Faraday Tomography of the local ISM with LOFAR



Detecting Magnetic Fields

Here's the processes involved:

- cosmic rays + magnetic field = polarized synchrotron emission
- polarization + free electrons + magnetic field = Faraday rotation

Change in polarization
$$\propto \lambda^2 \int_0^d n_e \vec{B} \cdot \vec{dl}$$

Faraday depth

(f)

Faraday Tomography

 Broad-band radio polarization cubes can be transformed into Faraday depth cubes:



Polarization processing

- Data requirements: 2+ channels/SB
- Apply correction for ionospheric Faraday rotation: RMextract (by Maaijke Mevius)
 + BBS/NDPPP Correct step
- Image Stokes Q/U for each channel
- RM synthesis: pyRMsynth

It's not that computationally intensive. Maybe your data is suitable? Ask me how to do it!

The (high-resolution) Faraday sky to date



The IC342 field



Tier 1 HETDEX



Tier 1 HETDEX



Colourized! Hue = Faraday depth, brightness = flux That green-to-purple gradient is very interesting.

Modelling the IC342 field



It's not just pretty, it's scientific. I did some modelling of the line-of-sight using the polarization data.



Take-home points

- LOFAR is amazingly sensitive to diffuse emission
- While polarization leakage is still unsolved, we can still get great polarization data out of LOFAR
- This is a very new way of exploring magnetic fields in our Galaxy, so there's still a lot to figure out.

For more information:

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Astronomy Astrophysics

Faraday tomography of the local interstellar medium with LOFAR: Galactic foregrounds towards IC 342*

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Exploring the Threefold Invisible Universe:



Coming this summer: my PhD thesis