

A spectacular view of the Toothbrush: filaments and inhomogeneous magnetic fields

Kamlesh Rajpurohit

Thüringer Landessternwarte Tautenburg

Matthias Hoelt and Reinout J. van Weeren



**The Broad Impact of Low Frequency Observing
Bologna, June 21, 2017**

Large scale diffuse radio emission in 1RX J0603.3+4214

- Radio observations by van Weeren +2012 :

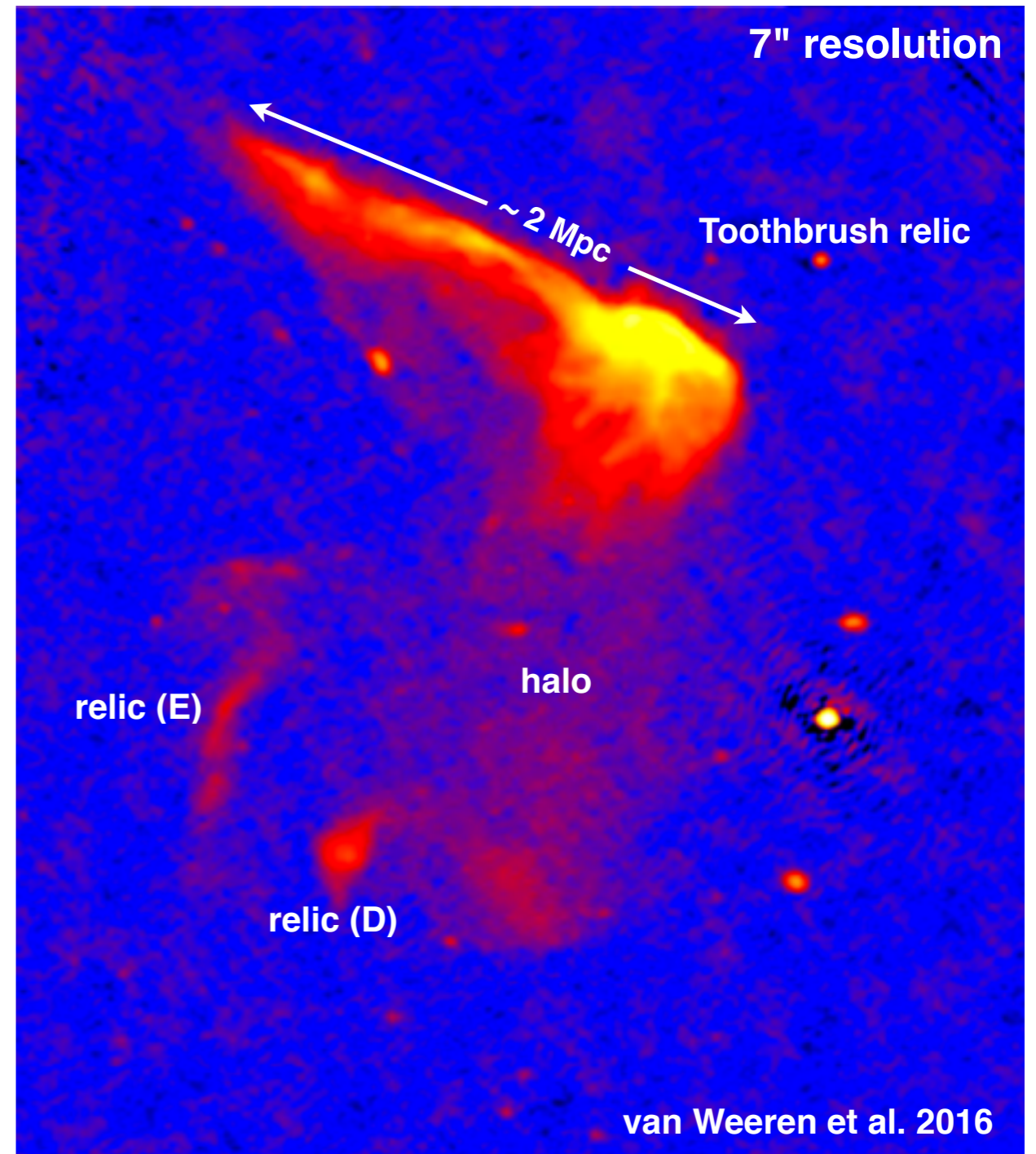
- cluster host ~ 2 Mpc relic
- additional fainter relics and halo
- $z = 0.225$

- Toothbrush relic:

- steep spectrum, $\alpha = -1.15 \pm 0.01$

↓
Mach number (M) ~ 3.78

- clear spectral index gradient towards cluster centre
- strongly polarized (up to 60%) at 4.9 GHz



LOFAR image (150 MHz)

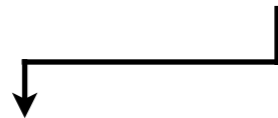
Large scale diffuse radio emission in 1RX J0603.3+4214

- Radio observations by van Weeren +2012 :

- cluster host ~ 2 Mpc relic
- additional fainter relics and halo
- $z = 0.225$

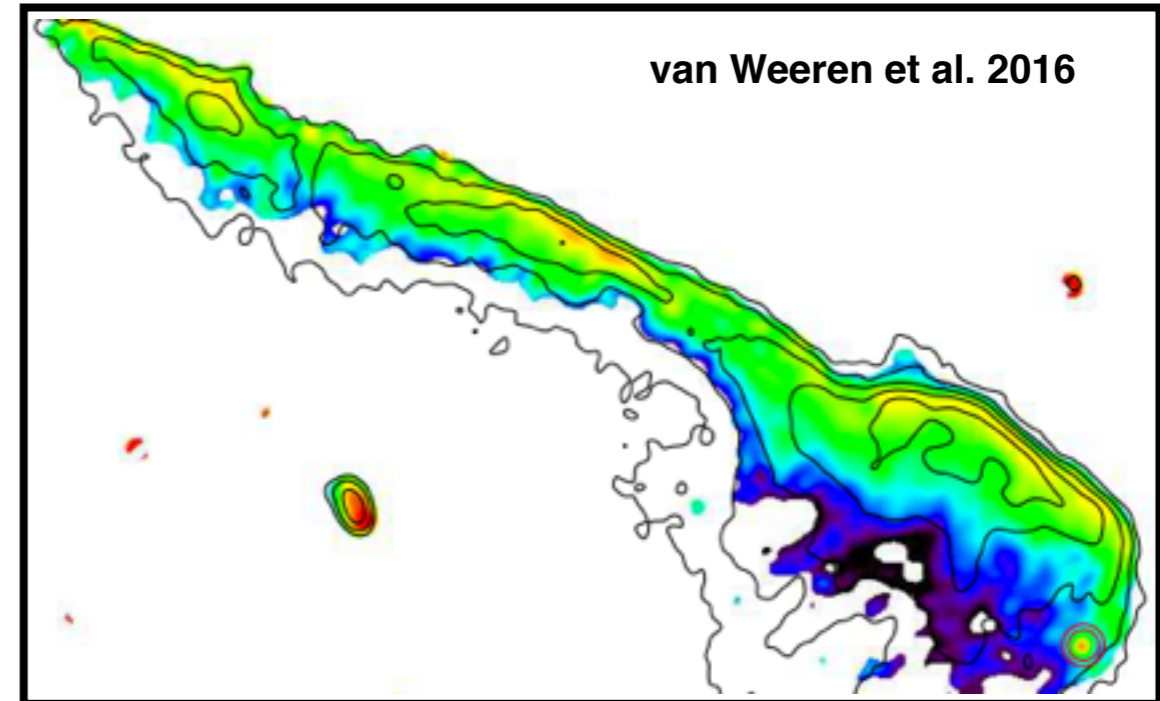
- Toothbrush relic:

- steep spectrum, $\alpha = -1.15 \pm 0.01$

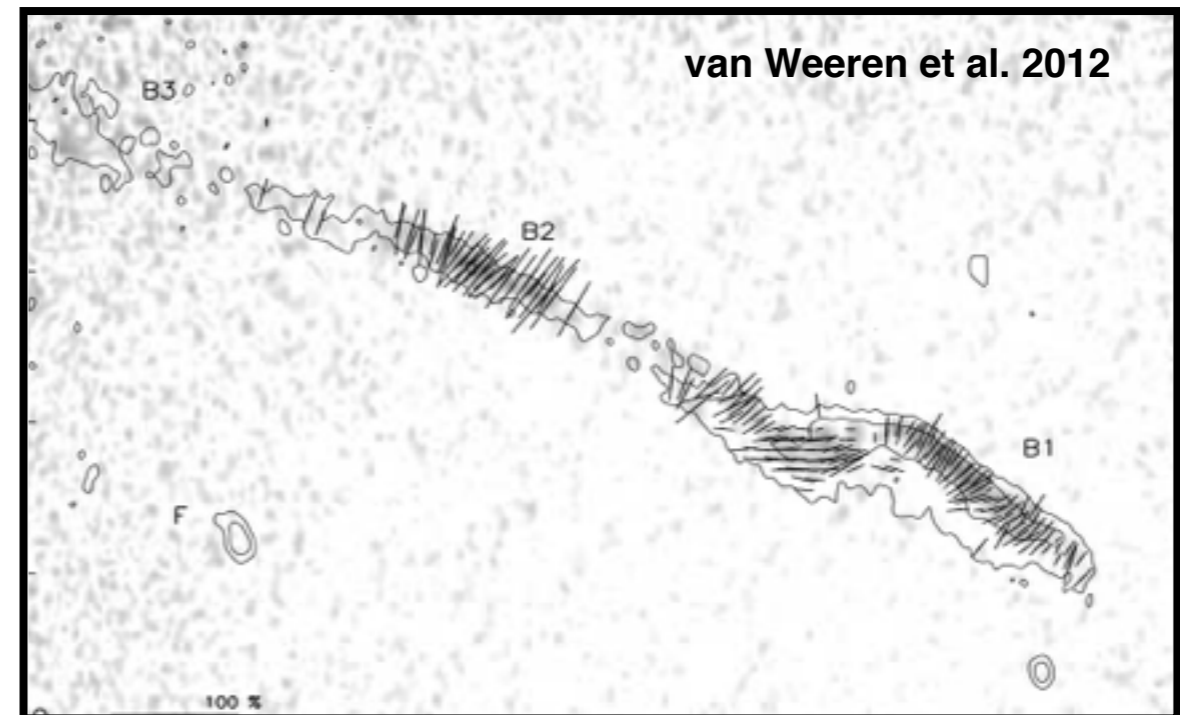


Mach number (M) ~ 3.78

- clear spectral index gradient towards cluster centre
- strongly polarized (up to 60%) at 4.9 GHz

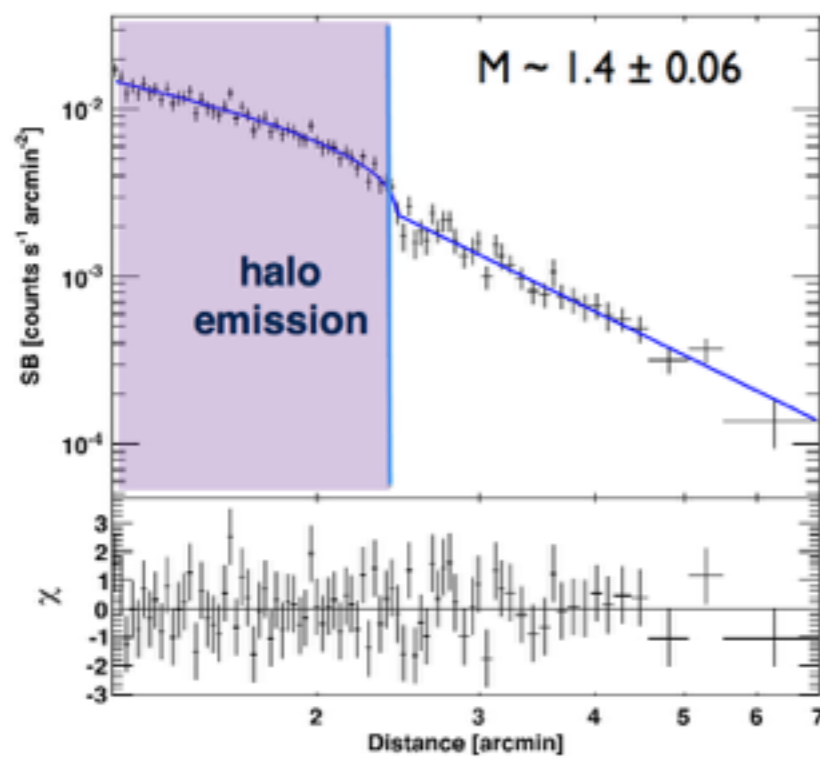
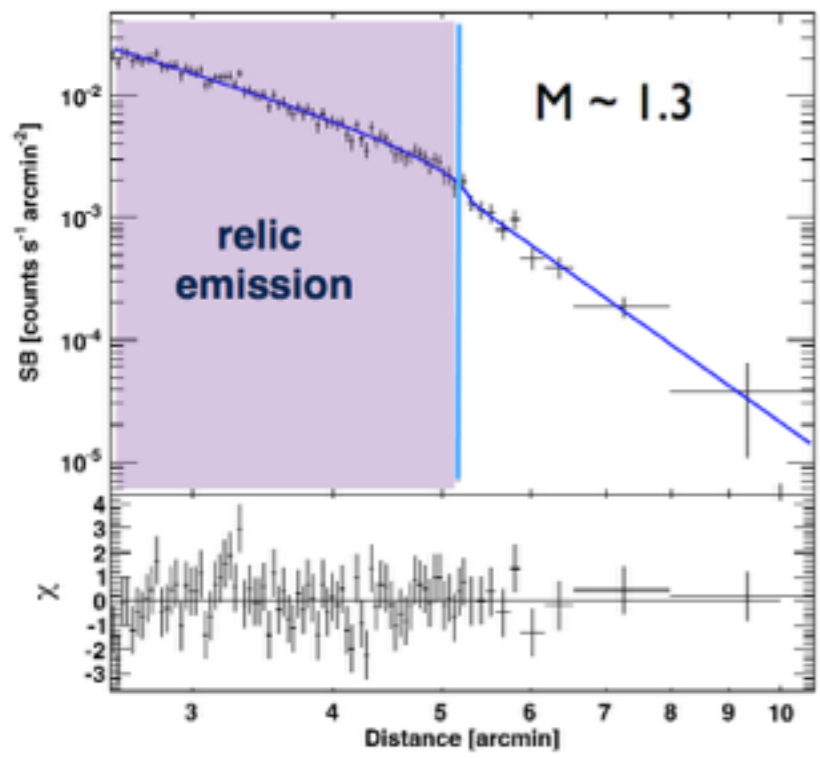


Low frequency spectral index map (150 - 610 MHz)

