



NWU LNA's for MFAA

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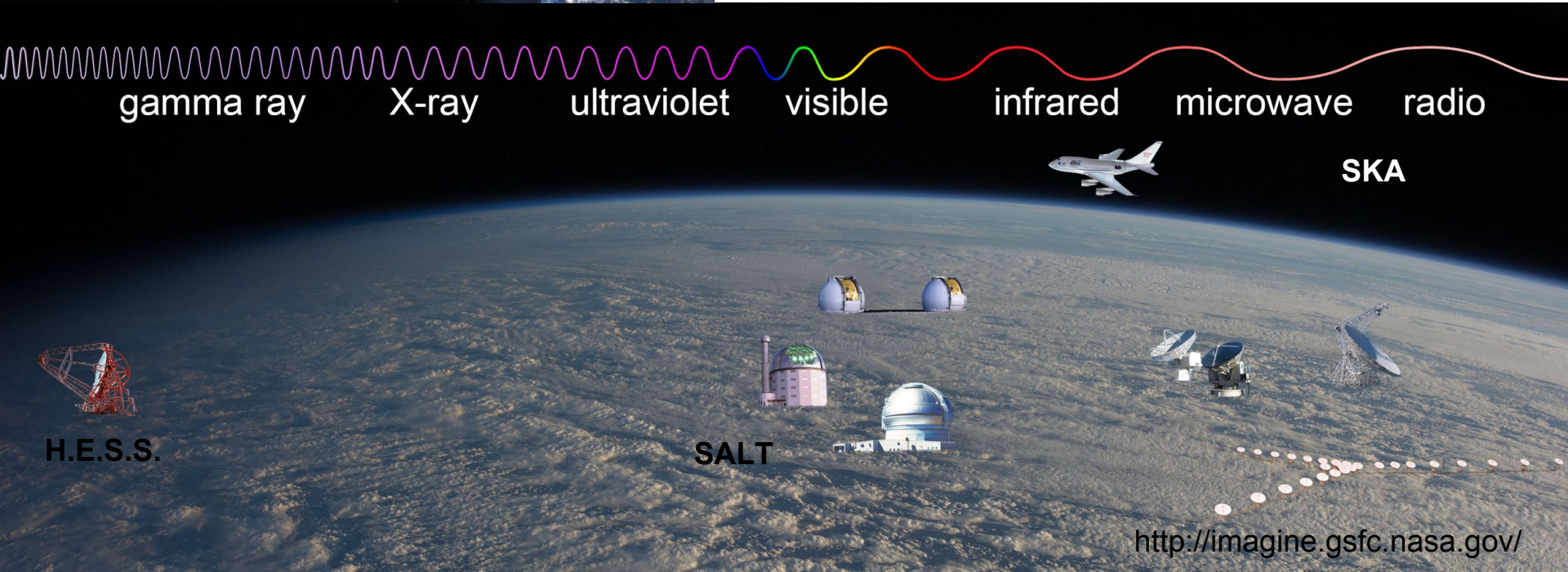
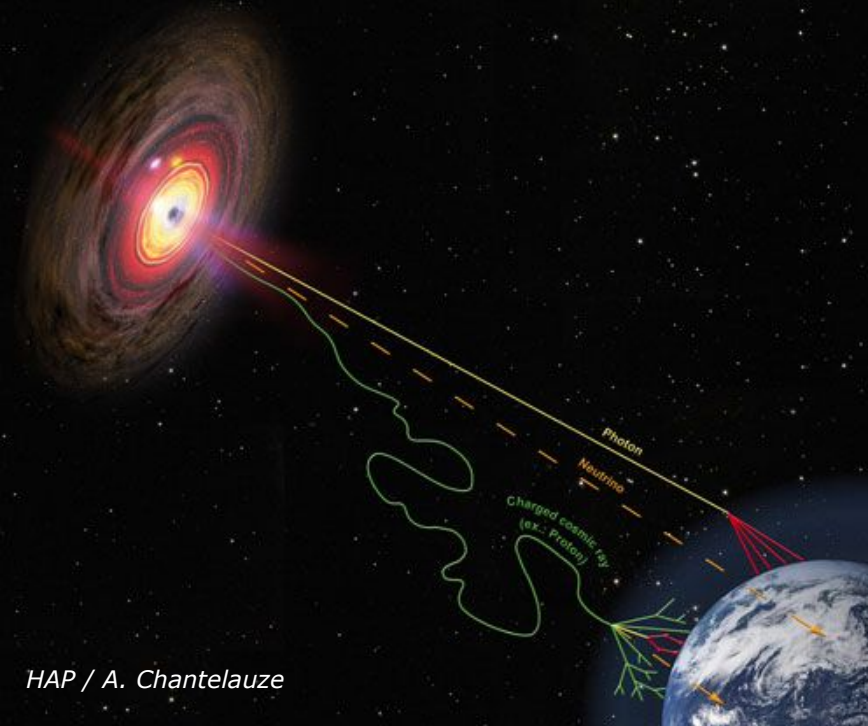


