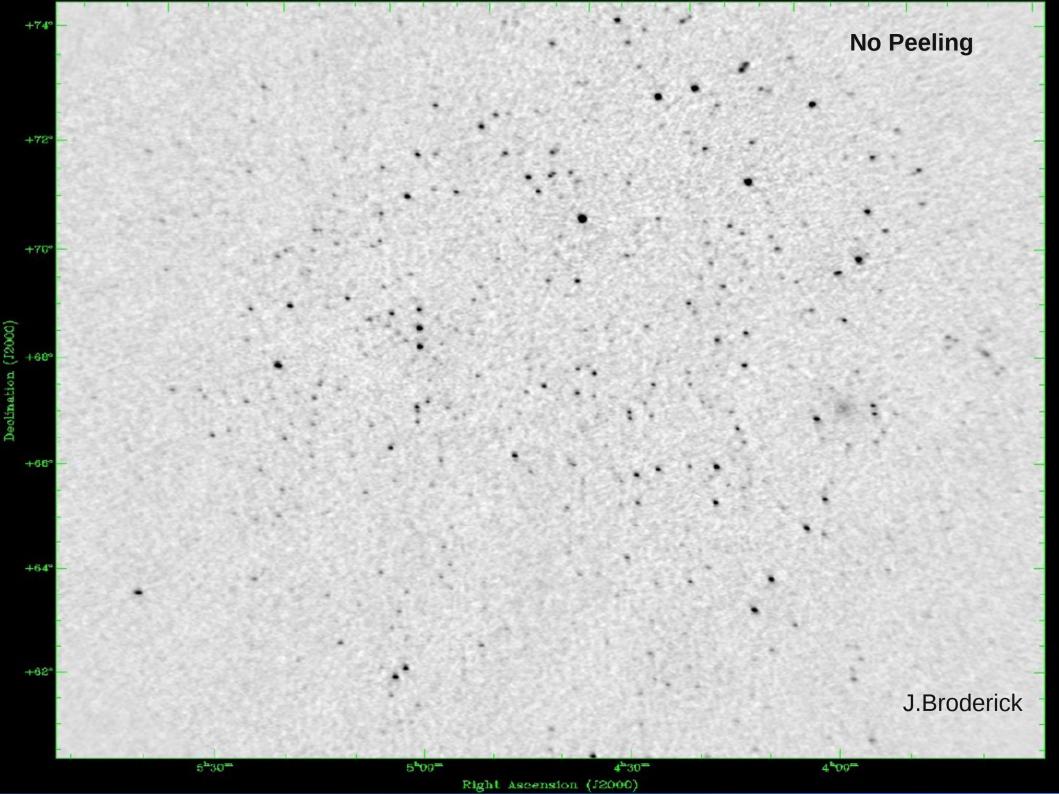
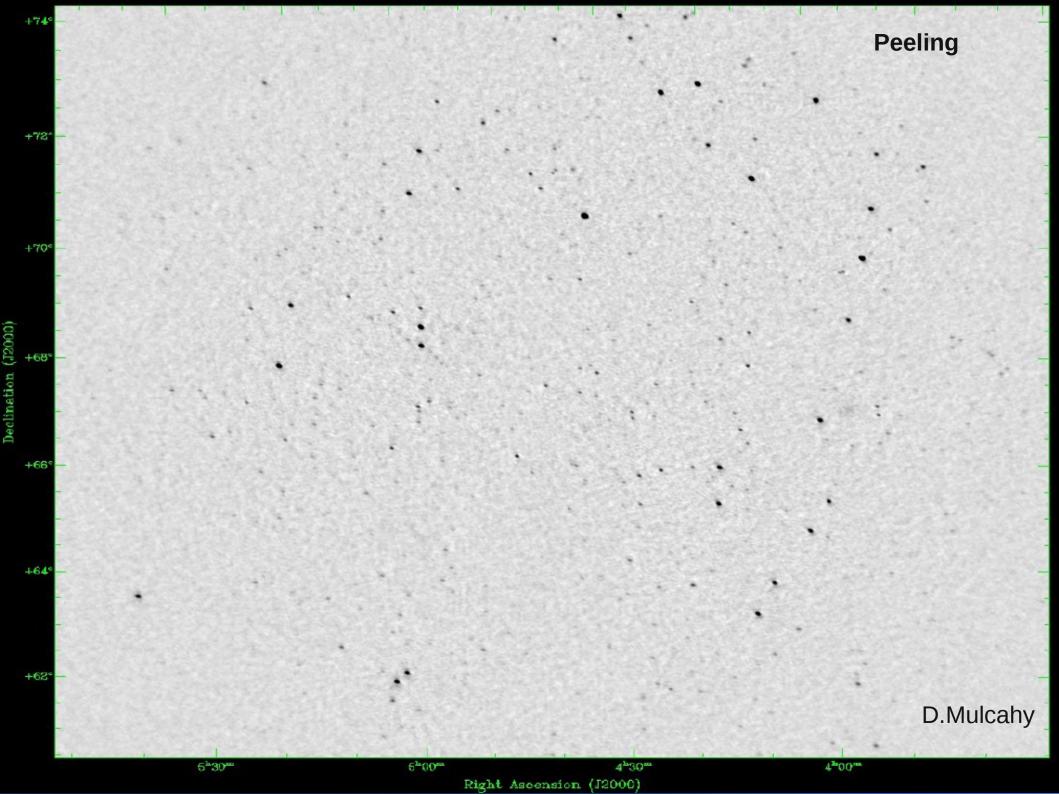
Recent Commissioning Progress on M51- Peeling methods