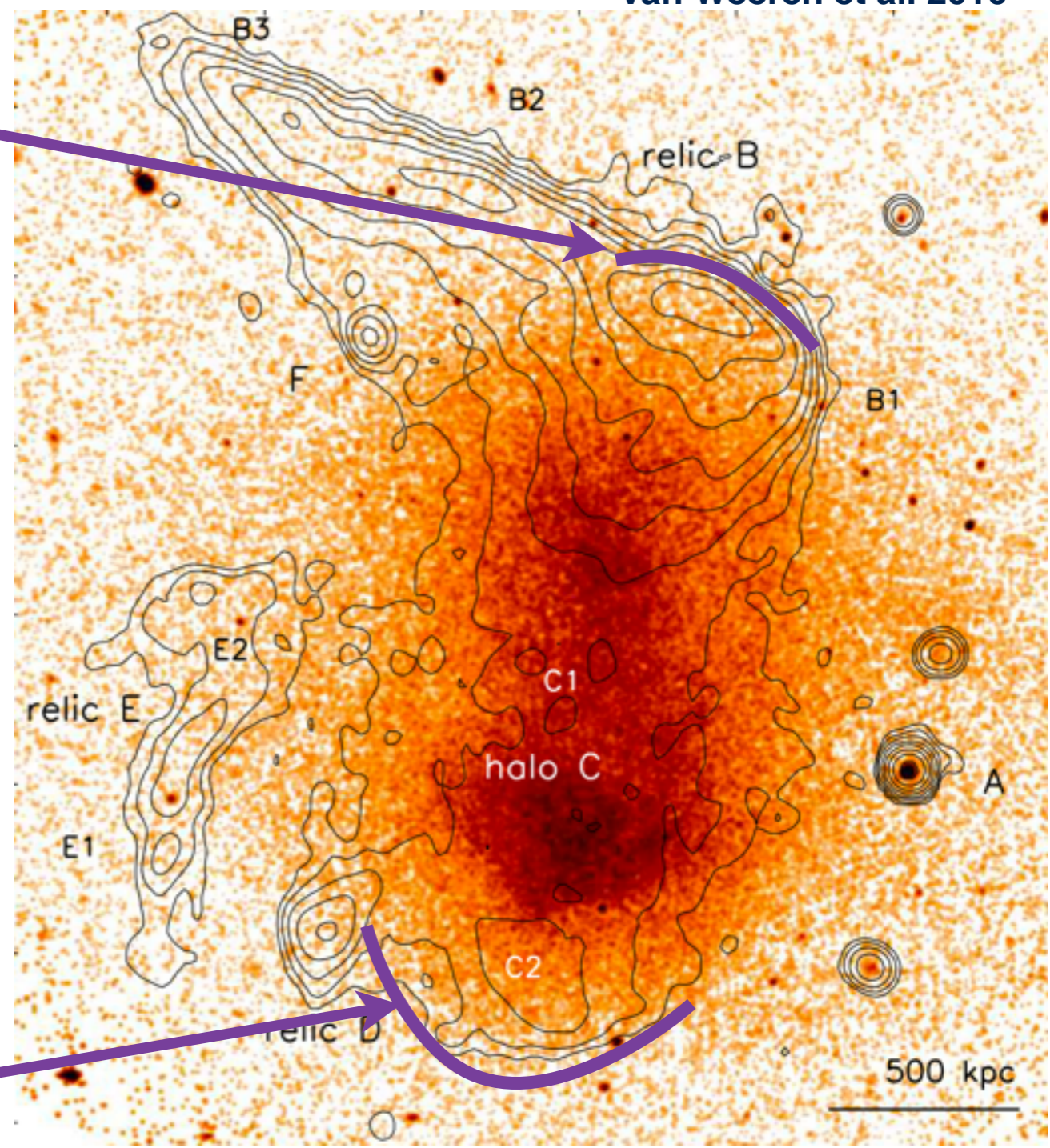


E-vectors distribution at 4.9 GHz

X-ray analysis: weak shock at northern edge

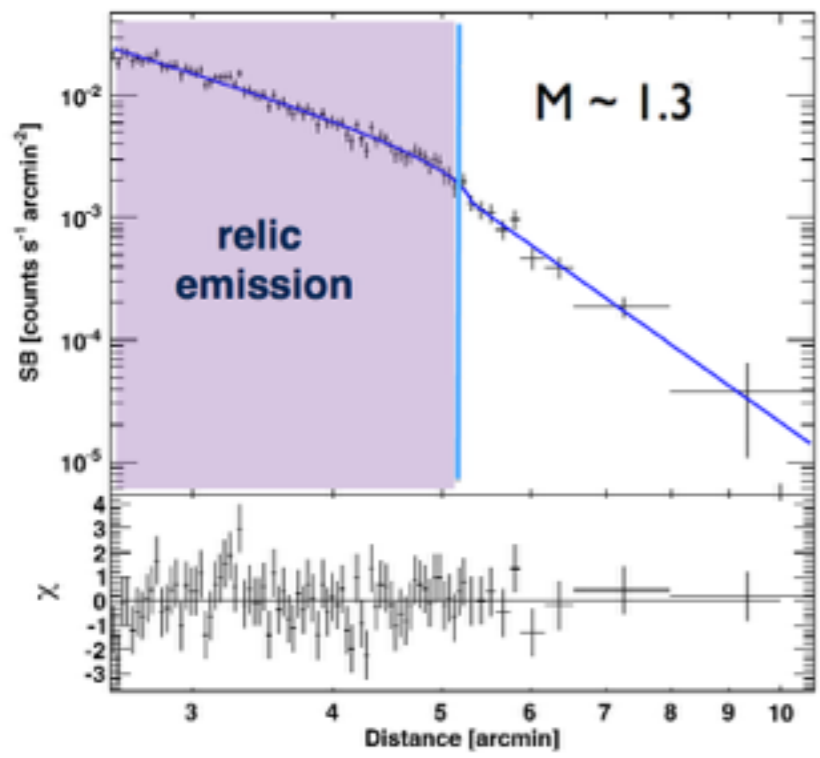


van Weeren et al. 2016

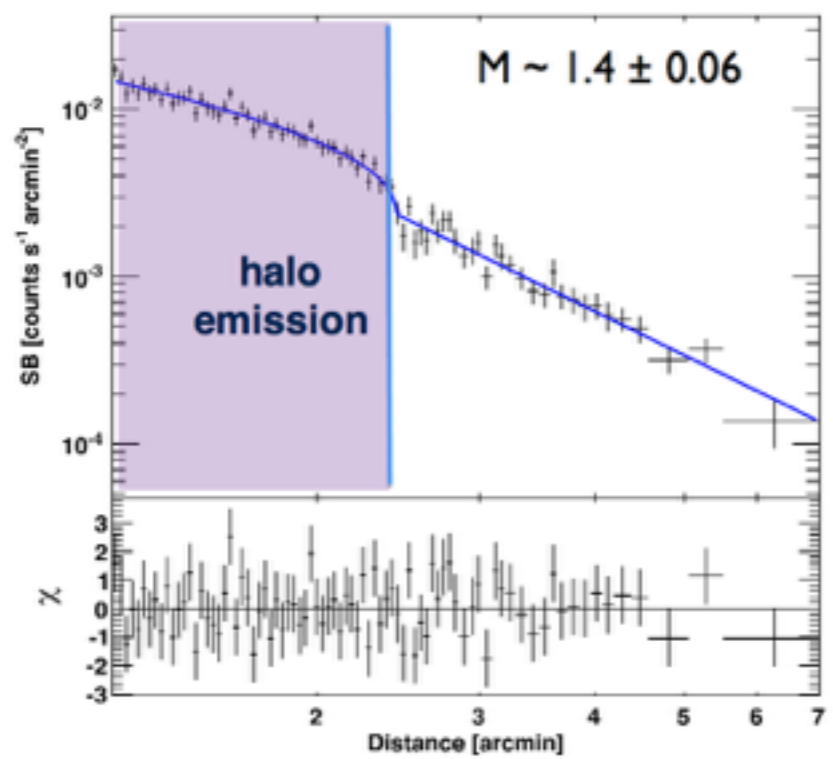


Chandra X-ray image

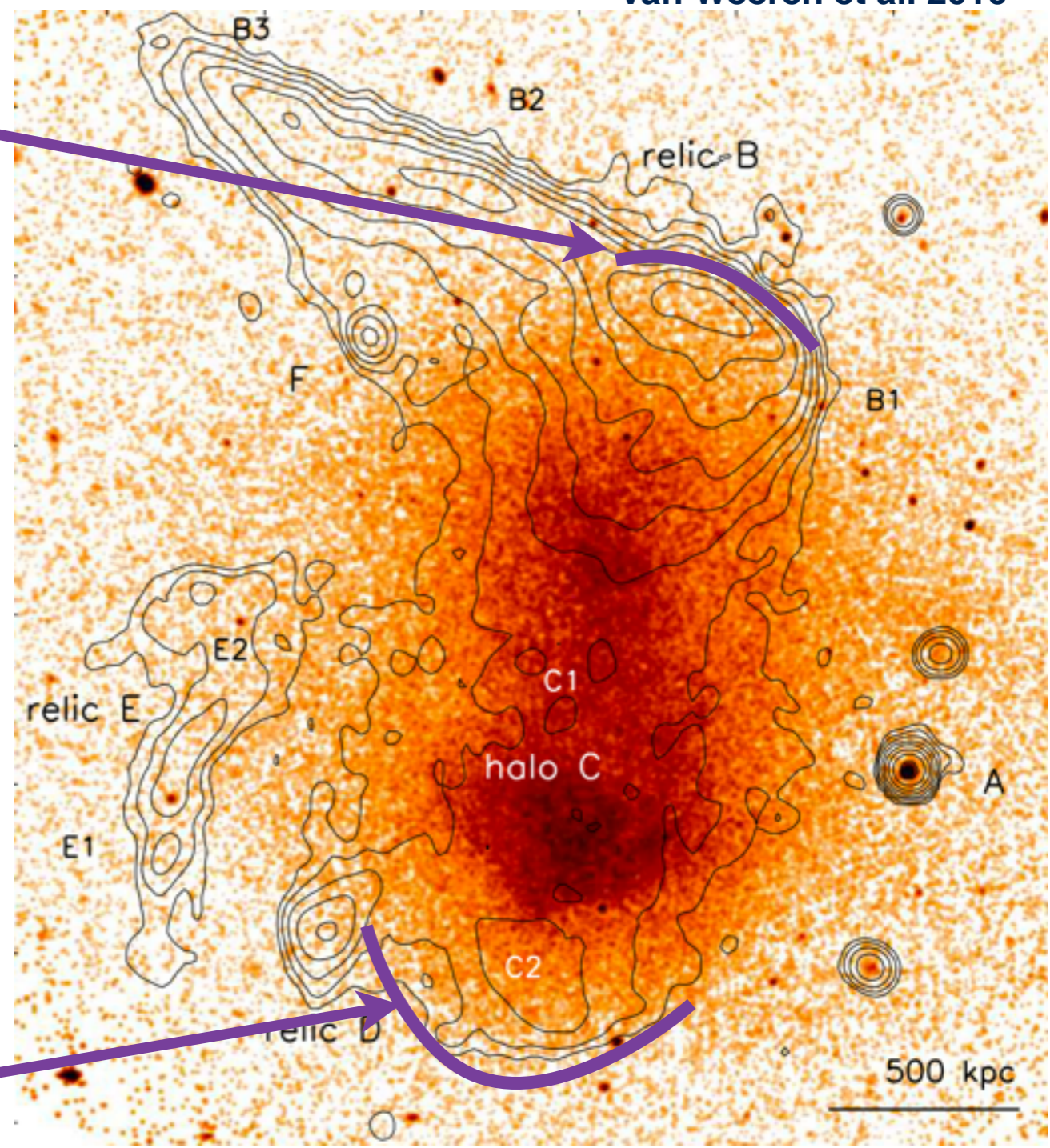
X-ray analysis: weak shock at northern edge



$M_{x\text{-ray}} \sim 1.3$ but $M_{\text{radio}} \sim 3.78$



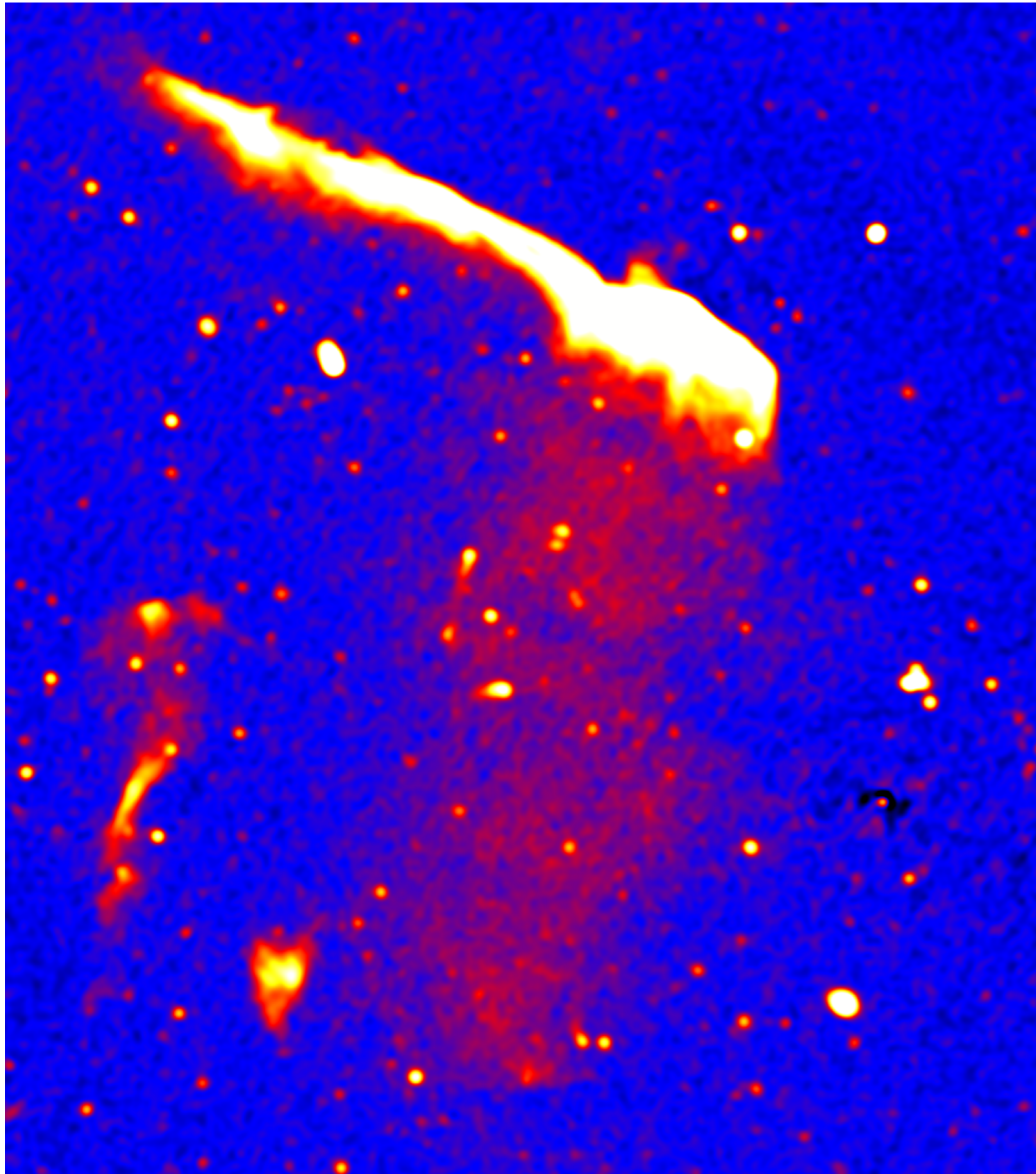
van Weeren et al. 2016



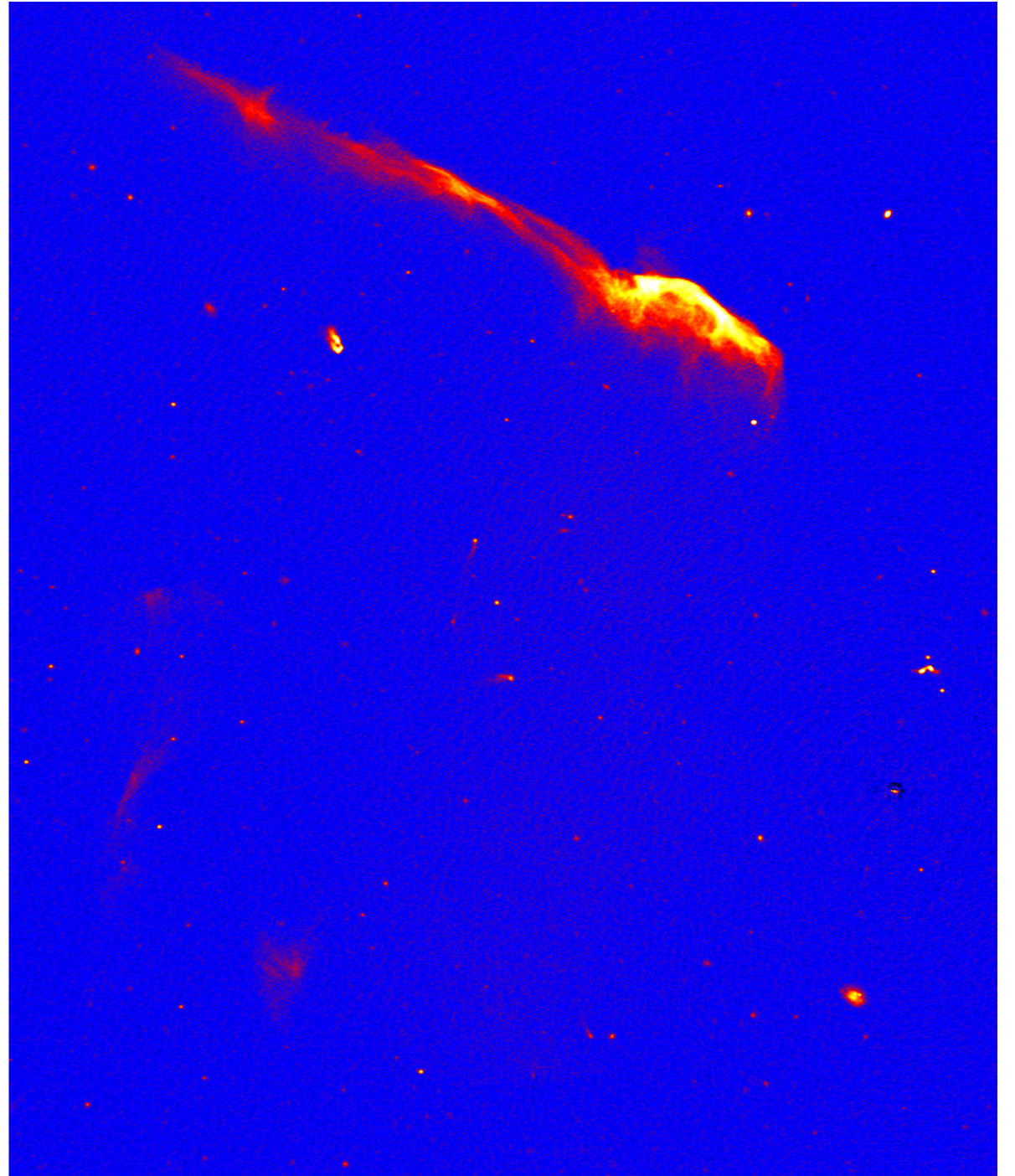
Chandra X-ray image

JVLA combined A+B+C+D array images

7" resolution, rms=9 μ Jy, frequency=1.5 GHz

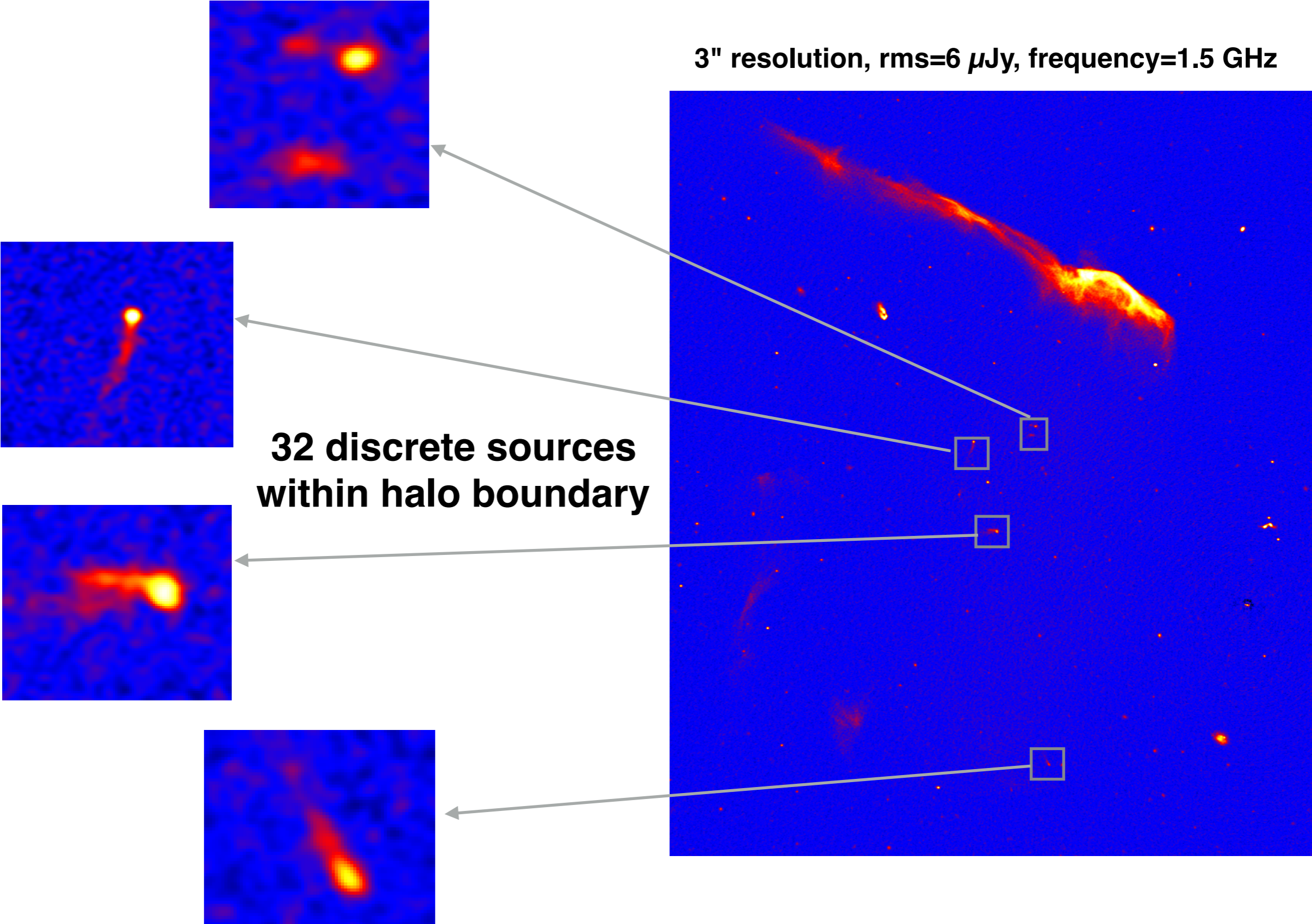


3" resolution, rms=6 μ Jy, frequency=1.5 GHz



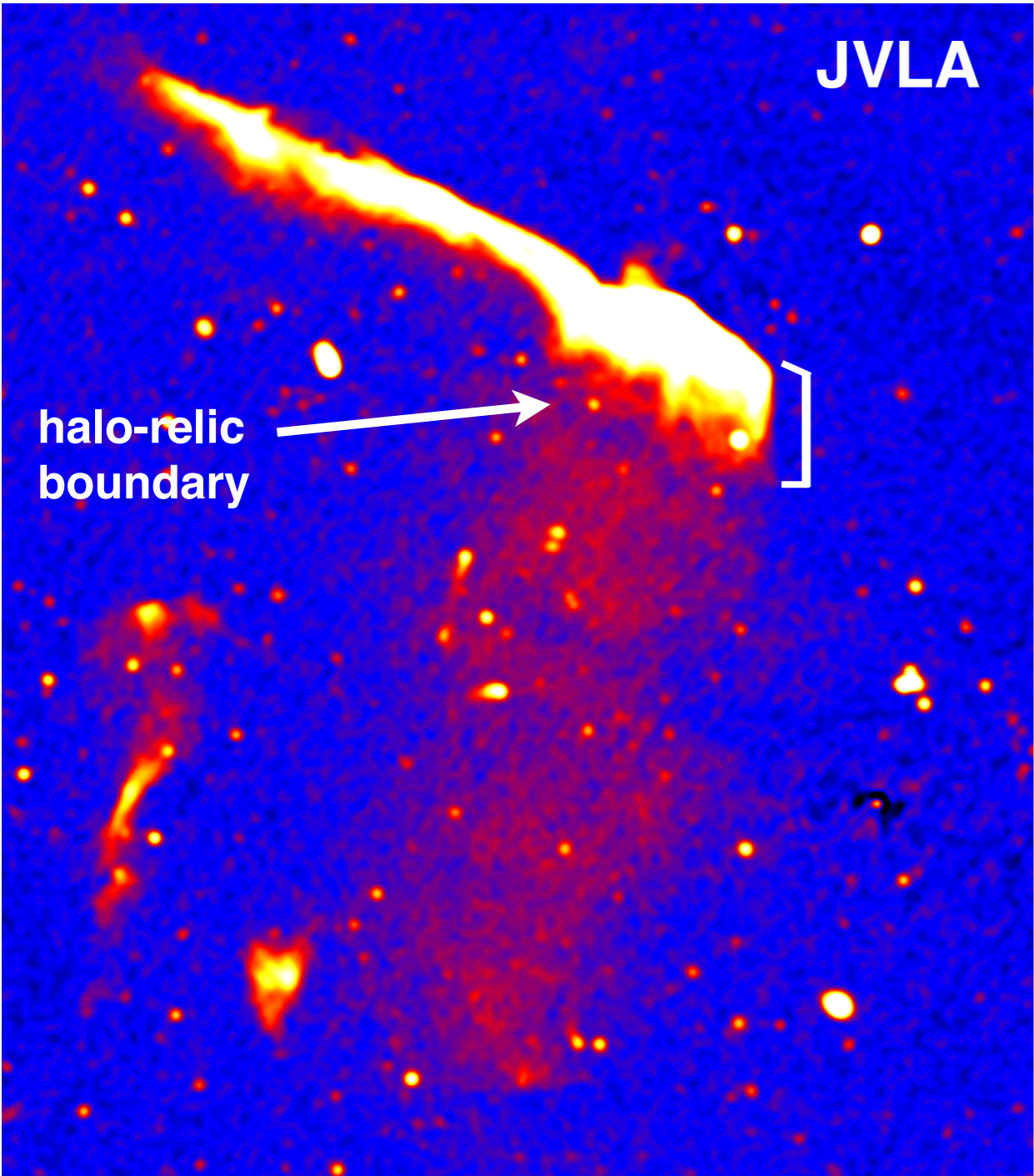
Detection of several head-tail radio galaxies

