

GLEAM: calibration, images, and catalogues

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International
Centre for
Radio
Astronomy
Research



Curtin University



MWA
MURCHISON
WIDFIELD
ARRAY



CAASTRO
ALL-SKY ASTROPHYSICS

The survey: some challenges

Calibration:

- Sky model not well known
- Beam model ok but not great
- No multi- ν surveys to cover the south

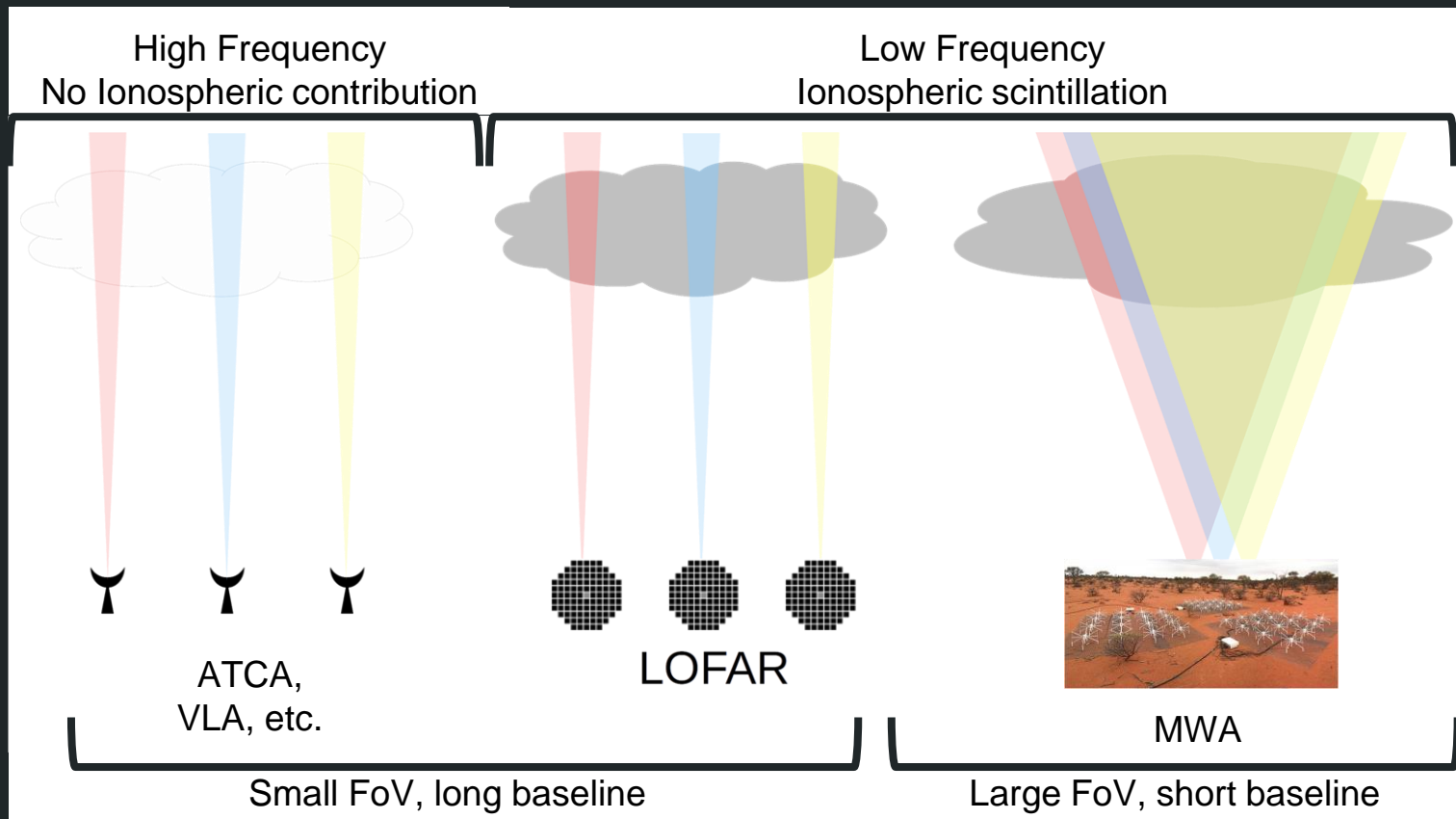
Imaging:

- Mosaicing distorted images causes blurring and loss of SNR

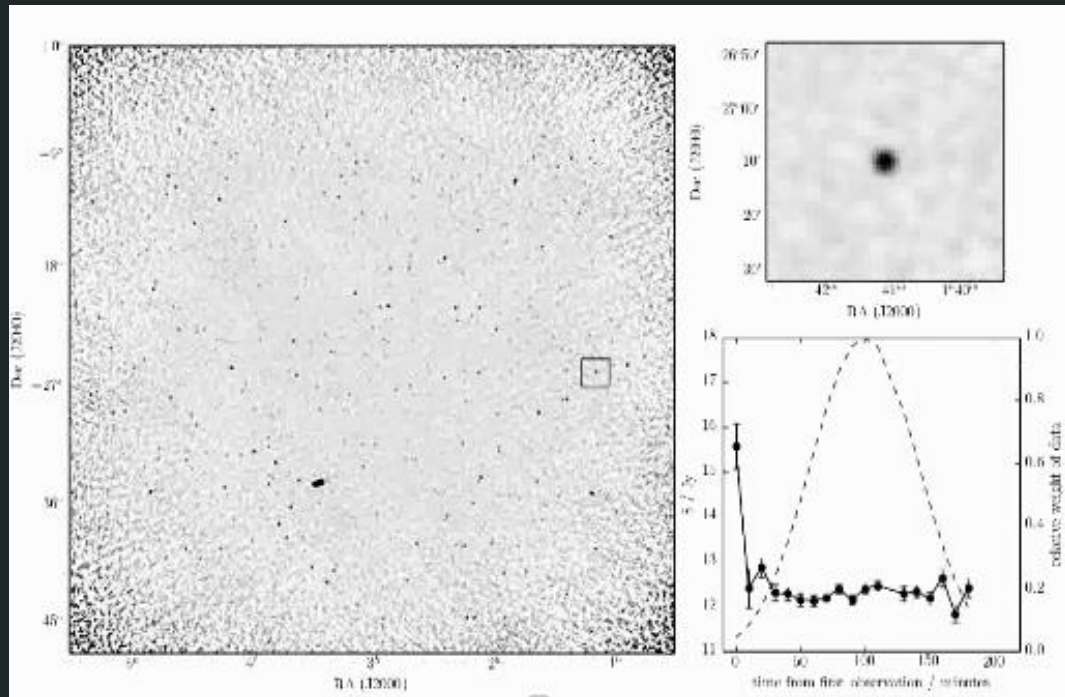
Catalogue creation:

- Large FoV - image properties change over the sky
- Large $\Delta\nu$ - changing PSF, and source morphology

More detail, catalogues, and images: Hurley-Walker, [Callingham](#), [Hancock](#) et al. 2017 MNRAS.464.1146H



The Ionosphere (+ Primary Beam)