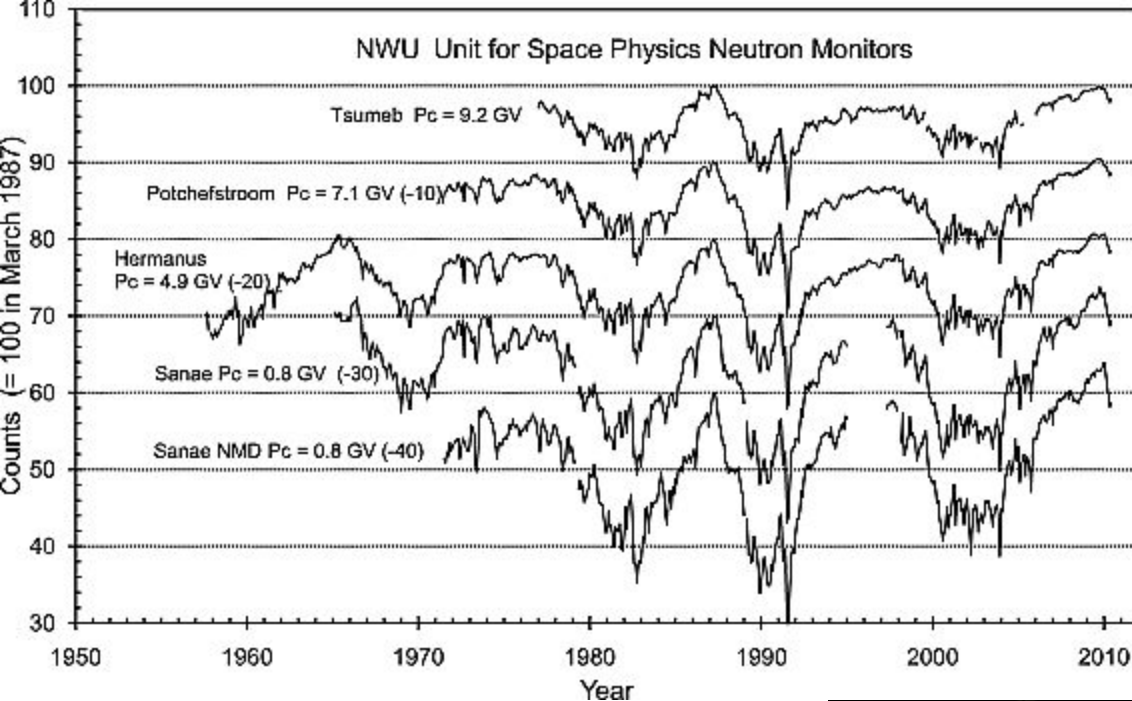
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®

NWU background

High-energy astrophysics
Gamma-ray astronomy
H.E.S.S. & CTA
Cosmic-ray research
Neutron monitors