3" resolution, rms=6 μ Jy, frequency=1.5 GHz

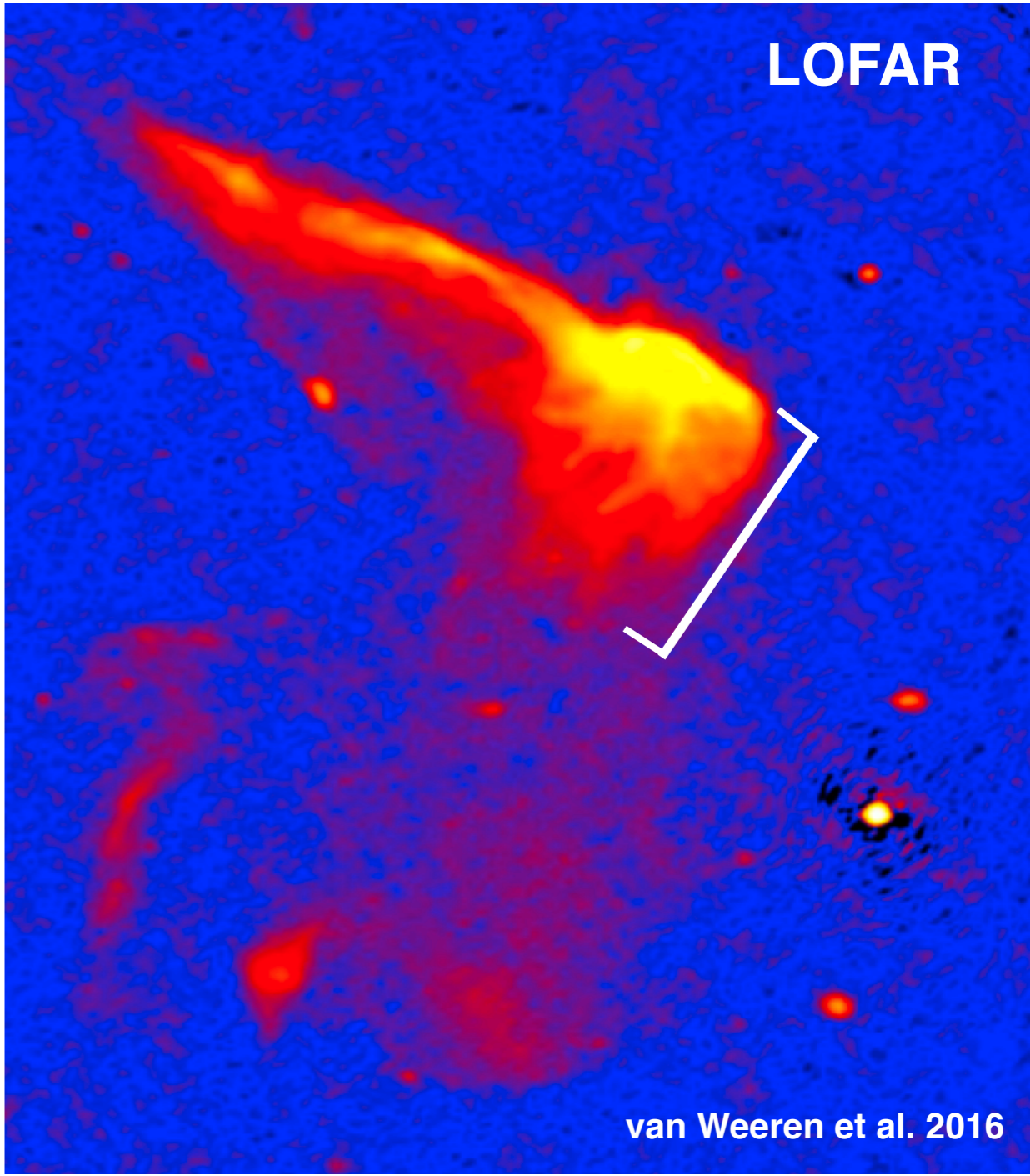


Comparison: Brush is extended at 150 MHz

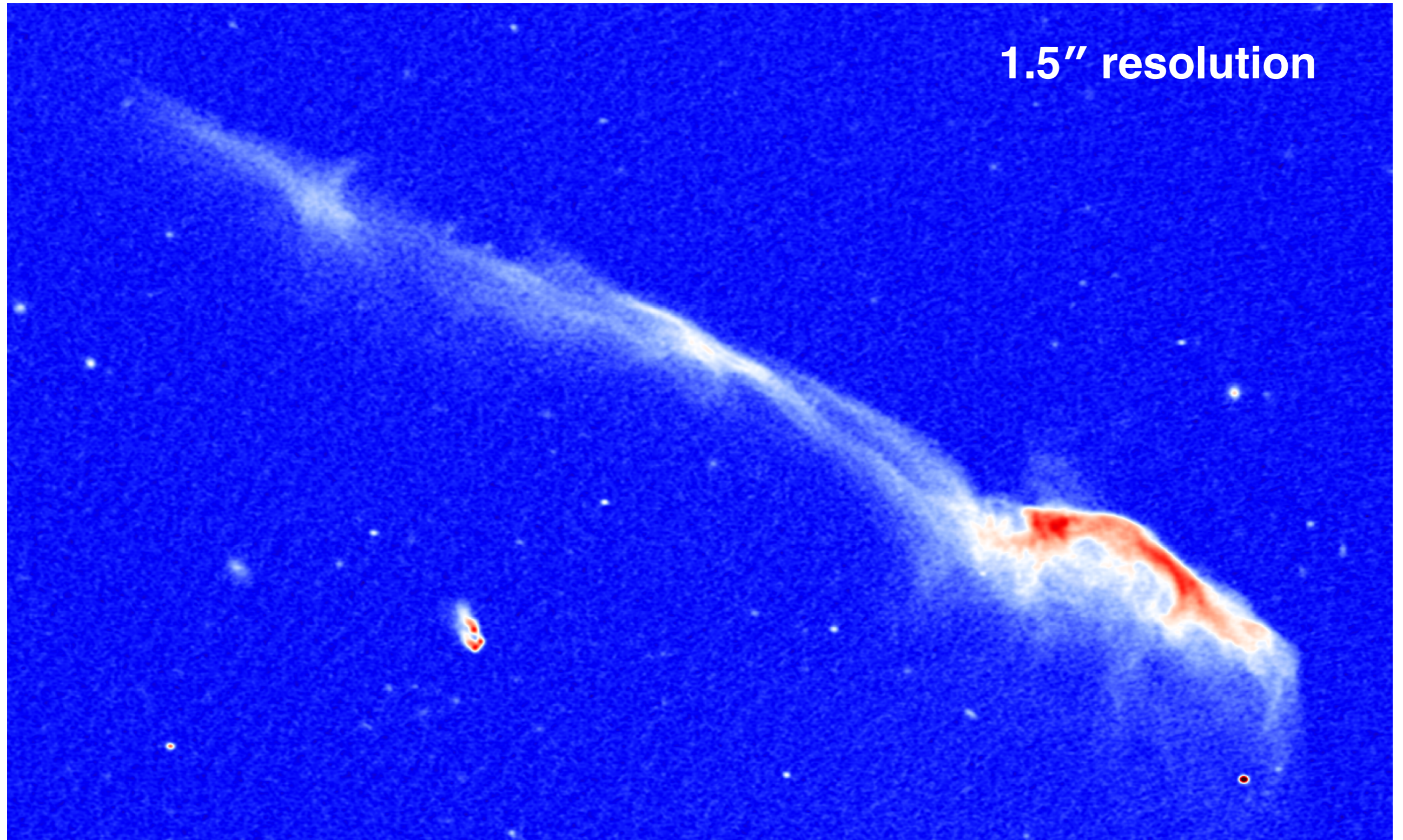
7" resolution, rms=9 μ Jy, frequency=1.5 GHz



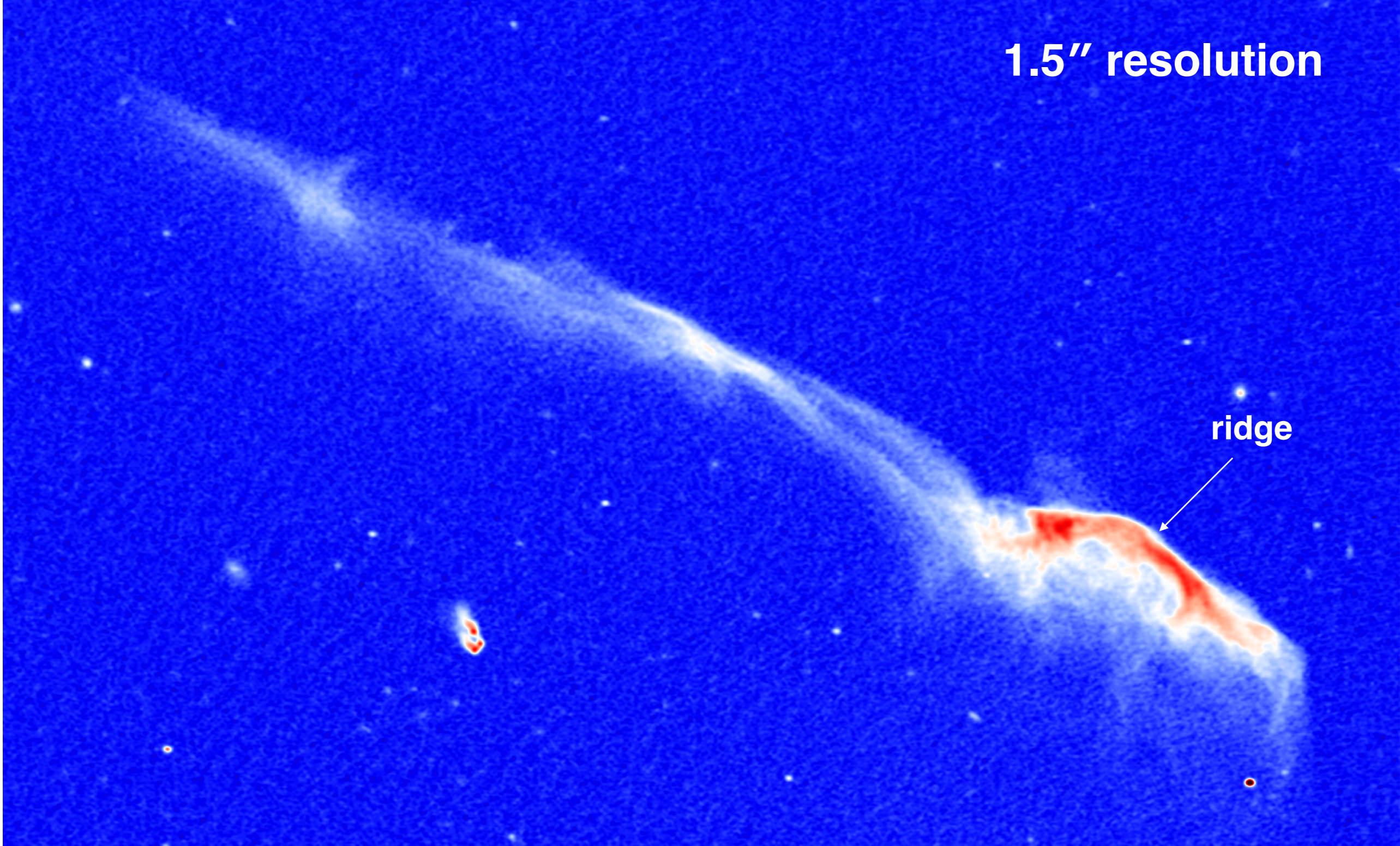
7" resolution, rms=93 μ Jy, frequency=150 MHz



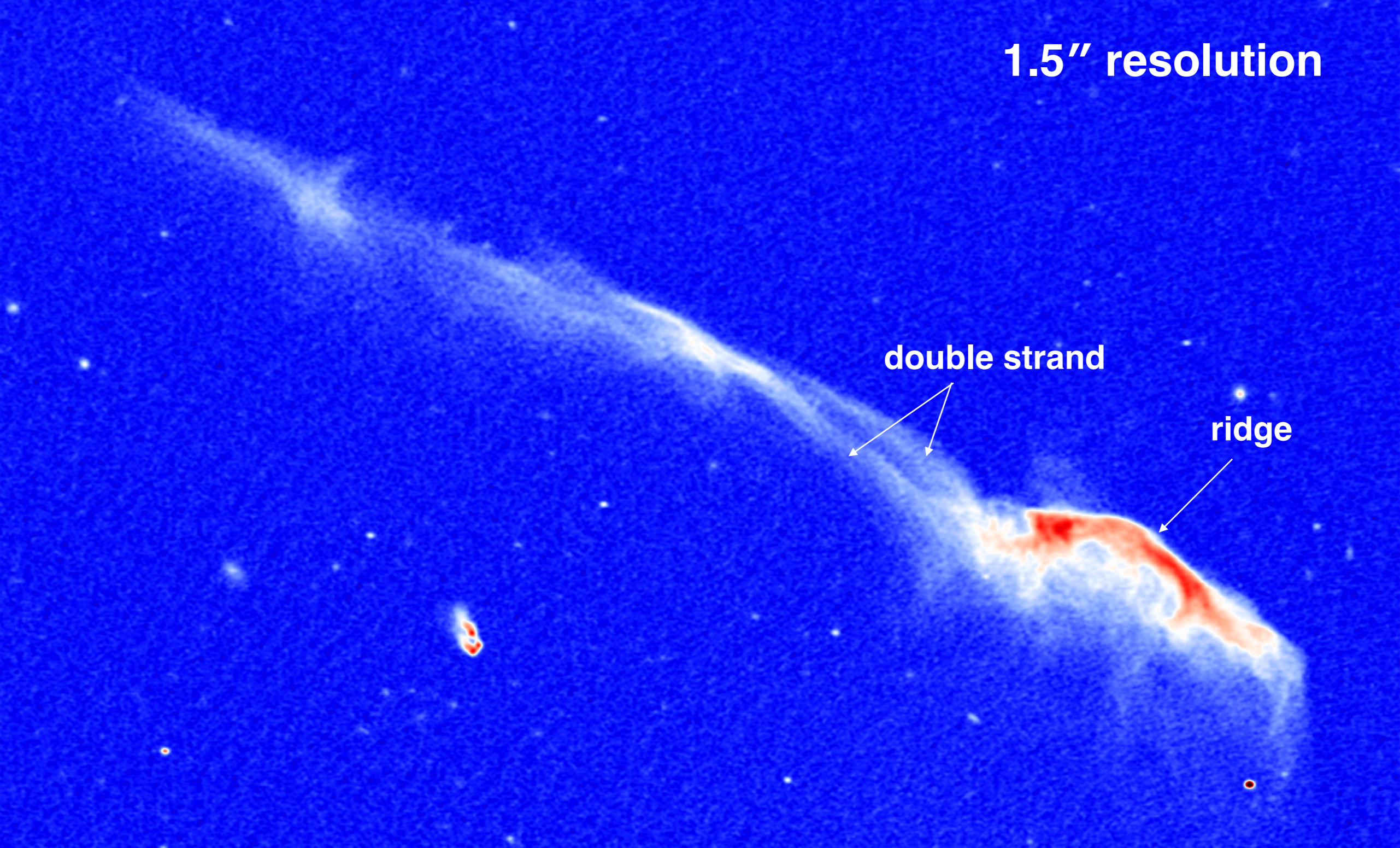
Toothbrush: enigmatic filamentary structures