David Mulcahy (MPIfR) MKSP

Some Background

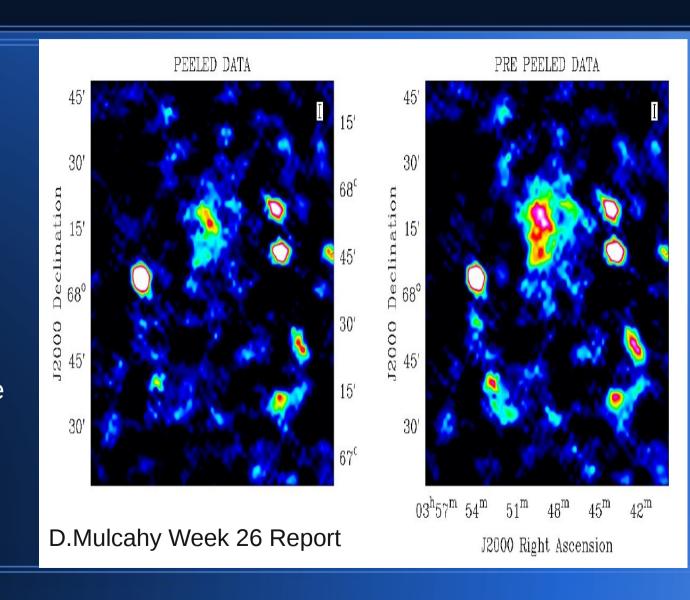
- Since July, MSSS have been using a very effective process to peel sources developed by Mike Bell.
- 20%-30% decrease in noise was seen and produced excellent images.





