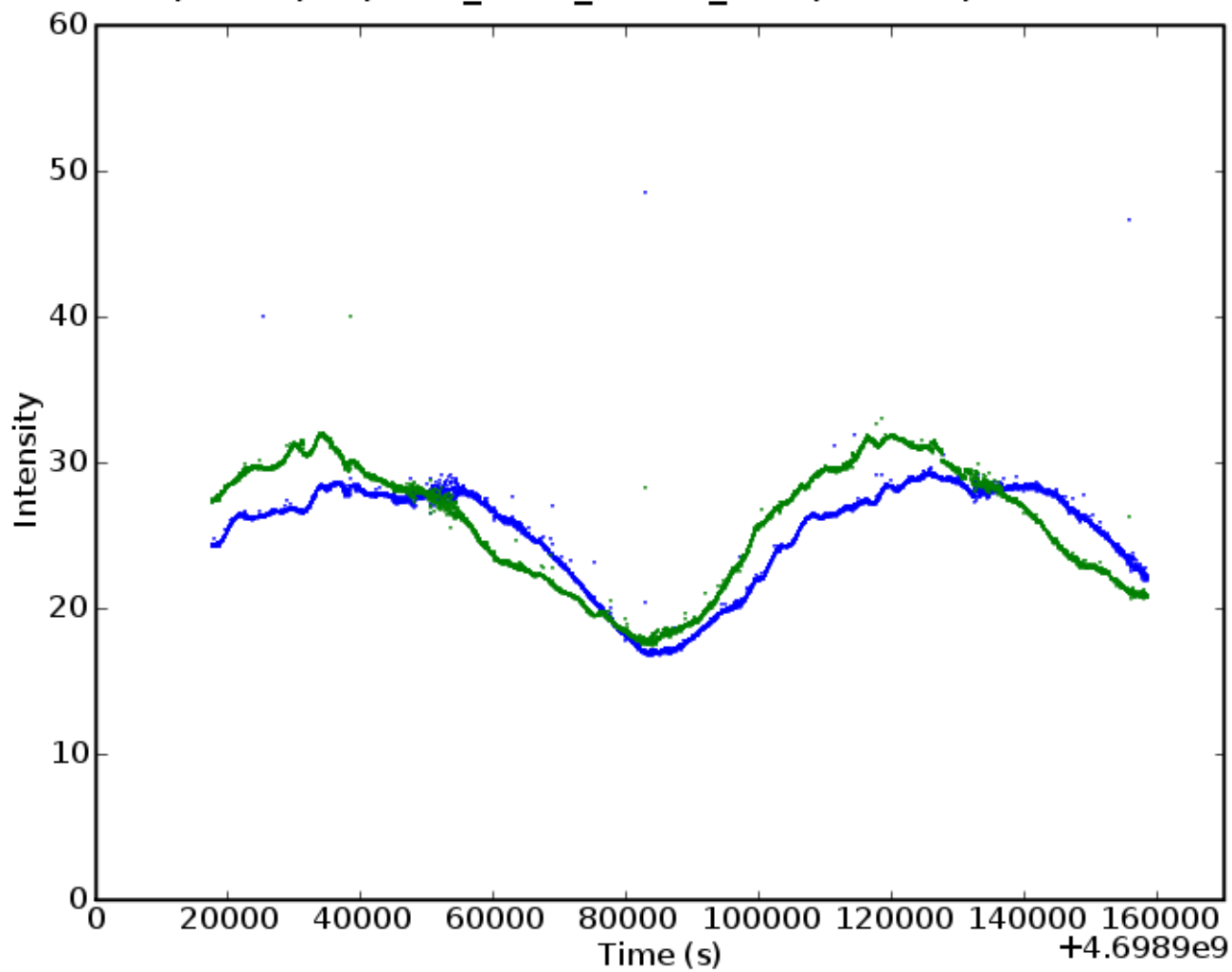
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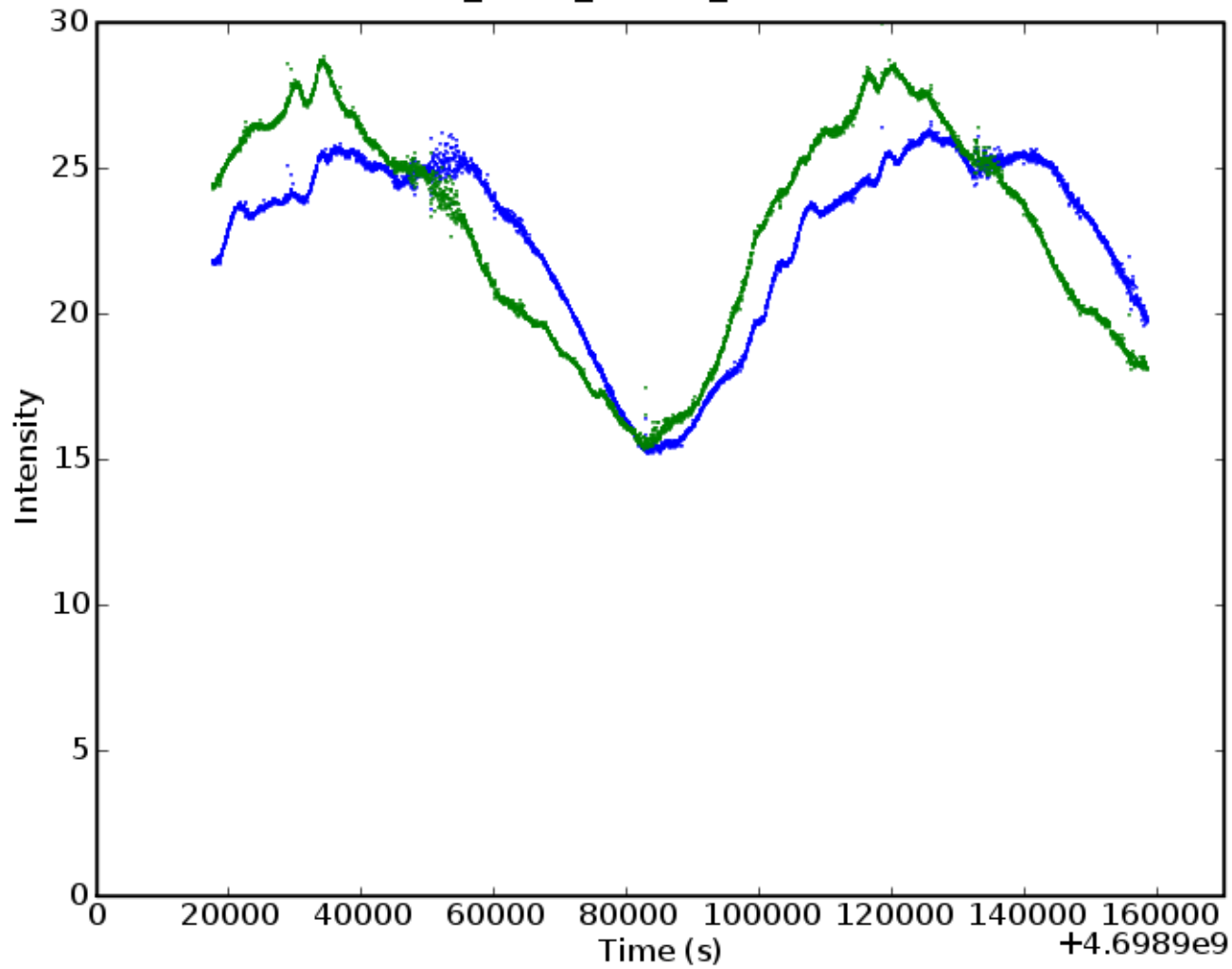
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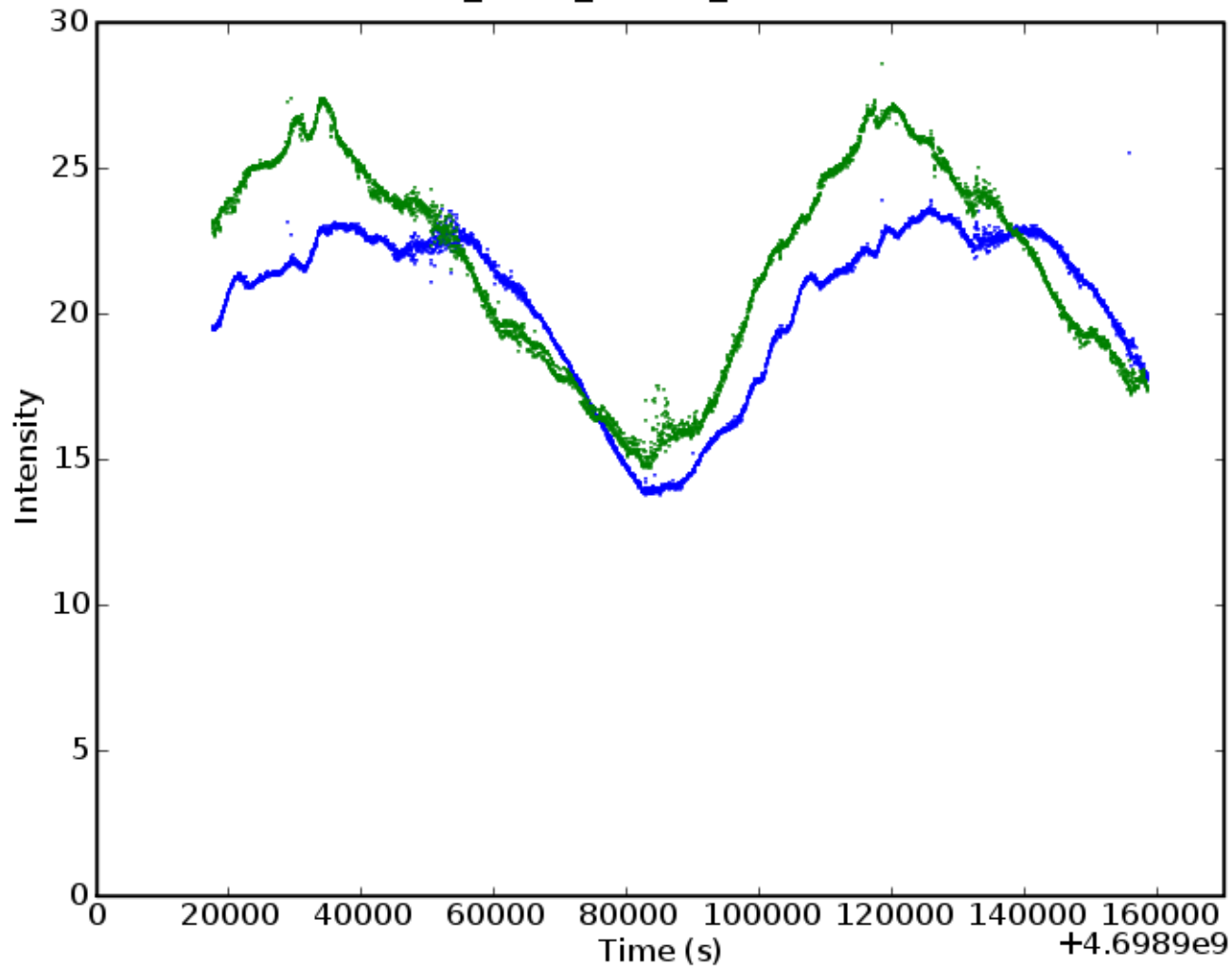
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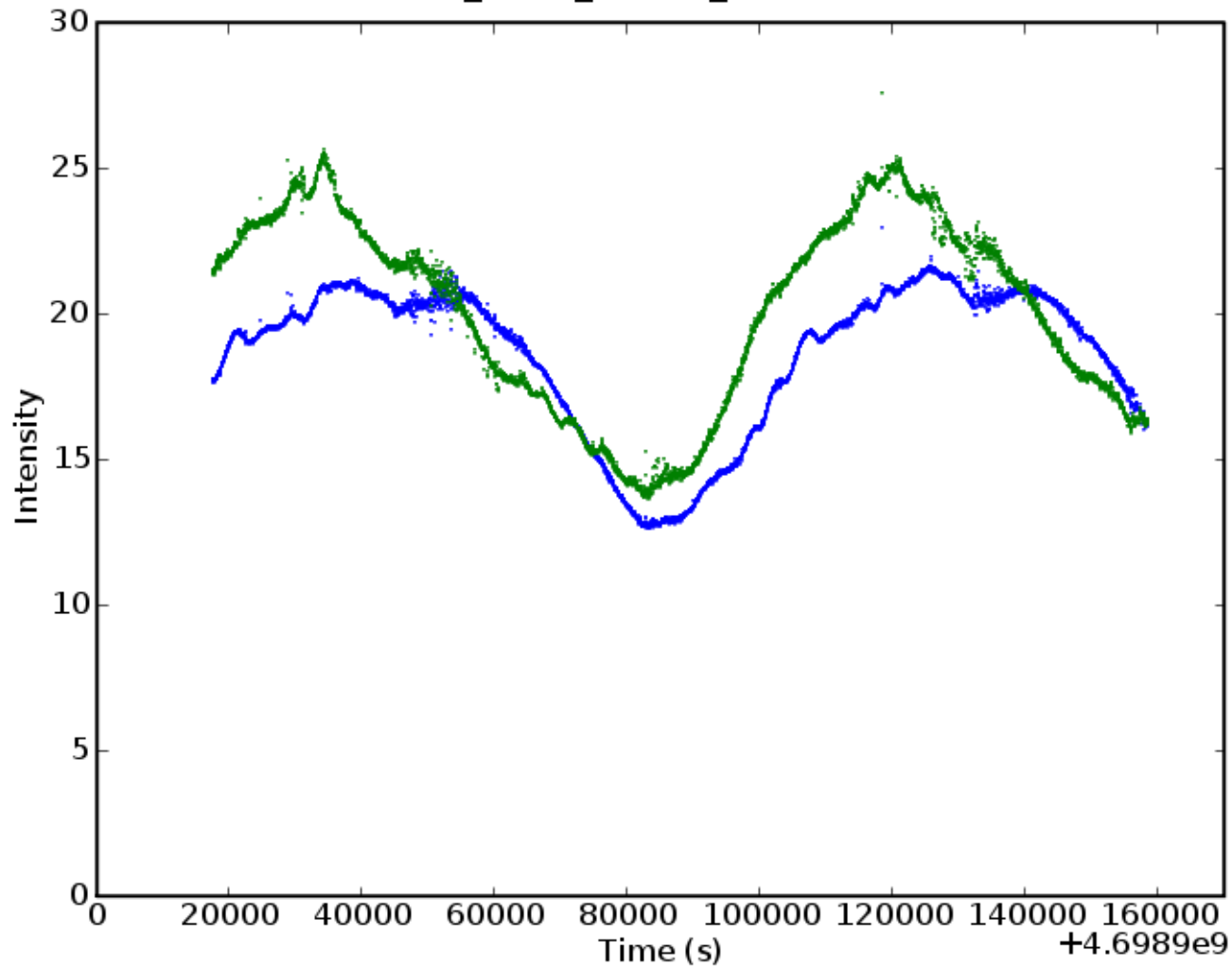
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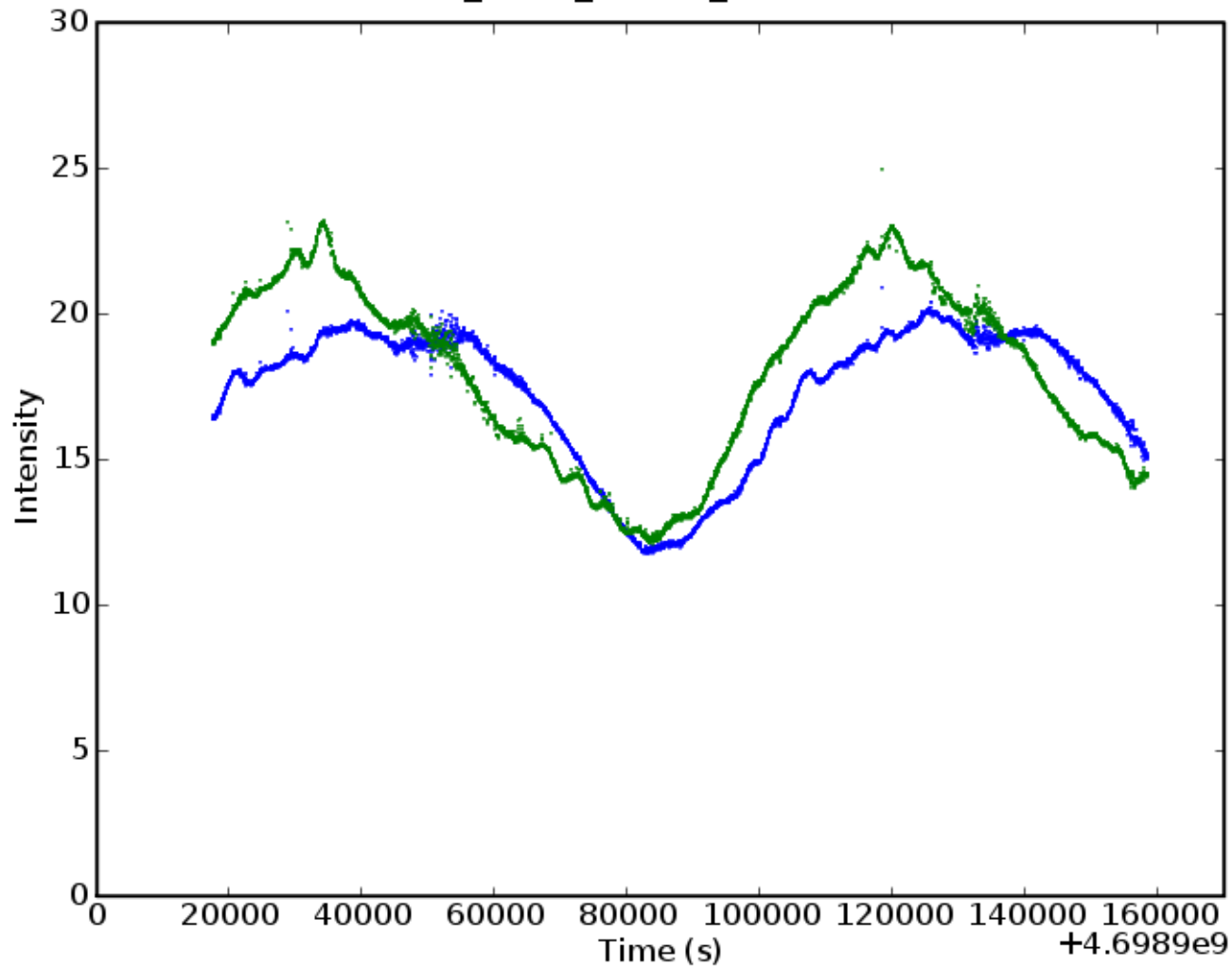