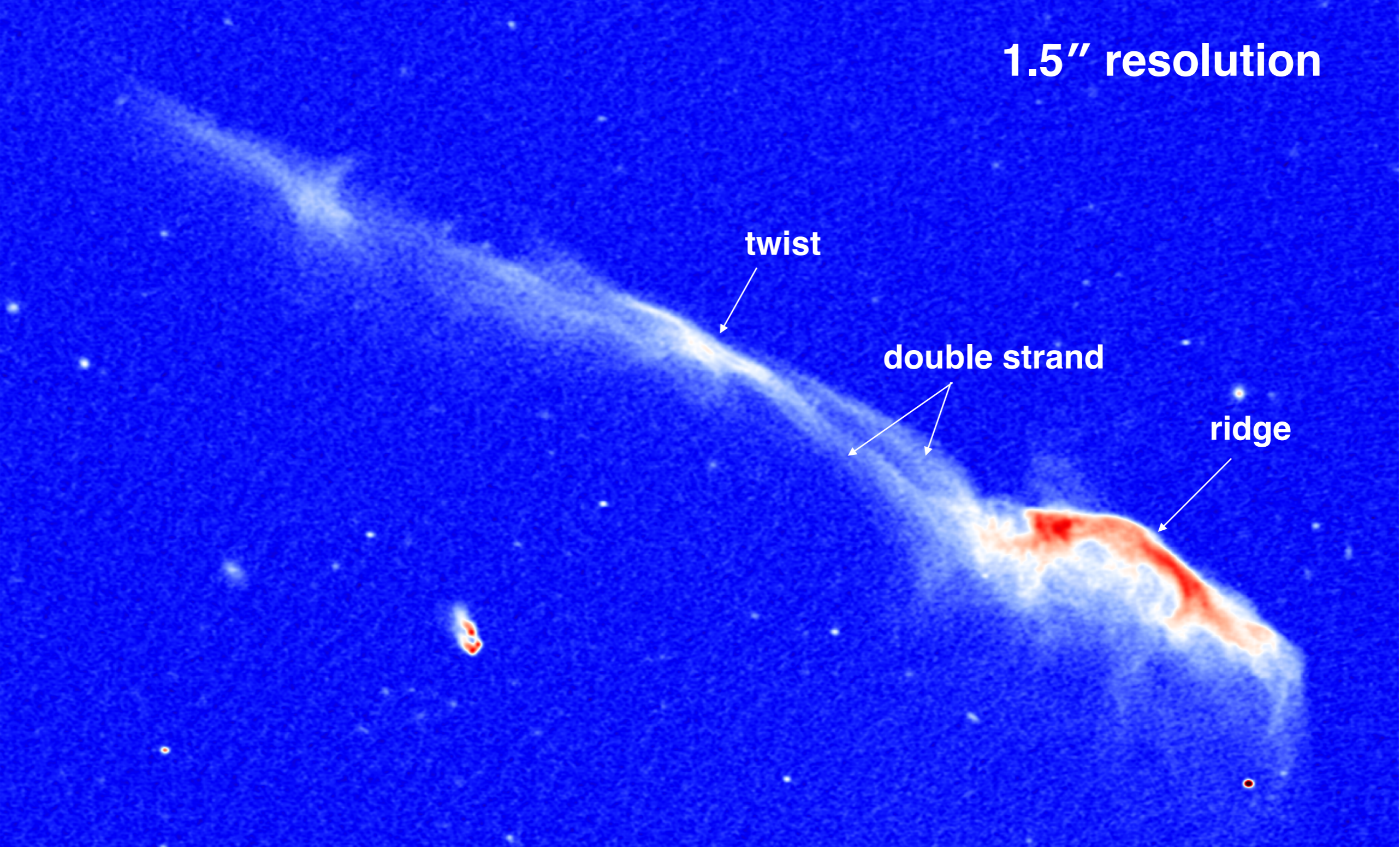
Toothbrush: enigmatic filamentary structures



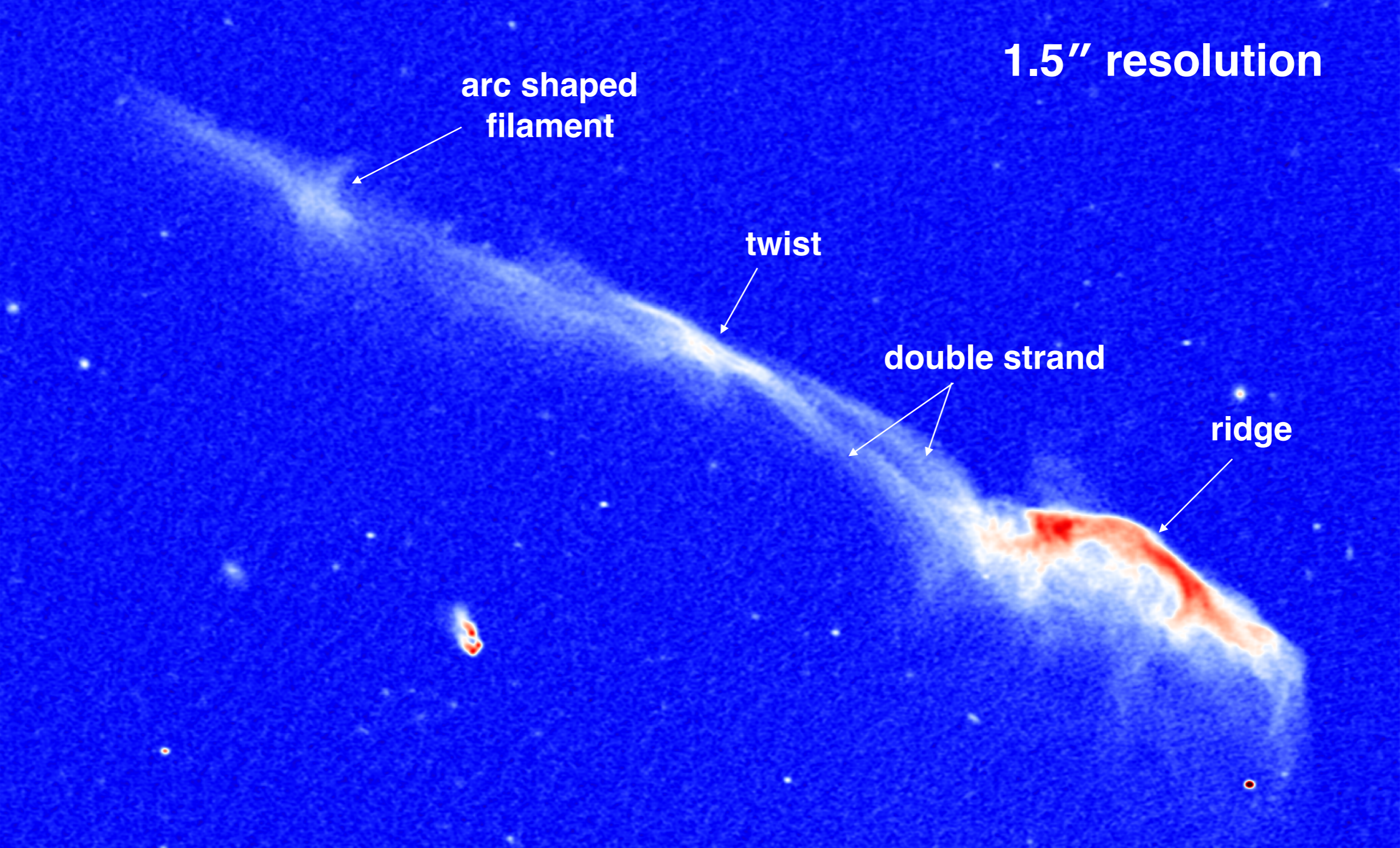
Toothbrush: enigmatic filamentary structures



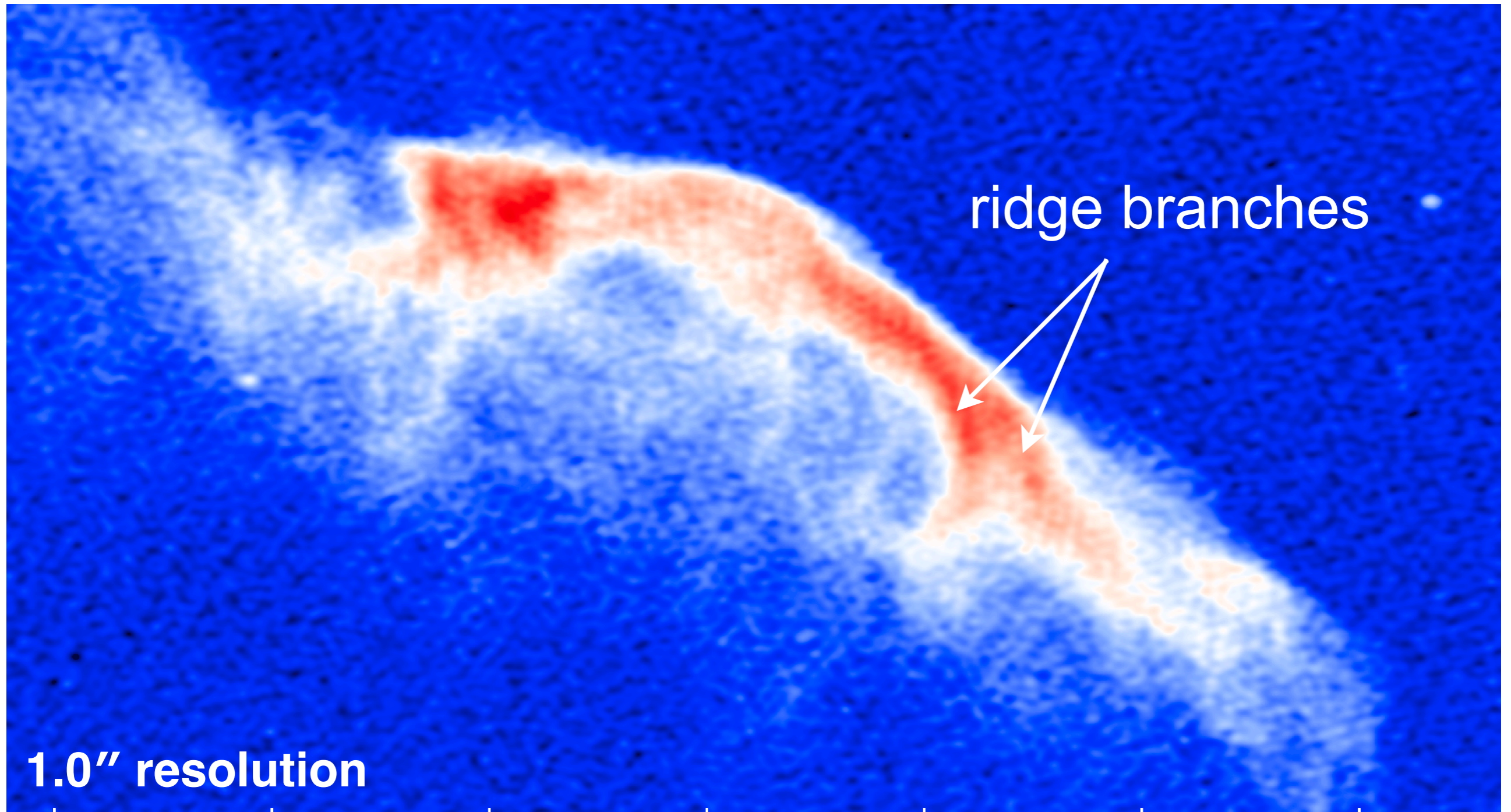
Toothbrush: enigmatic filamentary structures



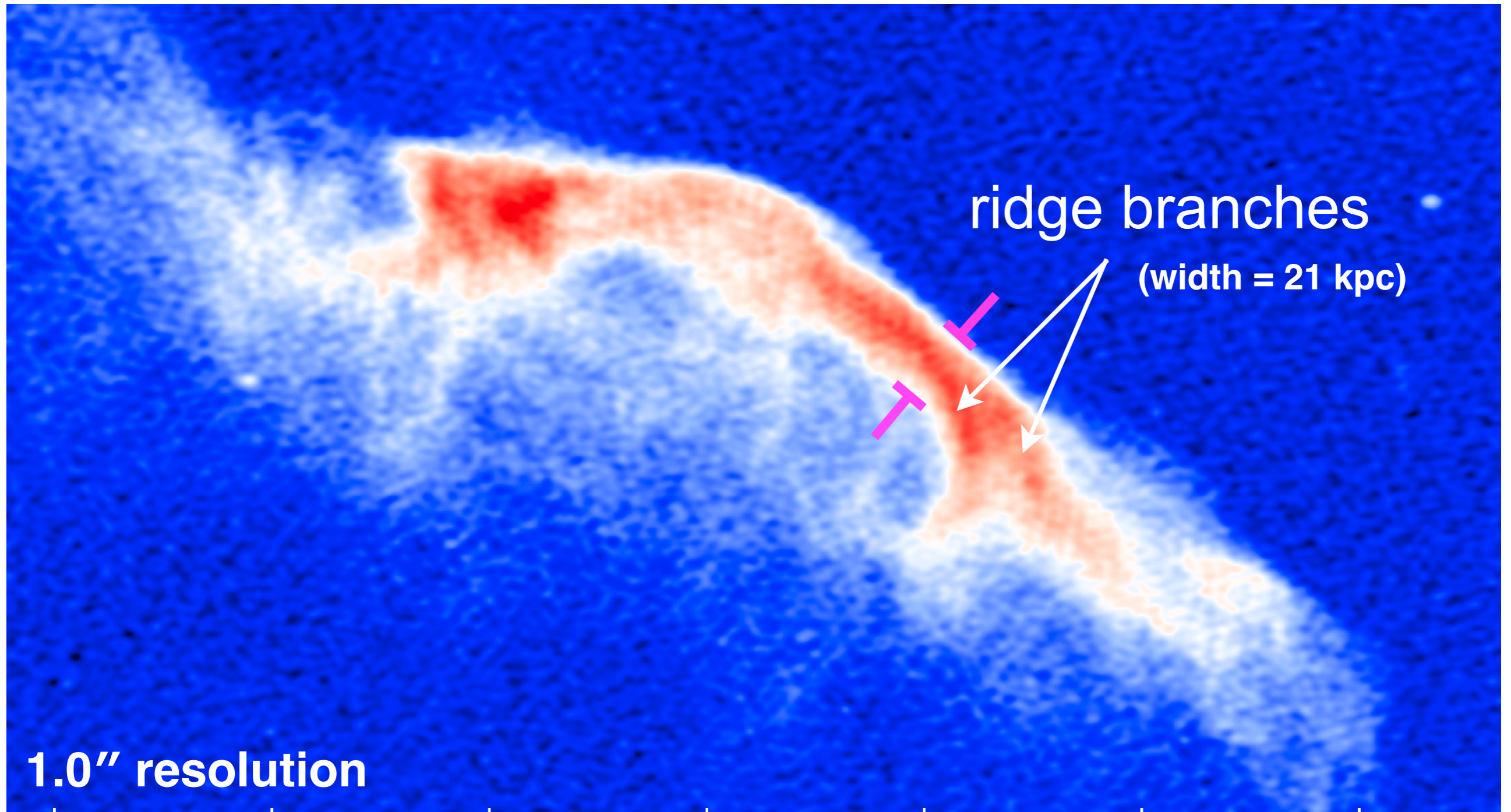
Toothbrush: enigmatic filamentary structures



