Report from Imaging Busy Week 9 ASTRON, 17-21 January 2011



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Imaging Busy Week 9

- 29 participants: 23 at ASTRON, 6 in remote through an EVO session
- 2 groups: 'beginners' and 'experienced' commissioners
- Beginners were initiated to Lofar data reduction
- The experienced commissioners worked on important imaging tasks:
 - production of good models for A-team sources;
 - subtraction of the A-team from the data;
 - production of scientific quality images

AVAILABLE DATA

Source	Observation ID	Band	Δν (MHz)	Dura tion	Multi beam	Status
3C196	L2010_21604_pizzo	LBA	30-90	10 h	NO	AOF+NDP3
	L2011_22793		15-30	10 h		AOF+NDP3
A-team (Cyg A, Cas A, Vir A, Tau A, Her A)	L2010_22120	LBA	30-90	1.5 h	YES	AOF + NDP3
	L2010_22121					
	L2010_22122					
	L2010_22123					AOF
	L2010_22124					
A2256	L2011_22663	LBA	10-58	6 h	NO	AOF + NDP3
Calibrators (3C147, 3C196, 3C286, 3C295, 3C298, 3C380)	L2011_23105 L2011_23124-40	HBA 0	110-190	1 m	NO	

3C196 field (30-90 MHz)



3C196 field (10-30 MHz)



• Calibration on 3C196 (using UVrange > 2.0 km)

• No Cas A/Cyg A subtraction

• Imaging with UVrange > 2.0 km

Cas A (30-90 MHz)



Cas A (30-90 MHz)



• A better removal of Cyg A from the visibilities is currently worked on

Cyg A (30-90 MHz)







• Cas A and Vir A are subtracted



Cyg A (30-90 MHz)



• Dynamic range ~ 7000

Virgo A (30-90 MHz)



Virgo A (30-90 MHz)



Virgo A (30-90 MHz)



Tau A & Her A (30-90 MHz)



Tau A , concatenation of 4 time-separated of the a multi-component model on the

1 SB (55 MHz), concatenation of 4 time-separated observation. Calibration with a multi-component model on the individual observations. No A-team sources removal.

Her A

 SB (53 MHz), concatenation of 4 timeseparated observation.
Calibration with a two-component model. No A-team sources removal. Right: uniform weighting. Left: natural weighting. The images have the same angular scale.





Abell 2256 (10-58 MHz)



• 5 SBs; data seems to be of good quality (i.e. little RFI)

• Data are strongly affected by CasA and CygA, located at distances of 30-40 deg

• Removal of Cas A & Cyg A, and correction for the central field. The subtraction seem to have worked fine

Calibrators (110-190 MHz)



- 3C196 observed 3 times at interval of 30 minutes
- BBS, no directional gains and using the Yatawatta HBA beam model.
- The solutions of 2 observations are plotted against the amplitudes found for third one
- The amplitude stability is quite good
- Next step: compare the amplitude solutions of 3C196 with other calibrators

NDPPP with 'rficonsole' option





Visibility Amplitudes (xx) for Baseline CS032LBA-CS004LBA

Courtesy of David Rafferty



LBA Global Bandpass (10-90 MHz)

Courtesy of D. Rafferty & L. Birzan



The Roadmap Document

Roadmap for progress with LOFAR imaging in the next 3-4 months

v.1 Neal Jackson 11.01.22 (after imaging BW9 and a meeting with John Conway, George Heald, Neal Jackson, Ronald Nijboer, Roberto Pizzo, Antonis Polatidis, Bas vd Tol, Reinout van Weeren).

Top priority: removal of the A-team from the data

a) Directional gains in BBS (time-intensive procedure: 1 hour observation, SB with 16 channels, 1 hour on 1 lce to solve in 2 directions);

b) As method a), but on compressed data -> access to D-factors in BBS (Joris vZ.);

c) Observing the A-Team sources in multiple beams and then cross-subtract (needs to know D-factors and beam models) (Cyril, George, G. vD.);

d) The observed visibility is written as sum of source visibilities in different directions, each having the appropriate beam/smearing corruptions. The corrected visibilities are recovered by matrix inversion (Bas vd Tol).

ANNOUNCEMENTS

• Models location (/globaldata/gsm)

incoming

· validated

When submitting your models, please inform G. Heald and R. Pizzo:

Quality control
Disk space control

• Parset library

Commissioning section on the Wiki



Lofar Imaging BW 9

Stay tuned...



3C196 field (30-90 MHz)



LBA Global Bandpass (10-90 MHz)



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A2256	L2011_22663	LBA	10-58	6 h	NO	AOF + NDP3
Calibrators (3C147, 3C196, 3C286, 3C295, 3C298, 3C380)	L2010_21732	LBA	30-90	1 m	YES	AOF
	L2010_21738					
	L2010_21739					
	L2010_21740					
	L2010_21741					
	L2010_21742					

3C196 field (30-90 MHz)