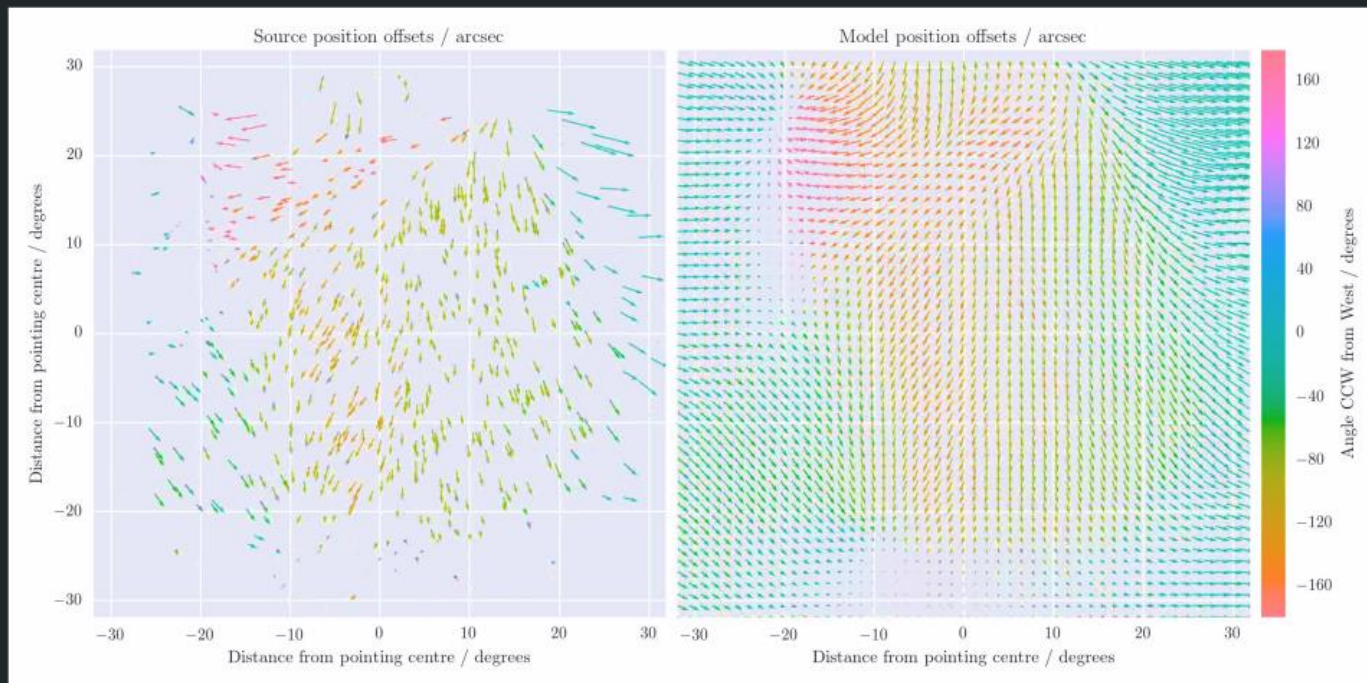


Track a source through multiple snapshots:

Position 'jitters' → Ionosphere

Flux changes → δ Primary Beam

Ionospheric Distortions



$$\nabla \cdot \delta \vec{x} \ll 1\%$$

Shifting but no
focusing

phase errors but
no amp errors

Ionospheric Corrections

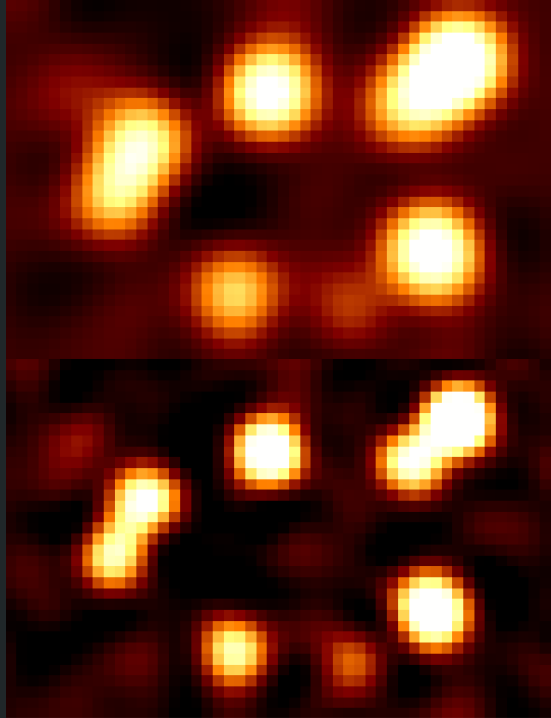
Normal stacking/mosaicing \Rightarrow

(10% loss of resolution)

image warping

+ stacking/mosaicing \Rightarrow

(full resolution)