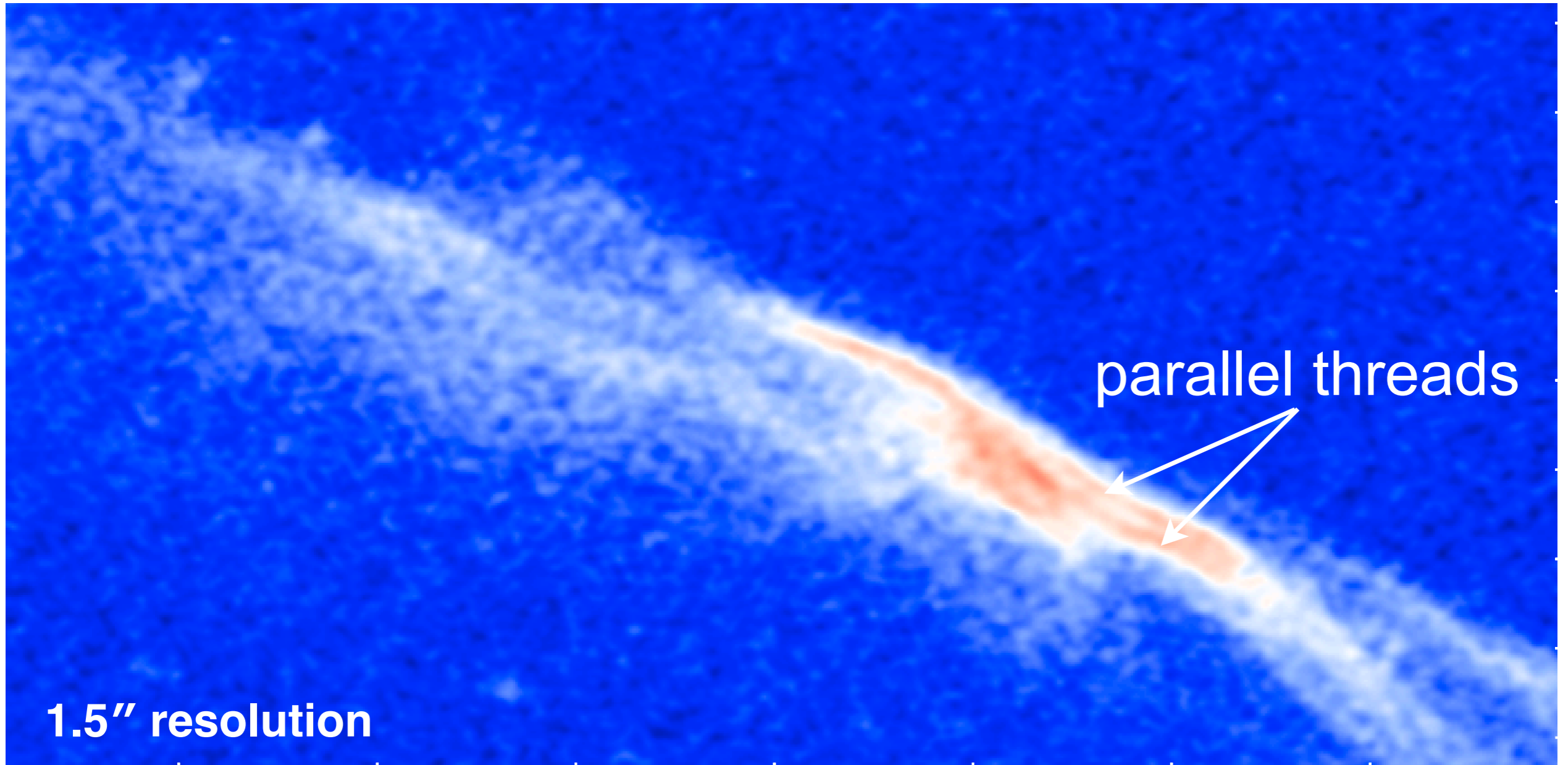
Ridge branches into two parts



Ridge branches into two parts

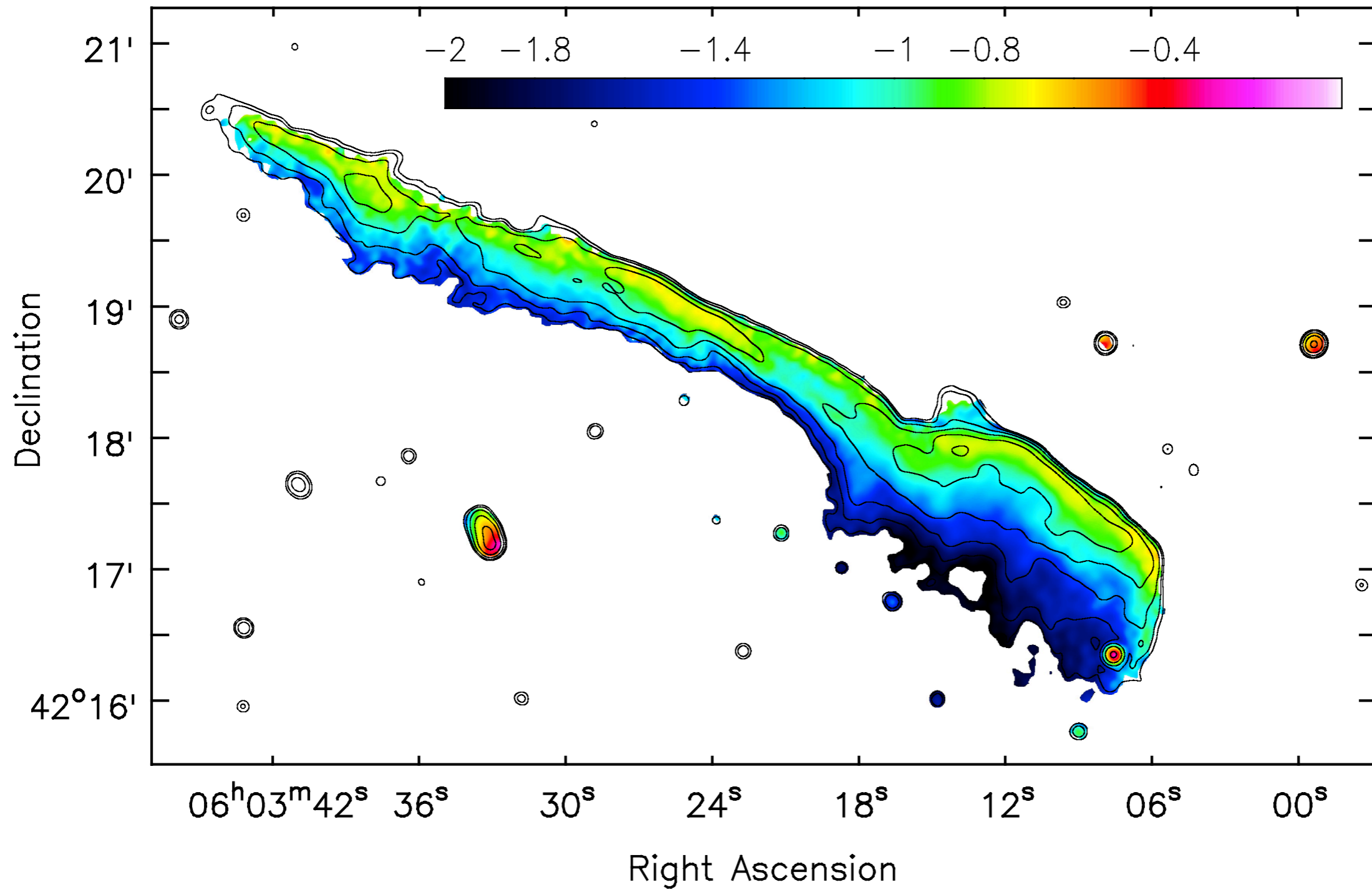


Two filaments separated by 5 kpc



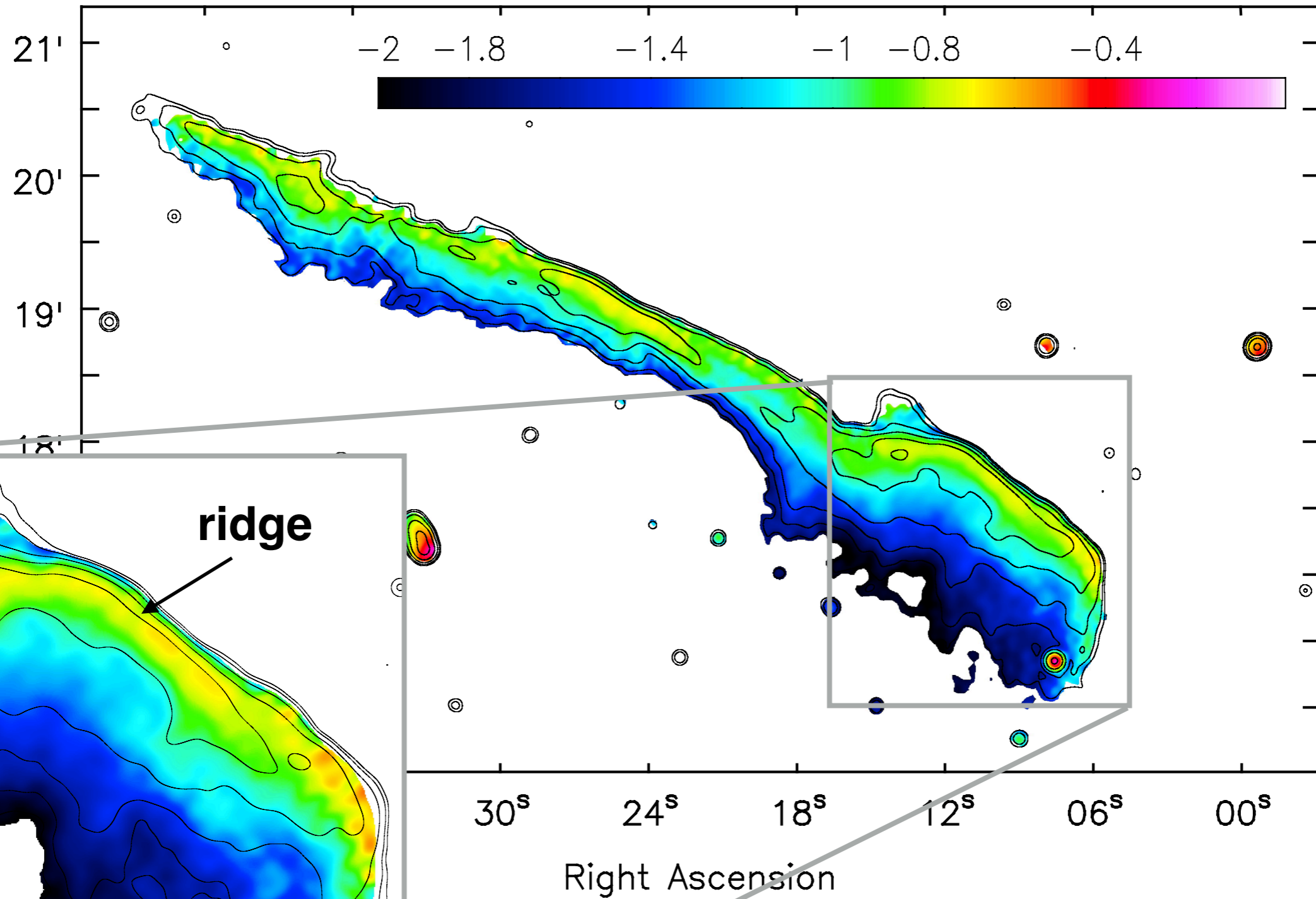
Spectral index at northern edge - $0.70 < \alpha < -0.75$

between 150 to 1500 MHz

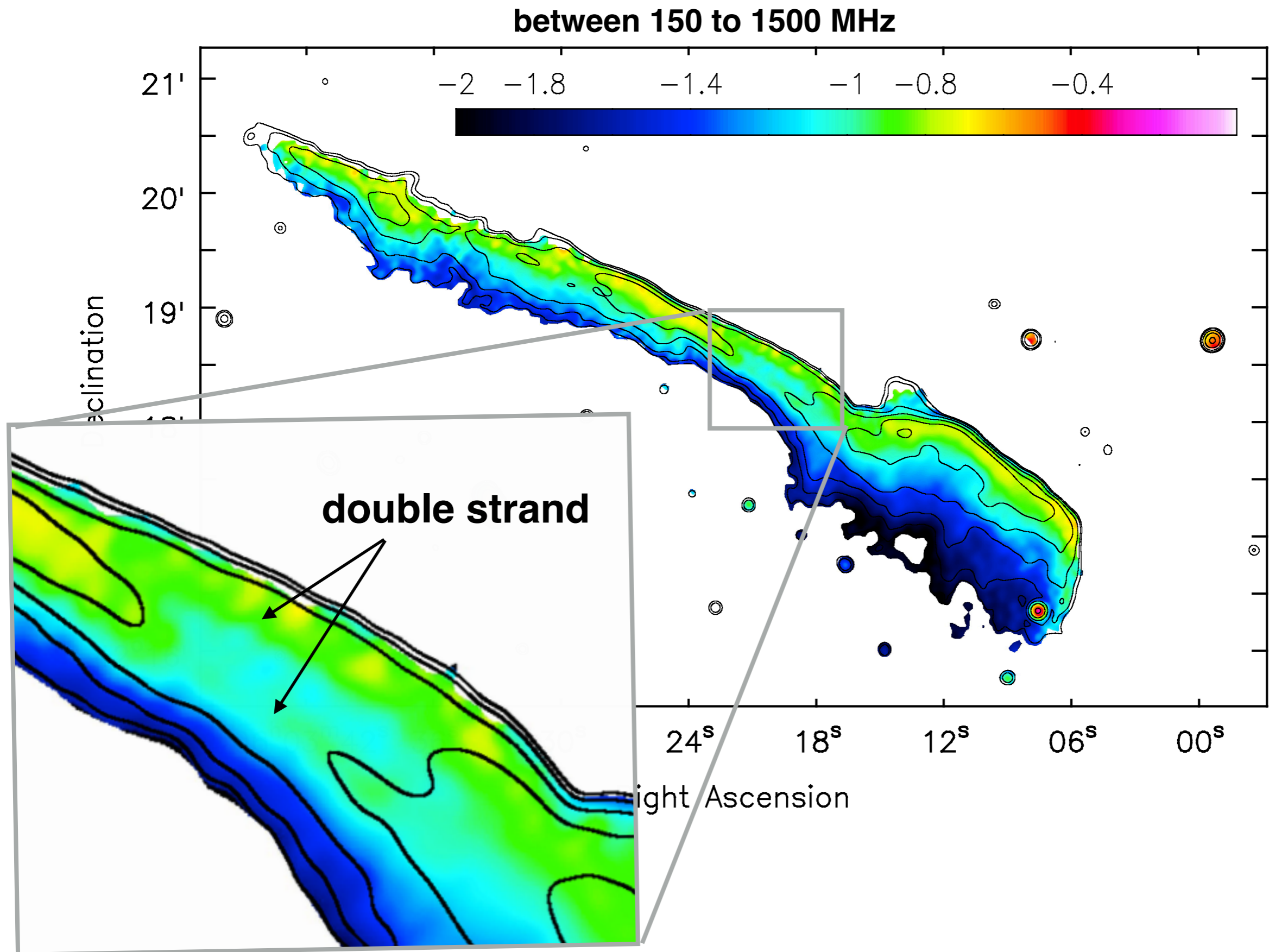


Spectrum steepens within the ridge

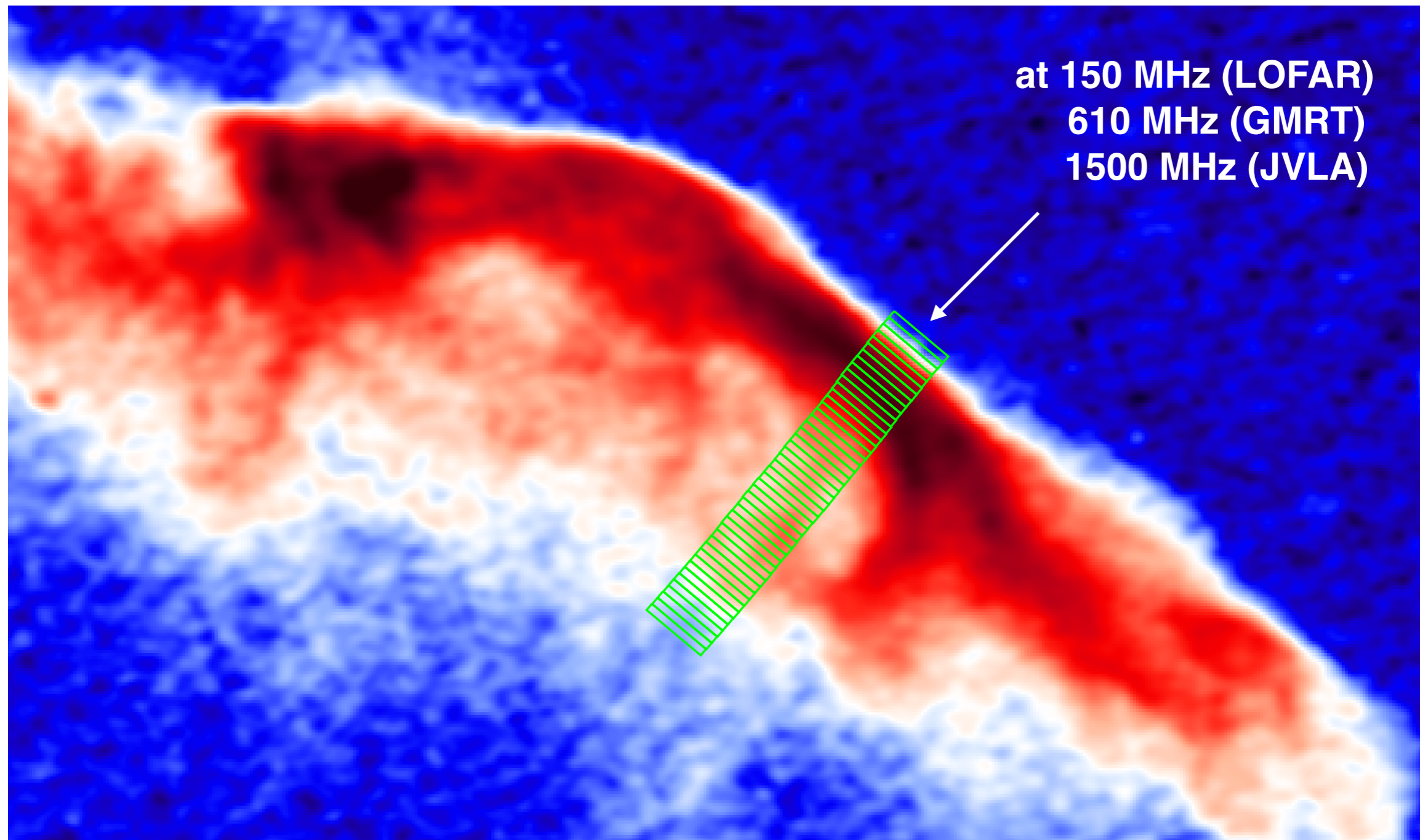
between 150 to 1500 MHz



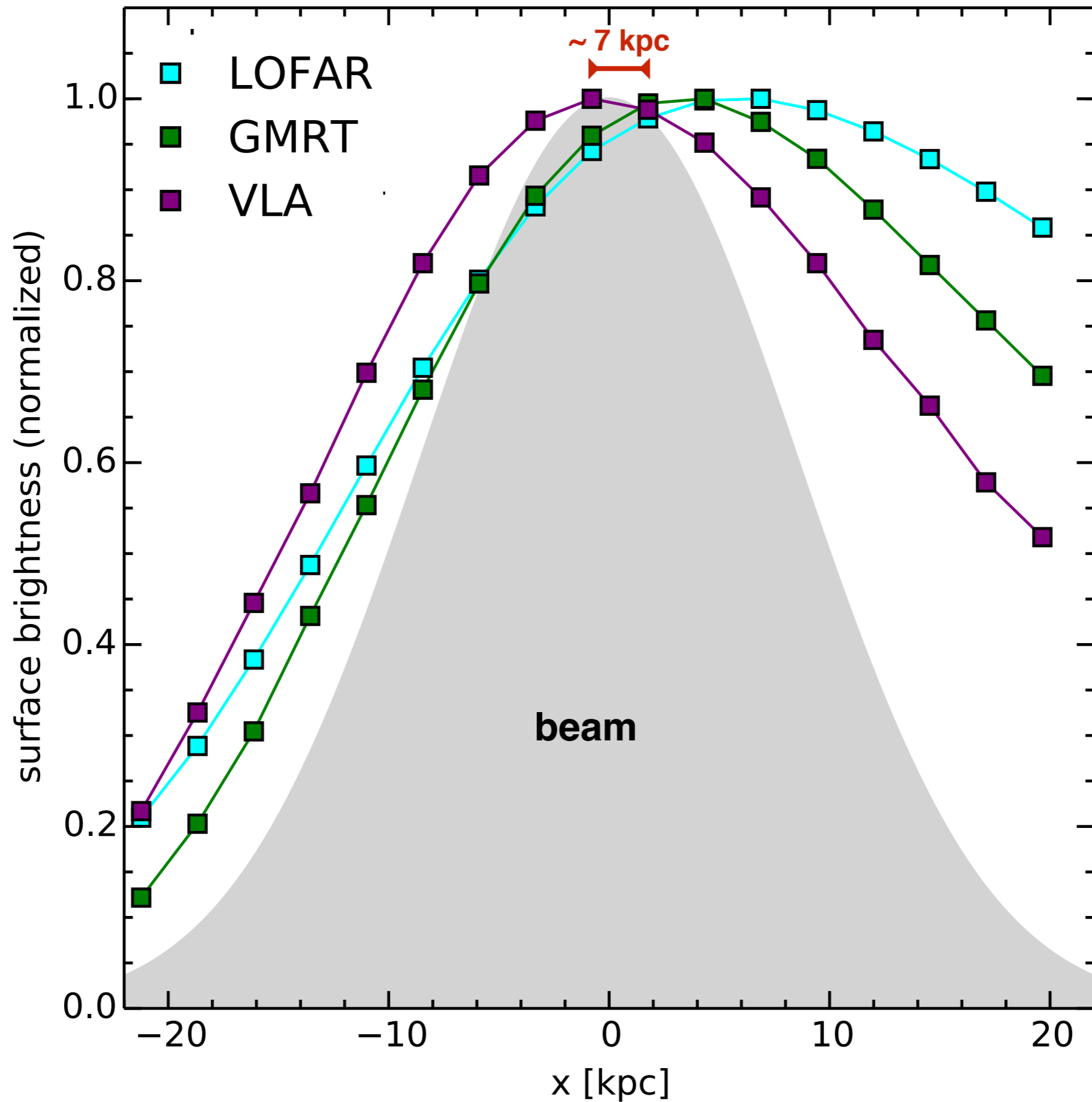
Spectral index across the double strand varies