Some Background

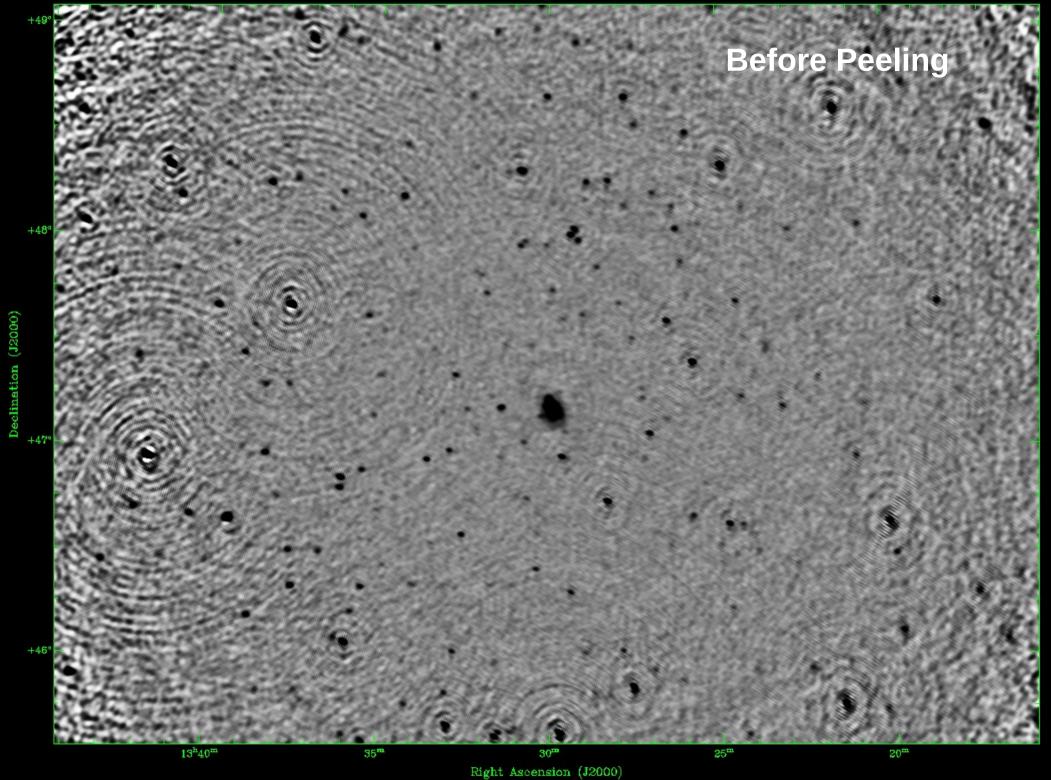
- However, it was seen that the flux in extended emission was decreased substantially. To the left shows IC0342 before and after peeling. See report by D.Mulcahy (Week 26)
- It was also seen by Blazej (week 37-38) that decreasing the time cellsize from 5 to 1 increased the flux by 6% (-40% to -34%).



Different Approaches to Peeling

- Case 1
 - Normal Global SkyModel (from gsm.py, with Time cellsize = 5 & 1)
- Case 2
 - Clean Component model taken from awimager
- Case 3
 - Clean Component model taken from awimager and casapy



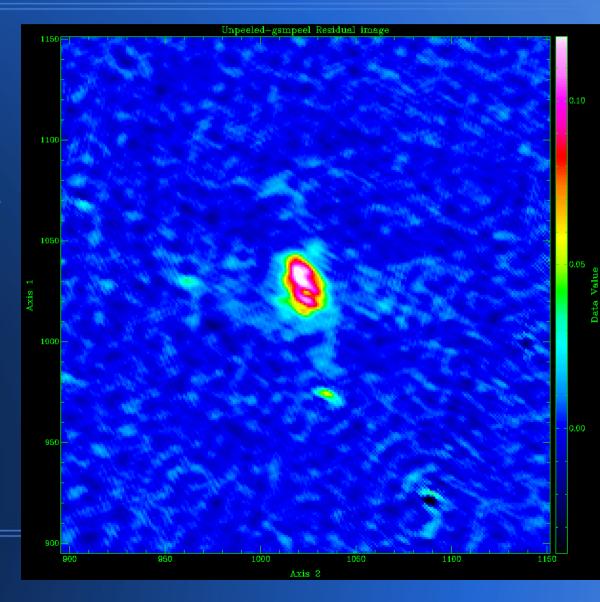


Case 1

- Use normal gsm model, peel 4 sources
- M51 described by a Gaussian only
- Decrease in Flux agreed with Blazej's previous work of 40% (Time cellsize=5) and 34% (Time cellsize=1)