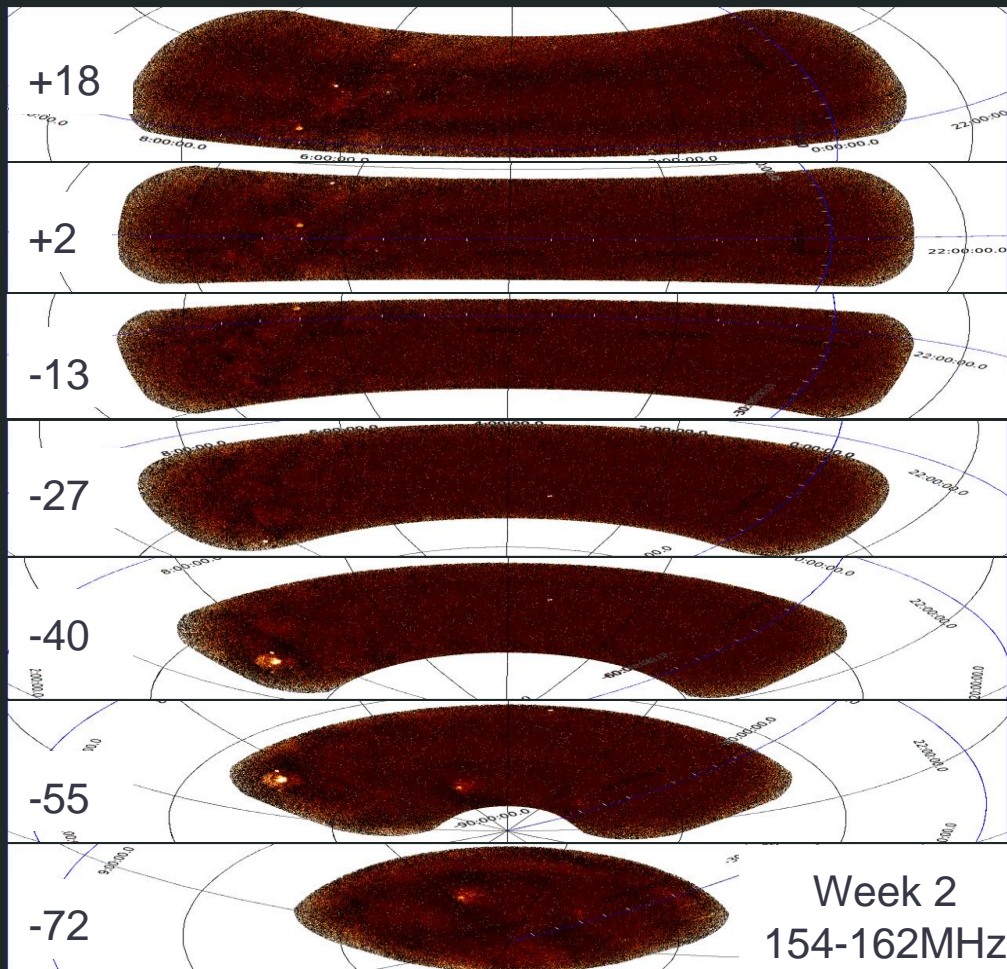
Mosaicking

7 Dec bands (1 per night)

4 RA slices (1 per 3mo)

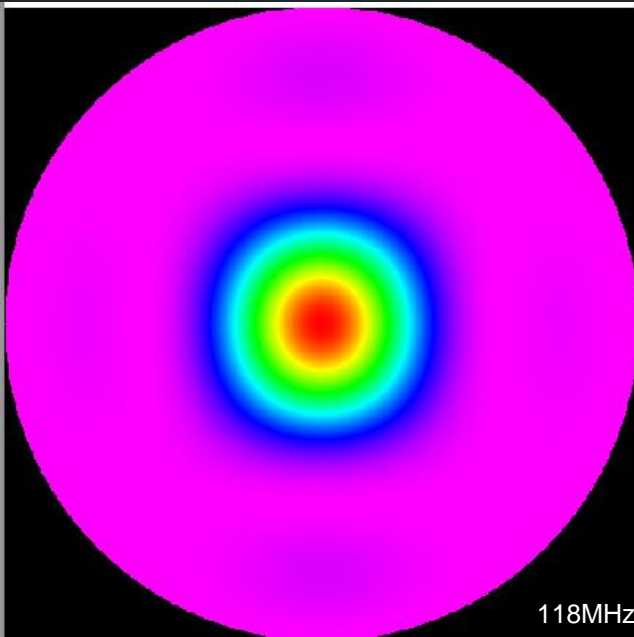
20 frequencies per night

560 mosaics!

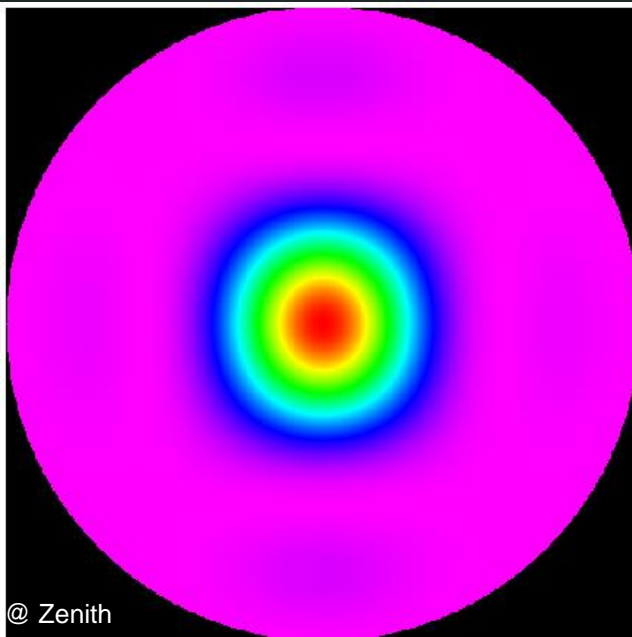


Primary Beam Model Errors

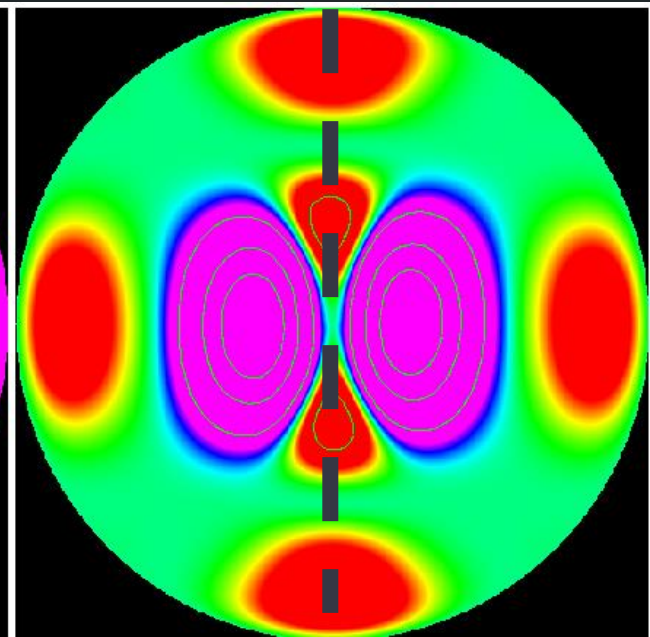
Average embedded element model
2014



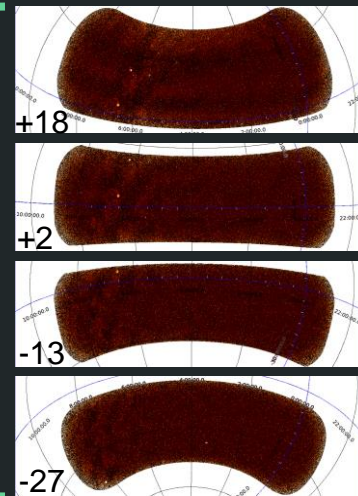
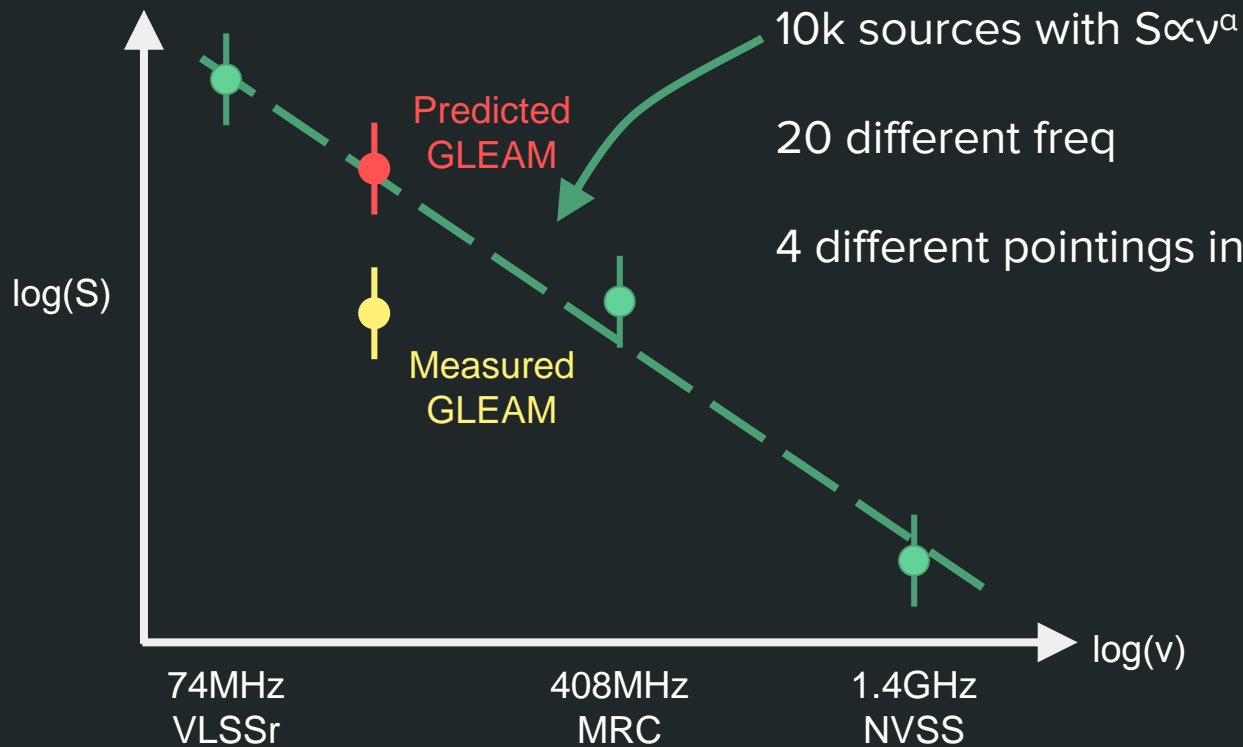
Full embedded element model
2017



