

Tracing multiple AGN outbursts at low frequency in cool-core clusters

Astronomical image showing a galaxy cluster core with multiple AGN outbursts. The image displays a central bright region with several smaller, distinct outbursts, likely representing the AGN activity. The colors range from blue to red, indicating different energy levels or emission lines.

Simona Giacintucci
Naval Research Laboratory

Maxim Markevitch (GSFC), Tracy Clarke (NRL), Matteo Murgia (INAF-OAC), T. Venturi (INAF-IRA)

Radio bubbles in cool-core clusters

Hydra A

MS 0735.6+7421

Perseus (3C84)

Kirkpatrick et al. 2009

McNamara et al. 2005

Fabian et al. 2005

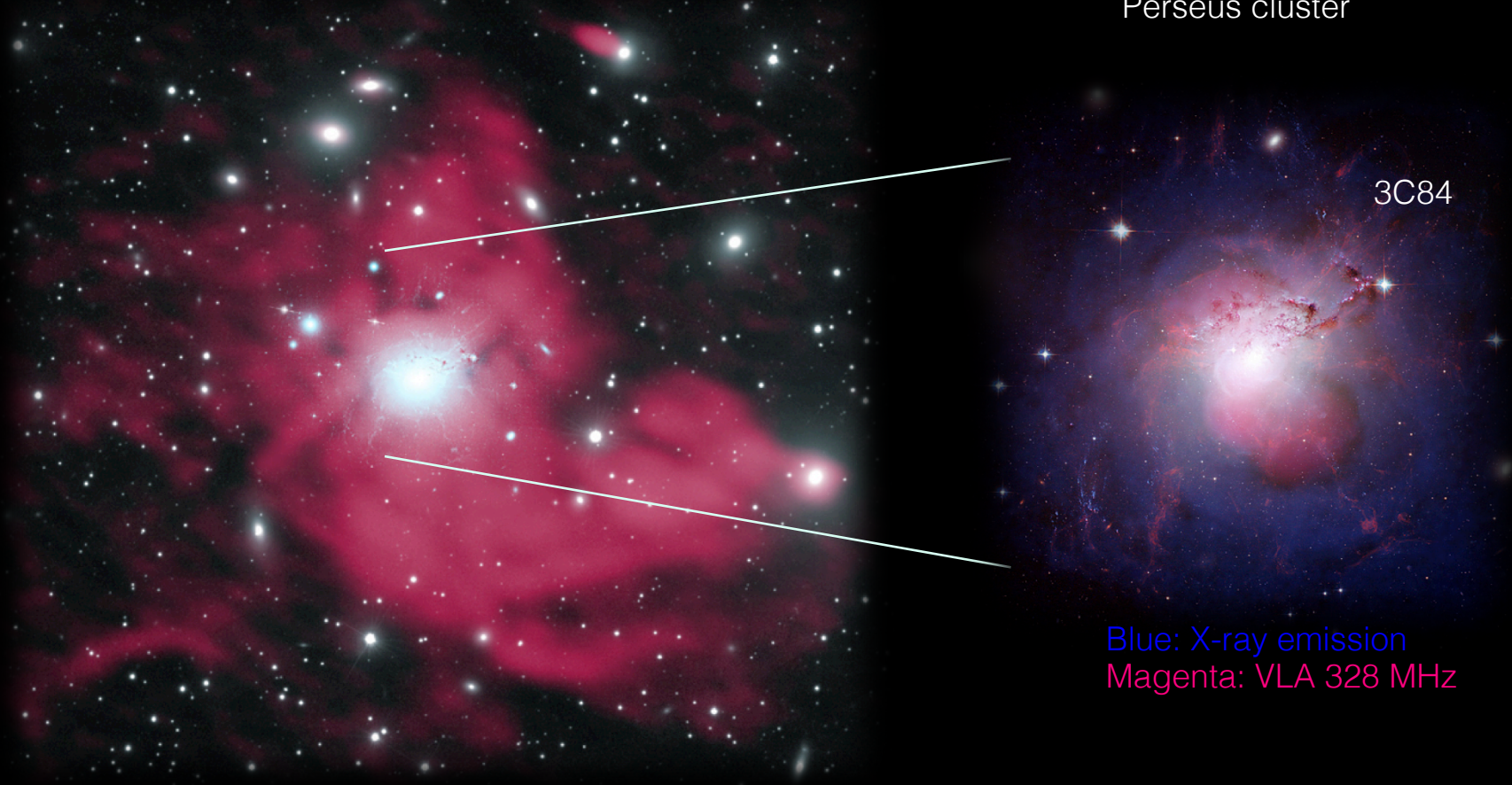
Blue: X-ray emission
Magenta: radio emission

Radio minihalos in cool-core clusters

Gendron-Marsolais et al. 2017

VLA 230-470 MHz

Perseus cluster



Diffuse radio emission on cluster core scale ($r \sim 50 - 300$ kpc), possibly linked to turbulence in the cool core (Gitti et al. 2002, ZuHone et al. 2013)