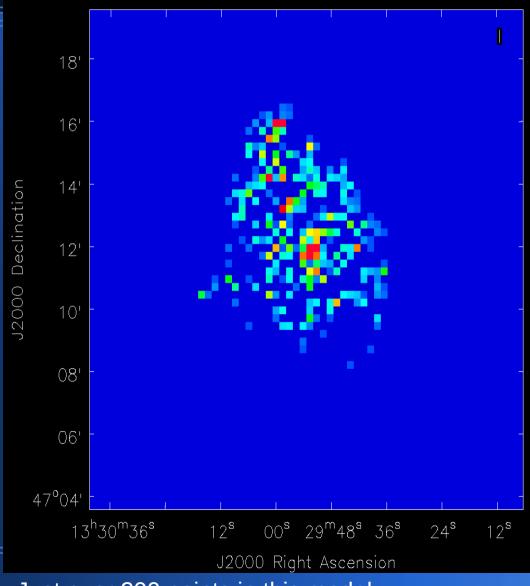
Case 1 : Global Sky Model Peeling

 Residual image to the left shows the difference in flux occurs over the whole region of the galaxy.

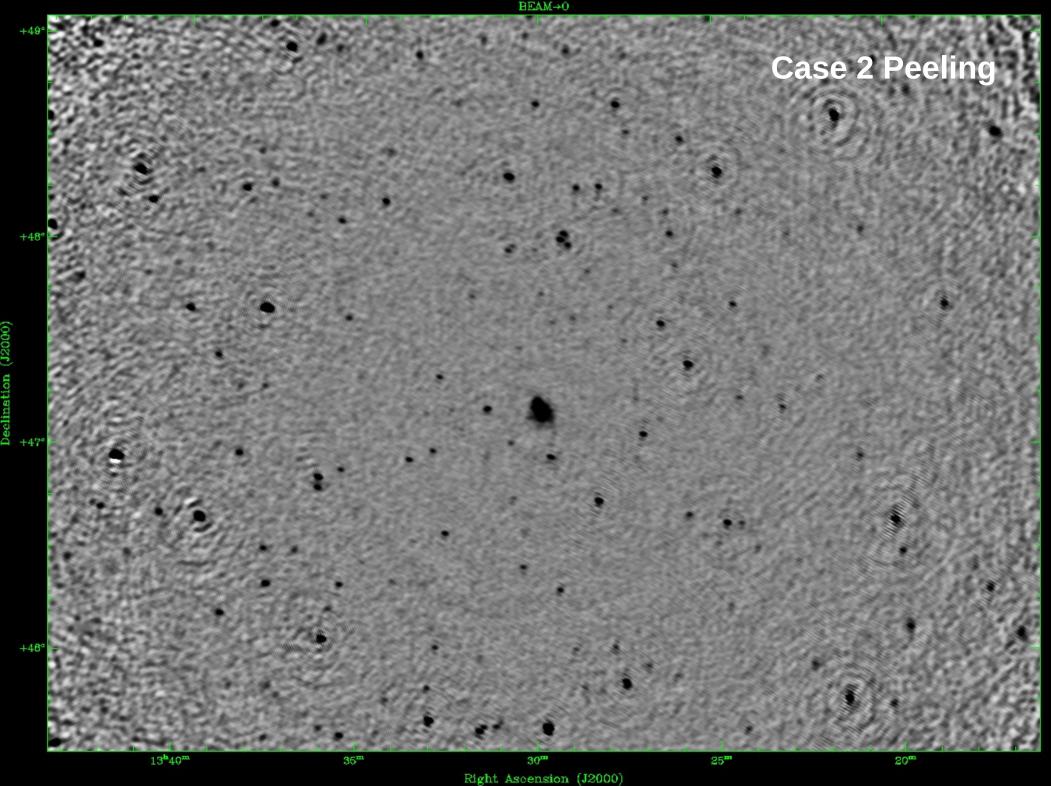


Case 2: CC model extracted from awimager

- Used normal default values when using awimager.
- Resulting cc model of M51 shown on the left.

