

# 115-185 MHz mJy-imaging with LOFAR

A first deep analysis of the NorthPole field



= Sarod Yatawatta & Ger de Bruyn

LOFAR Status meeting, 17-March-2010



Observations 20-22 Dec 2009

3C61.1 (Dec=+86°)

L2009\_16167

3s integration

HBA 115 - 185 MHz

8 Core Stations (x 2) (CS002,003,004,005,006,007,030,032)

4 Remote Stations out to 30 km (RS106,208,307,503)

MeqTrees calibration + Casa imaging on LOFAR subcluster

~ 50 subbands processed in 12 MHz wide groups

Expected thermal noise about 0.2-0.3 mJy/beam

Observed with lots of snow on tiles !

# Some ‘issues’ with this dataset

No station calibration turned on as yet

Different station beams (both 24-tile and 48-tile) hence 3 different effective interferometer beams

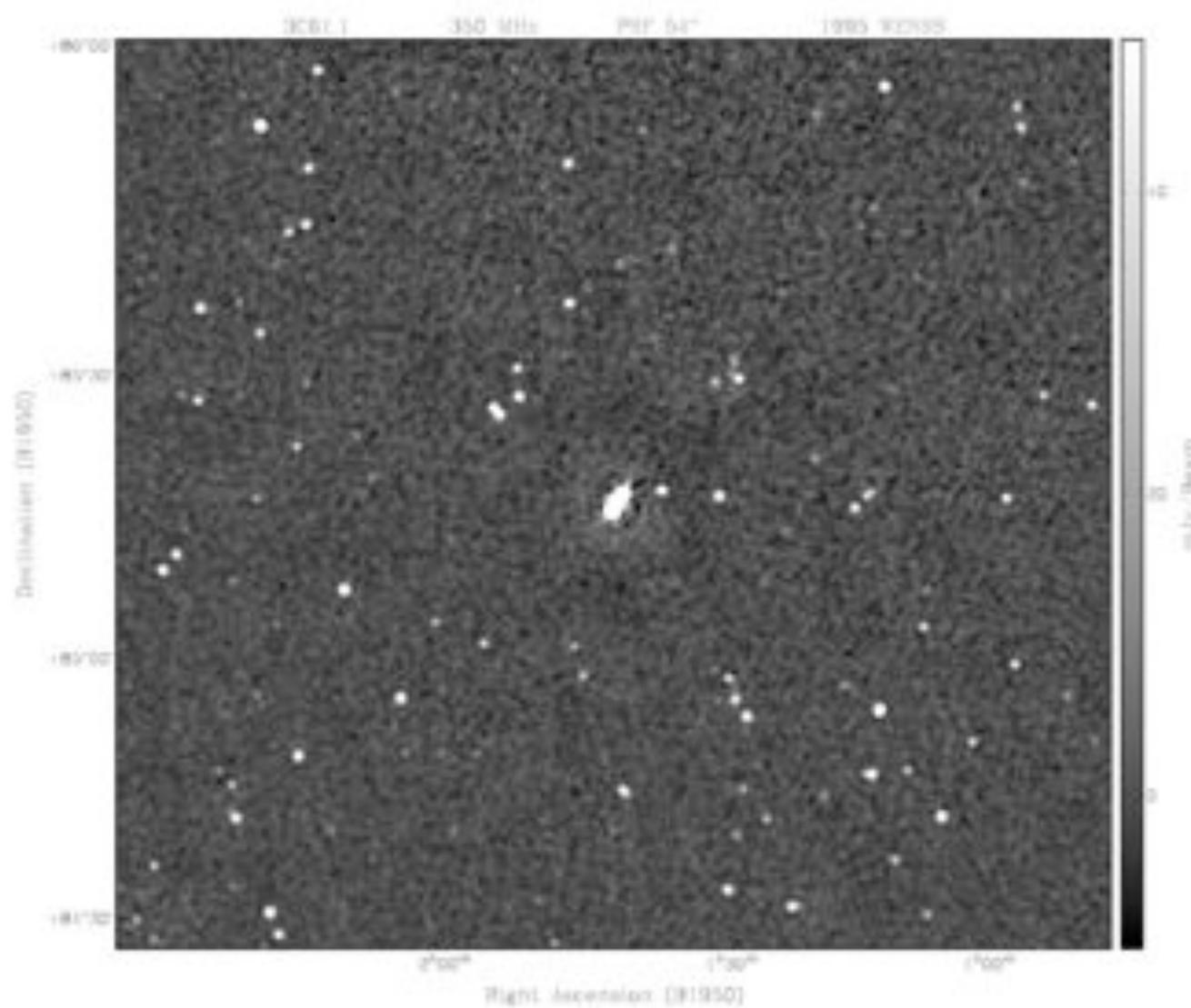
Extremely wide frequency range (115 - 185 MHz)