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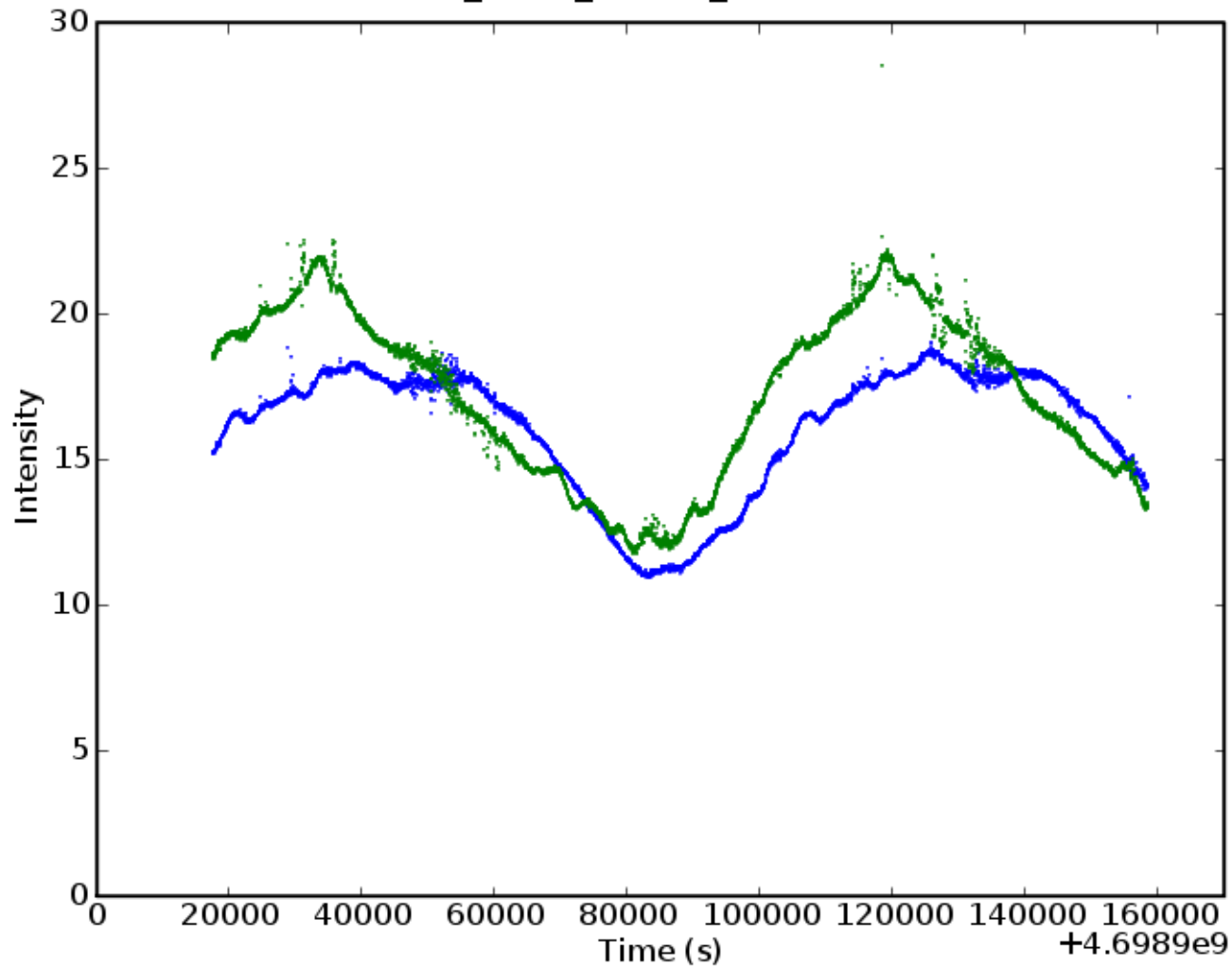
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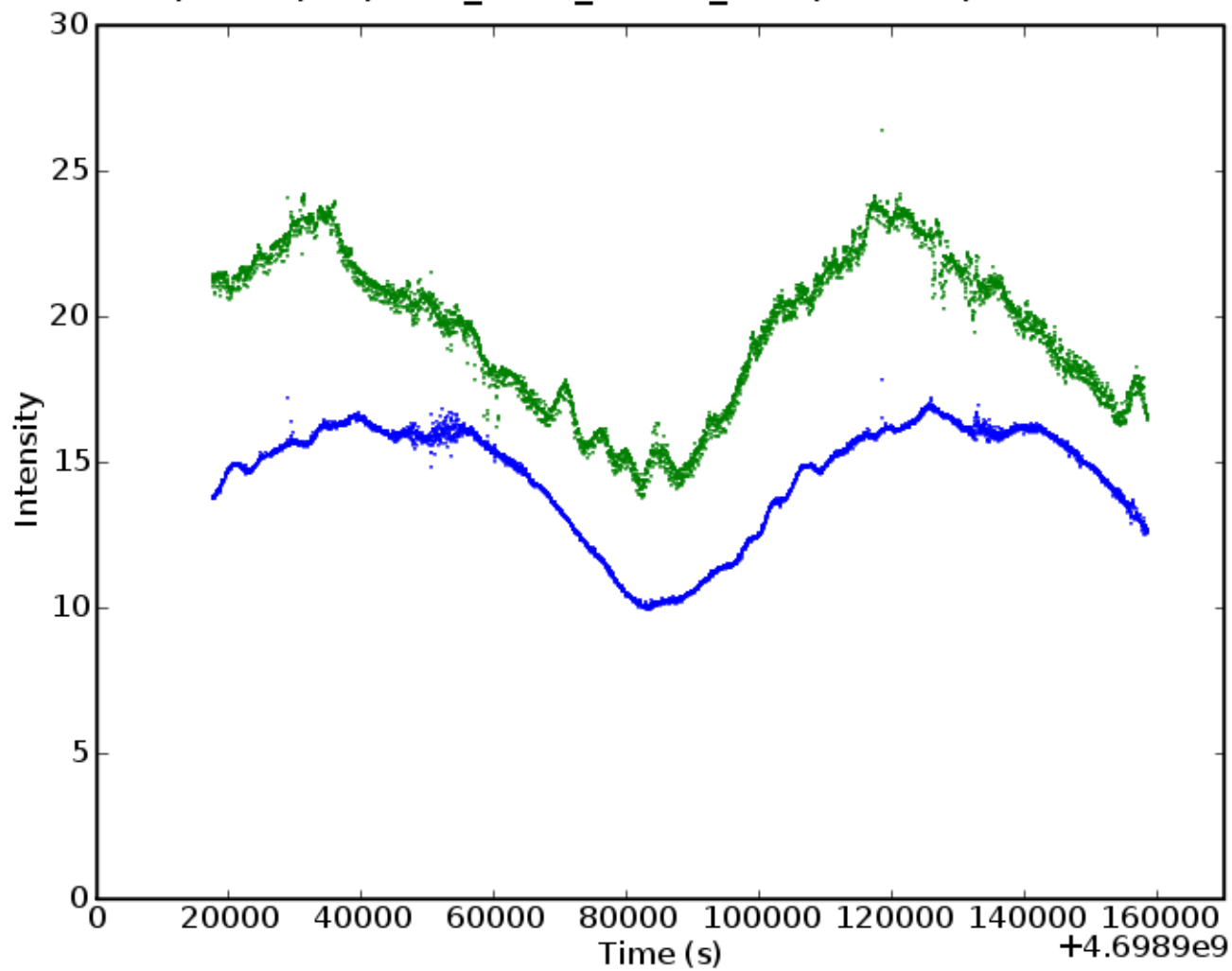
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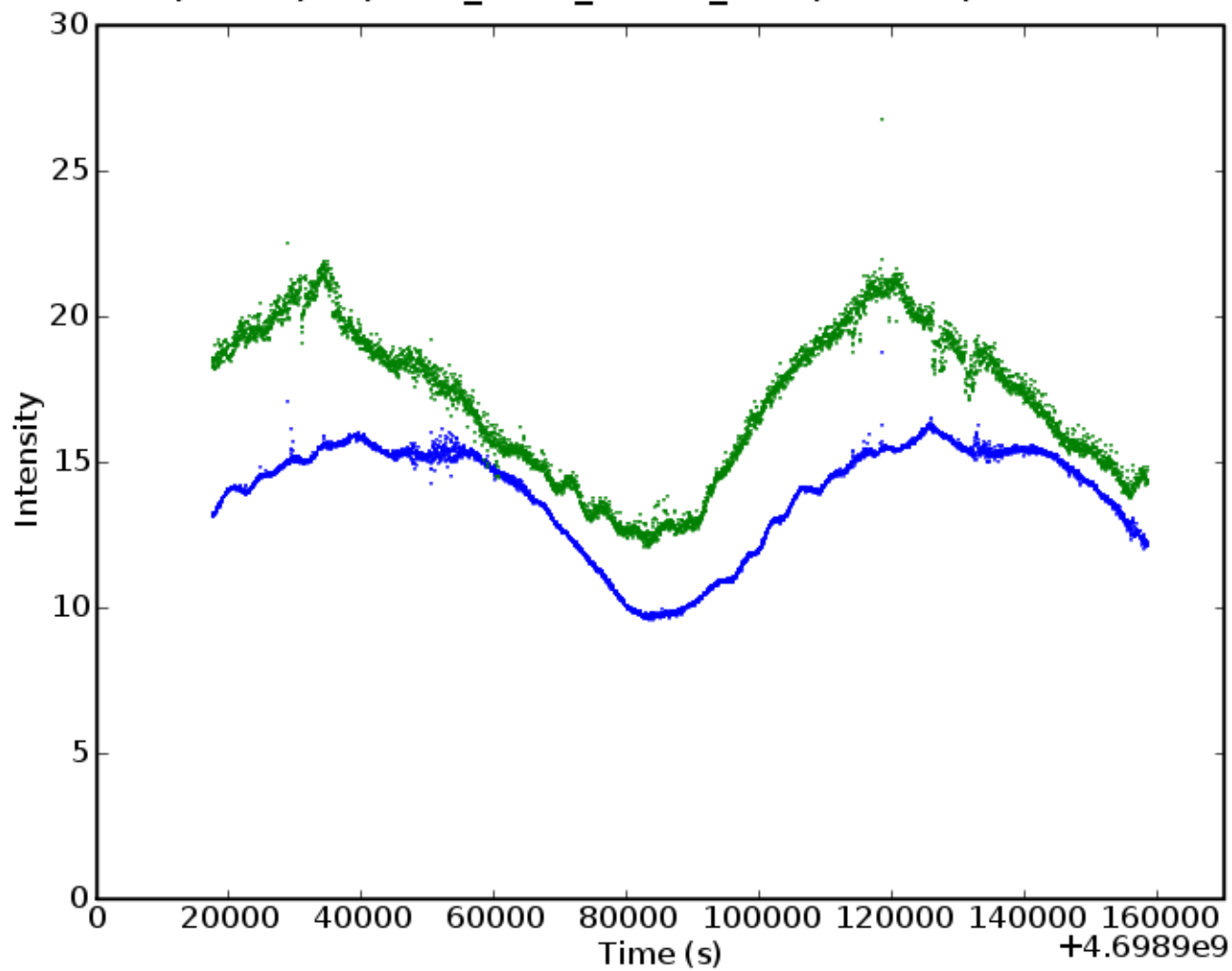
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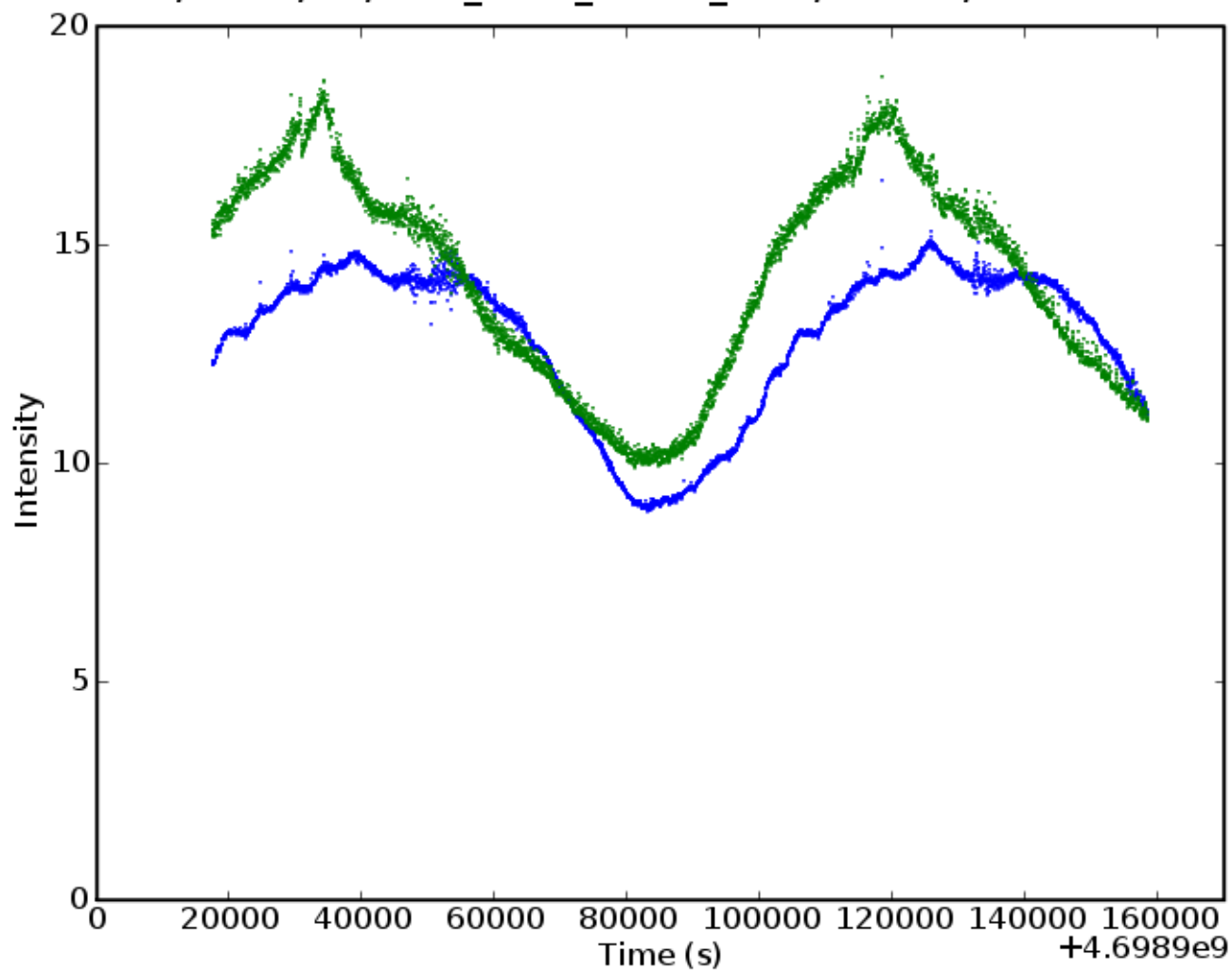
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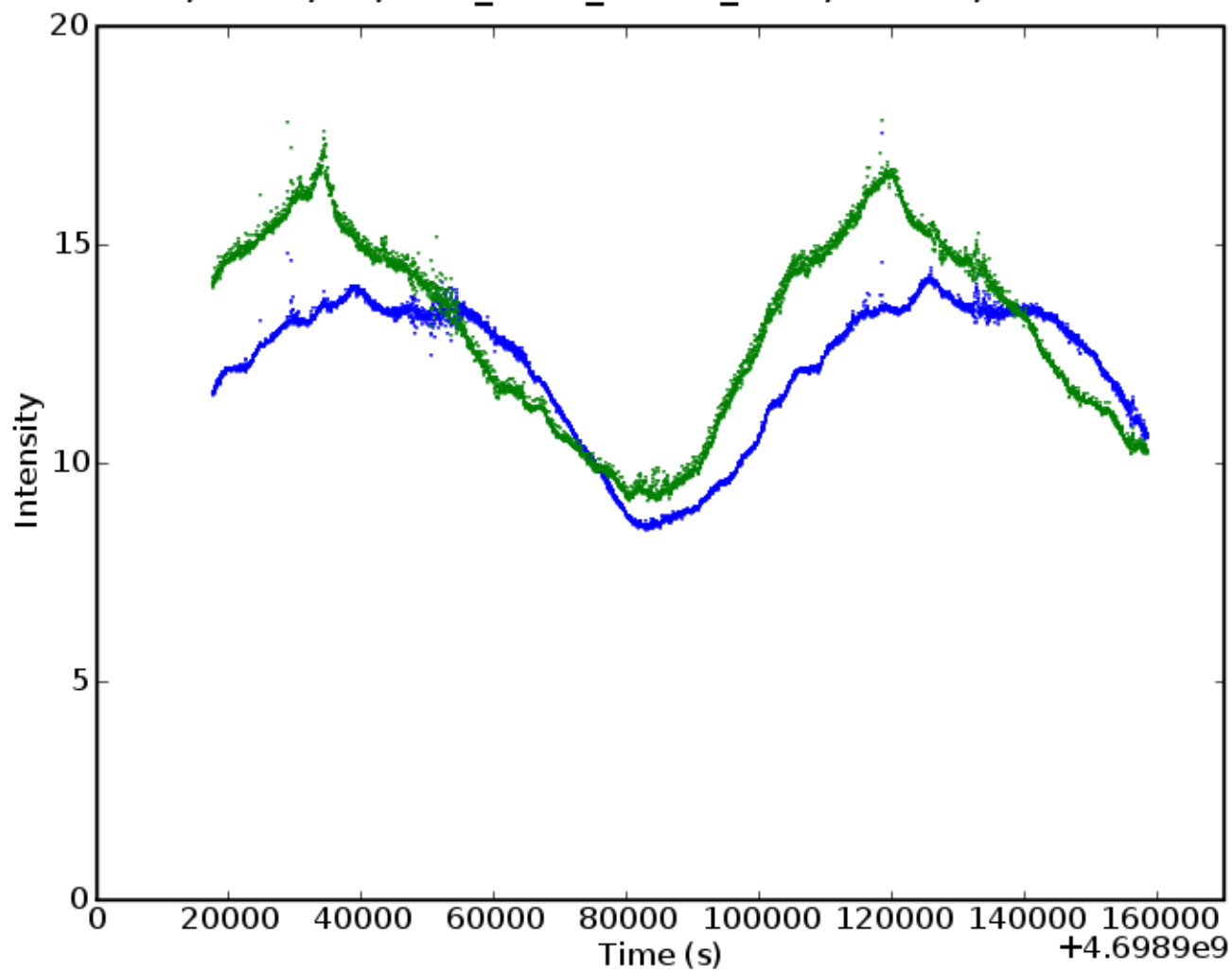