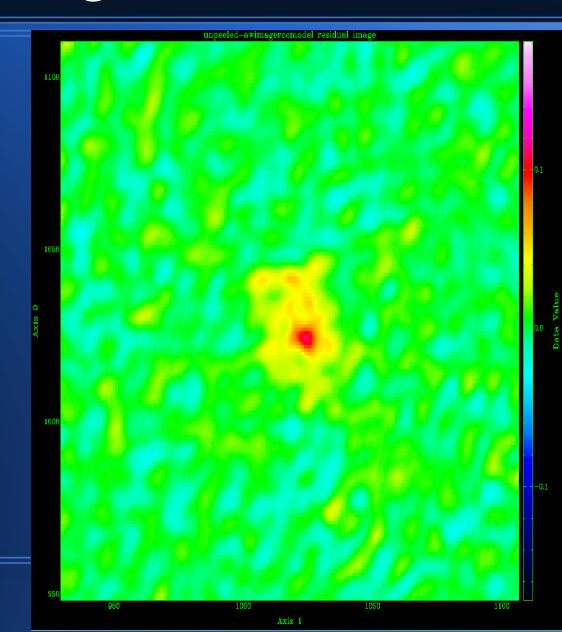


Just over 200 points in this model.



Case 3: CC model extracted from awimager

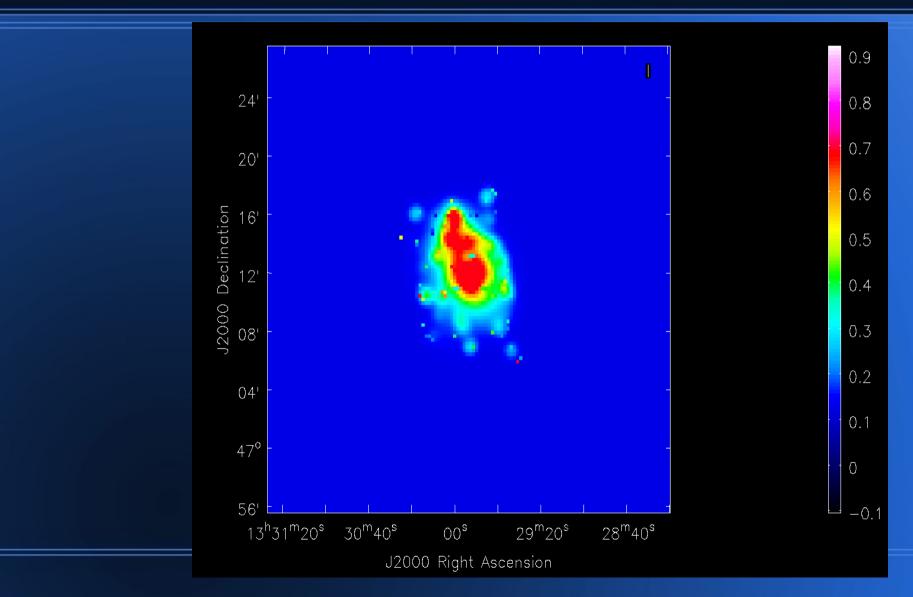
- Overall quality and noise of the image is better.
- Missing Flux is now only 20%. Must need to extract a more complete skymodel from cleaning.