Detailed investigation of the ridge

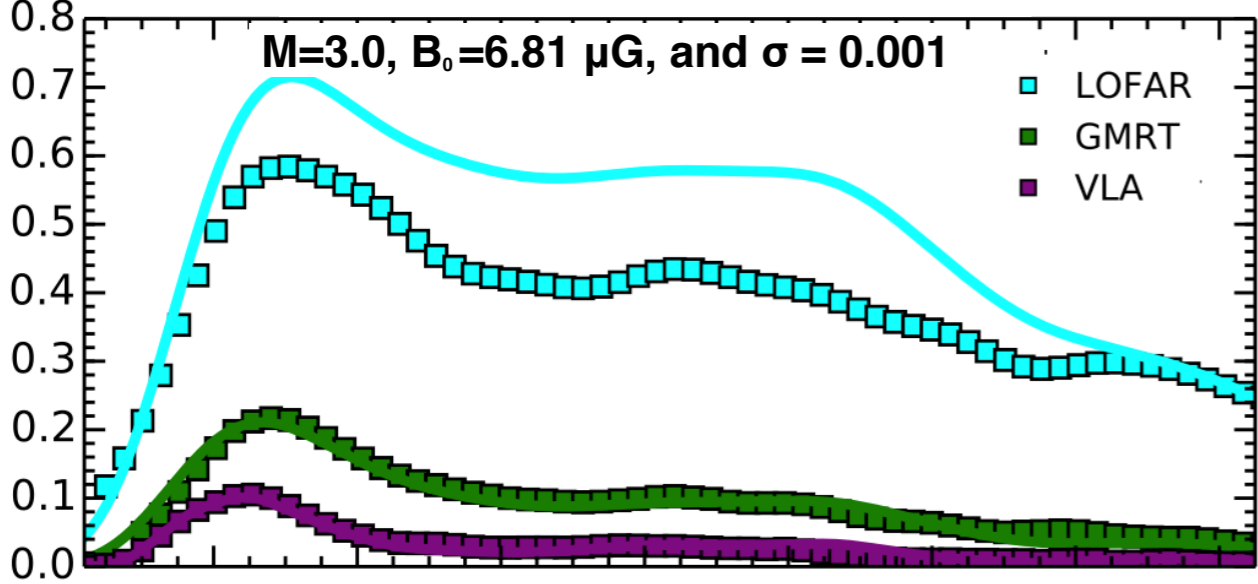


Ridge position shifts with frequency

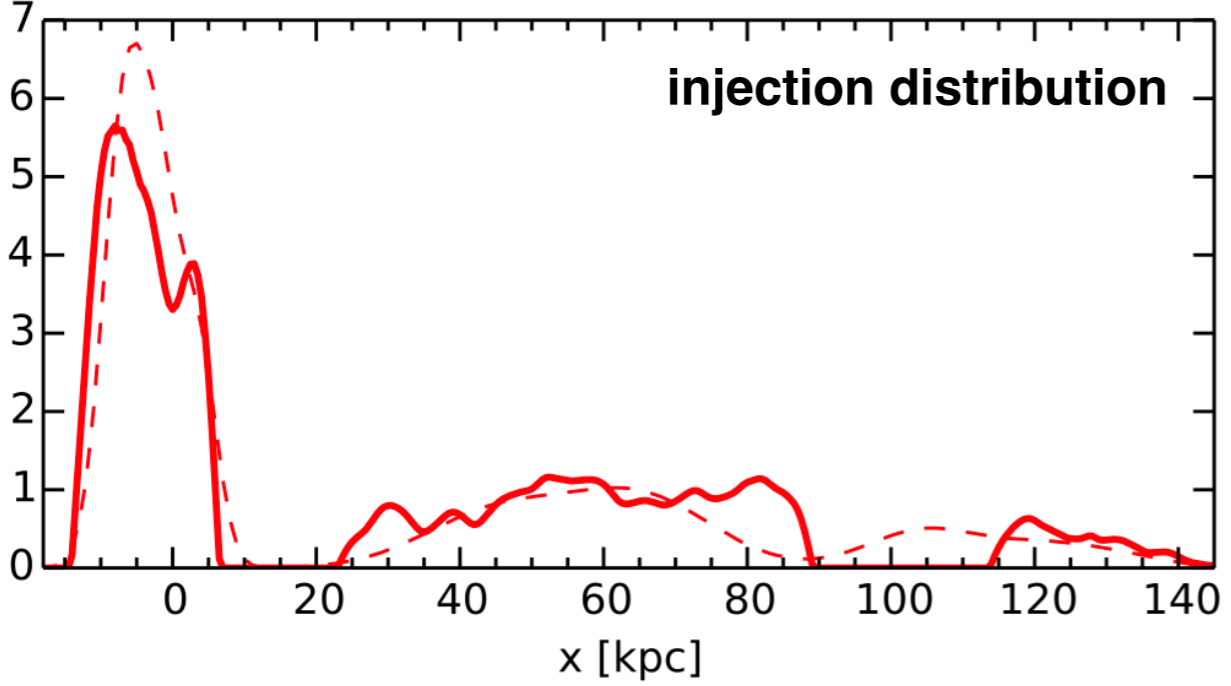


Surface brightness distribution:

homogenous magnetic field

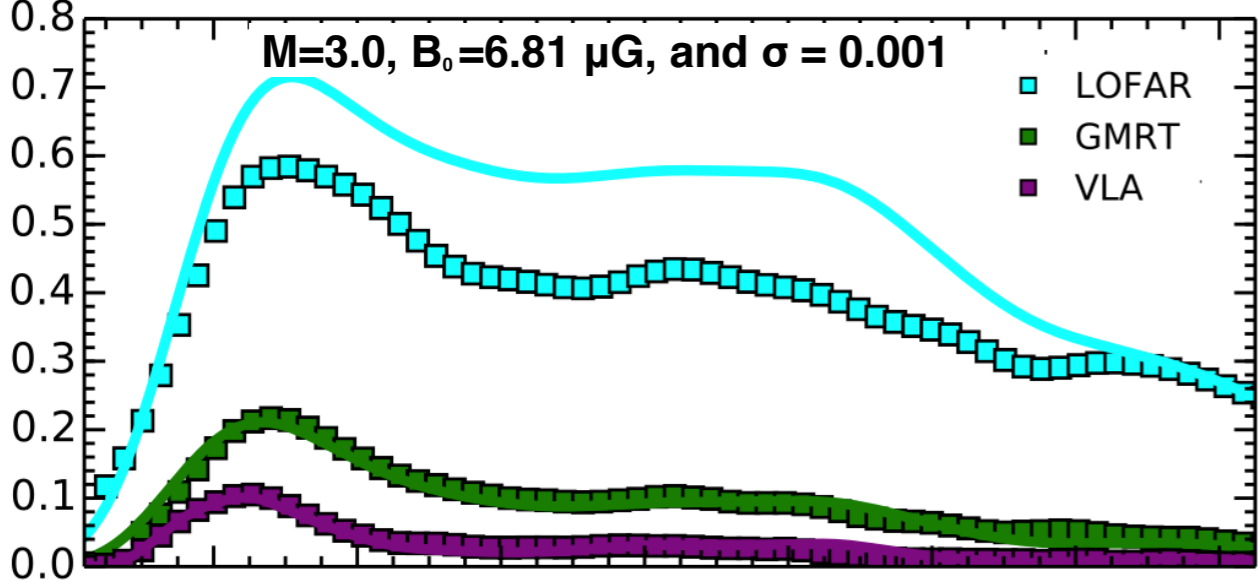


injection distribution

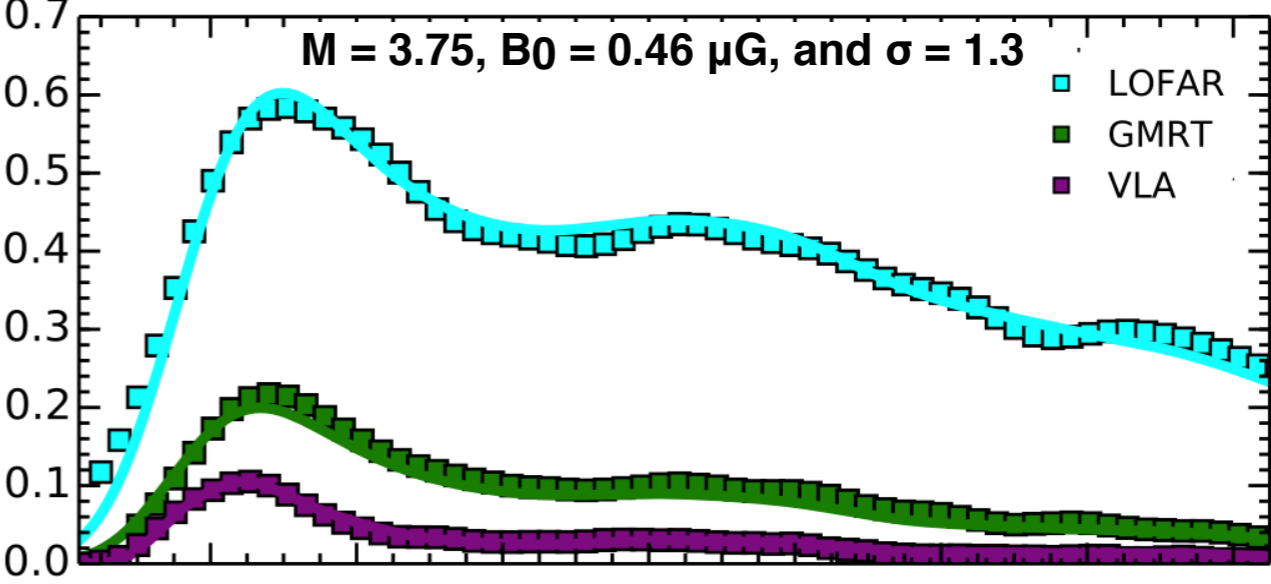


Surface brightness distribution:

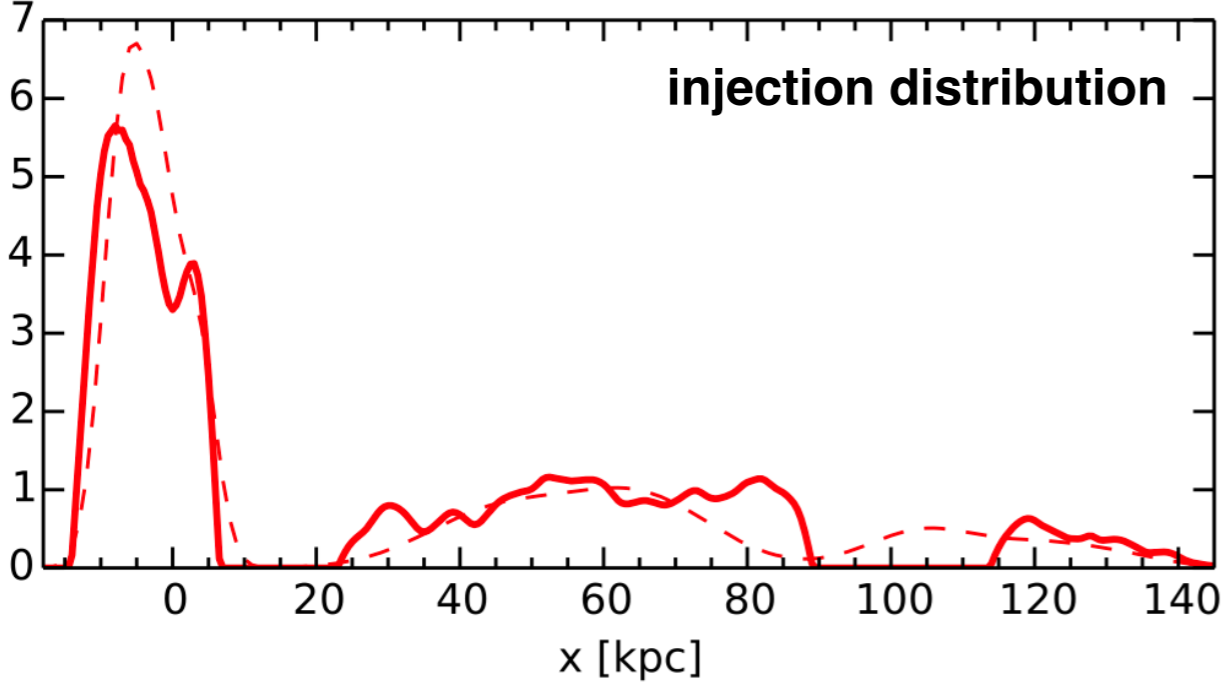
homogenous magnetic field



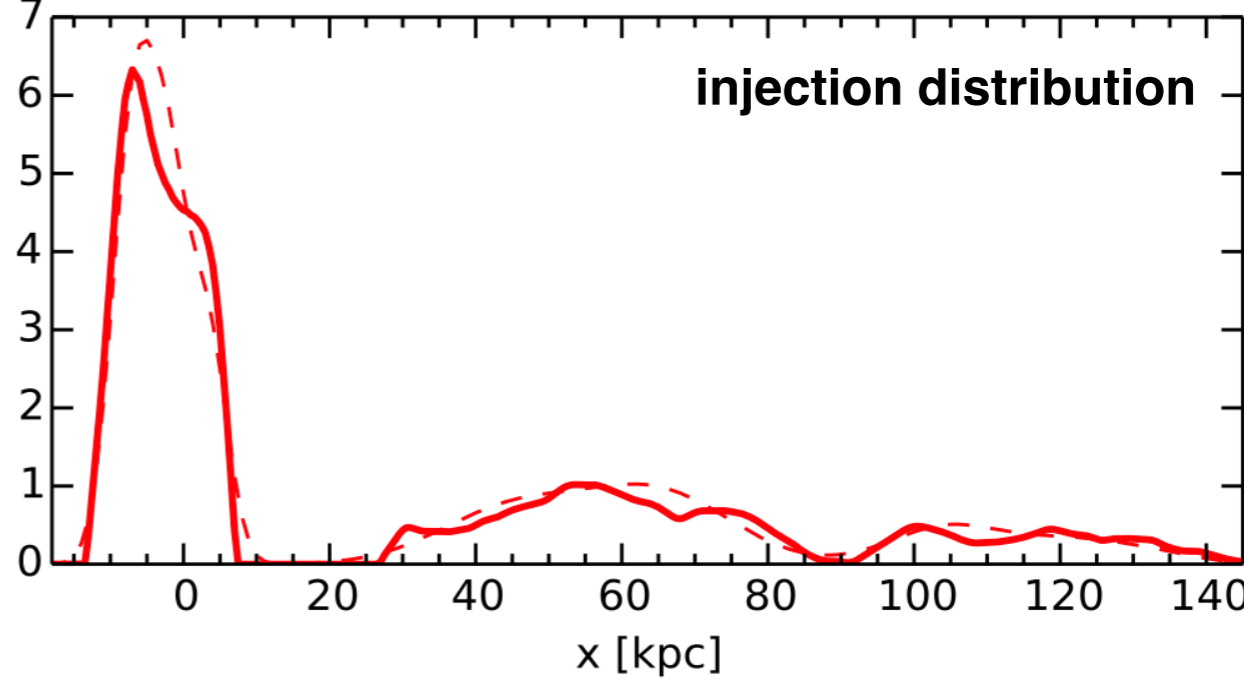
log normal distribution of magnetic field