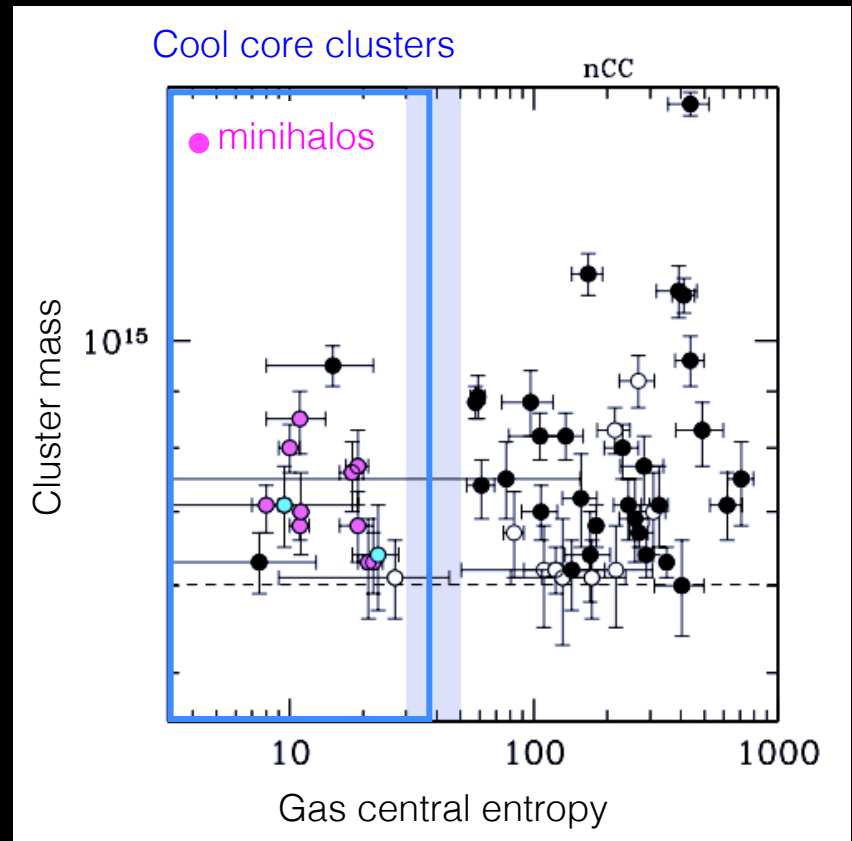
Radio minihalos in cool-core clusters

Gendron-Marsolais et al. 2017

VLA 230-470 MHz



Giacintucci et al. 2017

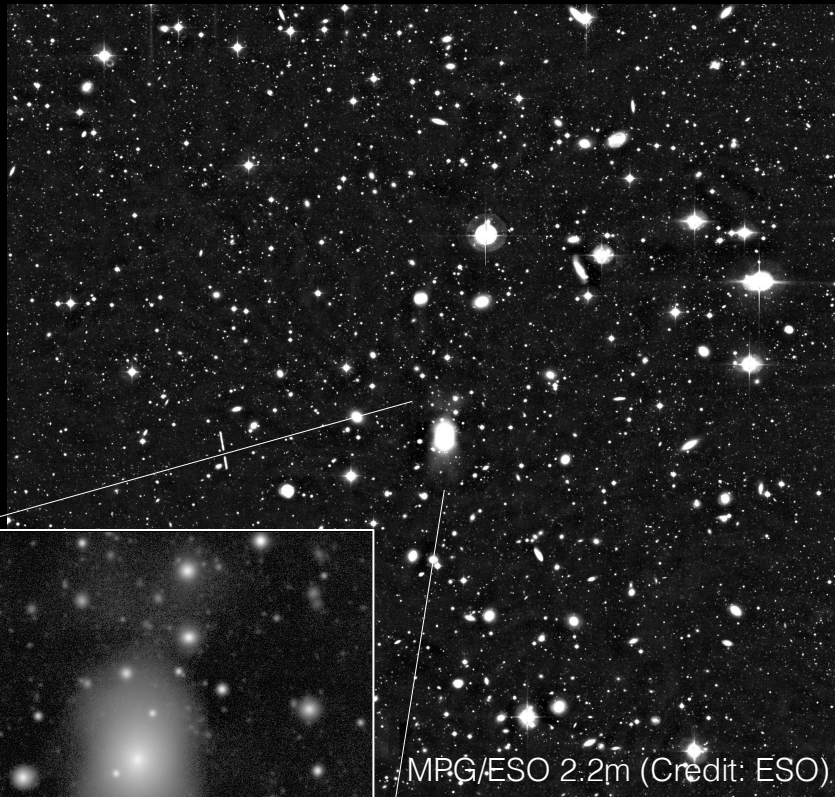


Almost all cool cores (~80%) in massive clusters possess a minihalo

The galaxy clusters A496 and 2A0335+096

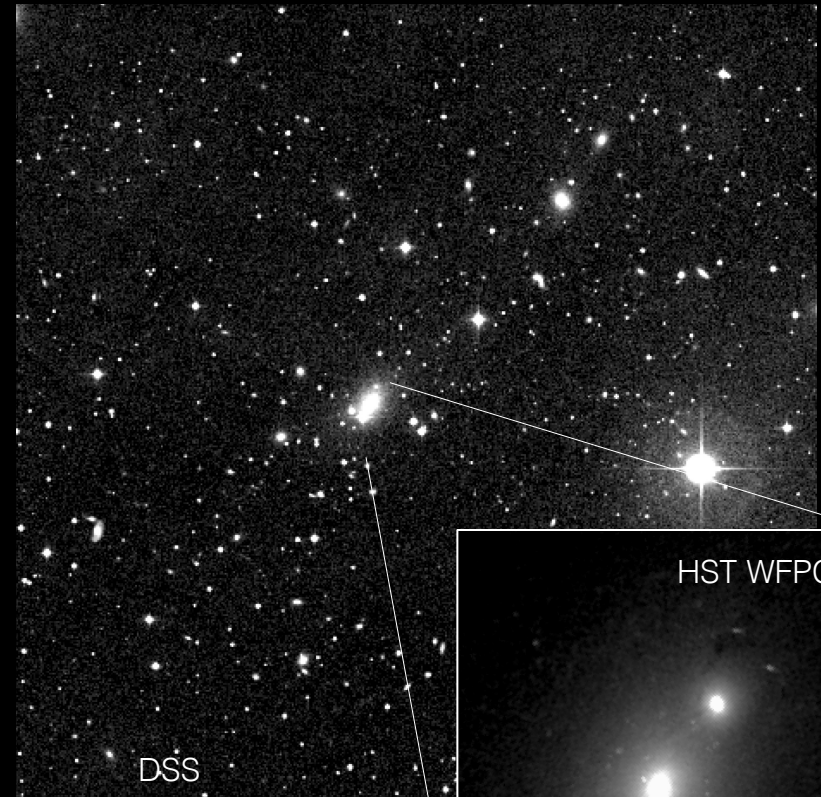
A496

$z=0.0329$, $M_{500} = 2.74 \times 10^{14} M_{\text{sun}}$



2A 0335+096

$z=0.0349$, $M_{500} = 2.27 \times 10^{14} M_{\text{sun}}$



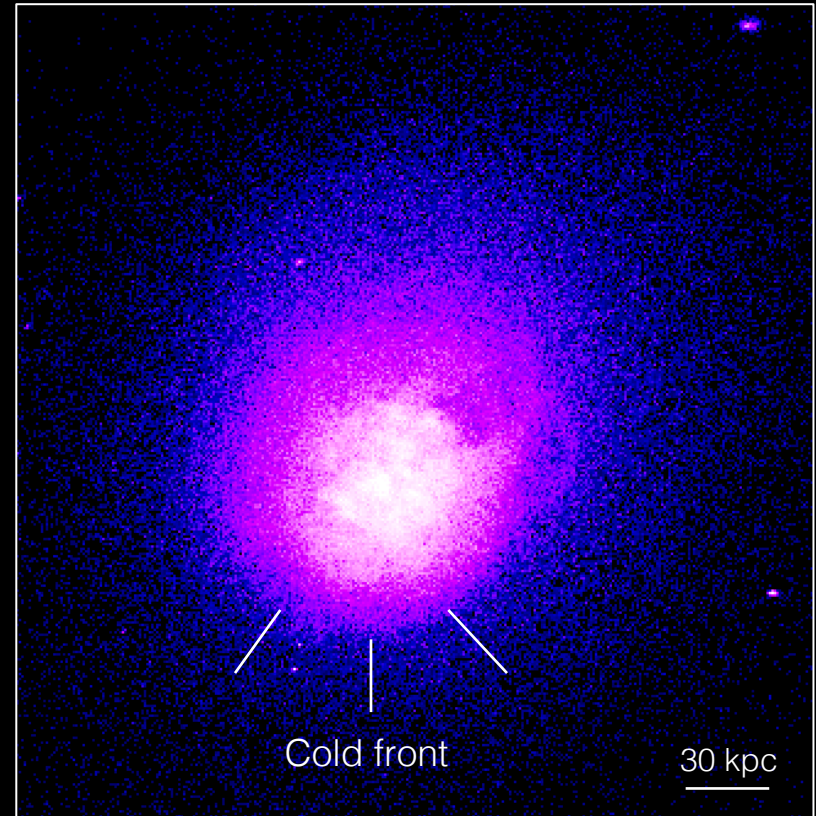
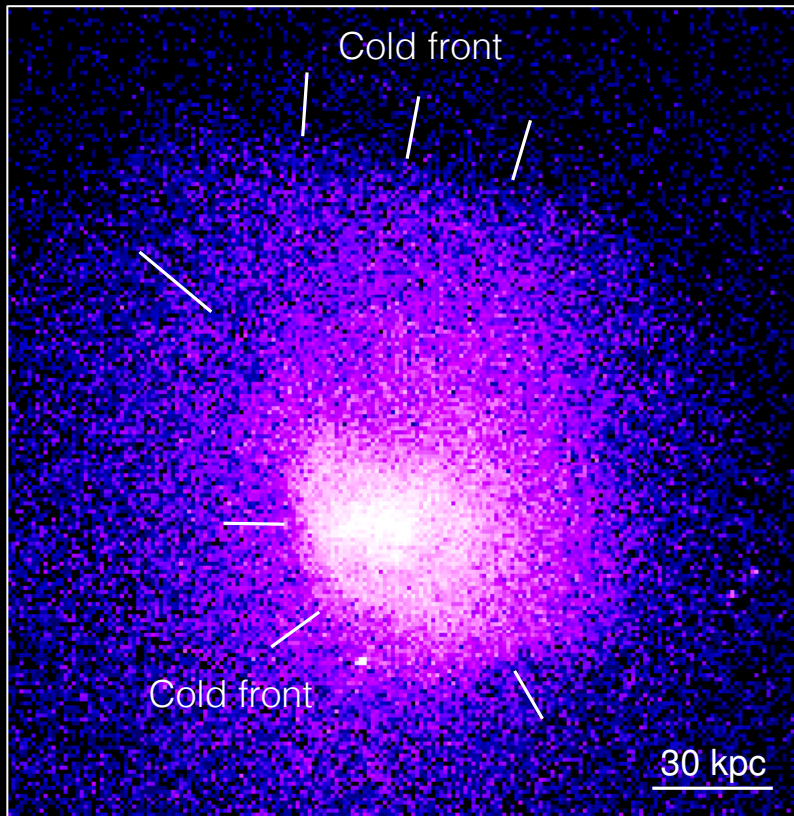
~1 Mpc