But:

good uv-coverage (24h)

only small variation in elevation ( $49^\circ$  -  $57^\circ$ )

# Area near 3C61.1 taken from WENSS 350 MHz



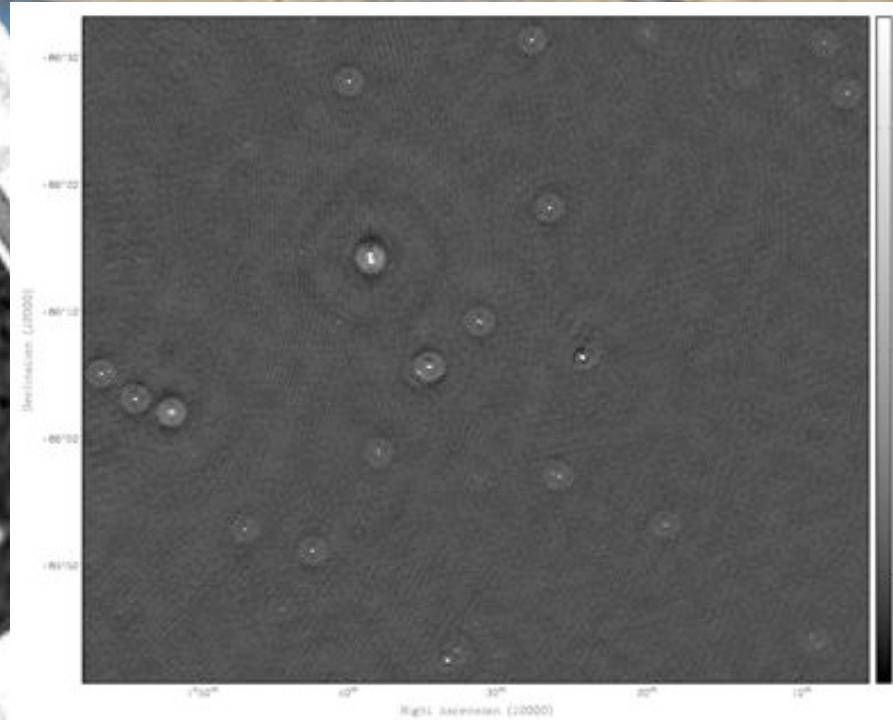
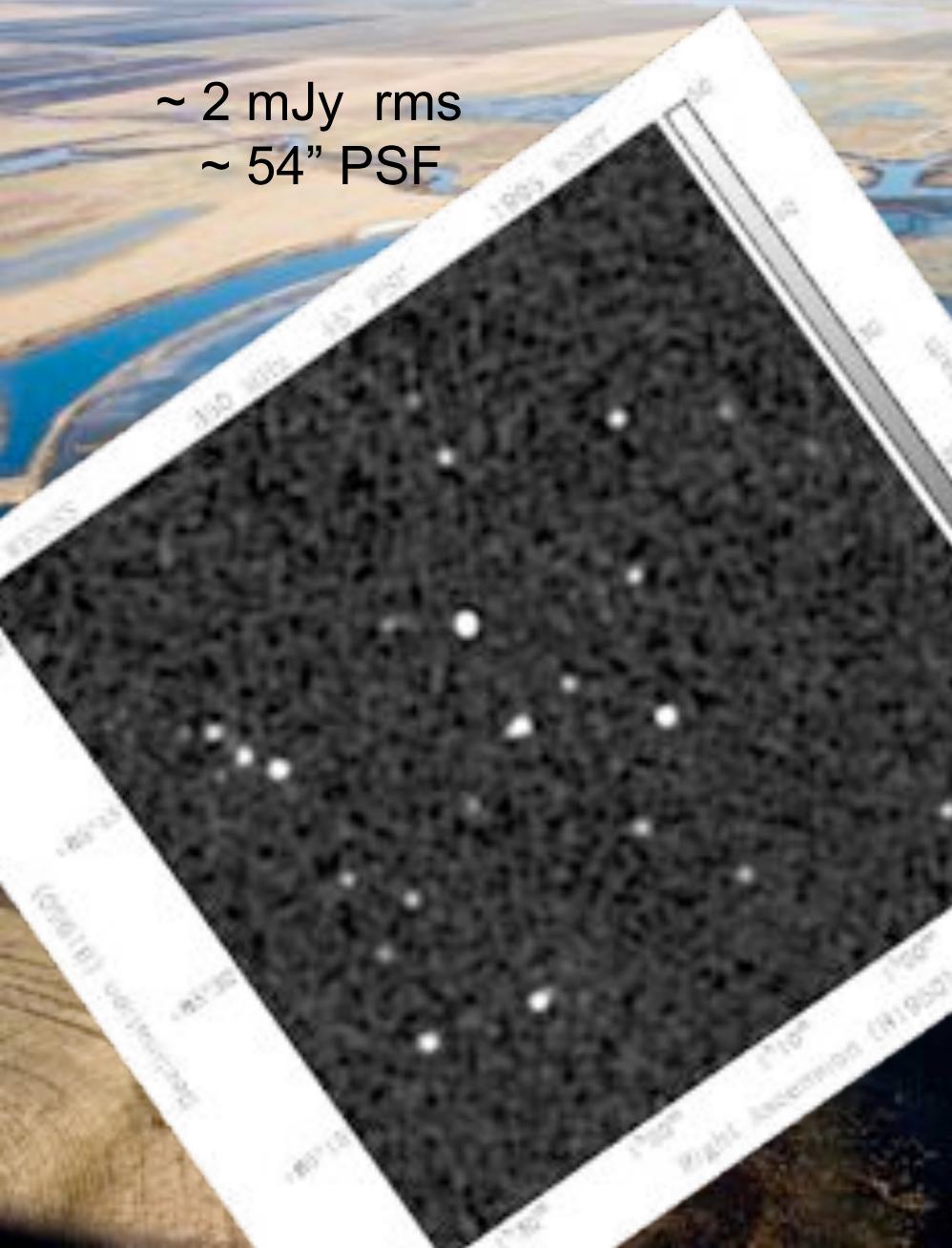
WENSS - 350 MHz