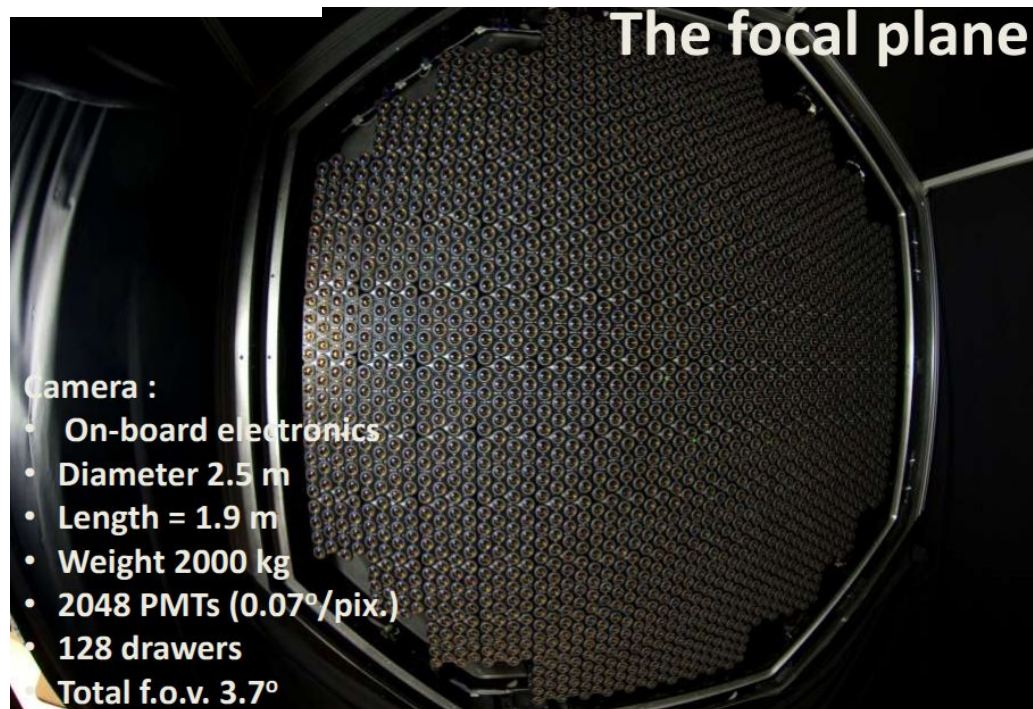
NWU background

Neutron monitors

Need room-temperature,
pulse (broadband) LNAs

Gamma-ray telescope

Need 1000x room-temperature,
pulse (broadband) LNAs

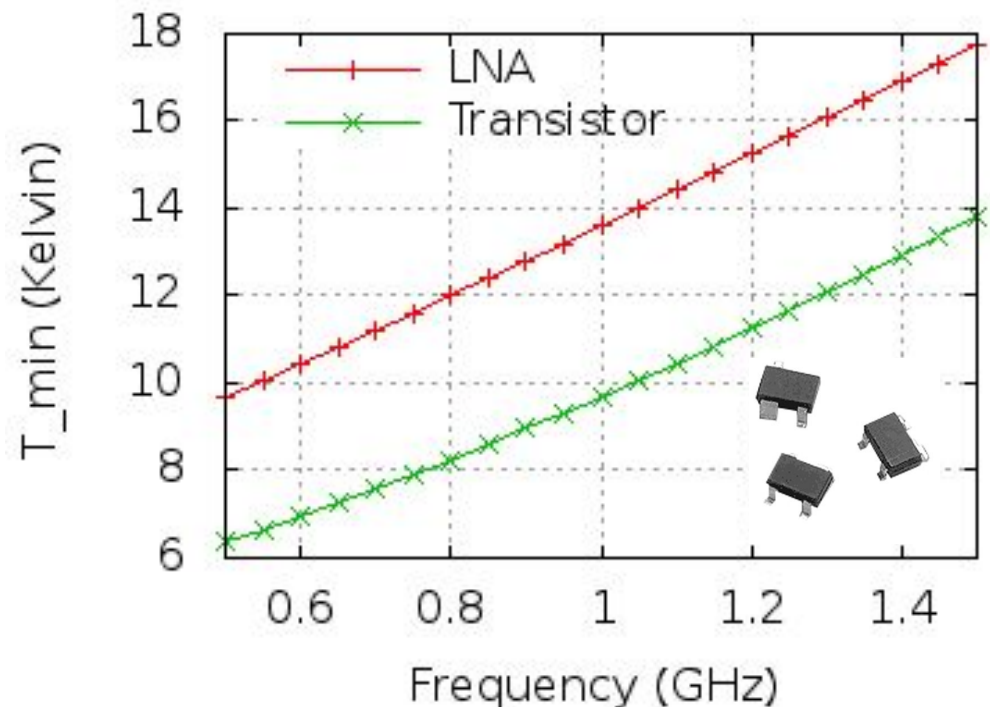


MFAA LNA requirements

- Broadband: 0.4 - 1.5 GHz
- Low noise: <30 K
- Little power consumption: <50 mW
- Inexpensive
- Fully differential

Commercial HEMT: ATF 35143:

- Broadband: 0.5 - 1.5 GHz
- Low noise: <30 K
 - **$T_{min} < 14K$ below 1.5 GHz!**
- Little power consumption: <100 mW (for 30 dB gain)
- Inexpensive
- Fully differential



Commercial HEMT: ATF 35143:

- Broadband: 0.5 - 1.5 GHz
- Low noise: <30 K
 - $T_{min} < 14K$ below 1.5 GHz!
- Little power consumption: <100 mW (for 30 dB gain)
 - 20mW / HEMT:
 - 2 HEMT = 30dB gain & 40 mW
- Inexpensive
 - <\$1 / HEMT
- Fully differential
 - 4 HEMTS



The Problem:

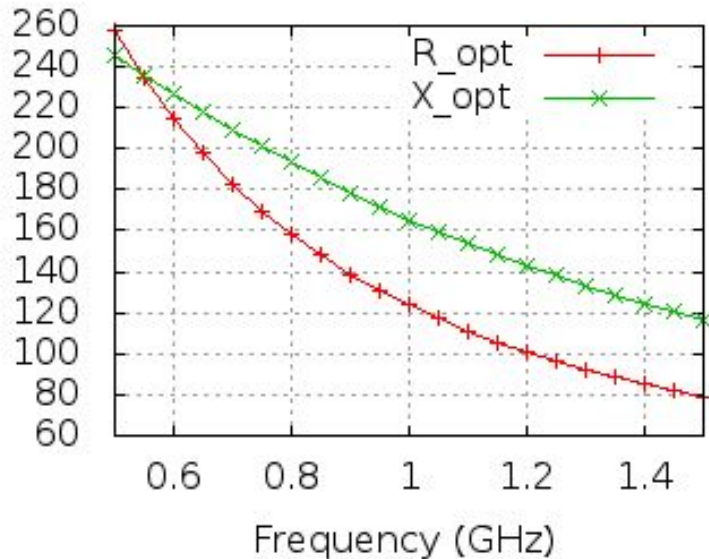
Very low T_{min} only at the optimum impedance, Z_{opt}

Z_{opt} is frequency dependent

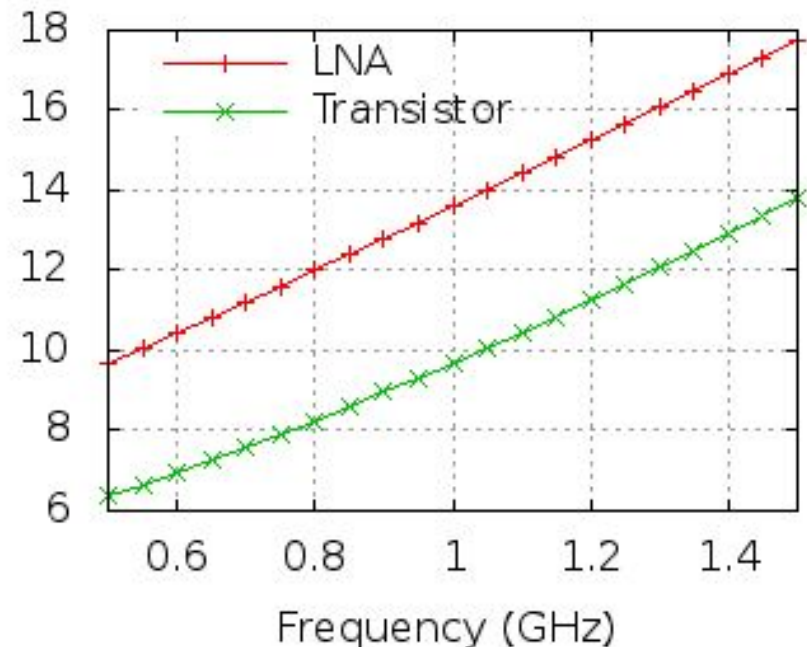


Avago ATF 35143:

Resistance / Reactance (Ohm)



T_{min} (Kelvin)

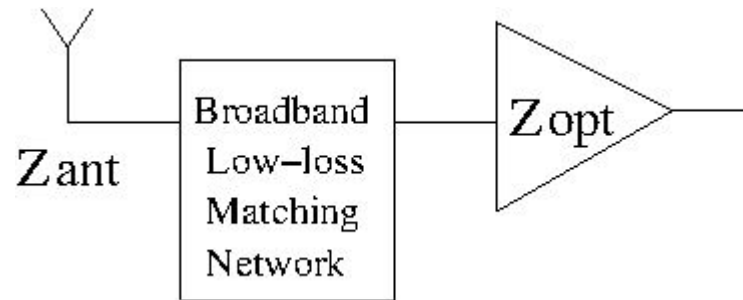


2 proposed solutions:

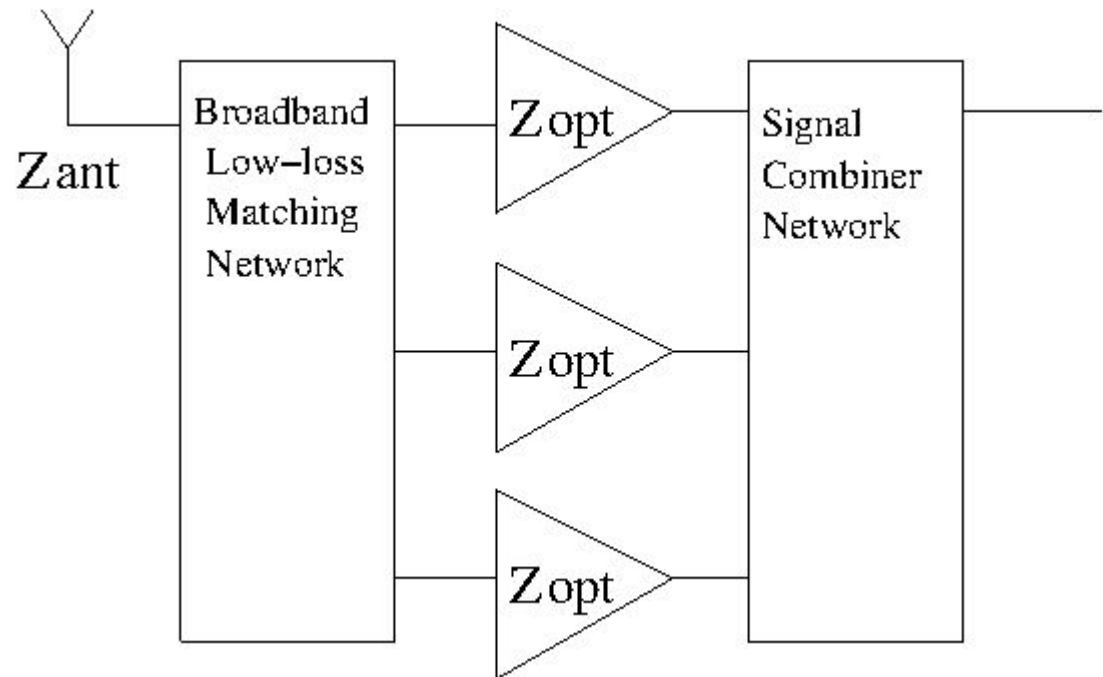
- Multipath amplifier
 - Low-loss matching technique to match antenna (Z_{ant}) to the optimum impedance (Z_{opt})
- Modify antenna
 - So that the antenna impedance is close to the optimum impedance, $Z_{ant}=Z_{opt}$

1. Multipath amplifiers

Difficult to match Z_{ant} to Z_{opt} over broadband with low loss:

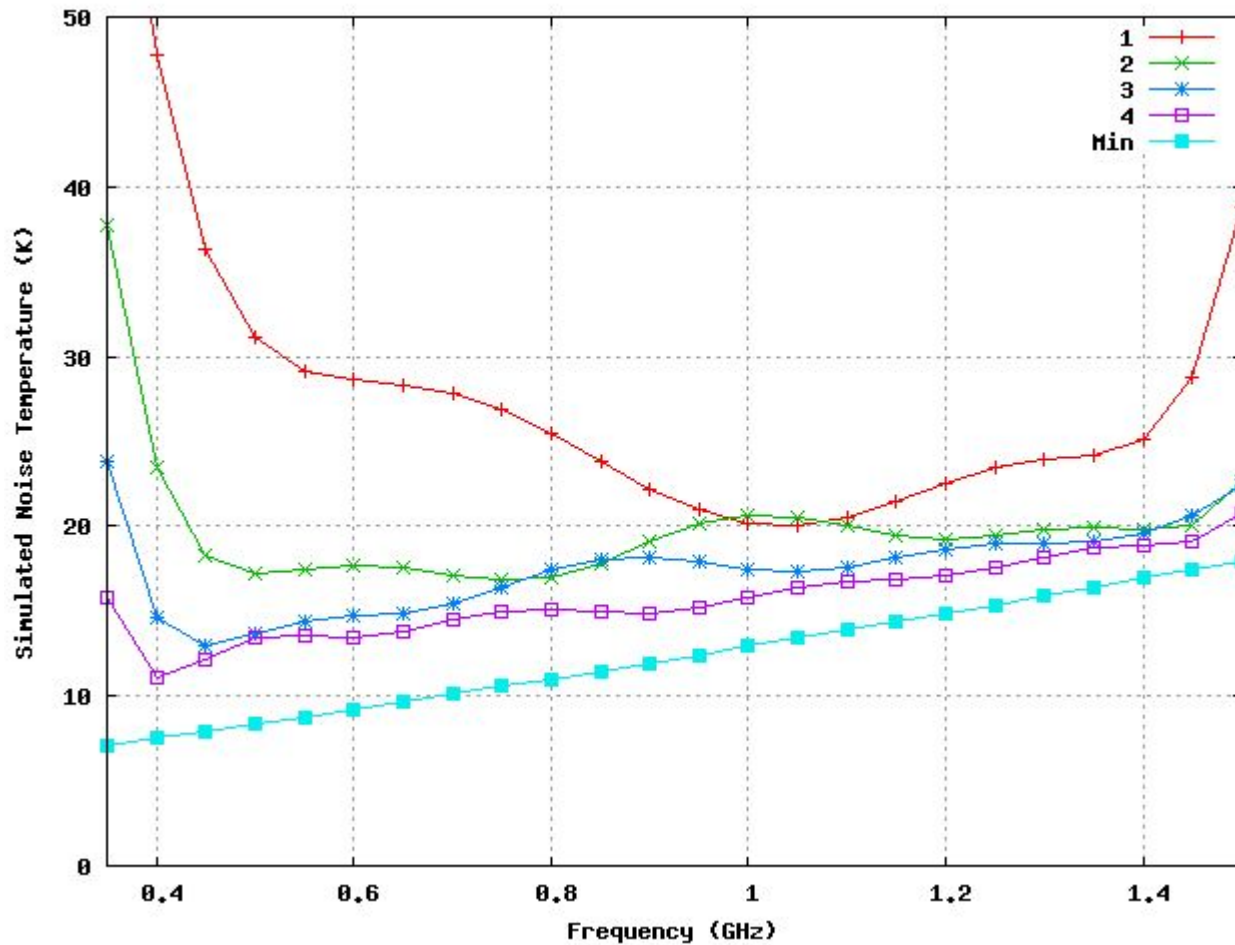


Much easier to match Z_{ant} to several amplifier's Z_{opt} :