Sloshing cold fronts in the cool core

A496

Chandra X-ray 0.5-4 keV images

2A 0335



Dupke & White 2003, Dupke et al. 2007
Ghizzardi et al 2014

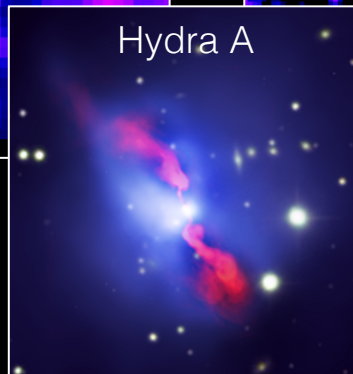
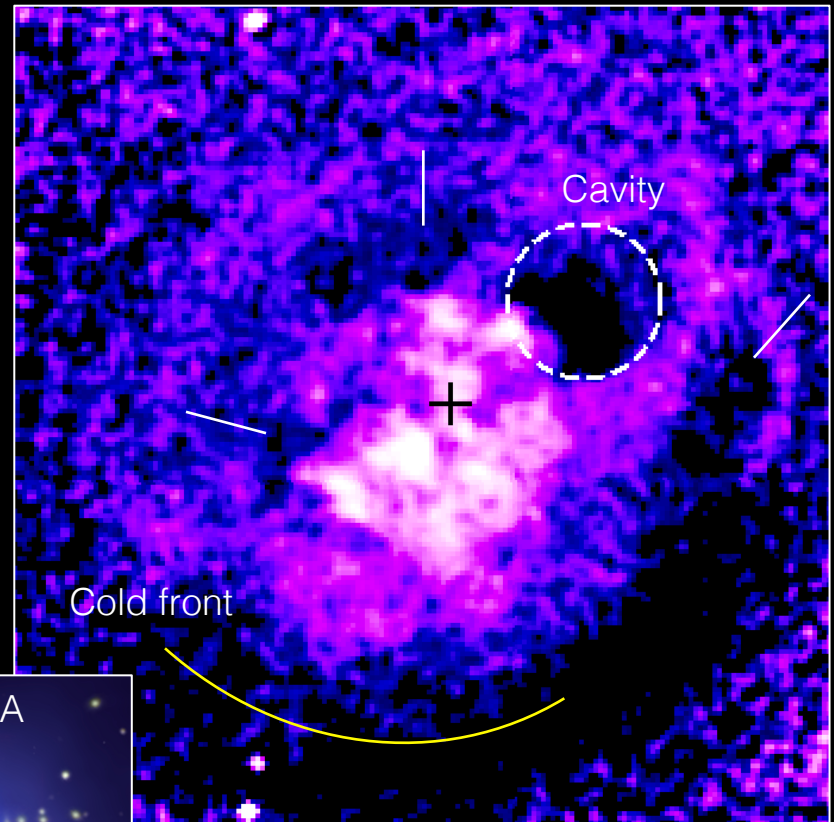
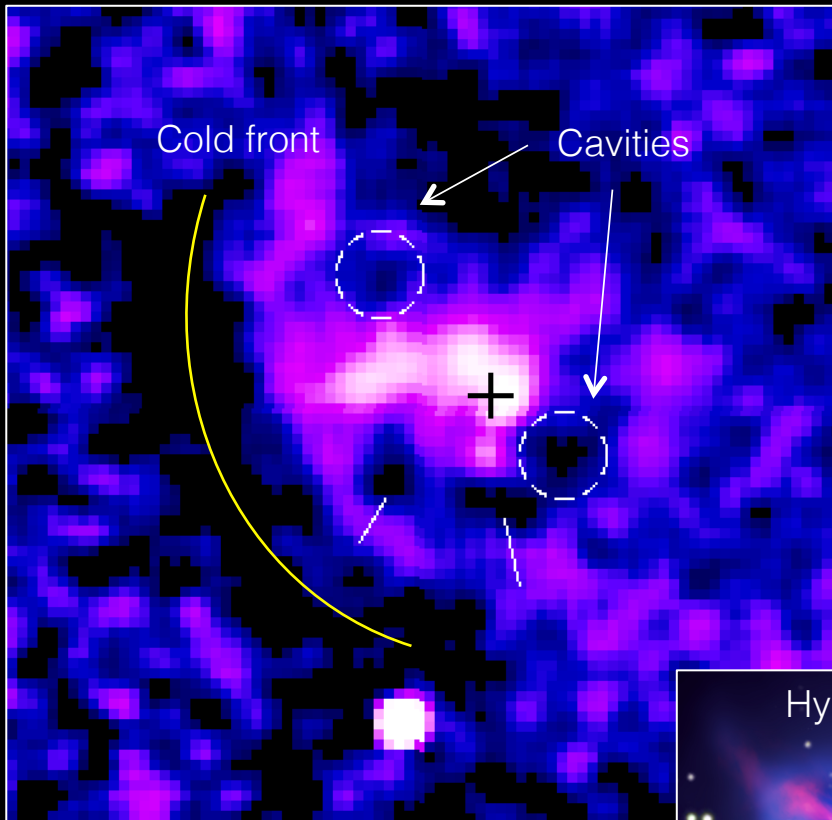
Mazzotta et al. 2003, Sanders et al. 2009

X-ray blobs and cavities in the cool core

A496

Chandra residual images

2A 0335



Giacintucci et al .

Mazzotta et al. 2003
Sanders et al. 2009

Central active radio galaxy

Chandra residual images

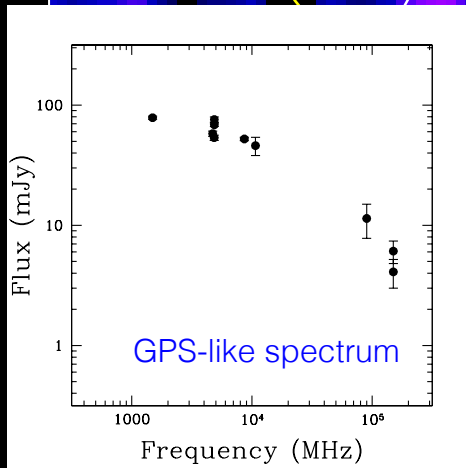
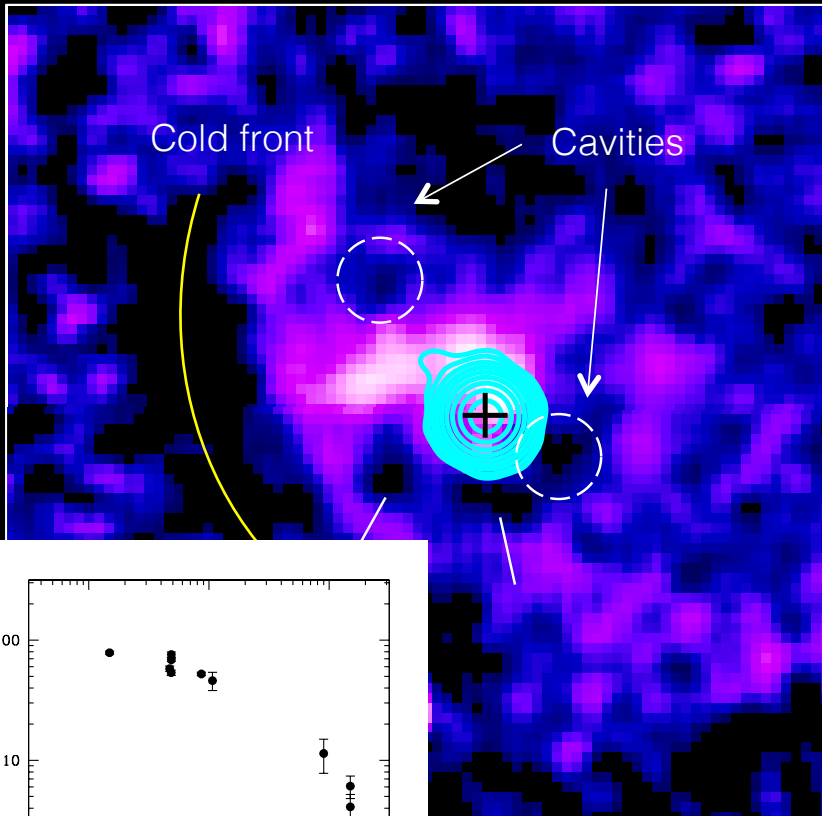
A496

2" res

VLA 5 GHz

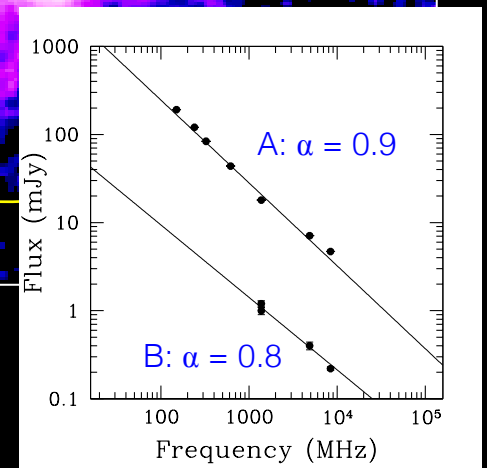
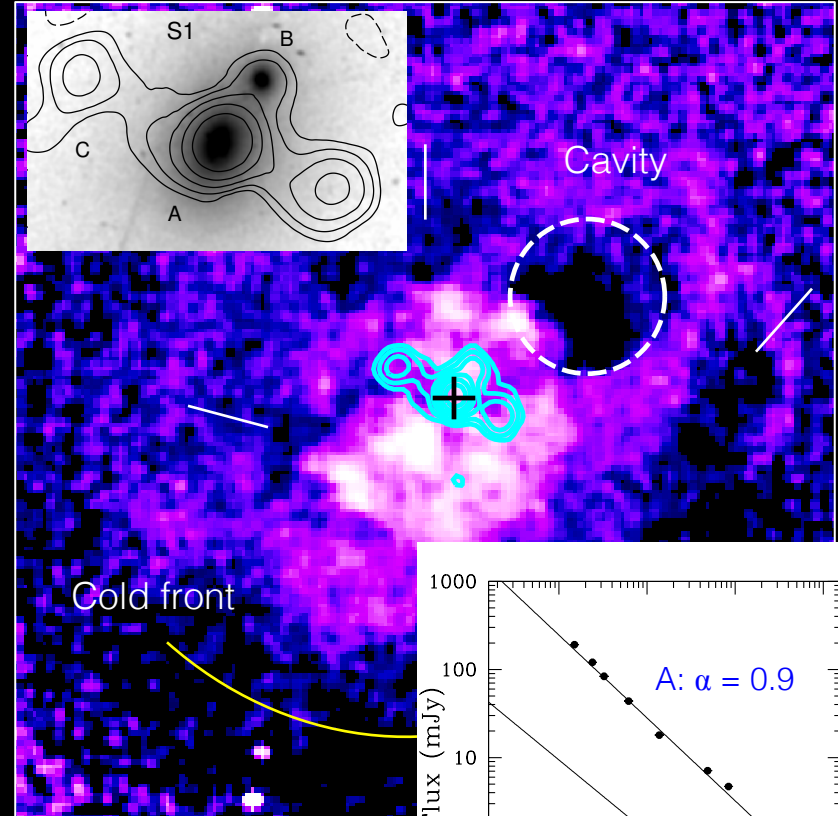
4" res.