vs

LOFAR - 150MHz

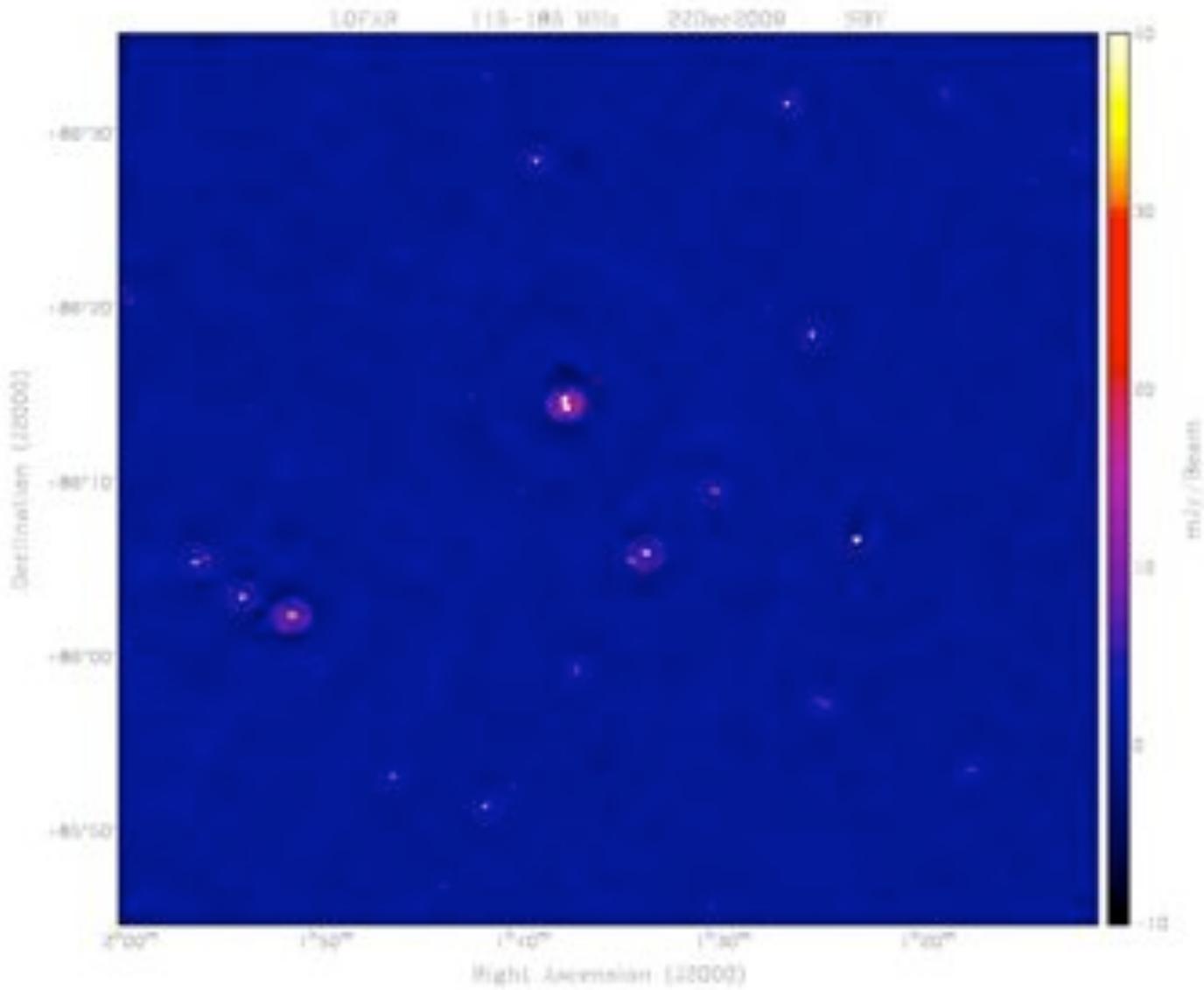
~ 2 mJy rms  
~ 54" PSF

~ 1 mJy rms  
~ 10" PSF



'cup-and-saucer' PSF

# Area of about $1 \times 1^\circ$ East of 3C61.1

