V. Heesen (U Southampton), J. Croston (U Southampton), J. Harwood (U Hertfordshire), E. Orru (ASTRON), A. Shulewski (U Amsterdam), F. de Gasperin (U Hamburg), R. Morganti (ASTRON) on behalf of the nearby AGN group and the LOFAR surveys team

Cycle o observations of the nearby FRI radio galaxy 3C31

Southampton

Nearby AGN group

HBA + LBA observations for the entire sample

Targets for LC0_012 Cycle0 LOFAR Using LOFAR for detailed studies of AGN, and AGN physics

		HBA	HBA		LBA	LBA		Pre-processing	Leading
		obs	comp		obs	comp			
-KI 🕻	3C31		10	30		10	15	Southampton?	Nearby 3CR - Croston, Volker Heesen
	3C223		10	30		10	15		Orru`+Croston
	3C452		10	30		10	15		Croston+Jeremy Harwood
-RII									
	B1834		10	30		10	15	Nijmegen?	DDRG - Orru`+
	3C35		10	30		10	15		Orru' (polariz.)+Shulevski
	4C33.33		10	30		10	15		Giant RG Jamrozy+
	3C237		10	30					LongBaselines group, Hardcastle et al.
	3C41		10	30					LongBaselines group, Hardcastle et al.
	M87		8	24		8	12		De Gasperin+
	3C48					10	15	Amsterdam?	RRL group - Oonk+
	Hydra A		6	18		6	9		Cavities - Rafferty, Wise+
	Hercules A		6	18					Cavities: Birzan+
	VLSS J1431.8+1331		8	24		8	30		Relics - Morganti, Shulevski, Kunert-Bajraszewska
	Cygnus A		10	30					McKean+
	6352								
	Total		118	354		92	156		
	Total observing		210				Allo	oc 210	
	Total computing		510					373	



Influence on the environment

AGN feedback
Galaxy lum. fct.
Shock heating
Entrainment

Radio continuum synchrotron emission

X-ray synchrotron emission



Centaurus A

Particle content in FRI/FRII galaxies

- FRI are under-pressured
- FRII are close to pressure equilibrium
- Cosmic rays (CRs)
 - Upper limits from IC radiation
- Magnetic fields
 - Upper limits from Faraday rotation
- Hot gas
 - Upper limits from X-ray emission
- CR electron injection spectral index

3C31 LOFAR observations

- 10 hrs observing time in HBA
- Interlaced 3C48 and 3C196 as calibrators
- HBA data pre-processed by ASTRON
- Initial NDPPP
- Calibrate calibrator and transfer solutions
- Combine sub-bands
- Phase-only calibration on each band
- Image with CASA or awimager

Pipeline in Soton by Adam Stewart



HBA imaging

- Imaged with CASA clean (multi-scale)
- Peak flux density: 5.2 Jy, rms = 0.7 mJy/beam
- Resolution: 17x12 arcsec, S/N = 7400
- First skymodel: VLSS
- Self-calibration in phase, no change!
- Directional dependent gains for 3C34



Directional dependent gains

3c31_band8.image

3c31_band8.image



Radio spectral index (SI)



Expected from shock acceleration: -0.5

HBA in-band spectral index Consistent 144 / 330 MHz spectral index



30 MHz – 1.6 GHz spectrum



- Spectrum of the inner, bright part only
- Spectral index about expected ~ -0.7
- In-band spectral index consistent

Particle models of 3C31 jet



Wide field of view!

Still need to primary beam correct ...



Summary

- First promising results
 - Angular extent 15% larger than at 330 MHz
- Flux scale is broadly consistent
 - Spectral indices agree with expected values
- Noise level 5-10x thermal level
 - Worse near bright sources
- Bright sidelobe structure near our target
 - Directional dependent gains on dominating source does not remove them



VLA contours + X-ray

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Virgo A, M87



HBA, published in A&A

LBA, de Gasperin, in prep.