2A 0335



Size < 0.3 kpc

see also Hogan et al. 2015



Size ~ 20 kpc

Is there any radio emission in the cavities?

Chandra residual images

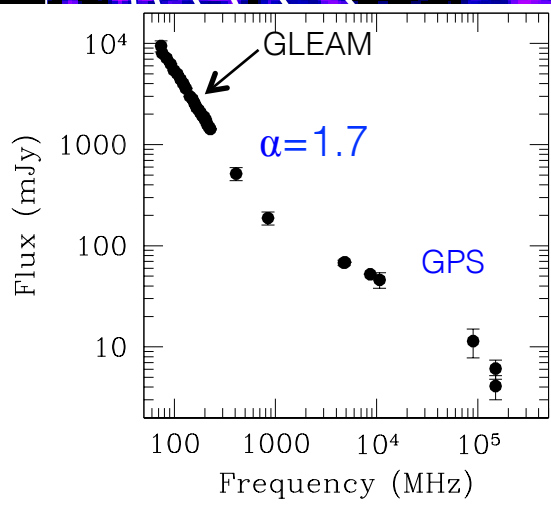
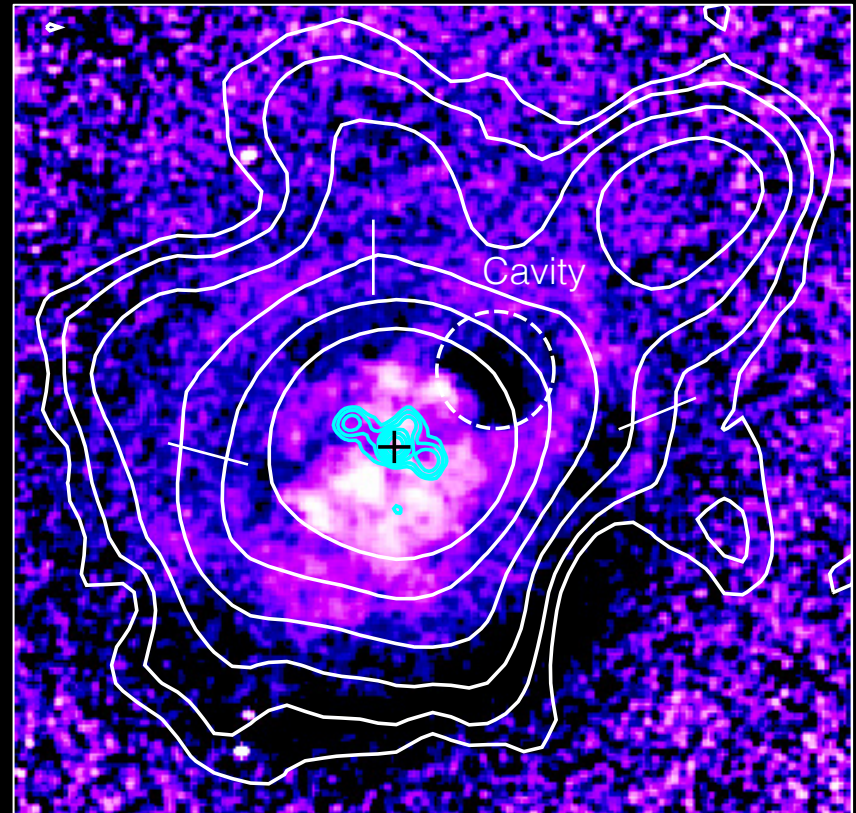
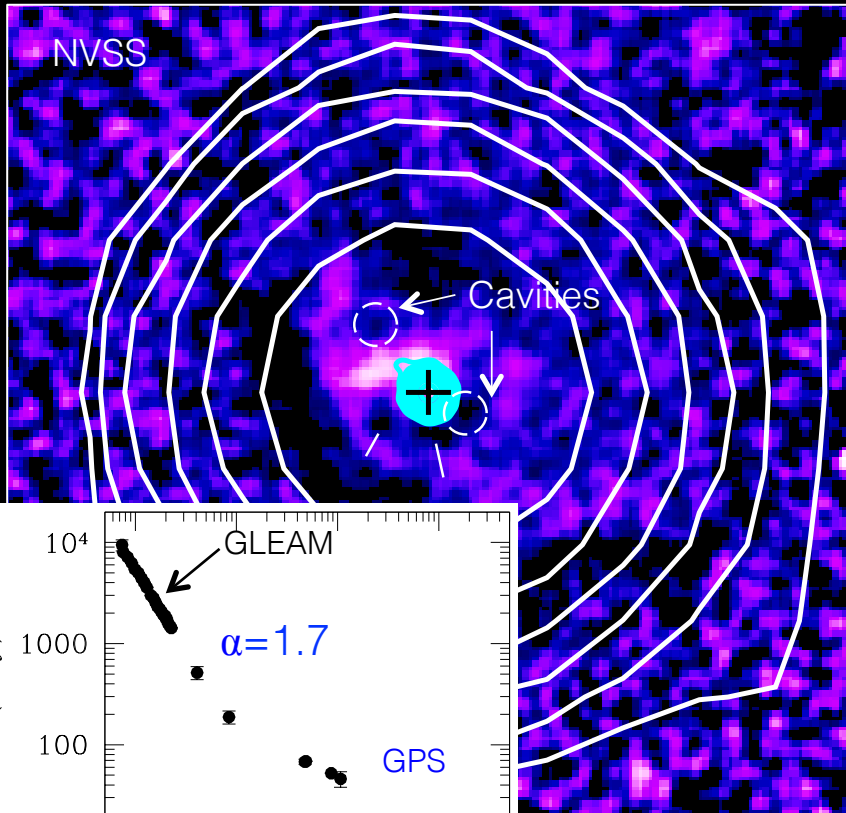
A496

45" res

VLA 1.4 GHz

30" res.

2A 0335



see also Hogan
et al. 2015

(Sarazin et al. 1995, Sanders et. 2009)

Radio minihalo

A496

GMRT 150 MHz (3 TGSS pointings)

GMRT 327 MHz

* VLITE 340 MHz

GMRT 610 MHz

VLA 1.4 GHz A, B and C configurations

VLA 5 GHz A, B and D configurations

2A 0335

GMRT 150 MHz (TGSS ADR)

GMRT 240 MHz

VLA 325 MHz

GMRT 610 MHz

GMRT 1.3 GHz

MERLIN 1.4 GHz

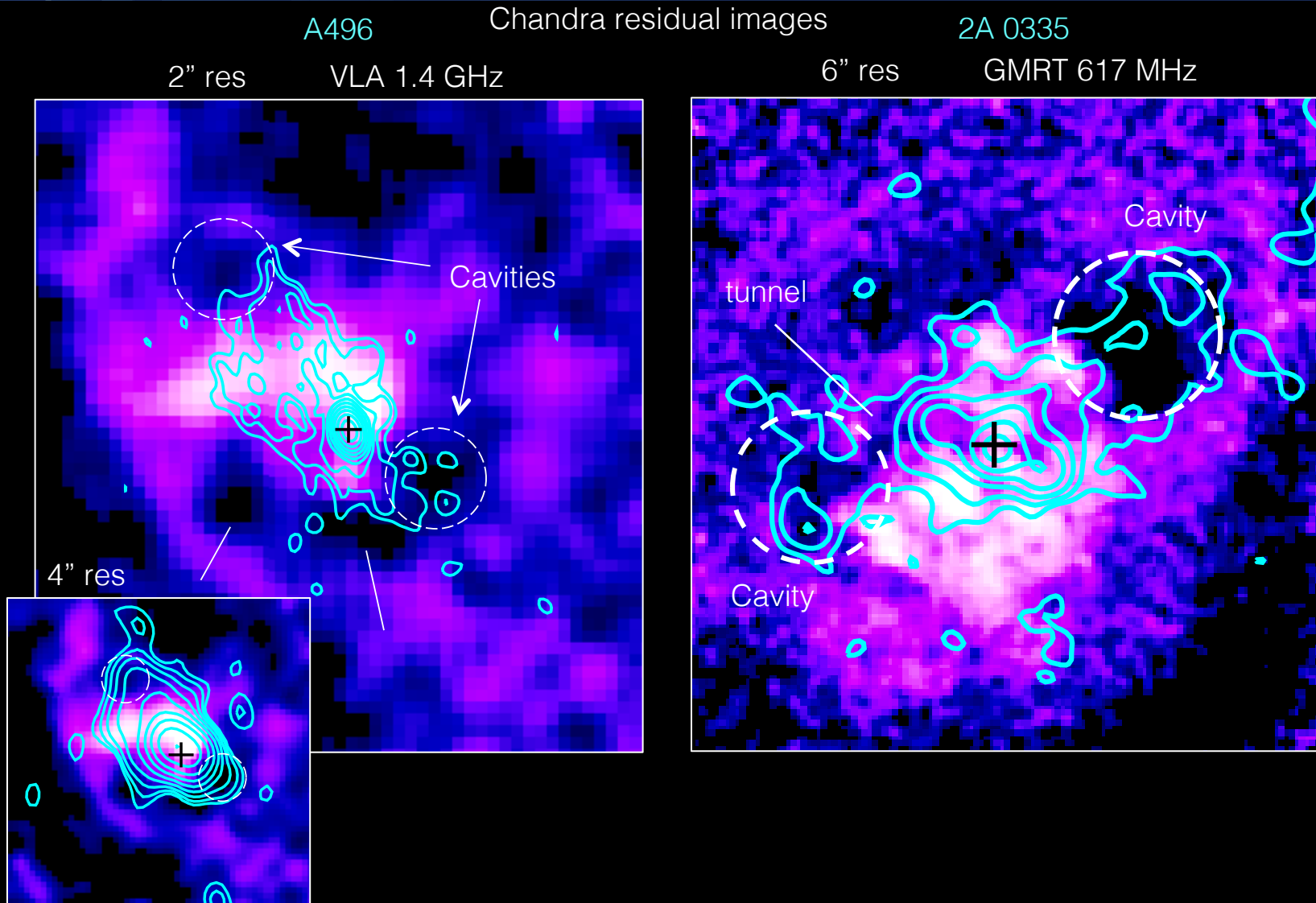
VLA 1.4 GHz A, B, C and D configurations