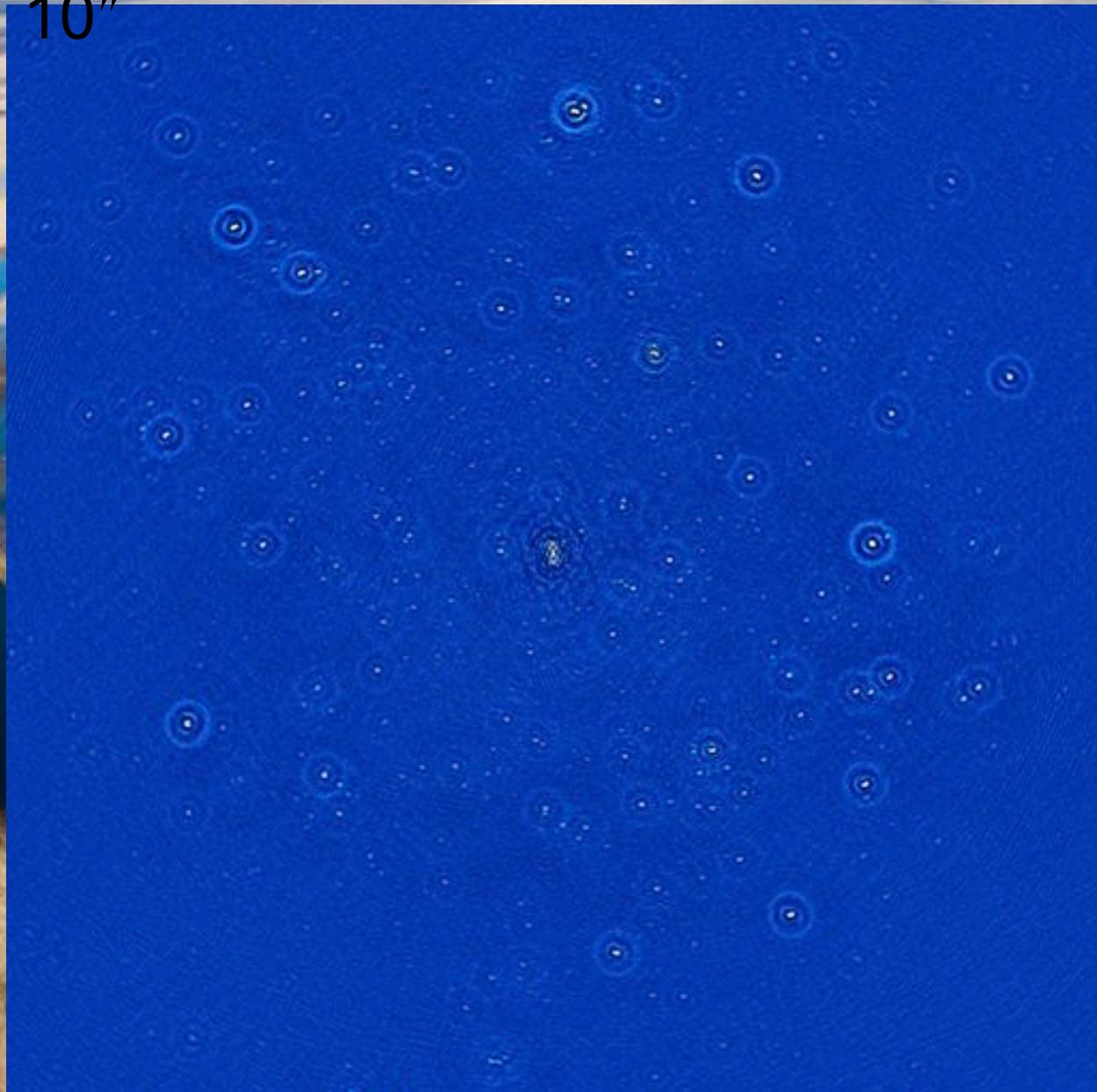


$8^{\circ} \times 8^{\circ}$

7200x7200 pixels ( $4''$ )

PSF ~

$10''$



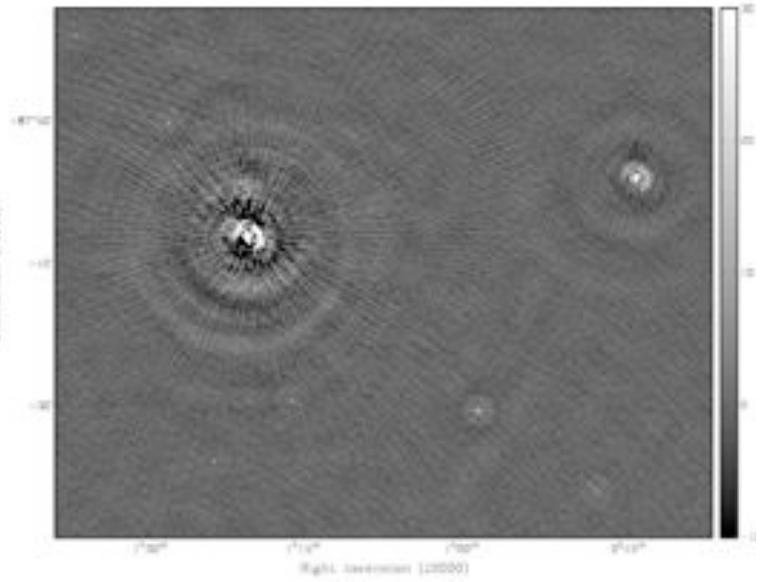
40 source  
selfcal, all  
subtracted  
+restored