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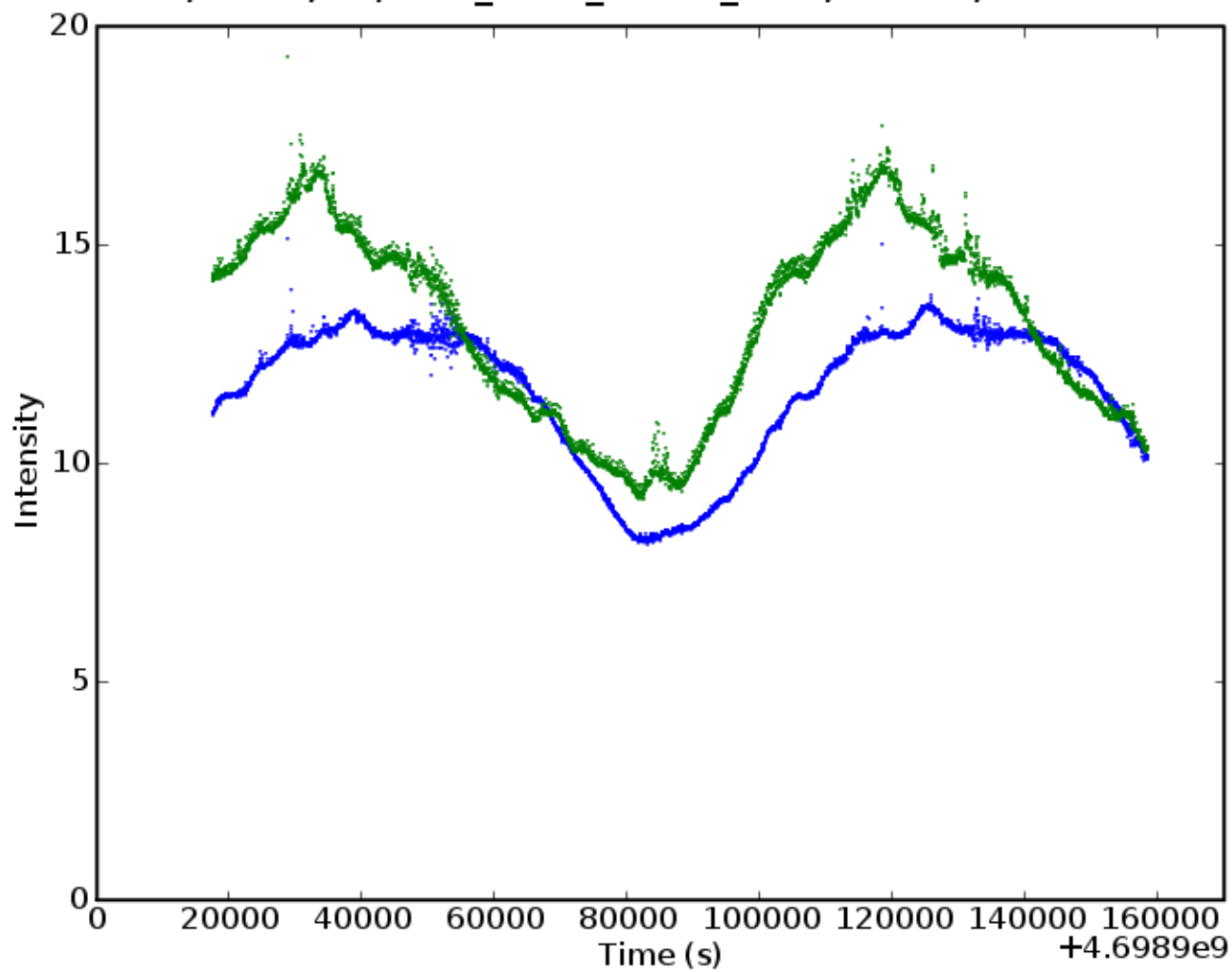
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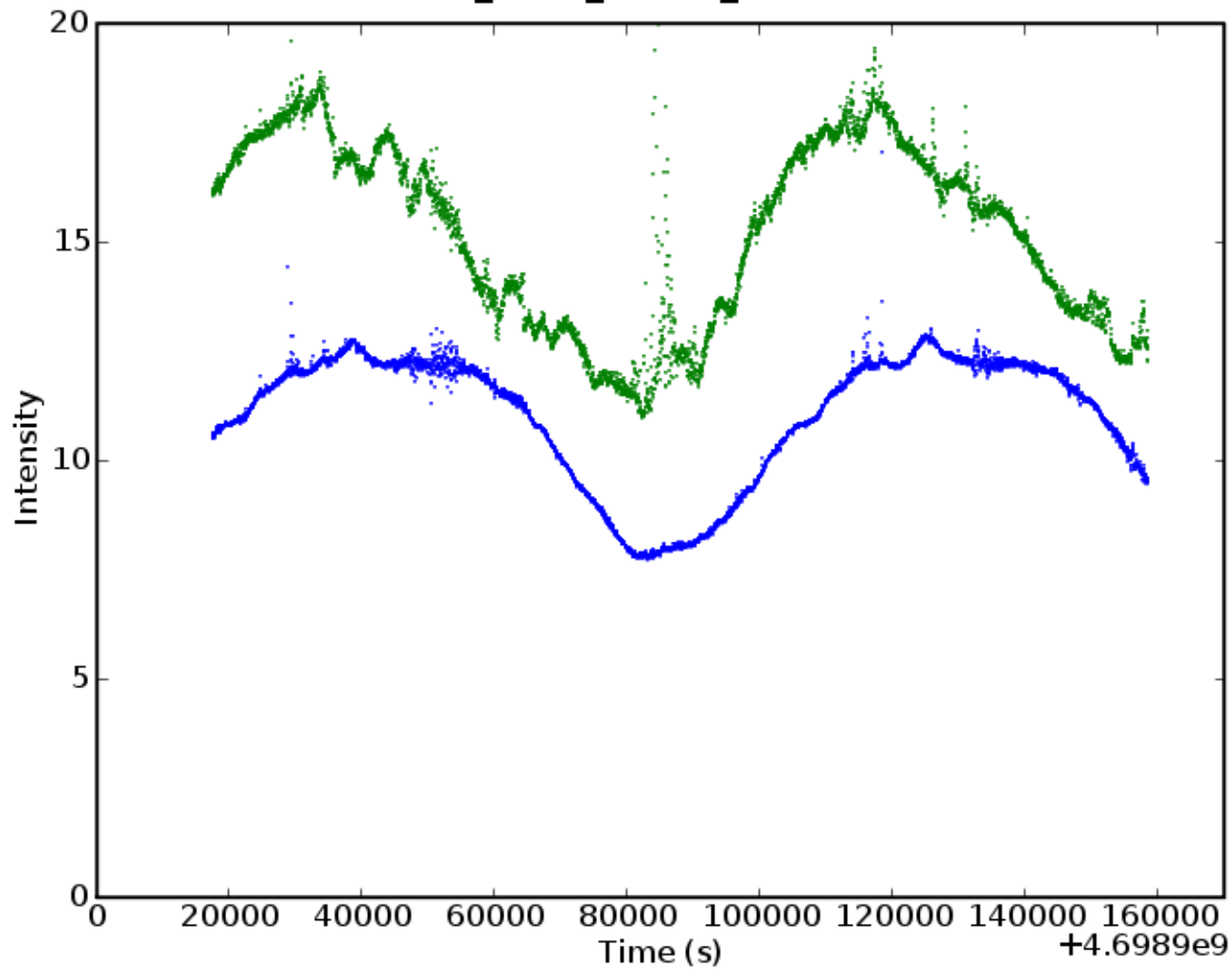
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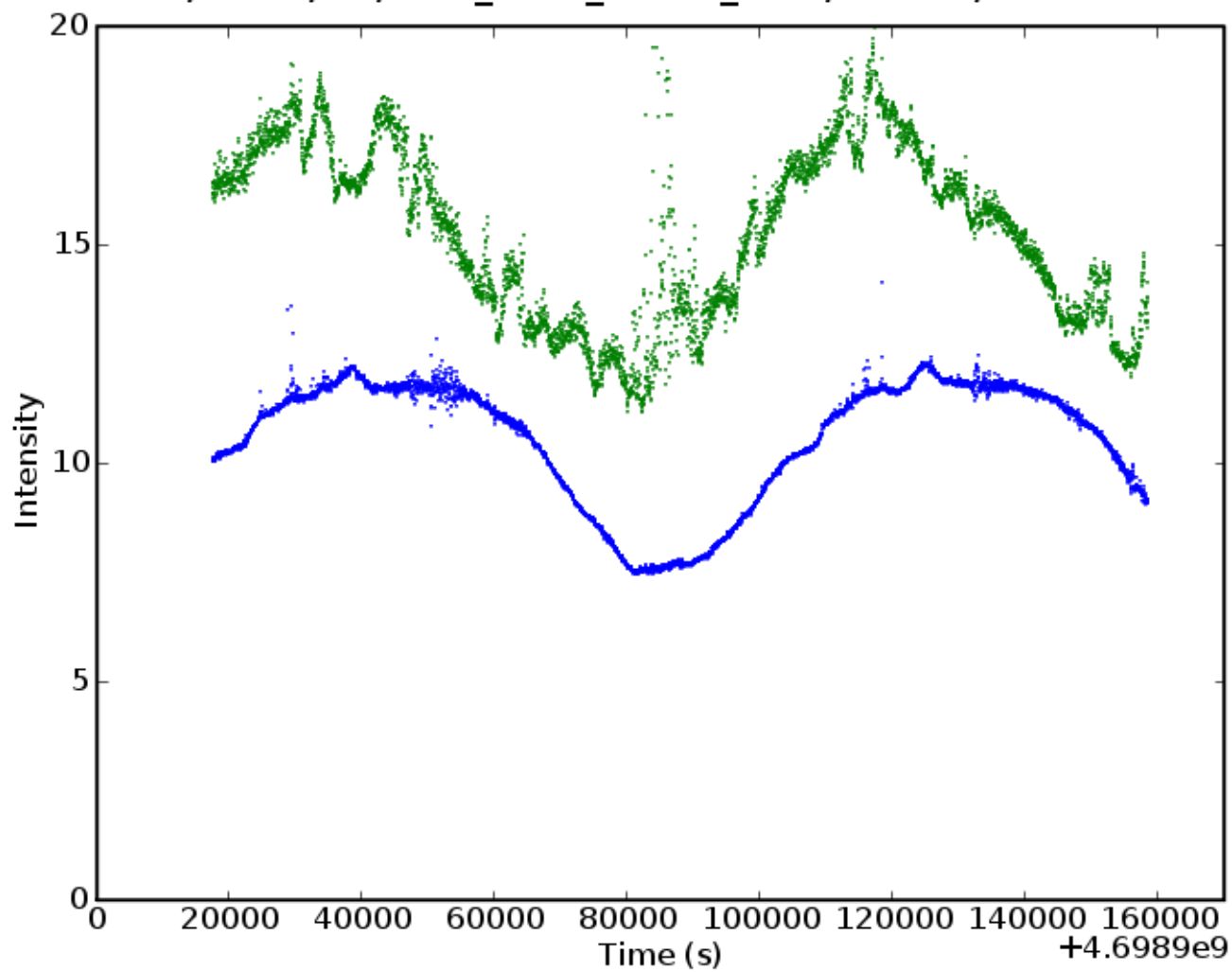
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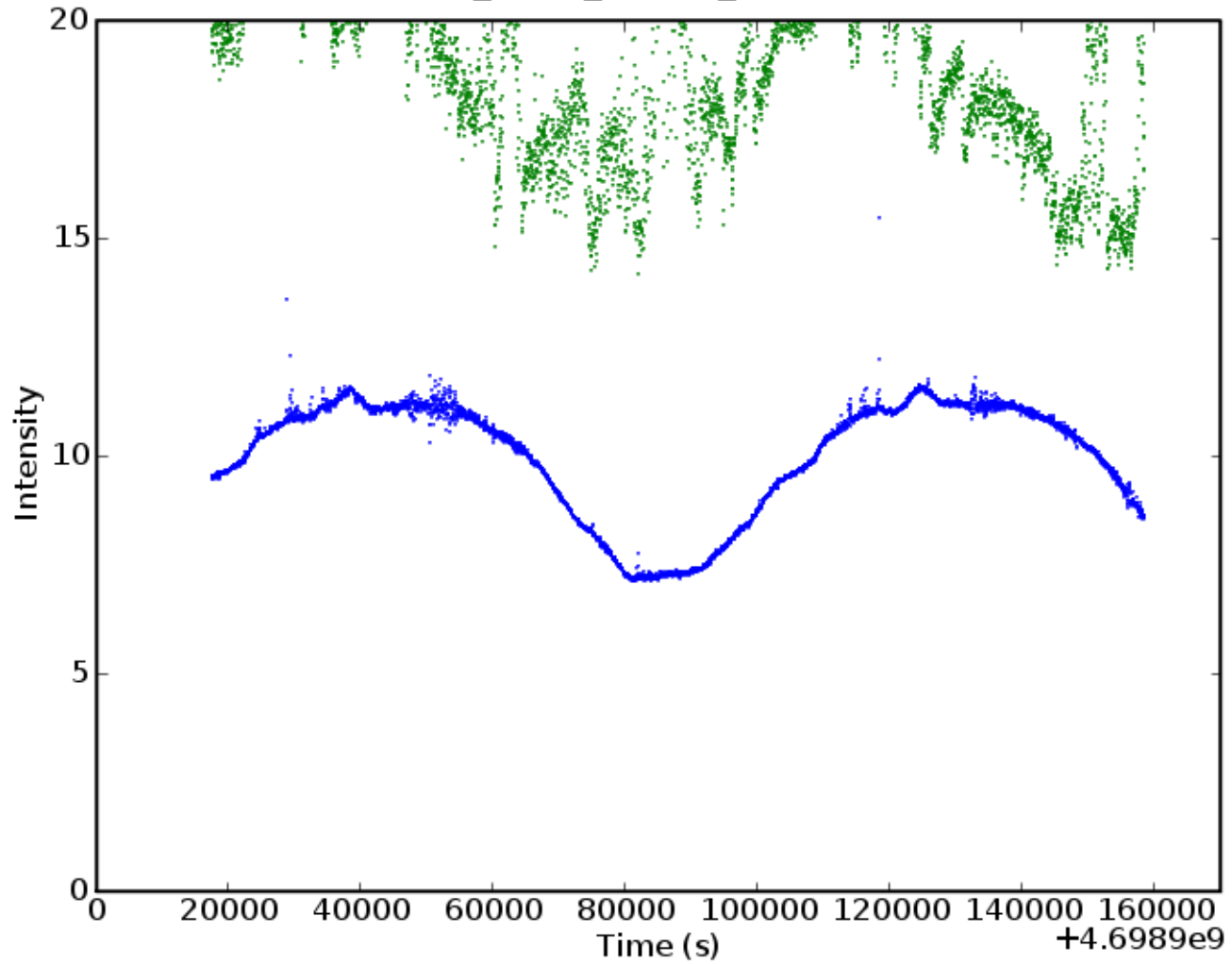