VLA 5 GHz C and D configurations

VLA 8 GHz B and C configurations

* VLA Low-band Ionosphere and Transient Experiment (see Tracy Clarke's poster)

Extended radio emission in the cavities

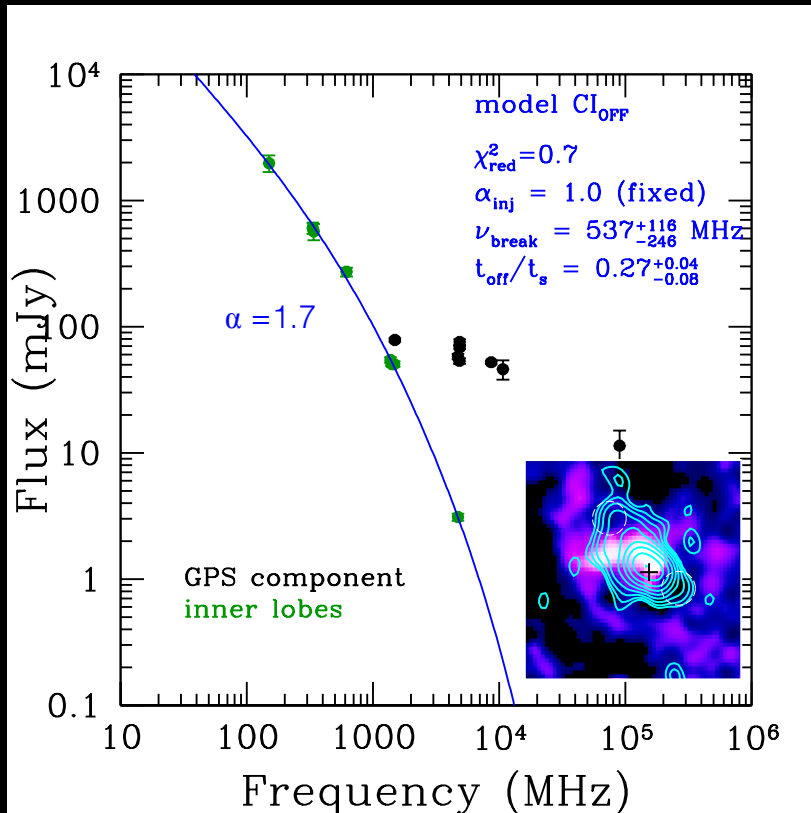


Relic lobes from past radio outburst

A496

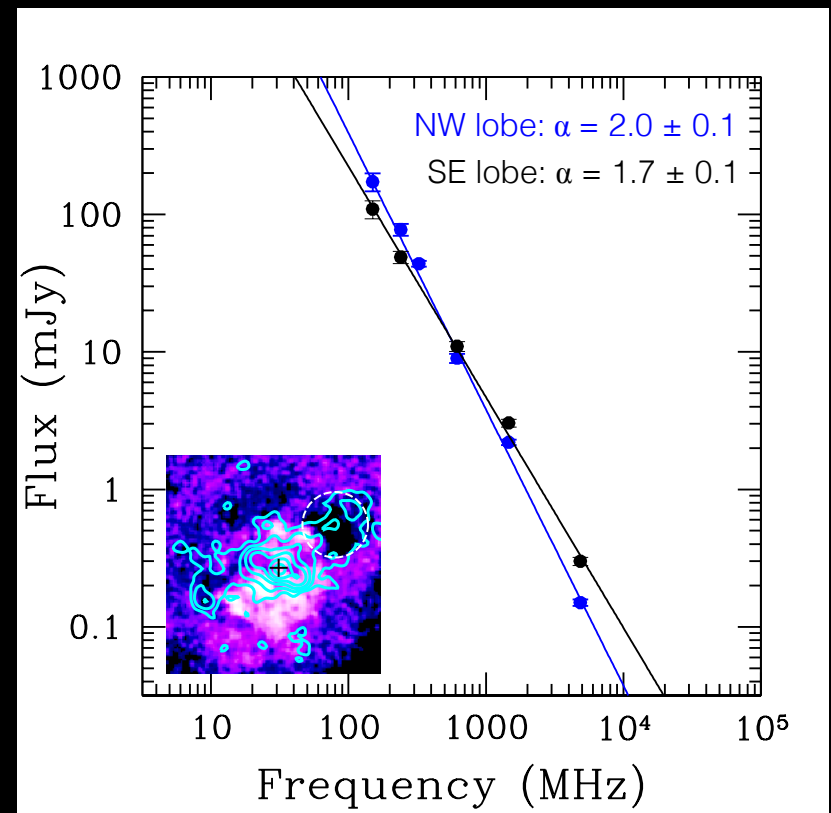
2A 0335

Radio lobe spectra



for $B_{min} = 14 \mu\text{Gauss}$:

$t_s \sim 23$ Myr



for $B_{min} = 24 \mu\text{Gauss}$:

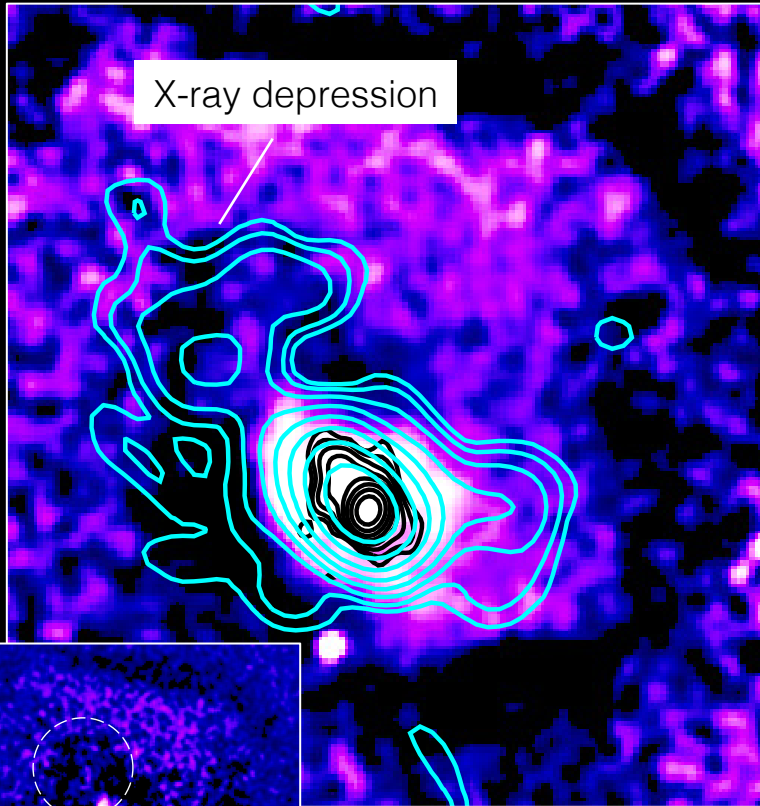
$t_s > 28$ Myr

Relic lobes from an even older radio outburst?