shapelets+  
component  
model 3C61.1

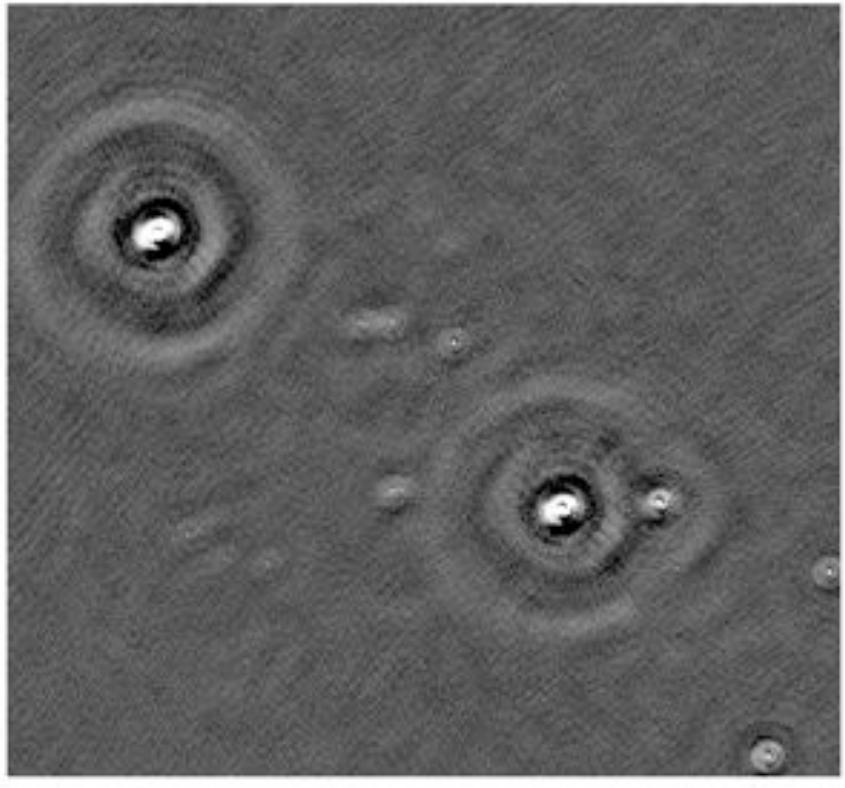
10 Jy peak

1 mJy noise