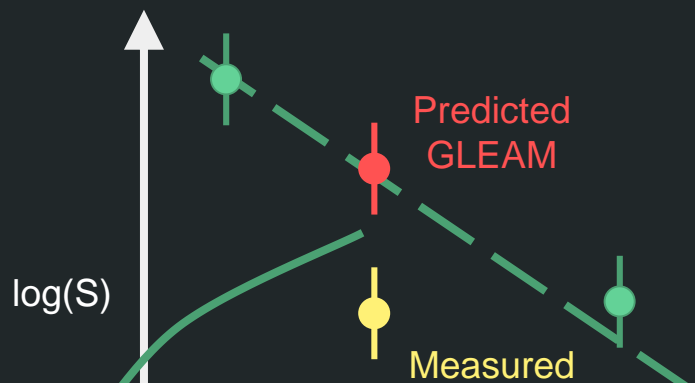
Difference
+/- 10%



Flux Calibration



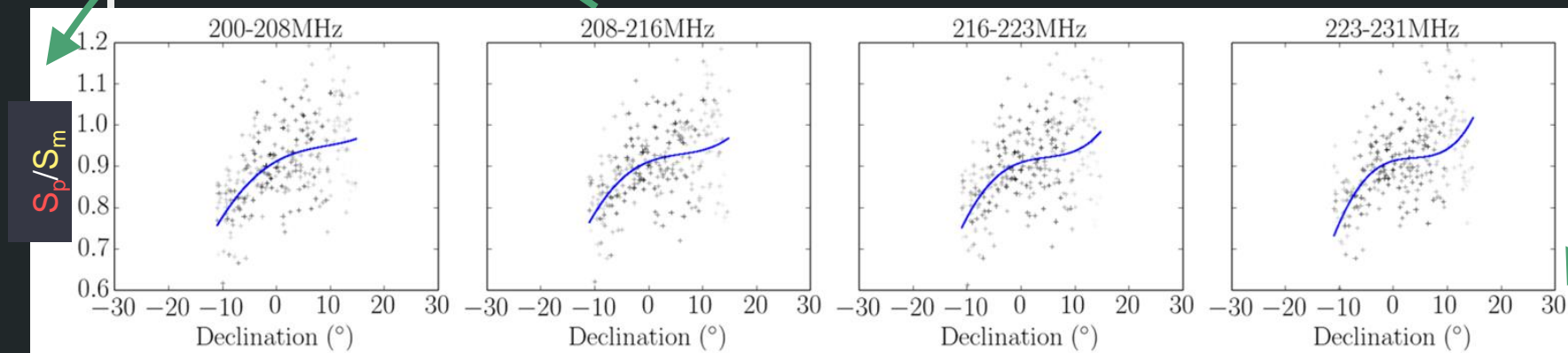
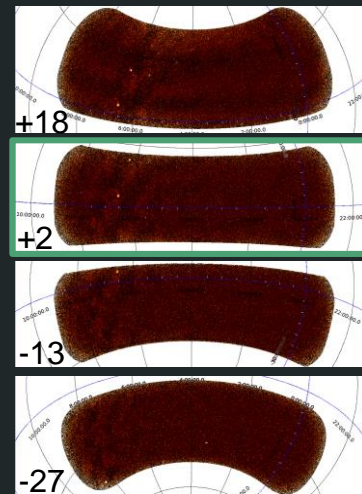
Flux Calibration