A496

12" res

GMRT 327 MHz



Chandra residual image

2A 0335

15" res

GMRT 617 MHz

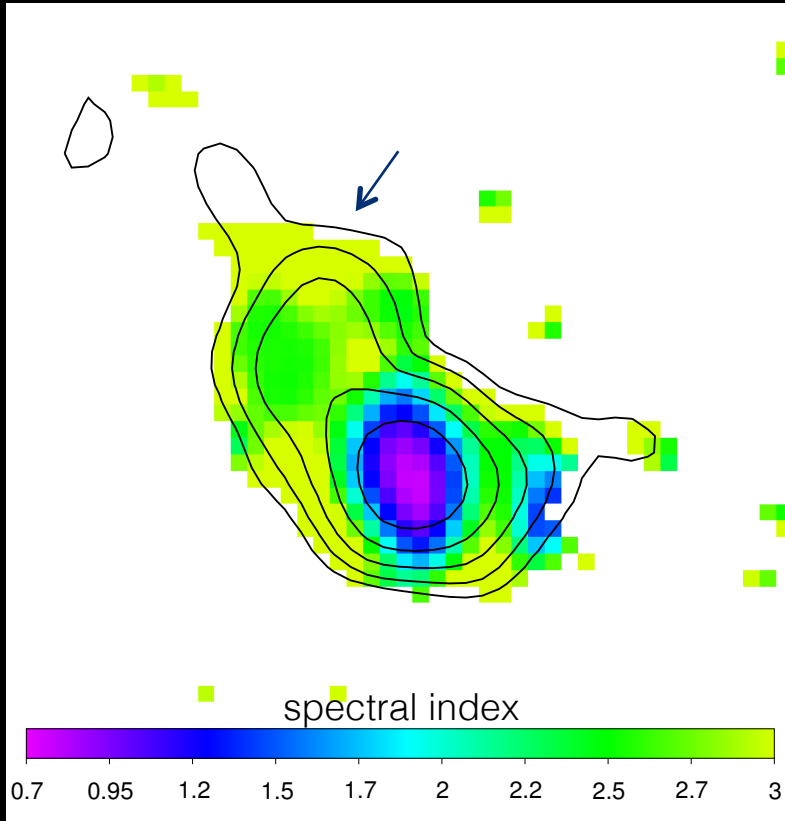


XMM+Chandra residual image

Relic lobes from an even older radio outburst

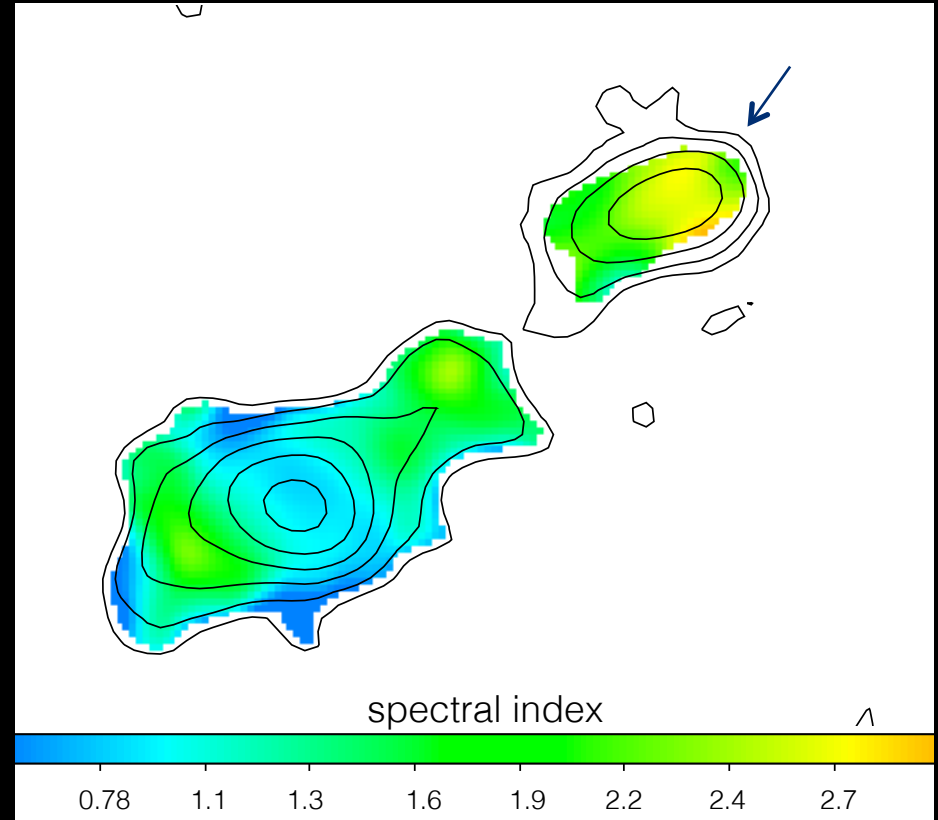
A496

GMRT 150-327 MHz



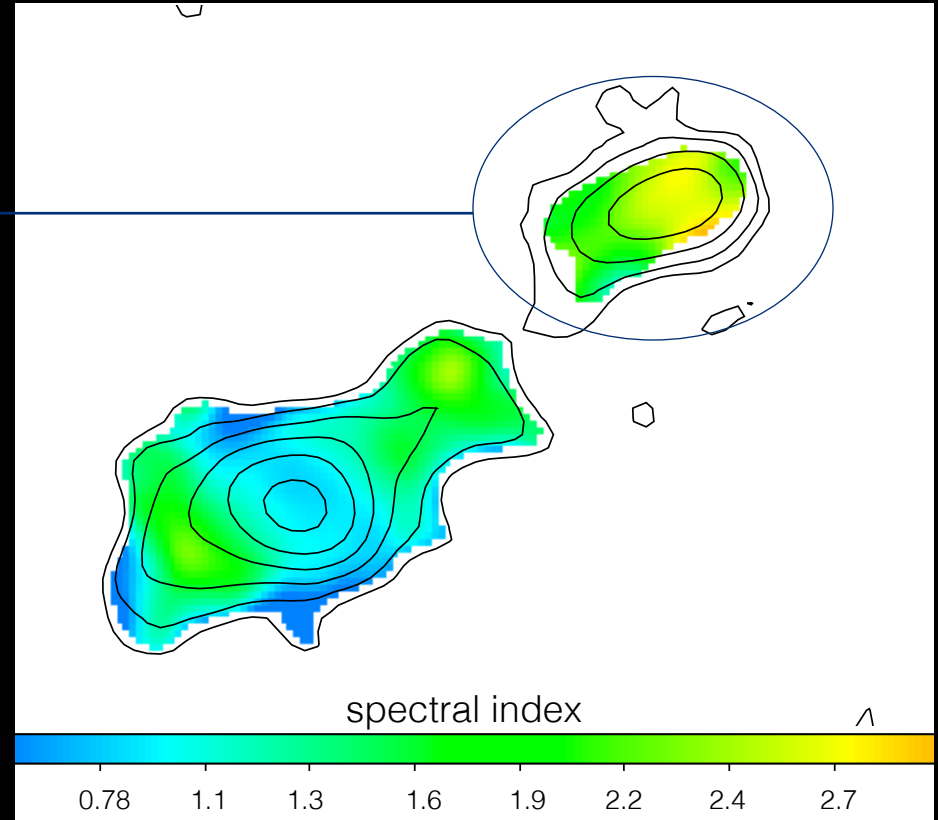
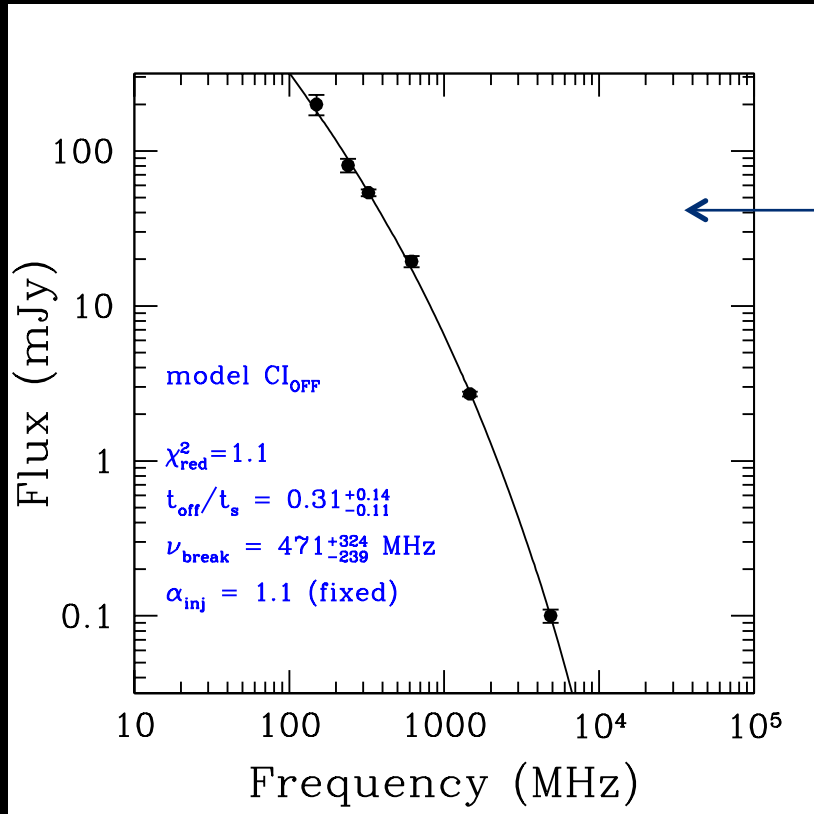
2A 0335