# Mild non-isoplanaticity for distant off-axis sources

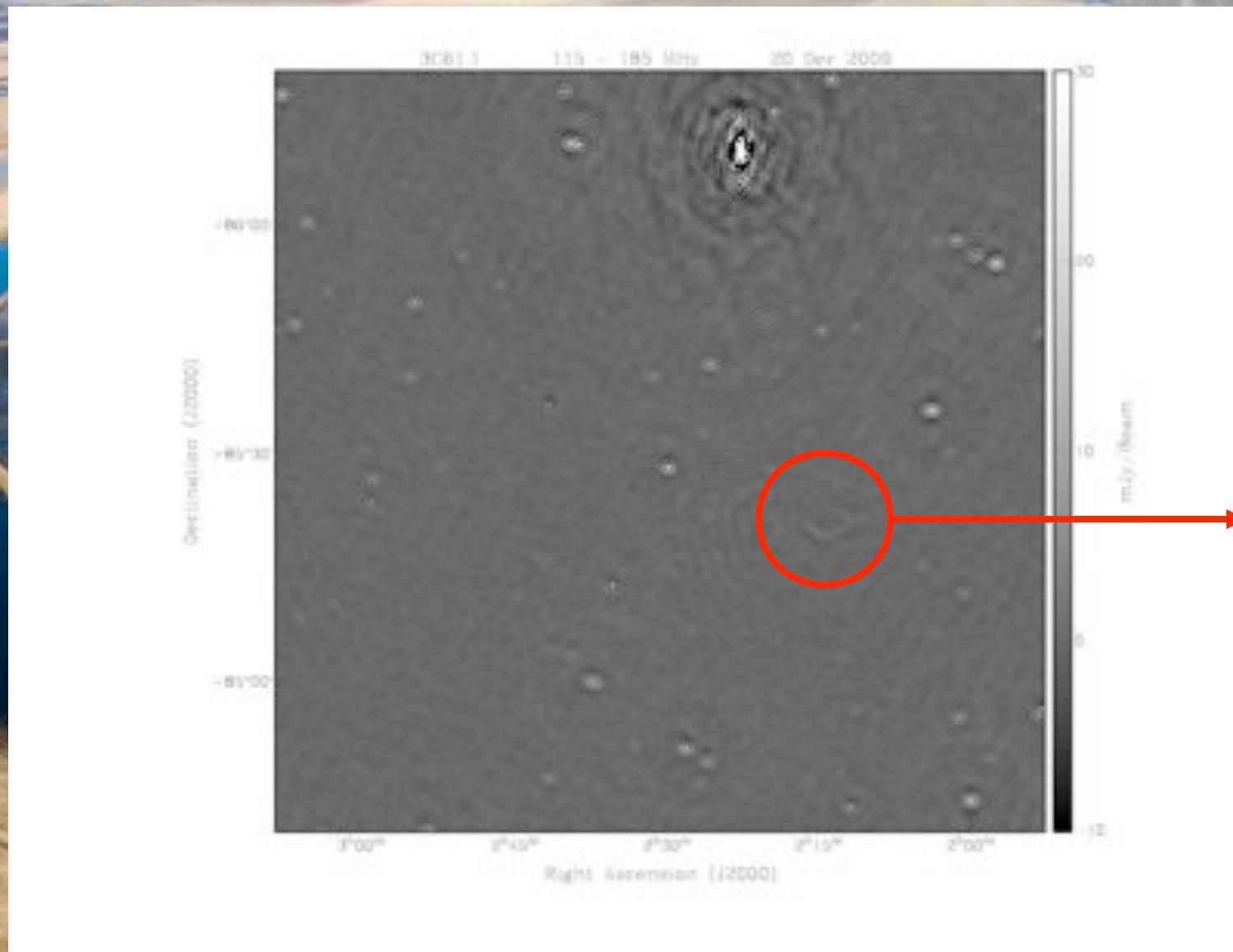


$\sim 2$  Jy peak ( $r \sim 1.5^\circ$ )