10k sources with $S \propto \nu^\alpha$

20 different freq

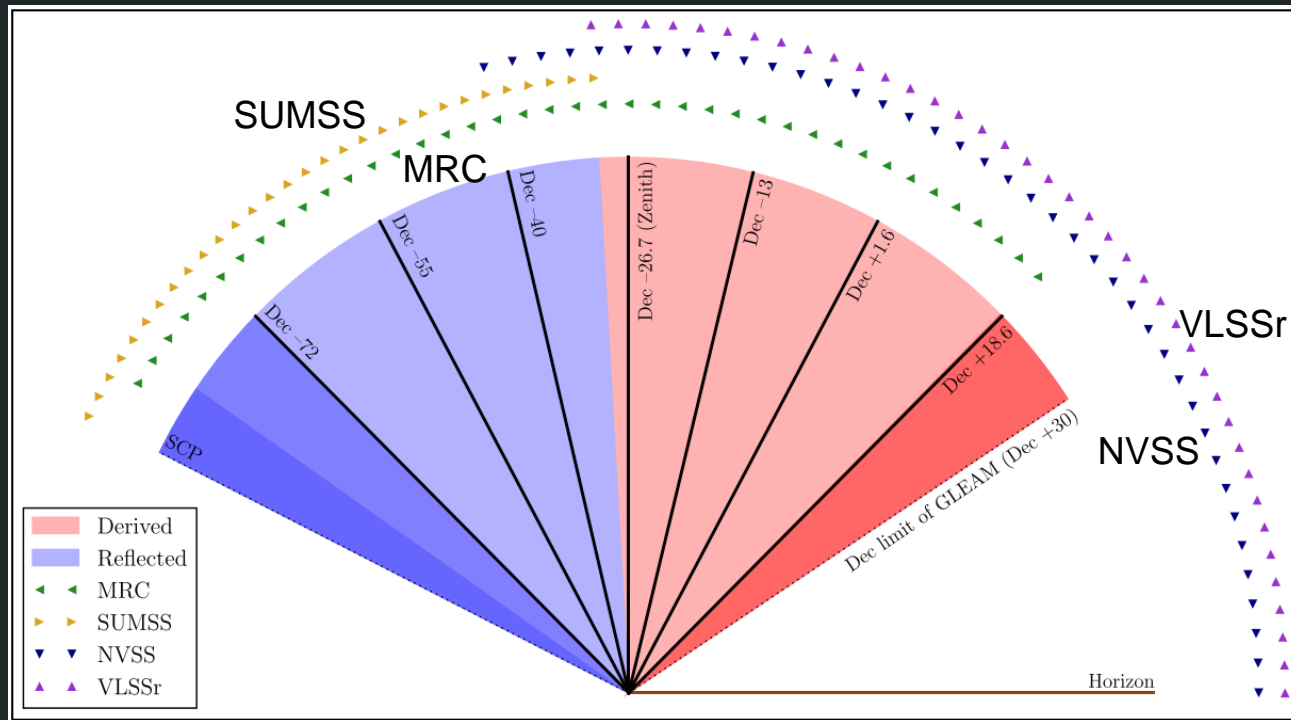
4 different pointings in Dec



Flux Calibration for the southern sky

Not enough
reference points
south of our zenith
for **interpolation**

Exploit the symmetry
of the beams

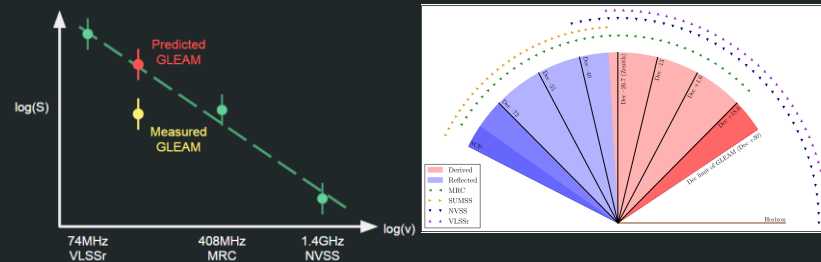
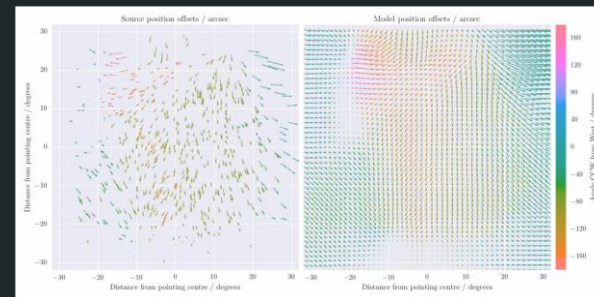