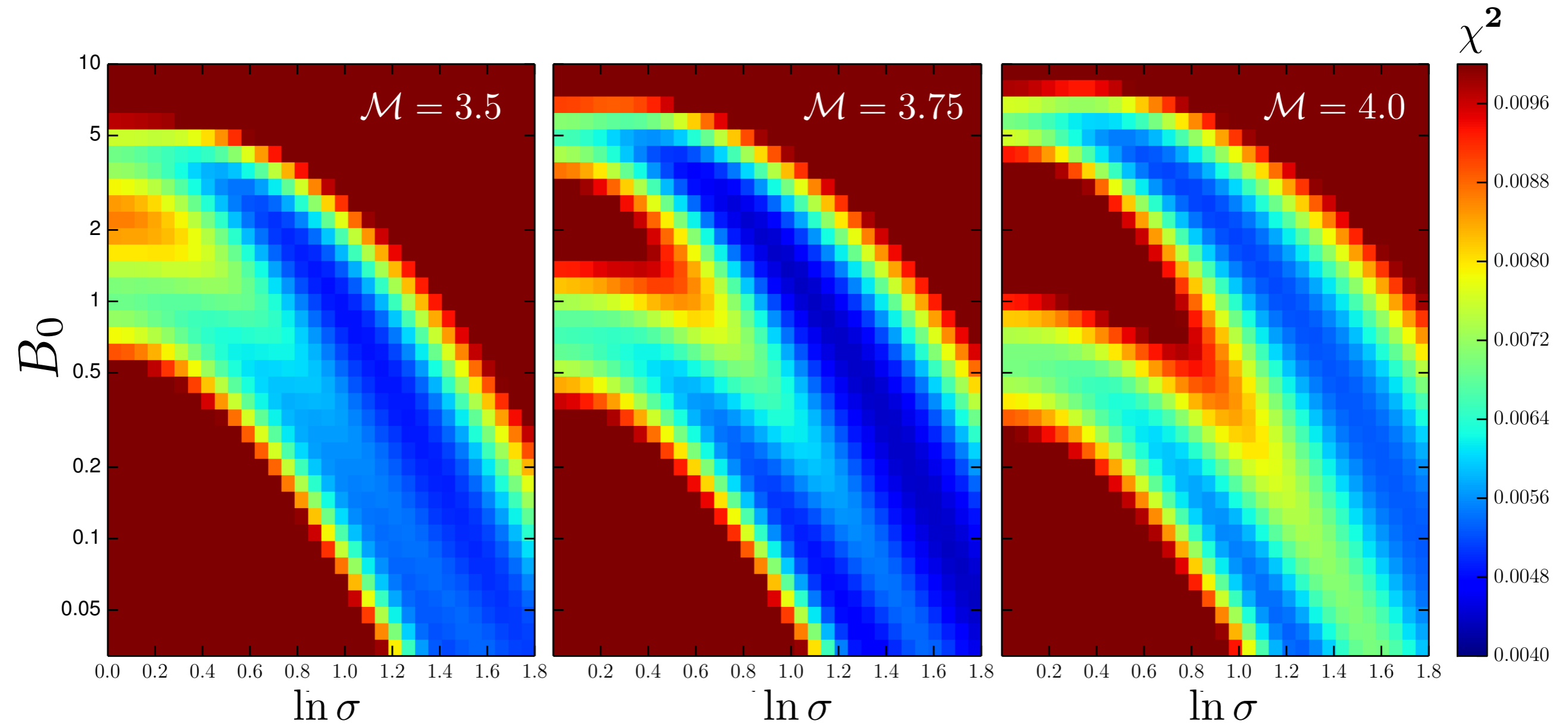
injection distribution



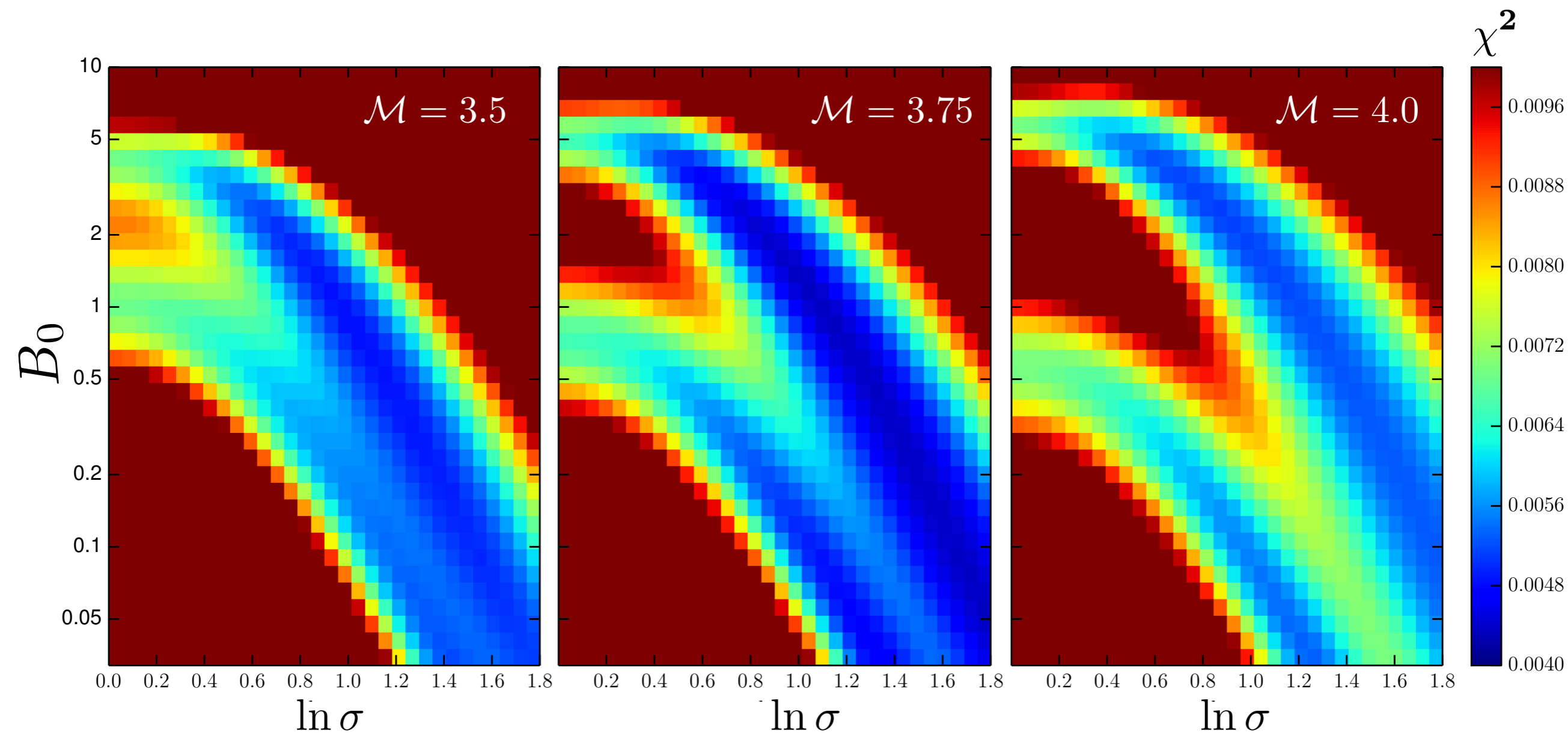
injection distribution



Best fit: $B_0 \leq 5 \mu\text{G}$, $\sigma \geq 0.7$ and Mach number = 3.75

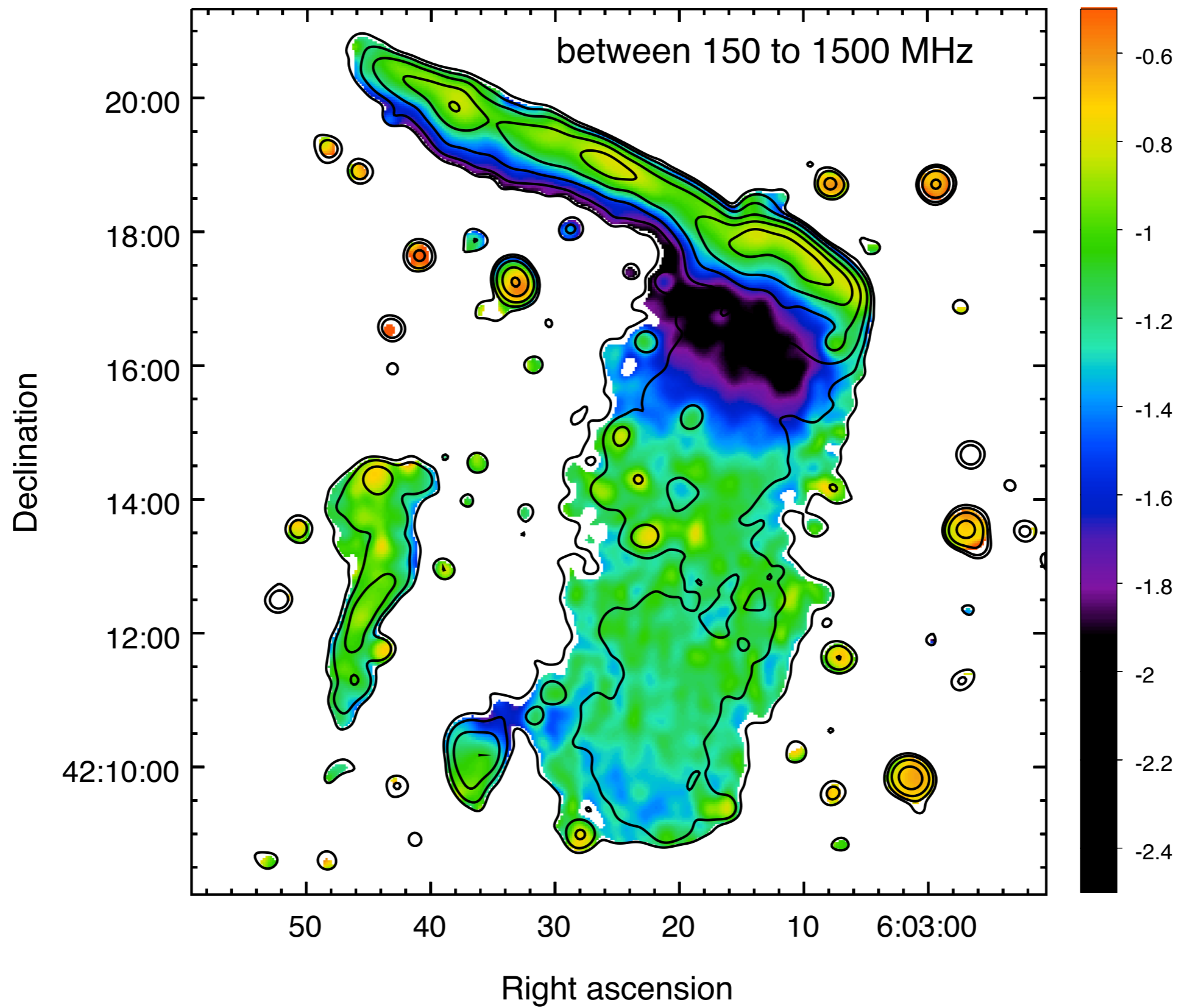


Best fit: $B_0 \leq 5 \mu\text{G}$, $\sigma \geq 0.7$ and Mach number = 3.75

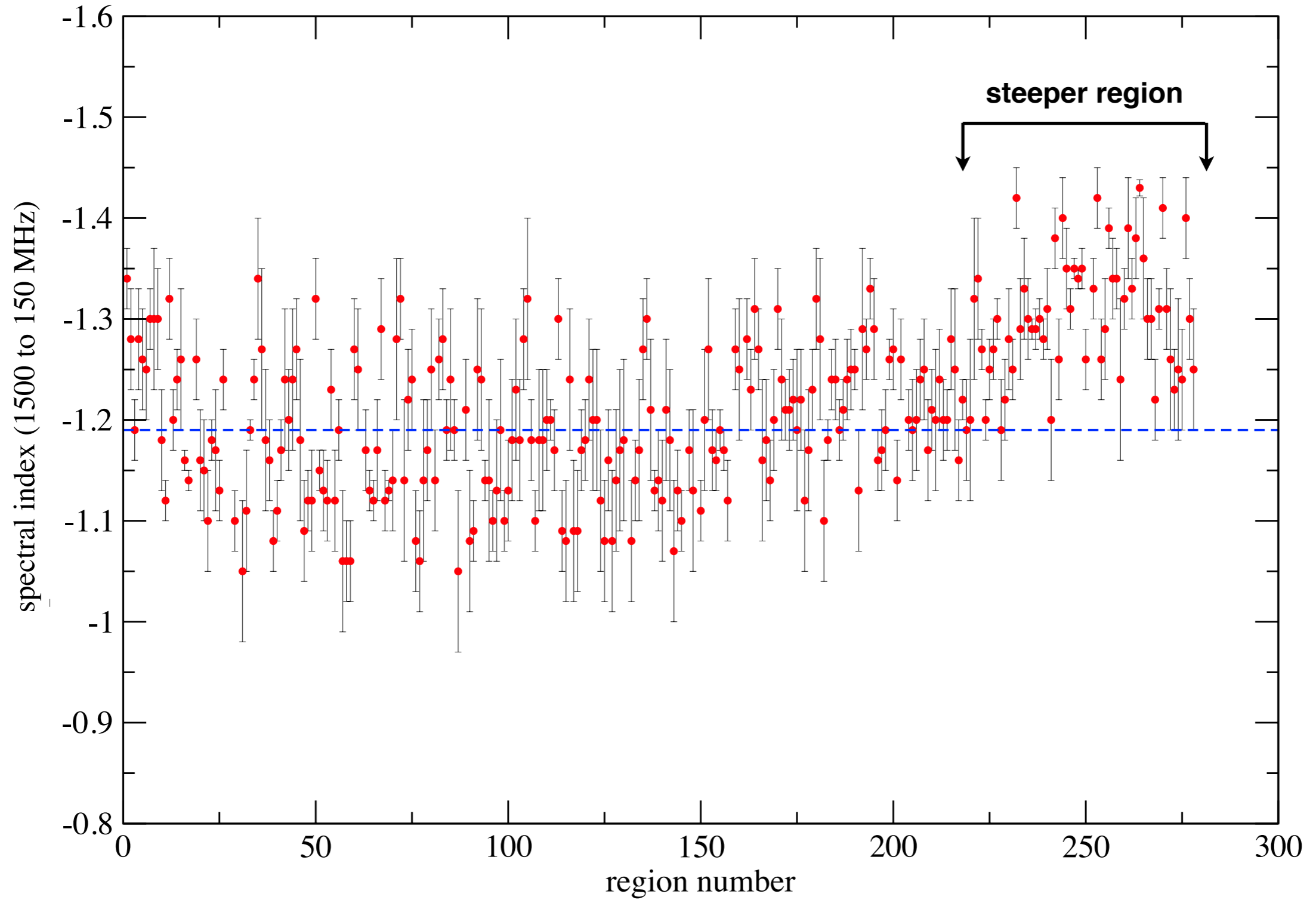


model ruled out field strength above 5 μG

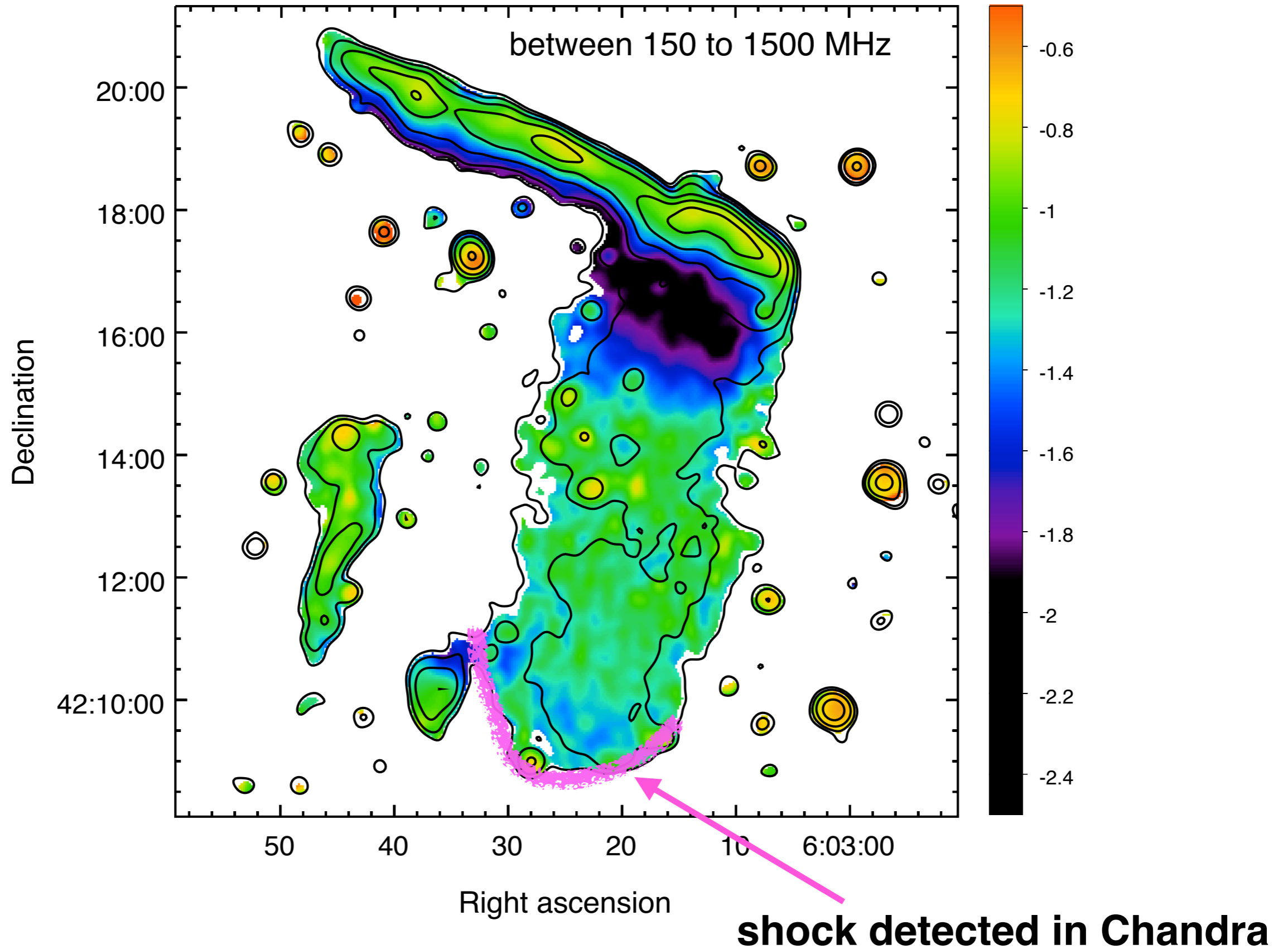
Halo: remarkable uniform spectral index



Halo: steep spectrum southern region

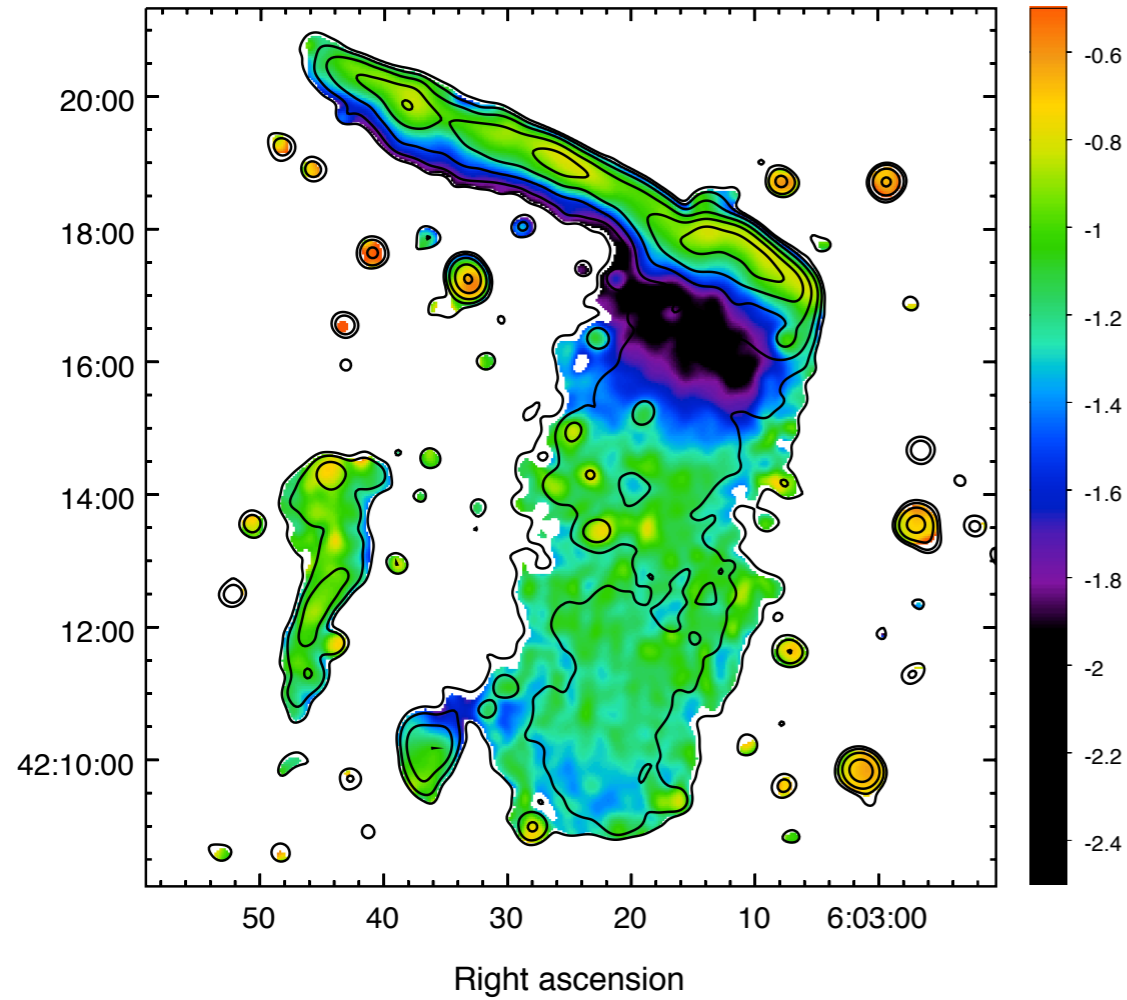


Halo southern most part: a fainter relic !

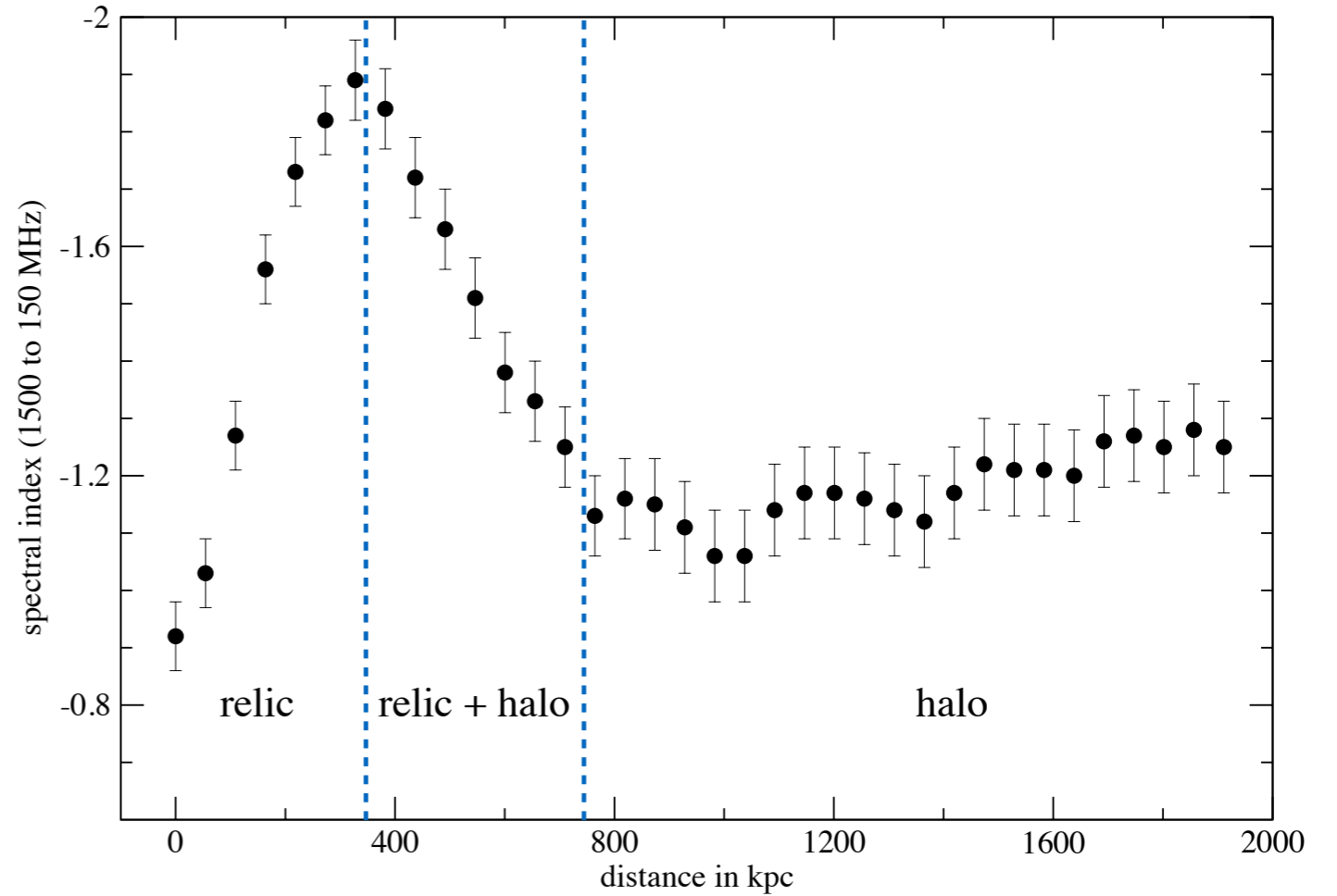


Relic and halo connection ?