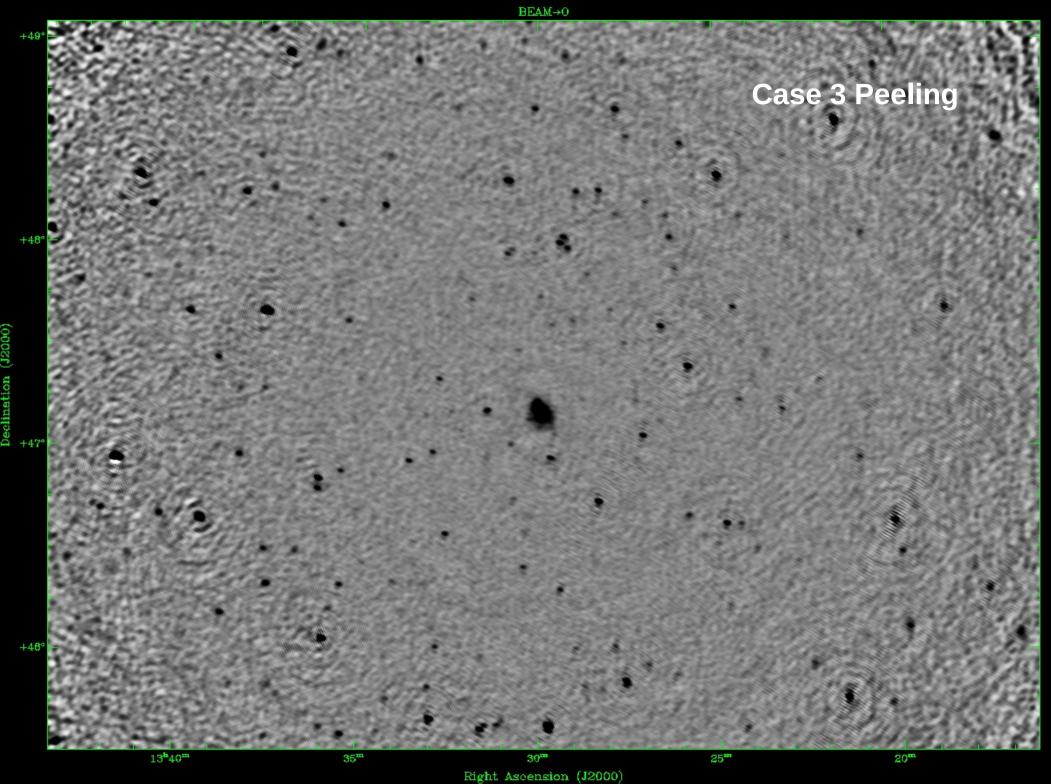


Case 3: CC model extracted from awimager & casapy

- Clean Components of point sources were taken from awimager and then modeled using casapy2bbs.py.
- M51 model was extracted from using casapy.
 Multi-scale cleaning was used.

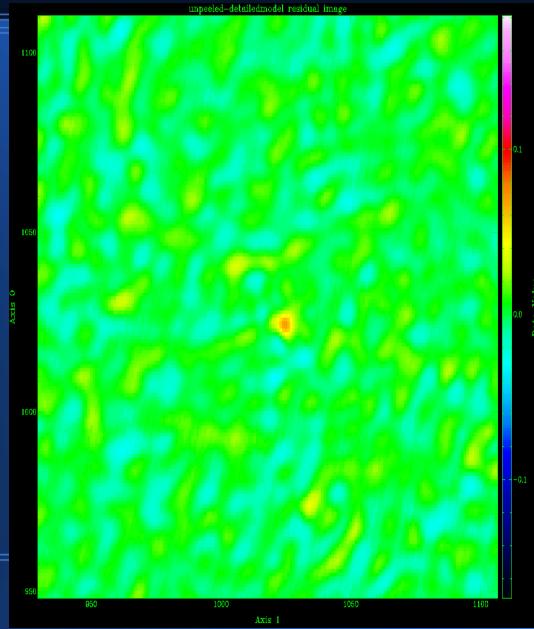
Case 3: CC model extracted from awimager & casapy





Case 3: CC model extracted from awimager & casapy

- Missing flux from extended emission is now only 2% which is from central region only.
- Possible to get this difference to less than 1%.



Conclusions

In order to avoid a reduction in extended emission when peeling, an extremely detailed model of the extended emission is needed.

Using awimager cc model for the point sources and a casapy cc model for the extended emission gives near perfect results (1-2% difference).

Multi-scale cleaning helps very well in extracting this extended source's skymodel.

Not sure how accurate this method will be for extended emission away from the phase center.