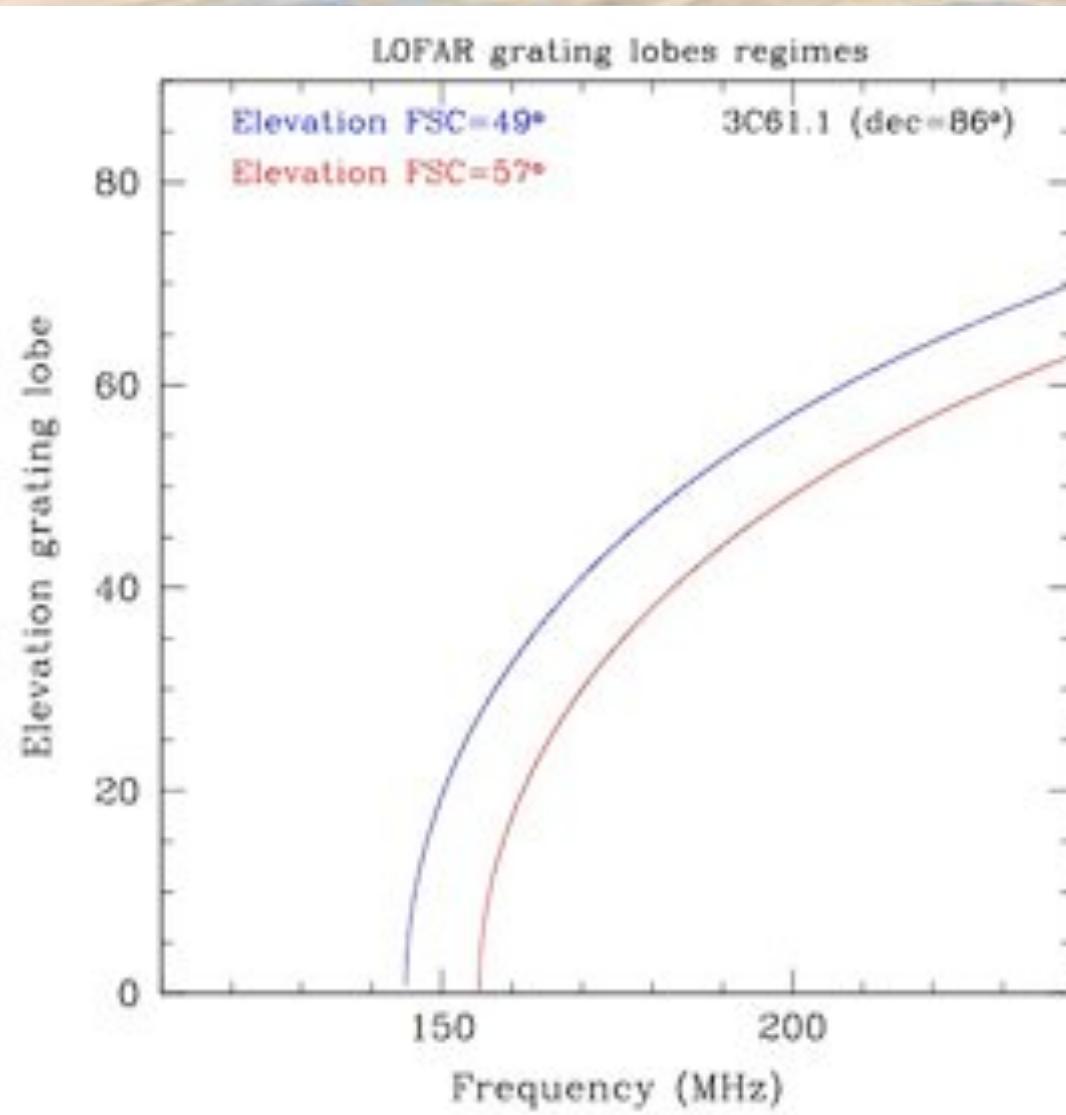


$\sim 0.2$  Jy peak ( $r \sim 3^\circ$ )

# Central area south of 3C61.1



# LOFAR HBA-tile grating lobes



# Some preliminary conclusions:

Some results:

- Imaging with good PSF over  $\sim 5^\circ$  diameter after single direction-independent selfcal
- Only 8x2 + 4 stations but wide bandwidth and 24h gave good uv-coverage
- Differential ionospheric motion ('seeing') (over 24h)  $< 10''$
- Noise level  $\sim 1$  mJy  $\Rightarrow$  DR  $\sim 10,000:1$  (within factor  $\sim 4$  from thermal)
- Very weak artefacts near NCP

Some remaining issues:

- Ghost structures
- Off-axis image 'distortions' (at  $r > 3^\circ$ )

Further improvements (mostly in new datasets) to be expected due to:

- Focused stations (+ 30-40% sensitivity , better/narrower station beams)
- Proper dealing with time/frequency variable beams
- Peeling / SPAM on off-axis sources
- Including spectral properties of sources in LSM
- Much better uv-covergae on long baselines
- Full polarization...