GMRT 617 MHz – VLA 1.4 GHz



2A 0335

GMRT 617 MHz – VLA 1.4 GHz

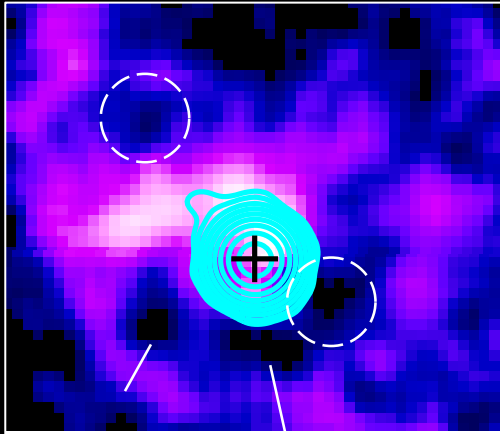


for $B_{min} = 10 \mu\text{Gauss}$

$t_s \sim 60$ Myr

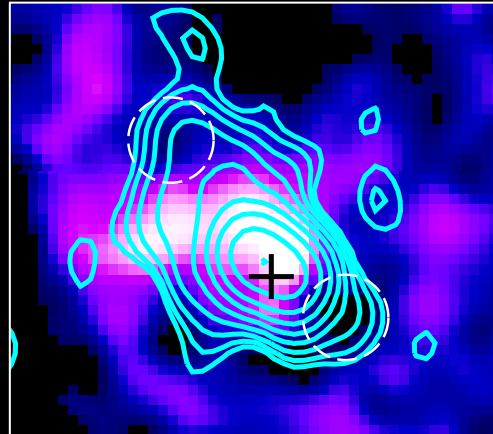
A history of 3 AGN outbursts

I. Current activity



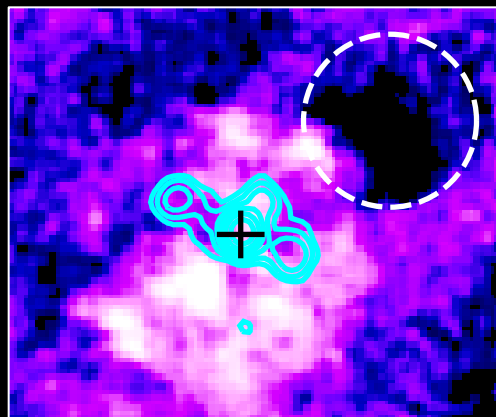
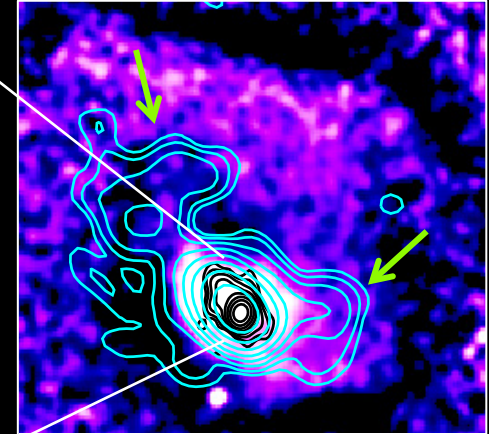
Size < 0.3 kpc

II. Past activity

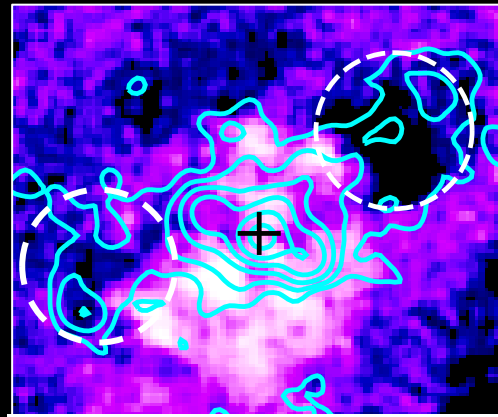


age ~23 Myr

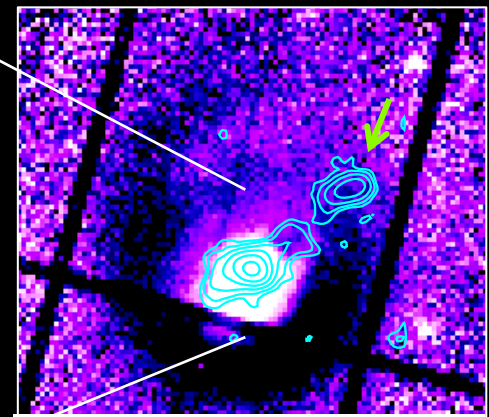
III. Even older activity



Size 20 kpc



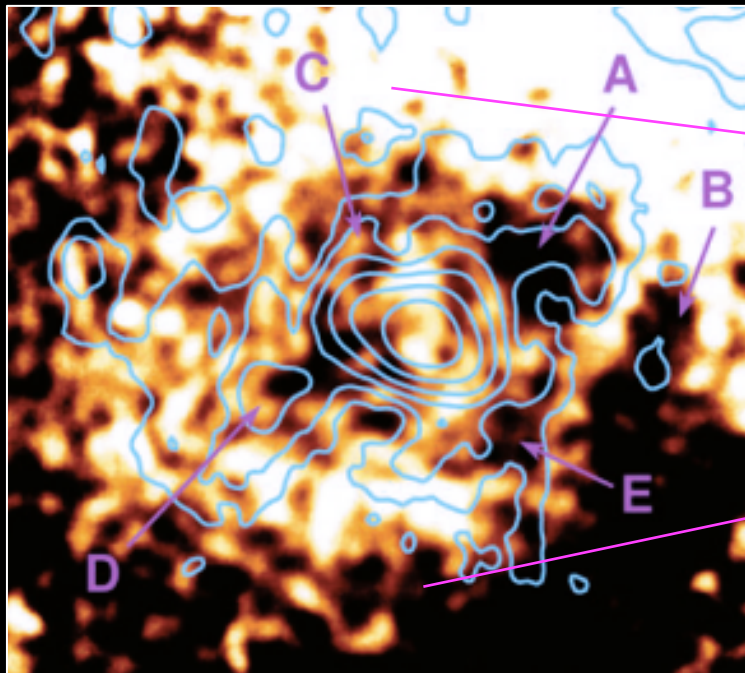
age > 28 Myr



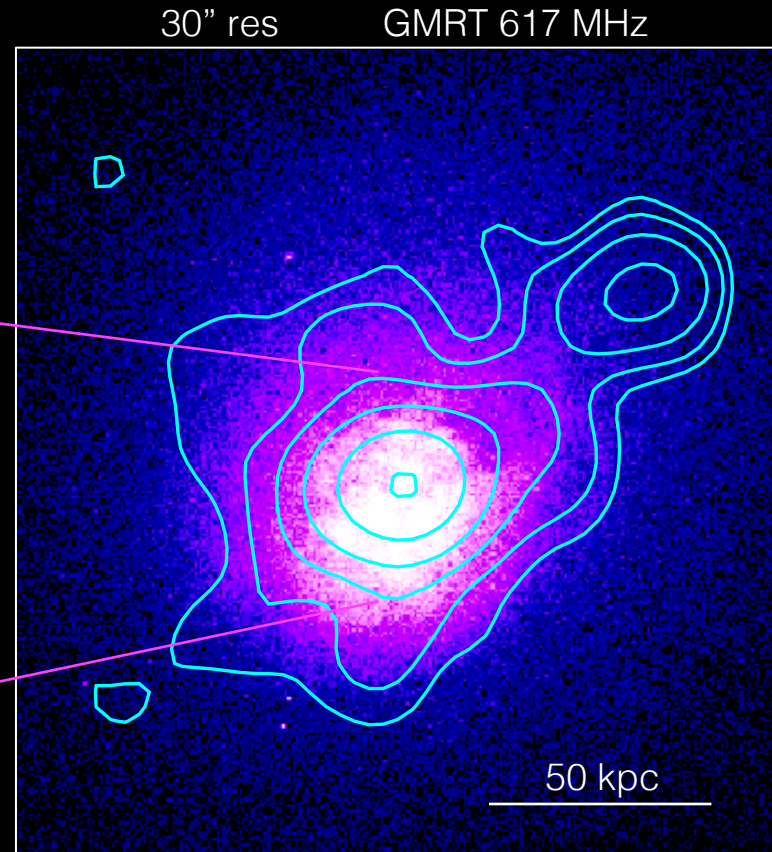
age ~60 Myr

The minihalo in 2A 0335+096

Is the minihalo formed by episodes of AGN activity?