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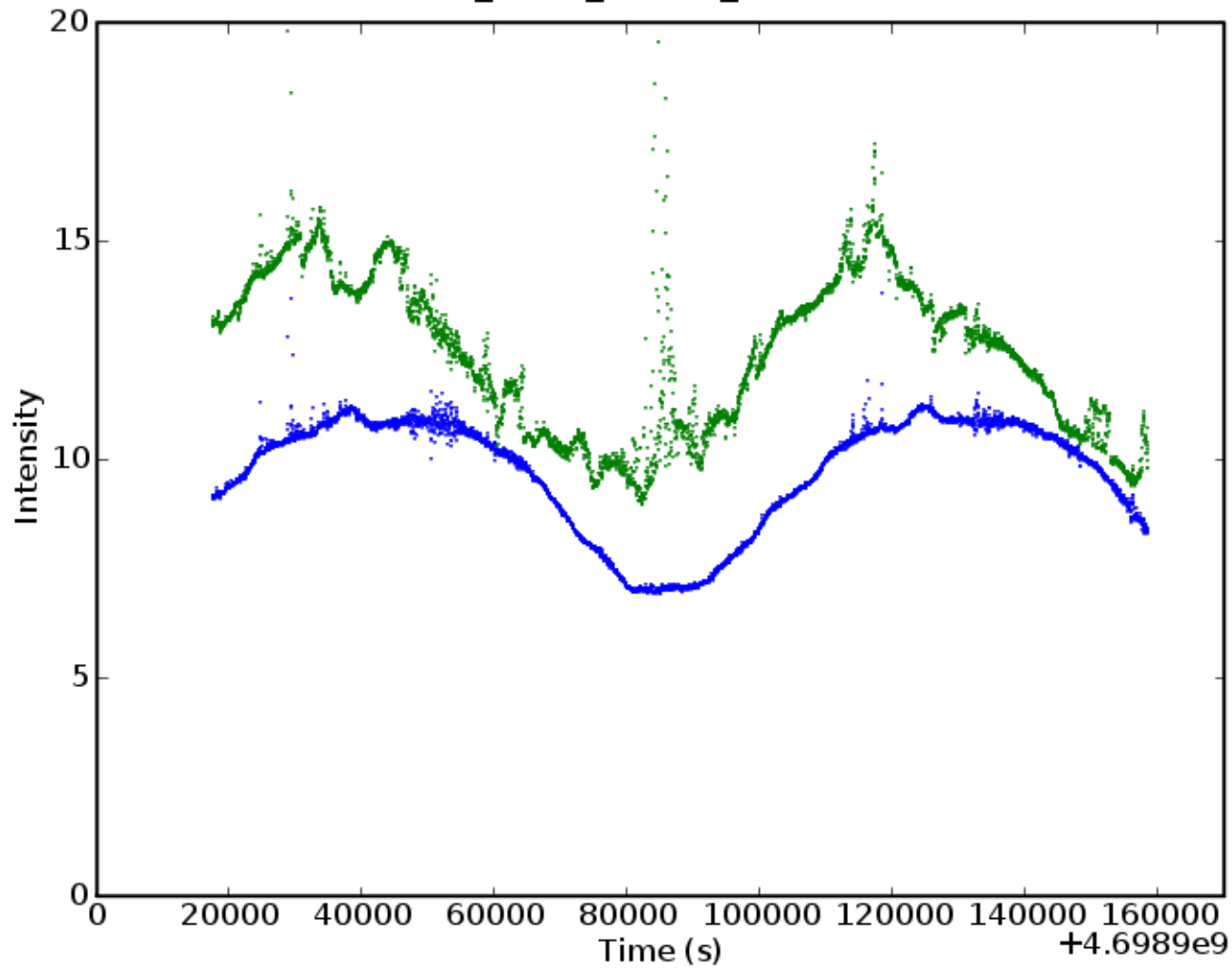
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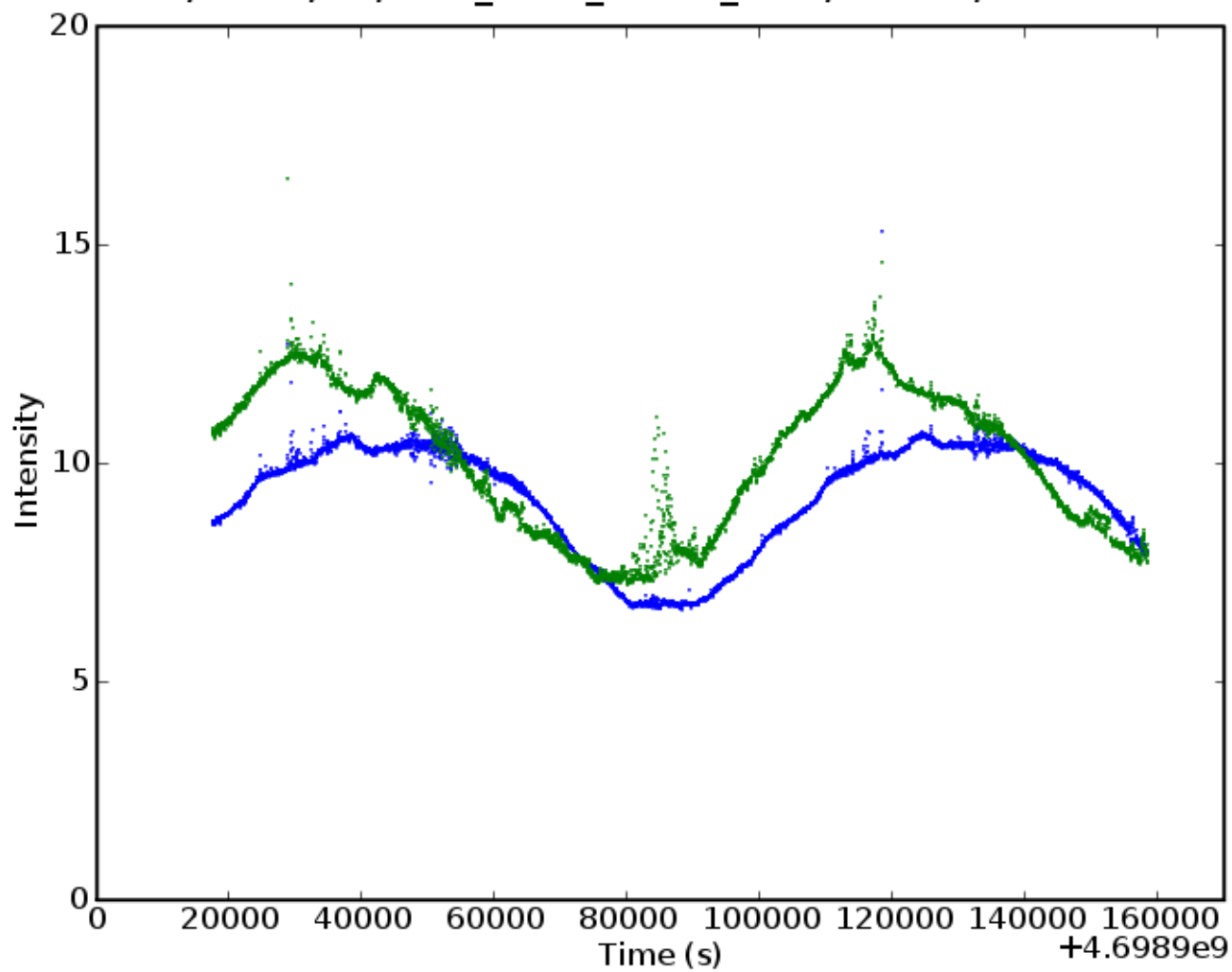
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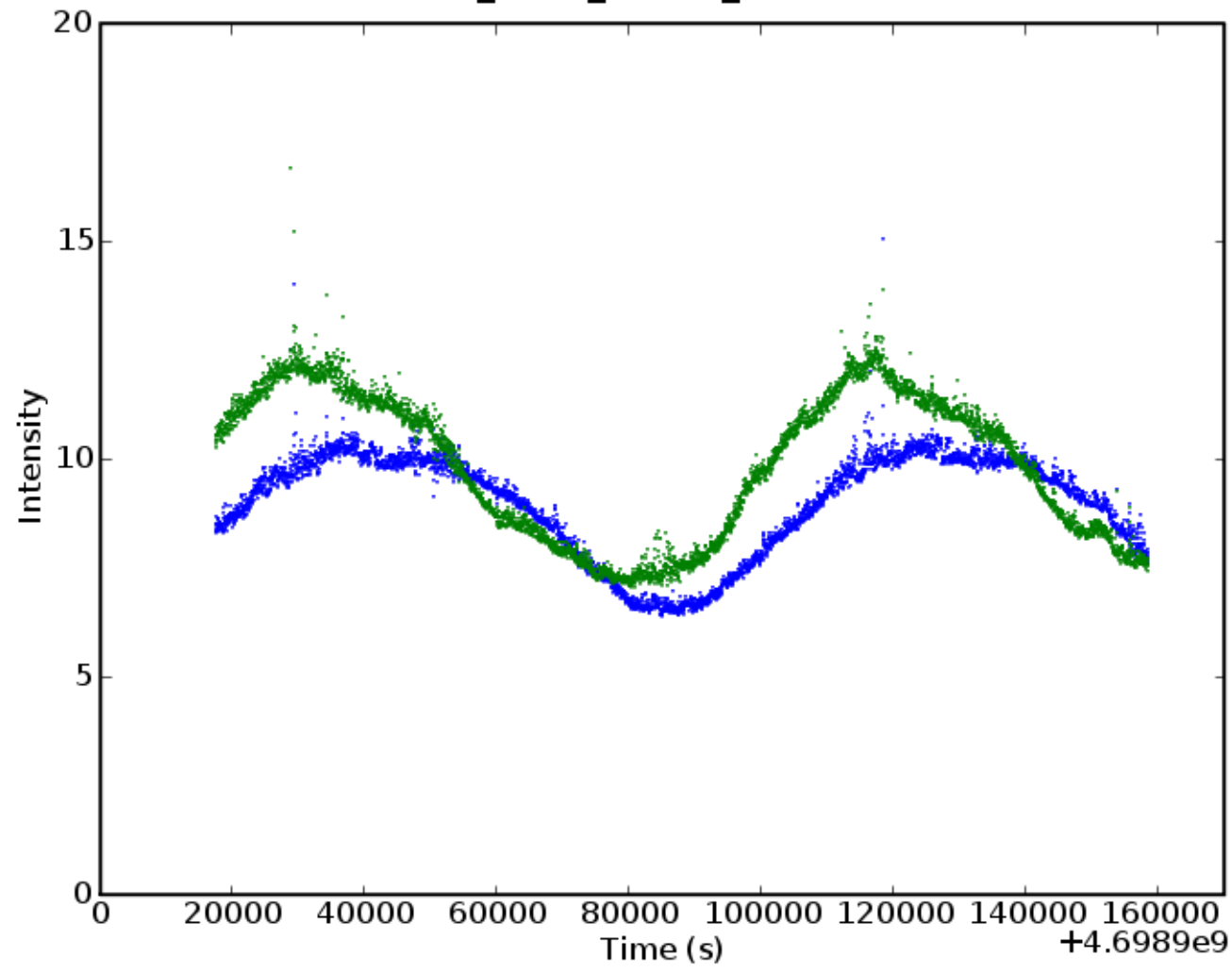
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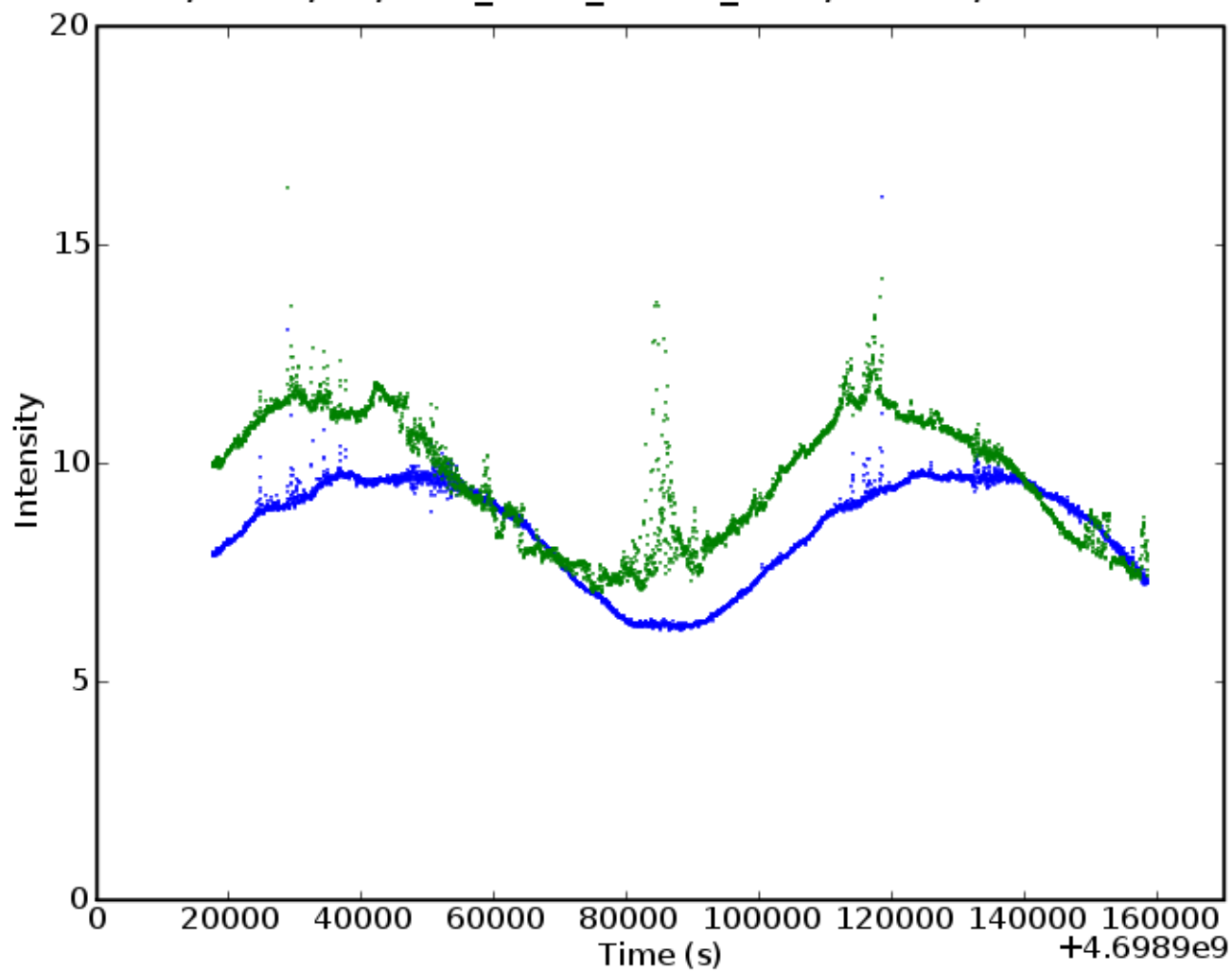
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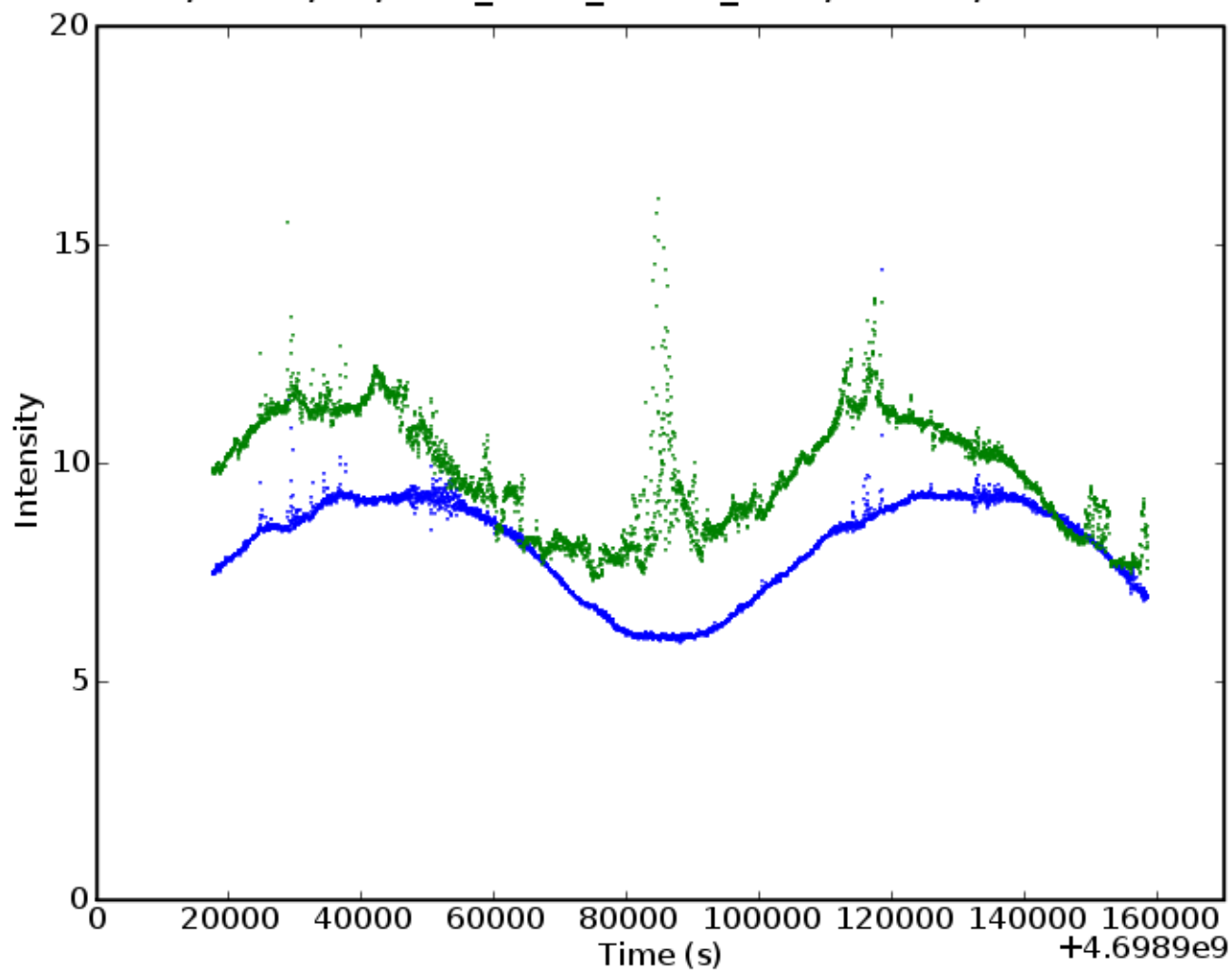