Calibration

Instrumental → point source cal obs, self cal on field sources

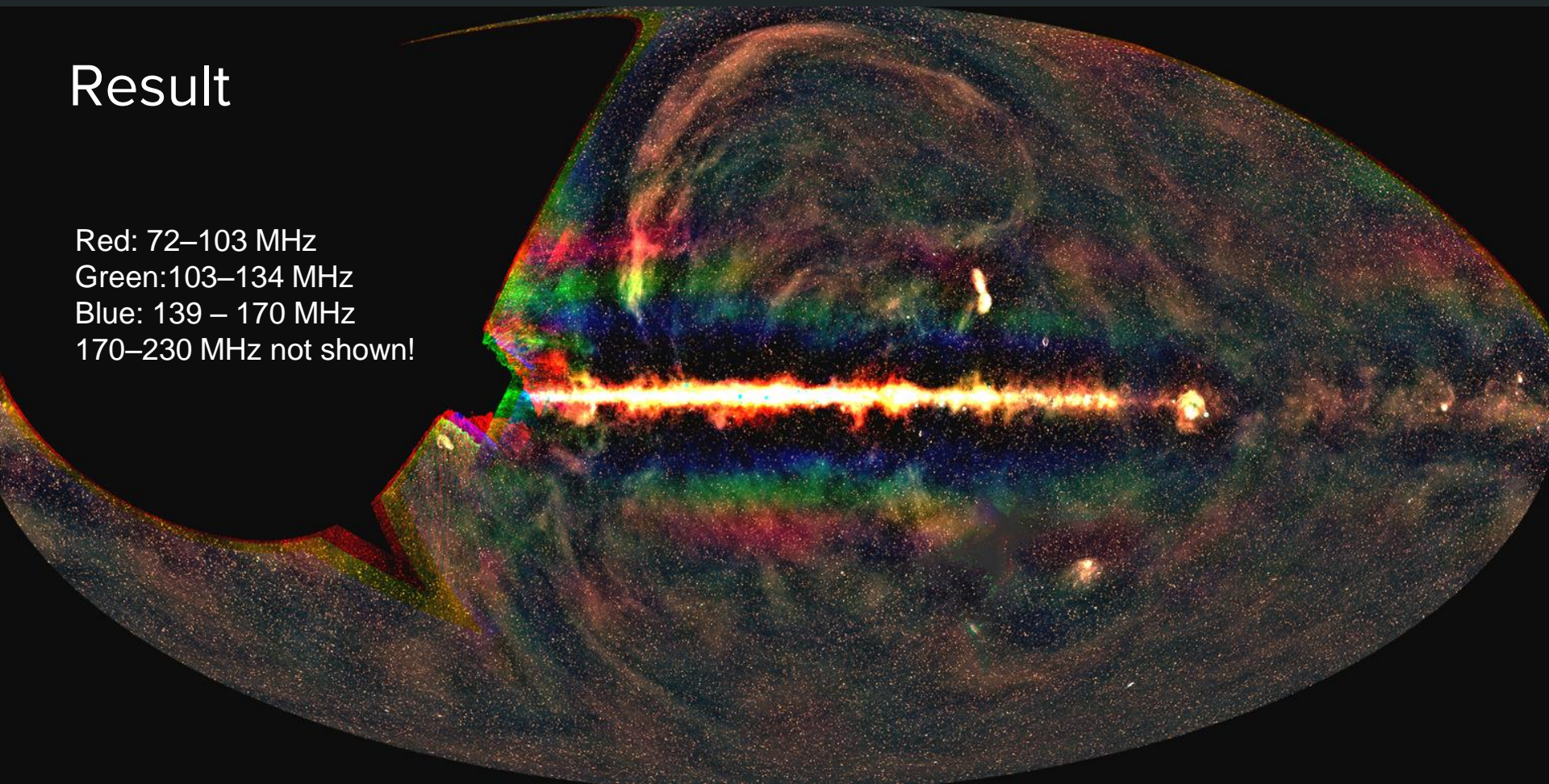
Ionospheric → warped images

Primary beam / flux → reference catalogues + symmetry

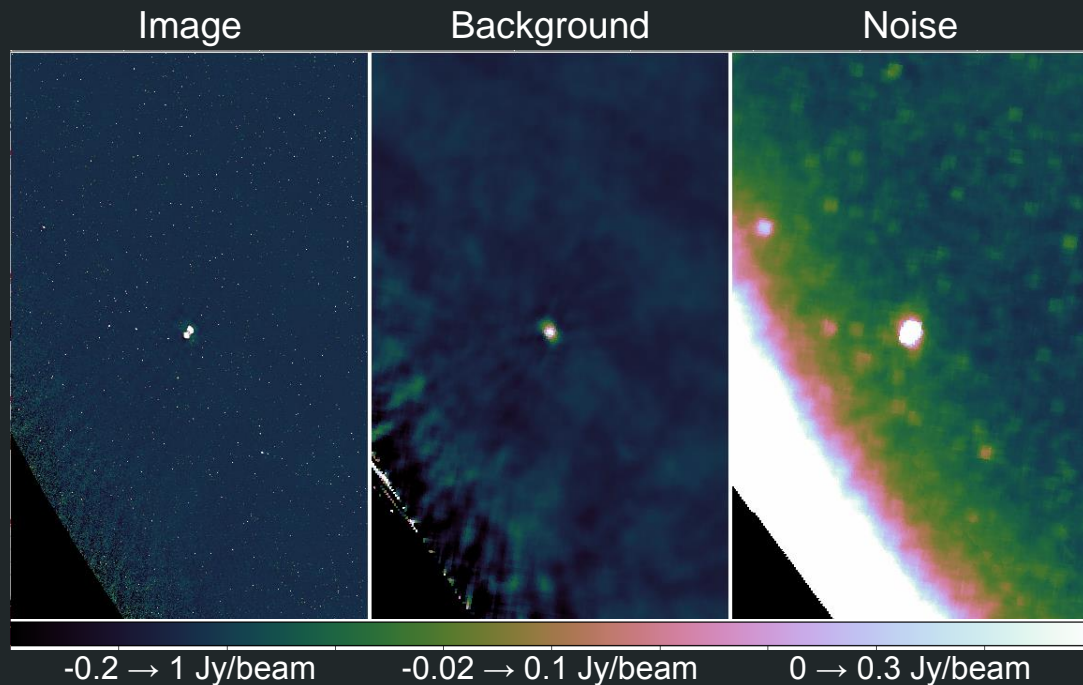


Result

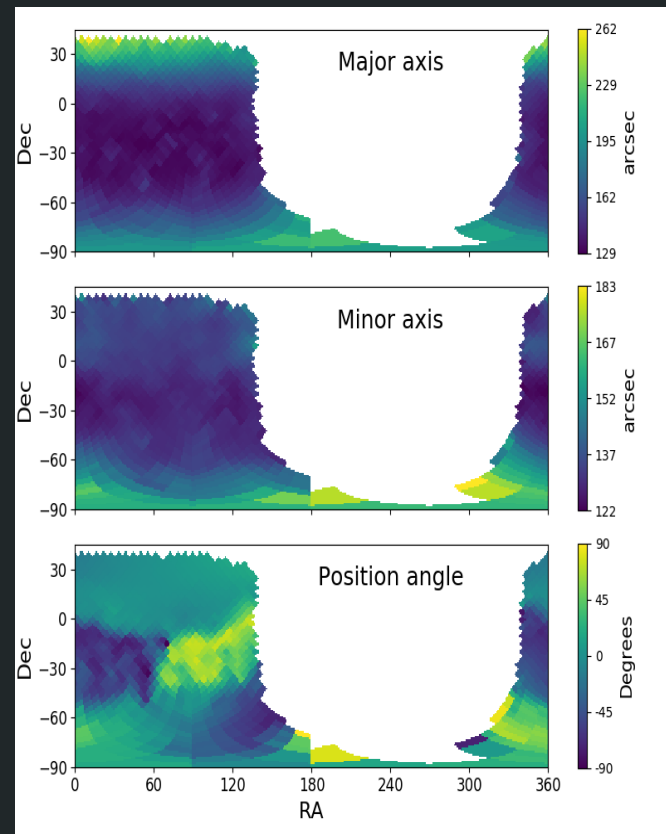
Red: 72–103 MHz
Green: 103–134 MHz
Blue: 139 – 170 MHz
170–230 MHz not shown!



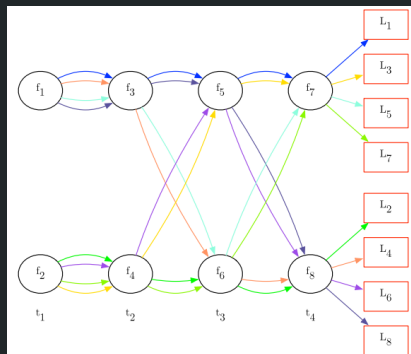
Catalogues at large FoV



Synthesized Beam

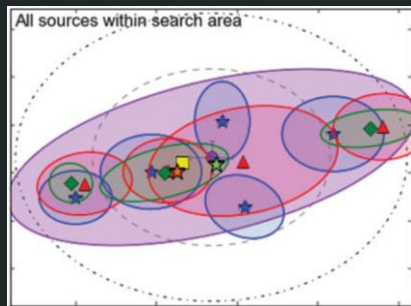


Catalogues at large $\Delta\nu$



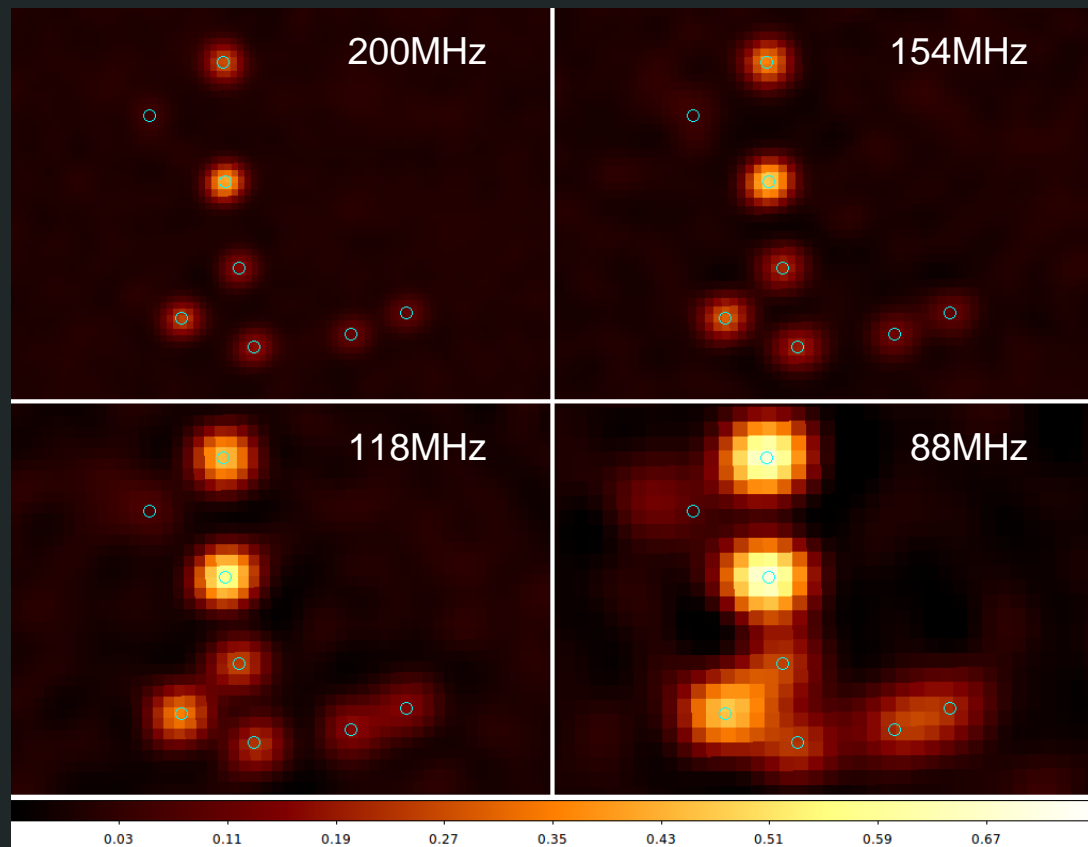
Catalogue
and X-
match?

Swinbank et al. 2015

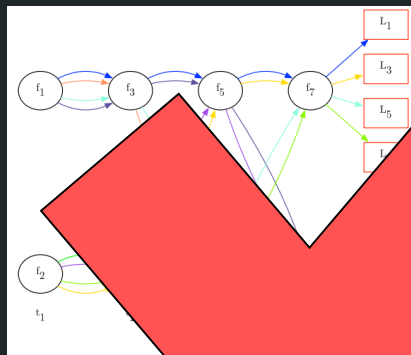


Hierarchical
association?