JVLA contours

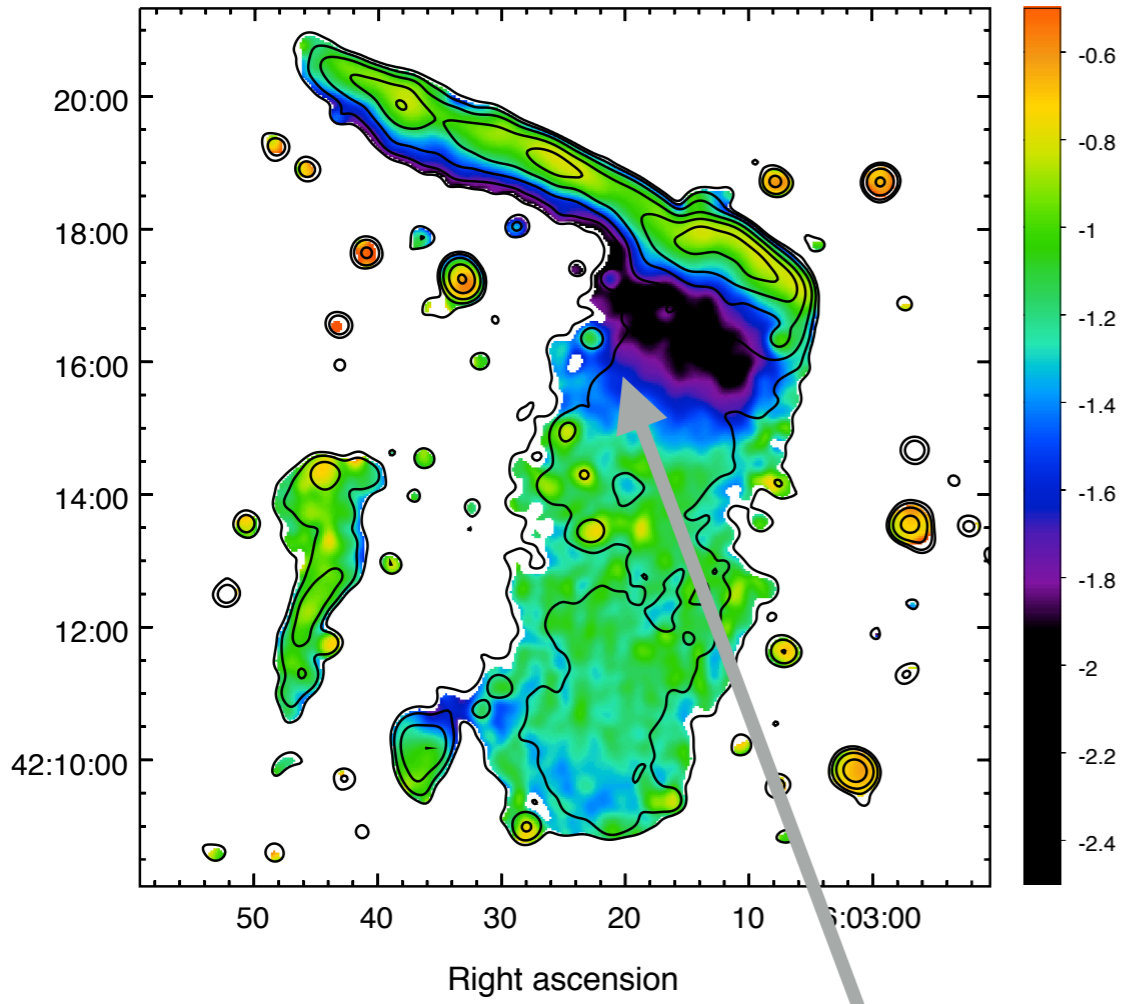


van Weeren et al. 2016 — gradual flattening is due to the re-acceleration by turbulence of “aged” electrons downstream of the relic

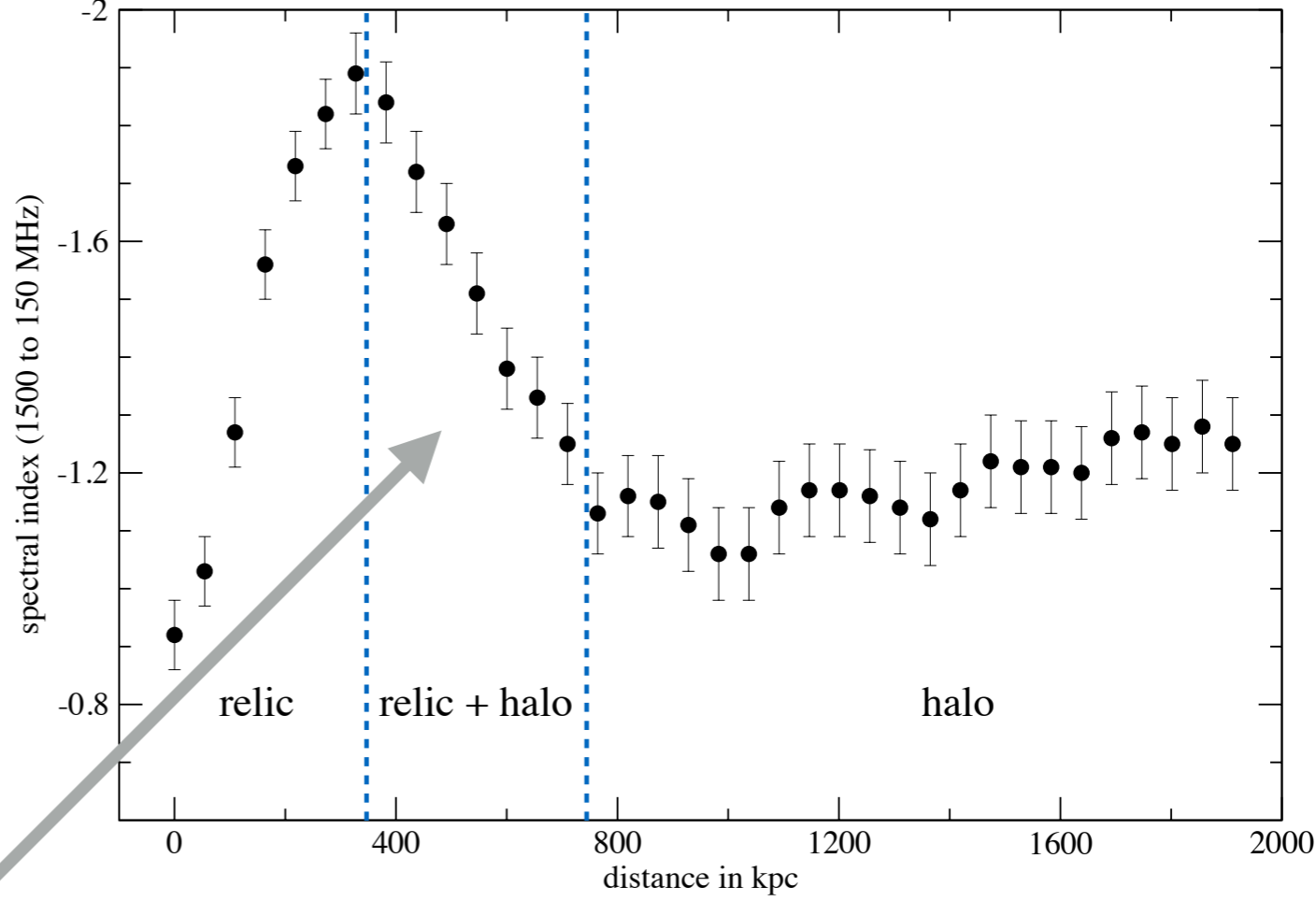


Relic and halo connection ?

JVLA contours

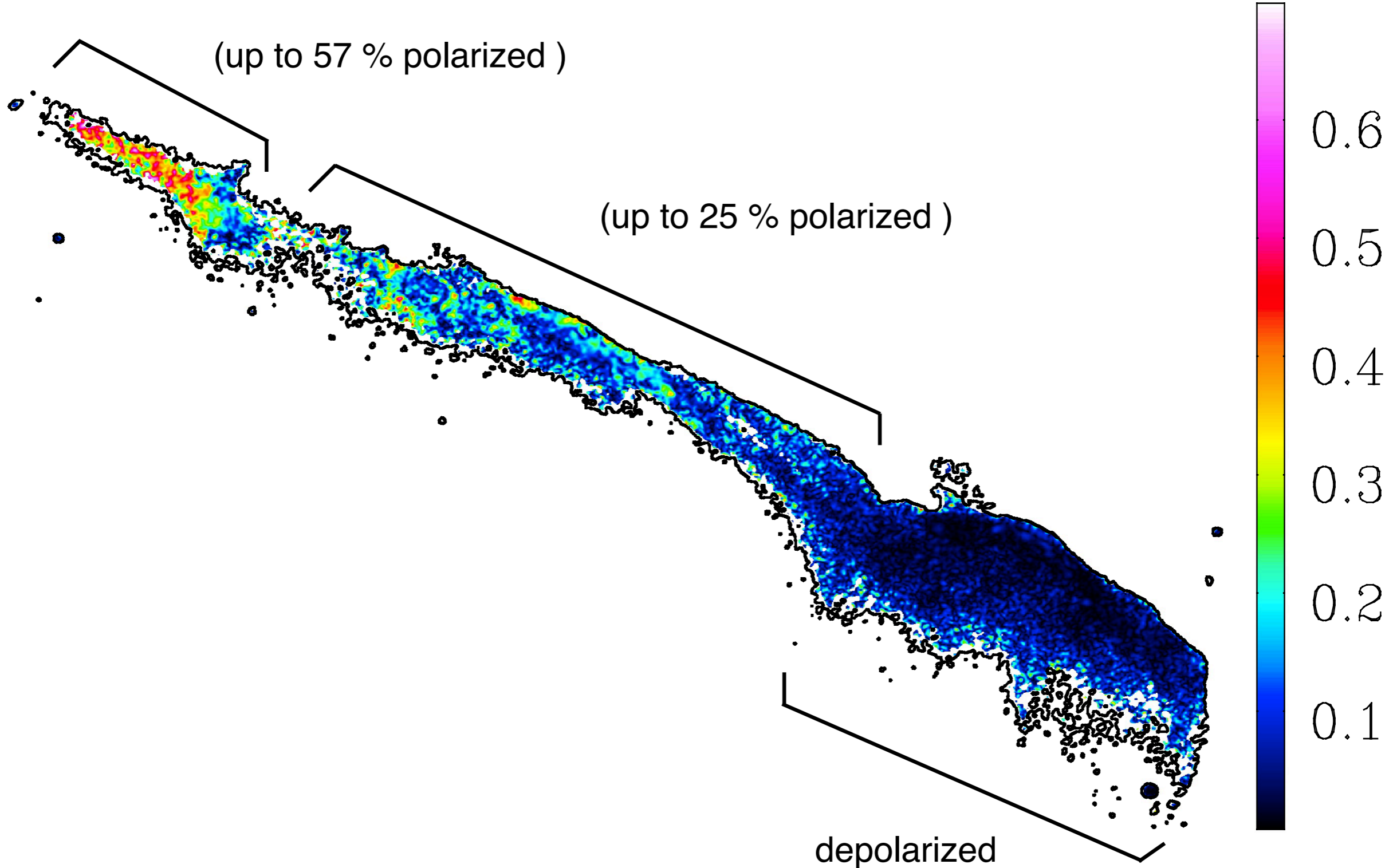


van Weeren et al. 2016 — gradual flattening is due to the re-acceleration by turbulence of “aged” electrons downstream of the relic

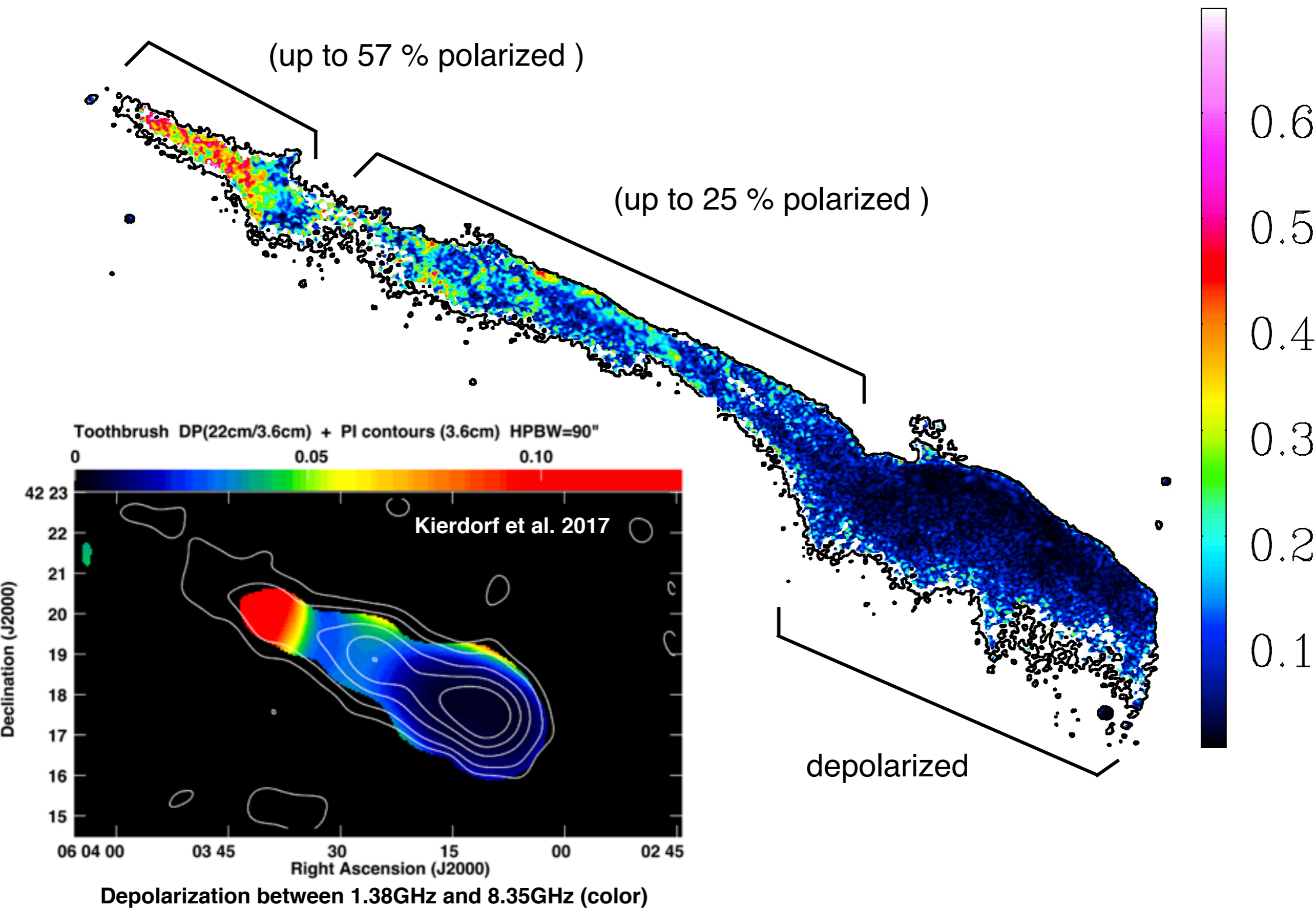


flattening is due to projection of halo and relic emission

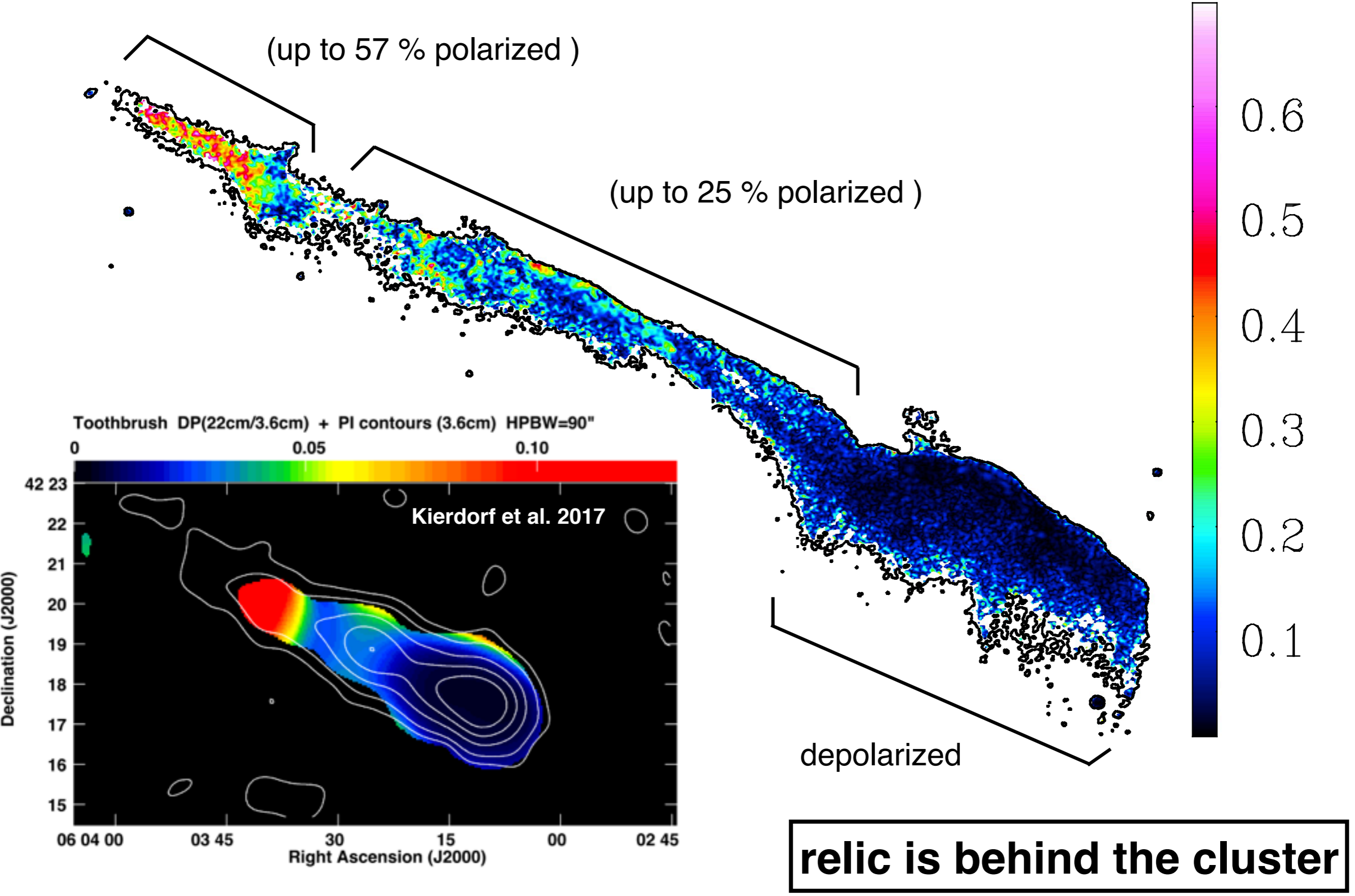
Degree of polarization: brush depolarized at 1.5 GHz



Degree of polarization: brush depolarized at 1.5 GHz



Degree of polarization: brush depolarized at 1.5 GHz



Summary

- Toothbrush is made up of filamentary structures
- Ridge peak shift with frequency, indicating intrinsic shape of the emission is frequency dependent
- Lognormal B distribution allows to approximate profiles significantly better
- Best fit: Mach ~ 3.75 , $B_0 < 5 \mu\text{G}$, $\sigma > 0.7$
- Southern part of the halo is steeper and flattens again at the edge
- Brush depolarized at 1.5 GHz

Radio
X-ray
optical