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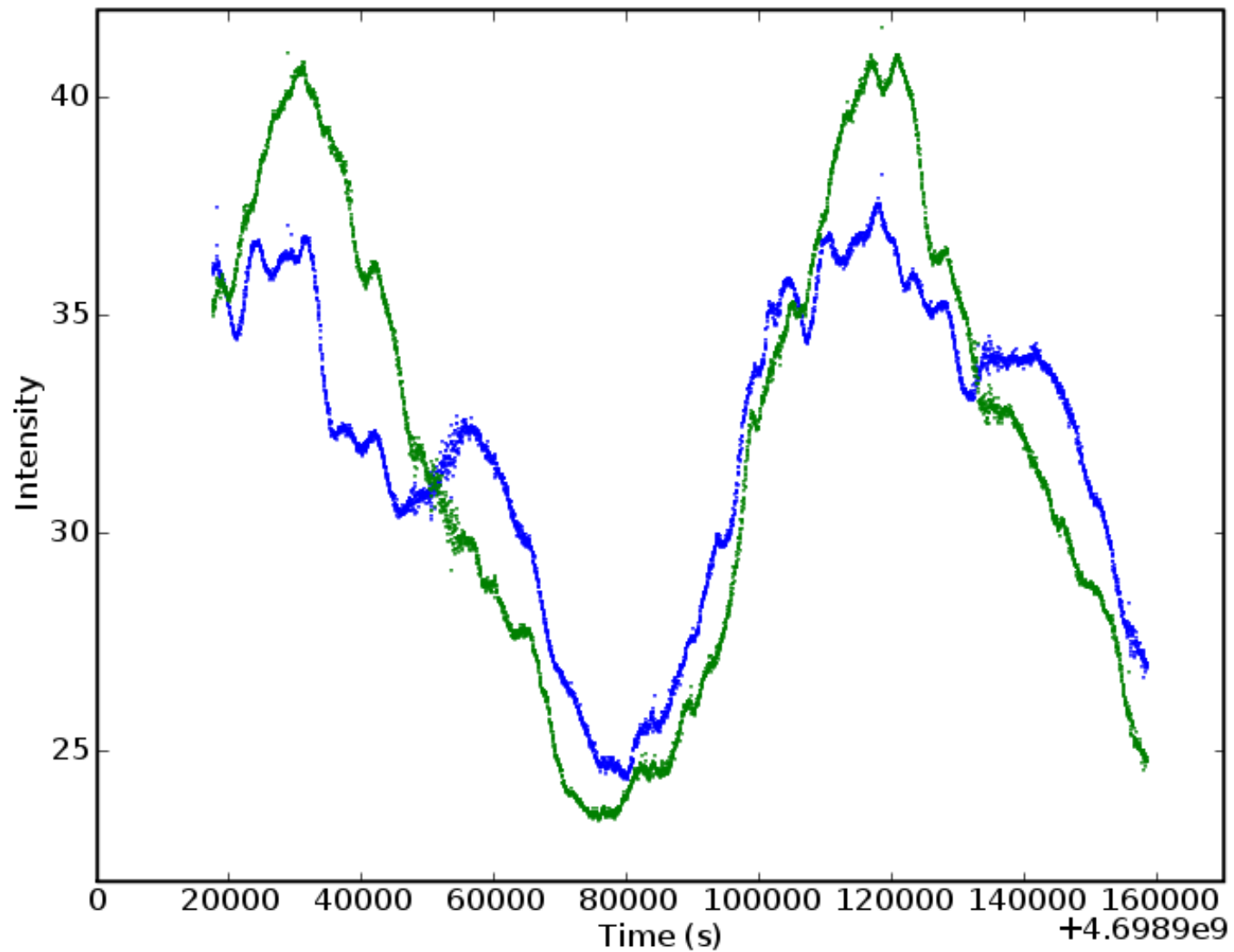
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Time vs. Amplitude, Baseline 0-0, Sub-band(0) Channel(1)
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Some conclusions

Can we **understand** temporal and frequency-dependent **fluctuations** in the total power of CS010 microstation 0 (=beamformed sum 48 dipoles) ?

Note that the station diameter ~ 65 meter, which at 5m wavelength (60 MHz) has a HPBW of $\sim 1.3 \times 5 / 65$ radian $\sim 5.7^\circ$ (**and 'full' beam about 10°**). Depends on spatial taper across station.

Observed fluctuations in Total Power on timescales of about **0.5 - 1h at 60 MHz**

1h corresponds to about $15^\circ \cos(58^\circ) \sim 8^\circ$ /hour movement on sky at the declination of CasA. But beamformer tracks CasA so we should **not see** any fluctuations from CasA and its environment (except the slow change due to the dipole beam response).

The main variations that we should observe are those caused by the convolution of the apparent sky (mostly diffuse Galactic structures and CygA) that moves through the changing station beampattern. These should take therefore place on timescales of about 30 - 60 min as indeed observed. This can and should be **simulated** to be compared in detail with the observed fluctuations. Use e.g. the Japanes 45 MHz all-sky image?

Can we nevertheless already say something about 'beamformer' stability ? Yes. The TP signals should repeat from day to day ! And they do not to a level of a few% which varies from subband to subband. We still do not understand where these frequency dependent variations come from.

To be continued...