Line et al. 2017



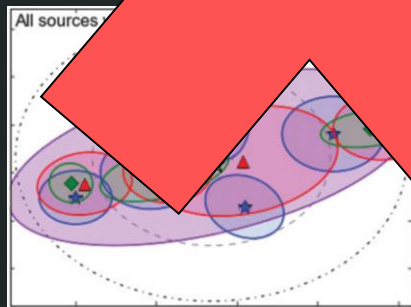
Catalogues at large Δv



catalogue
d X-
en?

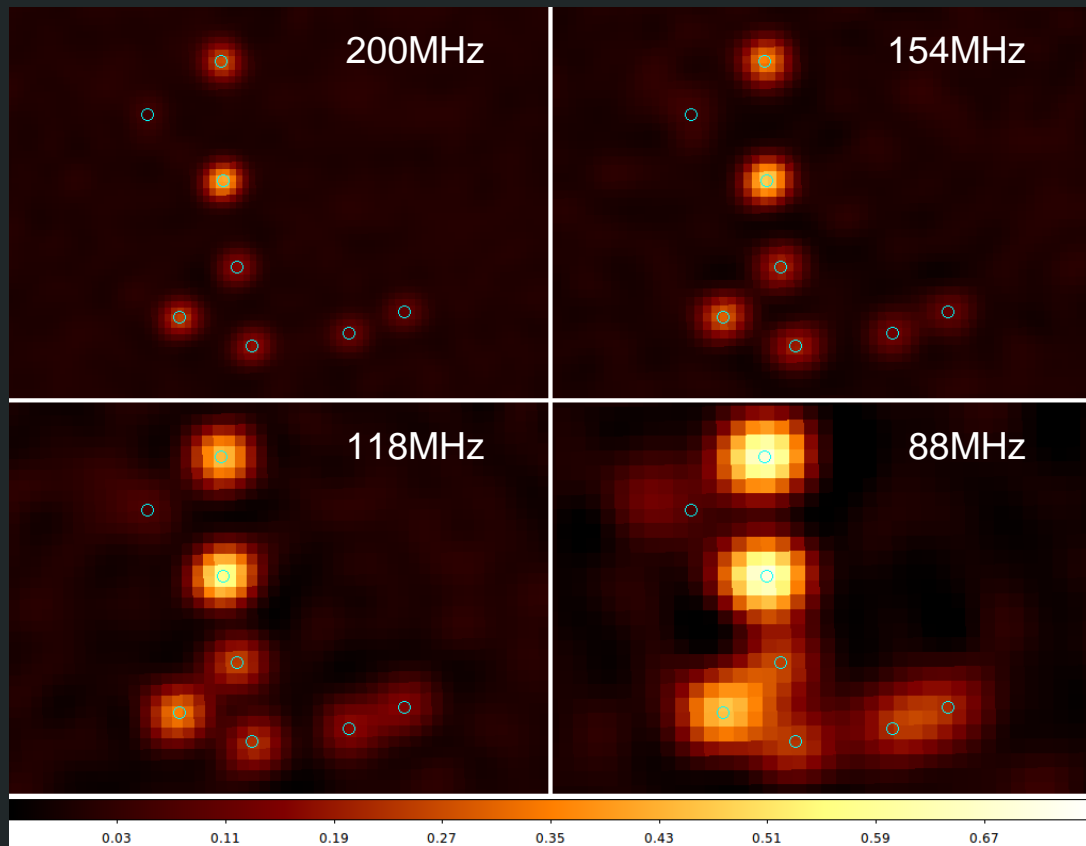
Swinbar

Too Hard!

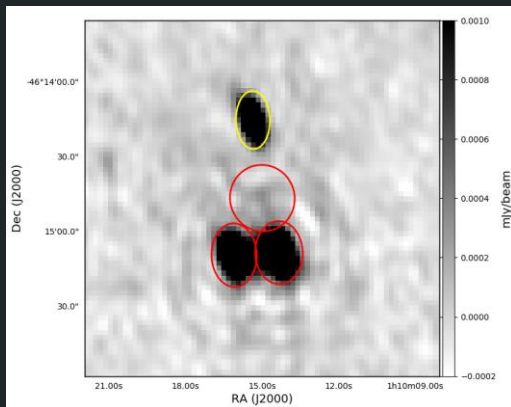


Architectural Association?