Sanders et. 2009



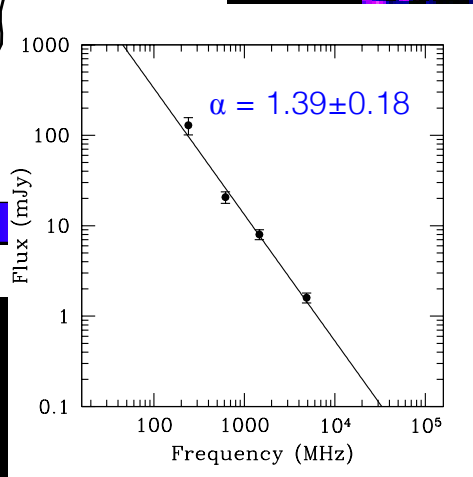
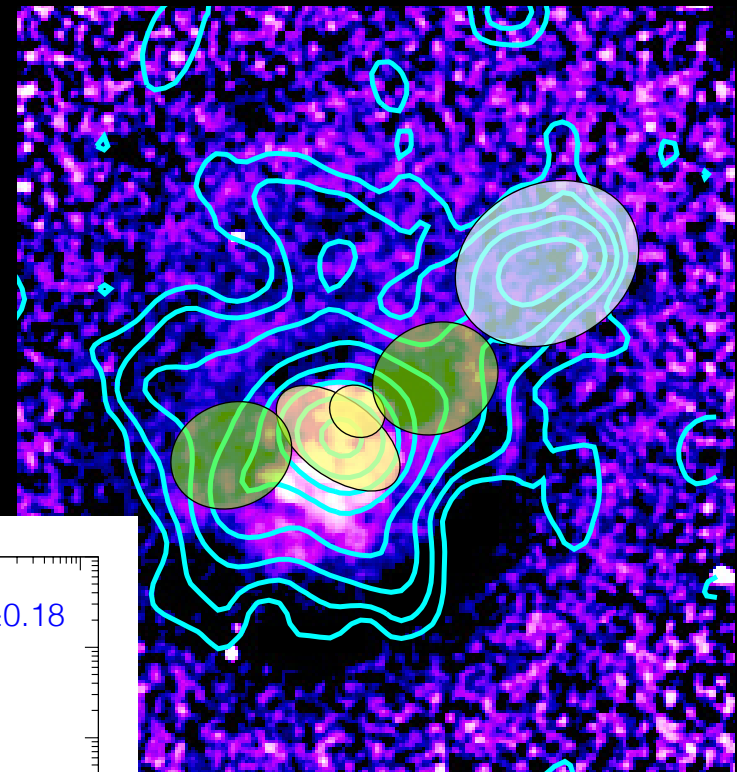
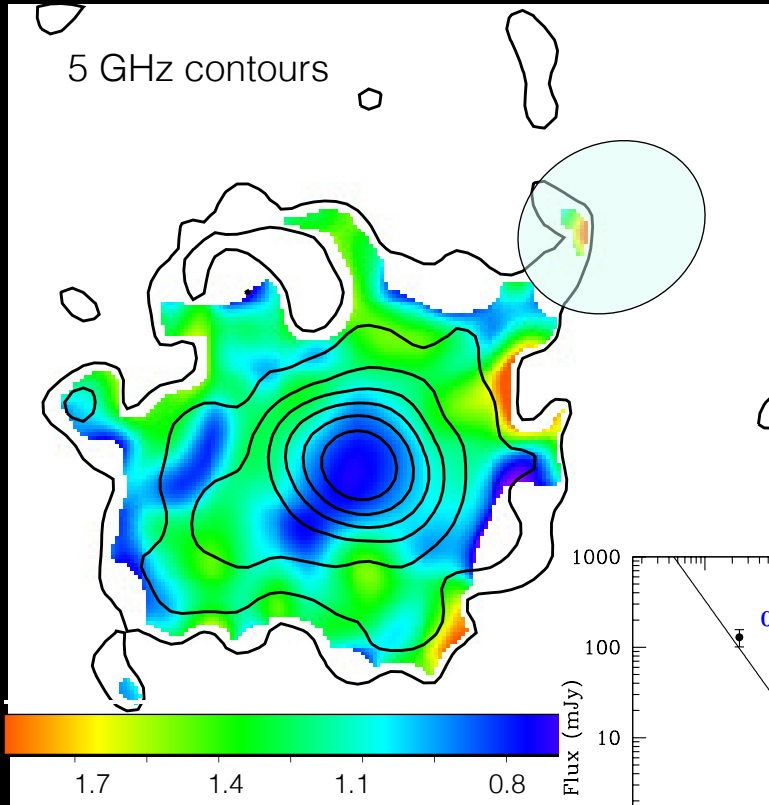
see also Sarazin et al. 1995

Minihalo spectrum

1.4- 5 GHz spectral index - 17" res

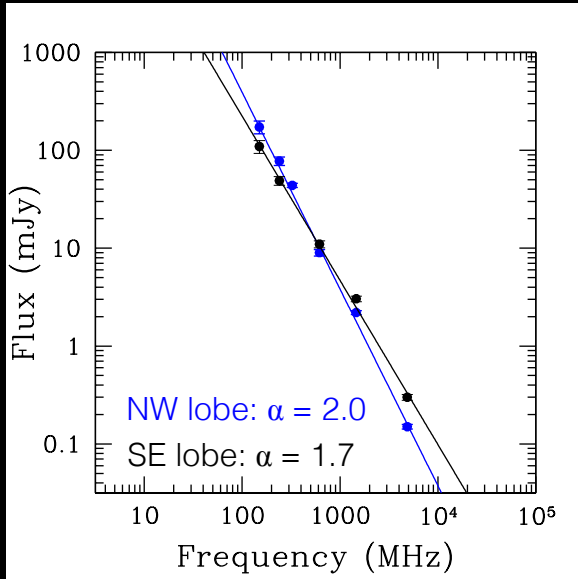
17" res

VLA 1.4 GHz

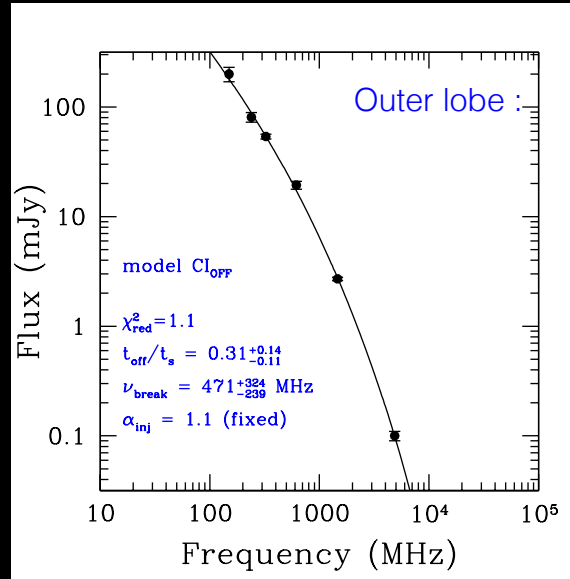


The minihalo spectrum is flatter than the spectrum of the relic lobes

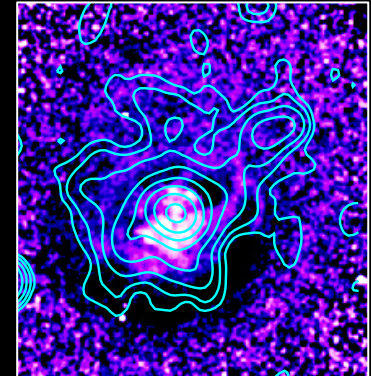
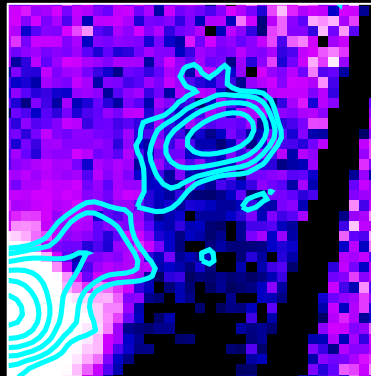
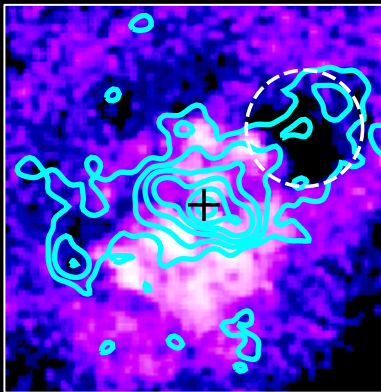
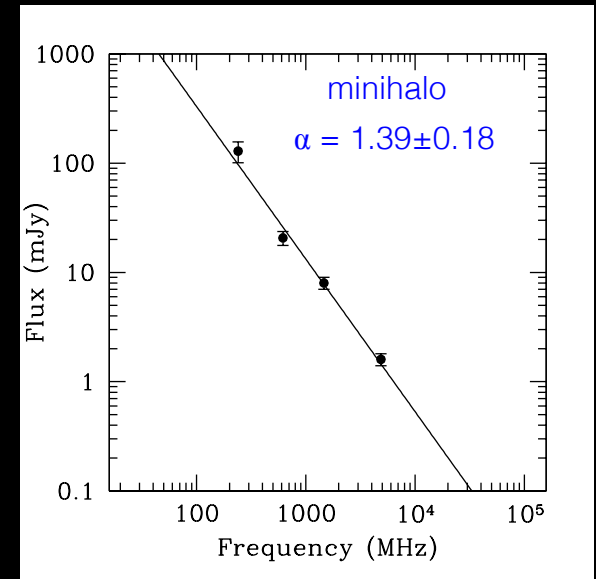
ageing electrons



ageing electrons



reaccelerated electrons



Thank you