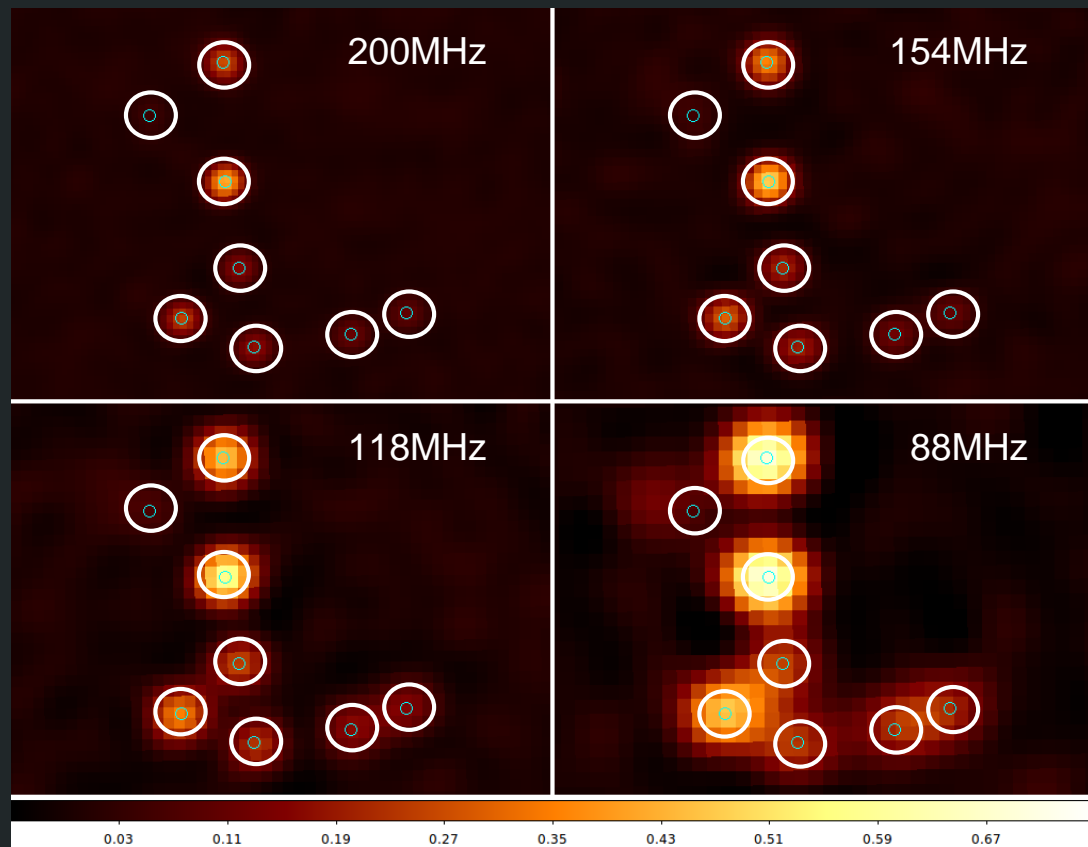
Line et al. 2017



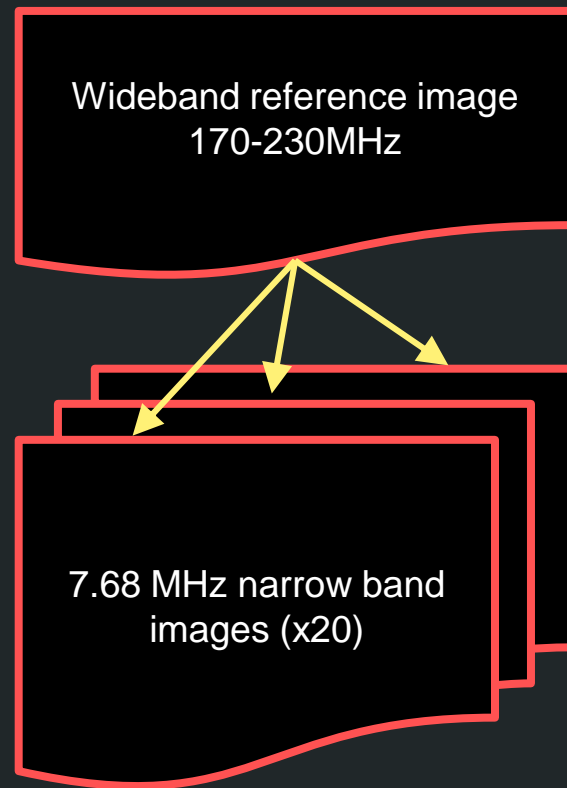
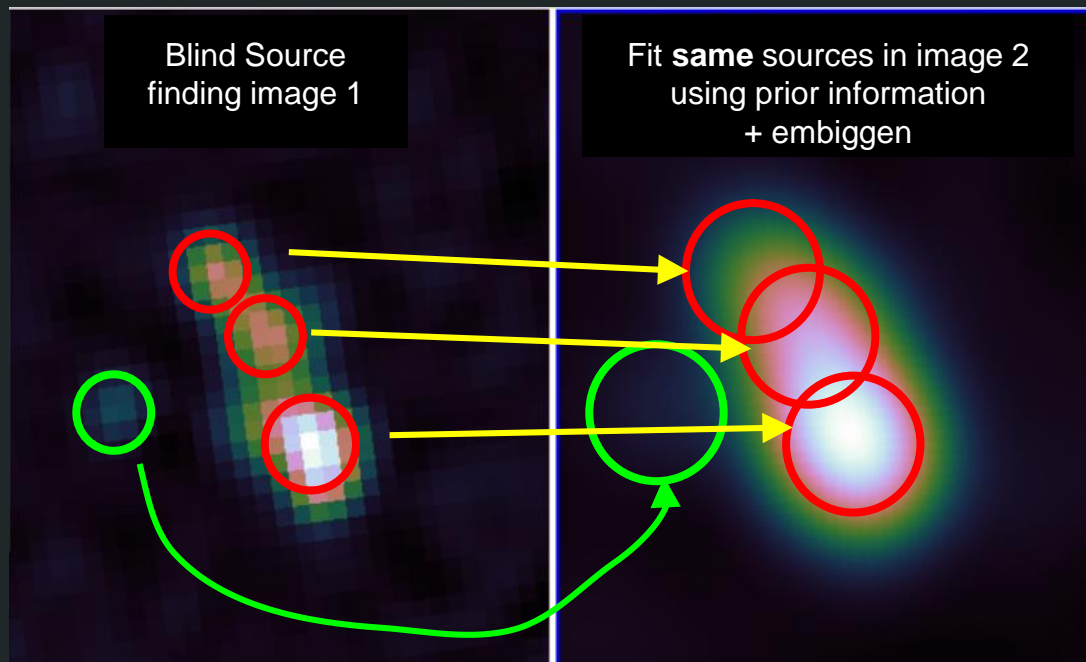
Catalogues at large $\Delta\nu$



Priorized fitting with Aegean
(Hancock et al. 2012)
(now also pyBDSF)



Catalogues at large $\Delta\nu$

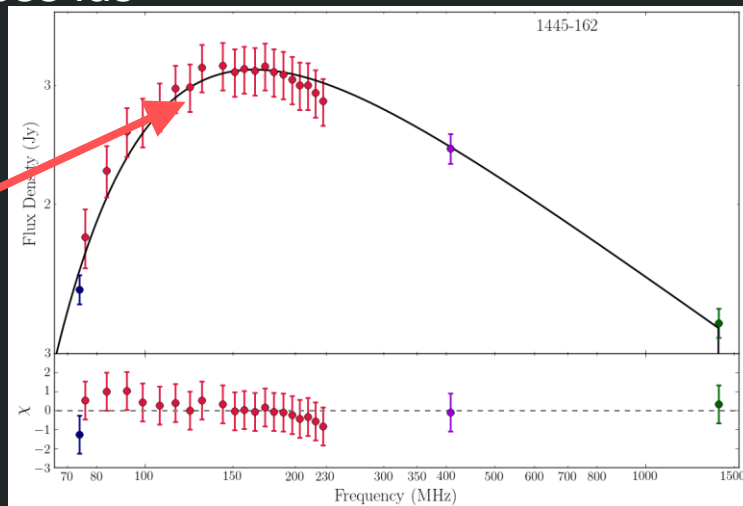


Catalogues at large $\Delta\nu$

Catalog contains

- **all** sources from deep image
- fluxes from each narrow band for each source
- sub-threshold fluxes
- ZERO false cross ids

GLEAM prioritized fits
at 20 frequencies



Callingham et al. 2017

Wideband reference image
170-230MHz

7.68 MHz narrow band
images (x20)

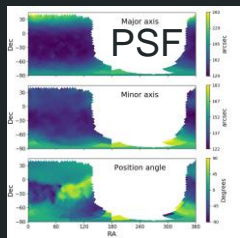
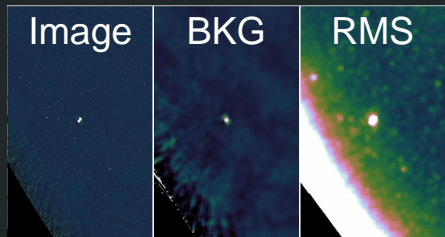
Source Finding Solution: Aegean

<https://github.com/PaulHancock/Aegean>



This repository

Search



Wideband
reference image
170-230MHz

7.68 MHz
narrow band
images (x20)

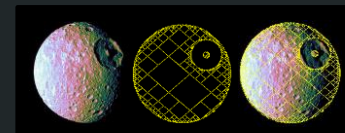
Aegean
Find sources
Characterise sources
Sources can be
gaussians or
blobs



BANE
Characterise
background
Characterise **noise**
Do it **right**
Do it **fast**



MIMAS
Describe regions
Combine regions
Mask images
Constrain Aegean
Write **MOC** files



Summary (what **we** did)

Flux calibration:

- Ensemble solution to bootstrap from other frequencies
- Rely on the N/S symmetry of the MWA beam

Ionospheric calibration

- Image warping based on positions from other frequencies

Catalogues at large FoV and $\Delta\nu$

- Create 60MHz deep master image
- Create background/noise/PSF maps
- Priorized fitting in each of 20 narrow band images
- SEDs for all $\sim 300k$ sources.

Summary (what **you** should do)

Flux calibration:

- Use GLEAM fluxes
- (Invest in a good beam model)

Catalogues at large FoV and $\Delta\nu$

- Use Aegean and BANE

Ionospheric calibration

- Use GLEAM positions