Multipath amplifier: Effect of the number of paths

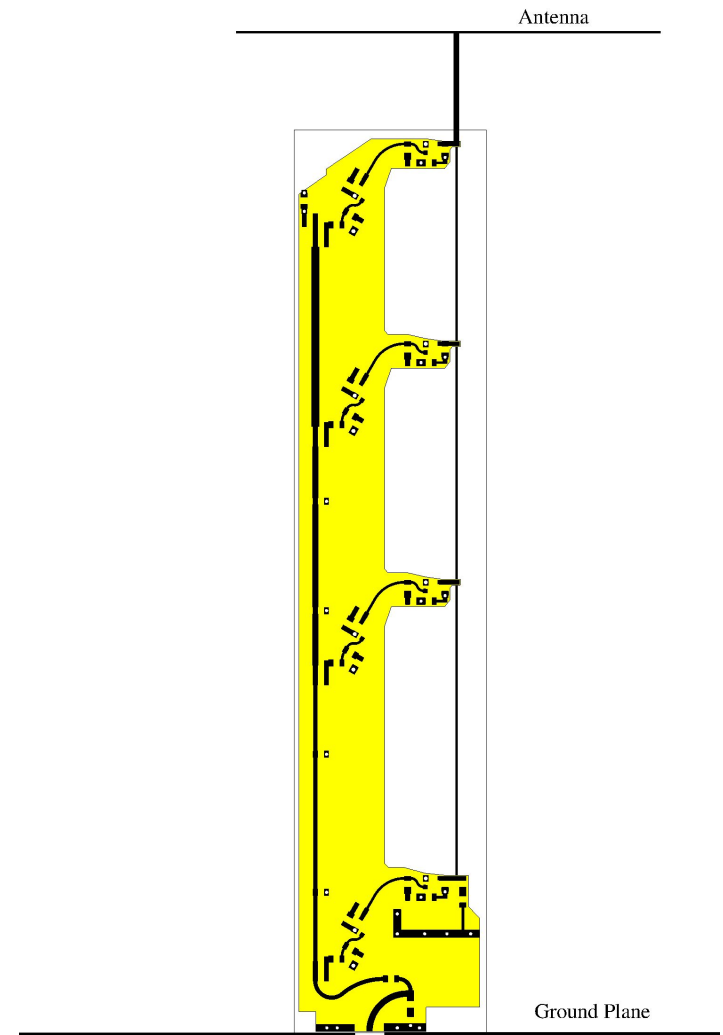
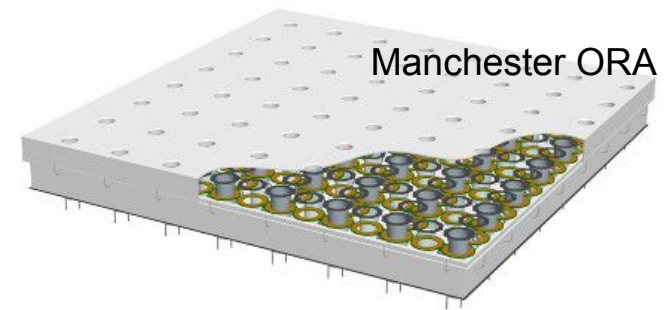
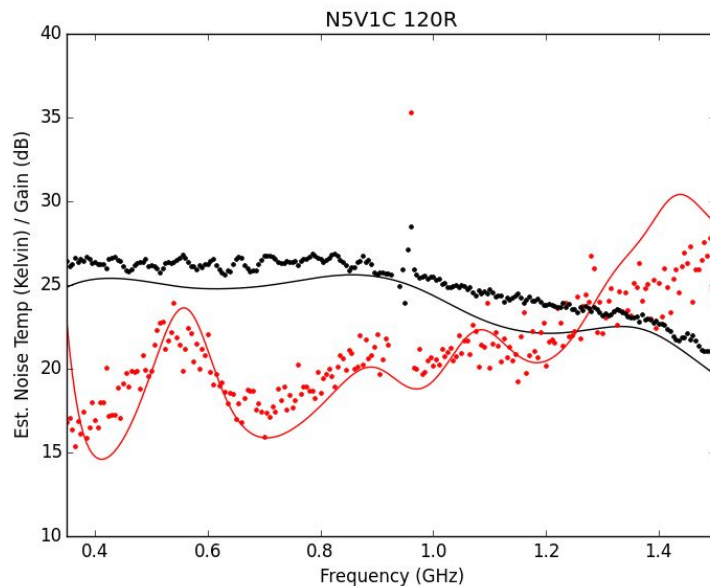
Matching ATF35143's Z_{opt} to 60 Ohm from 0.35 - 1.5 GHz:



4-path LNA demonstrator

Fully Differential LNA

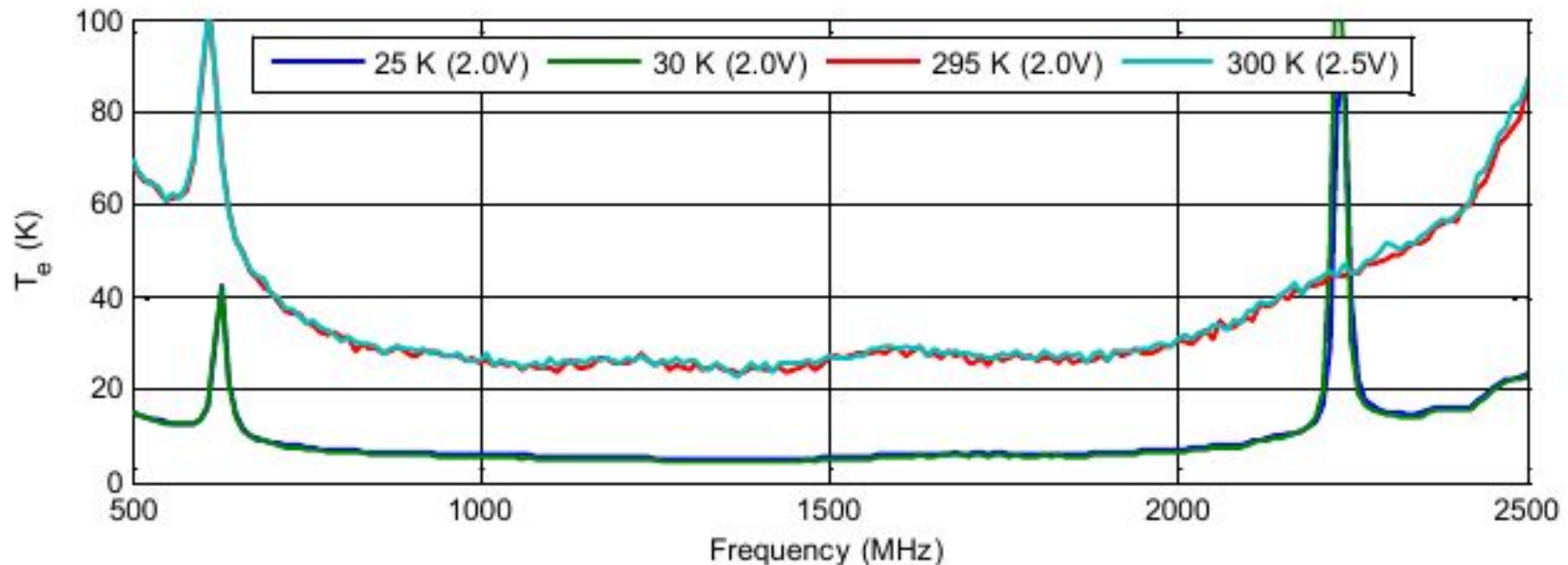
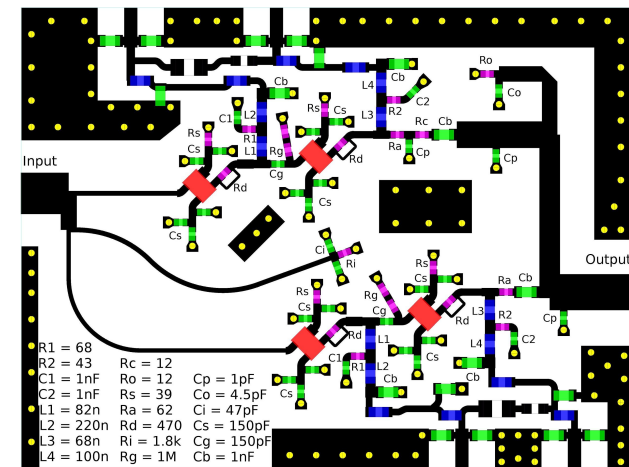
- + Differential Input: 120 Ω
- + Bandwidth: 0.35 - 1.5 GHz
- + Measured noise temperature: 20-25K
- + Power: 400 mW (16x ATF35143)



2-path LNA demonstrator

Single-ended LNA

- + Input: $50\ \Omega$ ($S_{11} < -10\text{dB}$)
- + Bandwidth: 0.8 - 2 GHz
- + Measured NT: $<30\ \text{K}$ @ room temp.
- + Power: 60 mW
- + 4x ATF35143



2. Design antenna having Z_{opt}

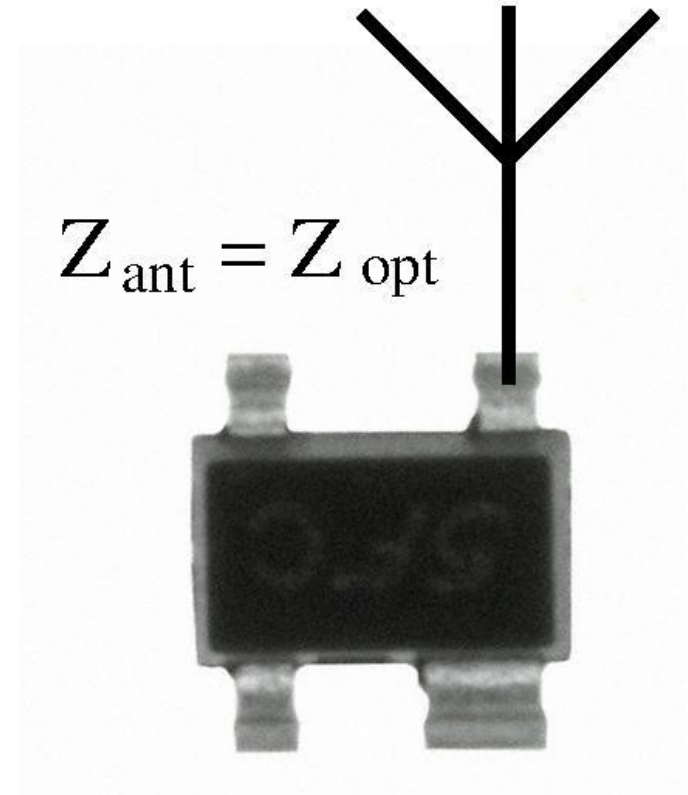
Antenna designed to have impedance close to Z_{opt}

Only 1 amplification path::

Reduce semiconductor cost

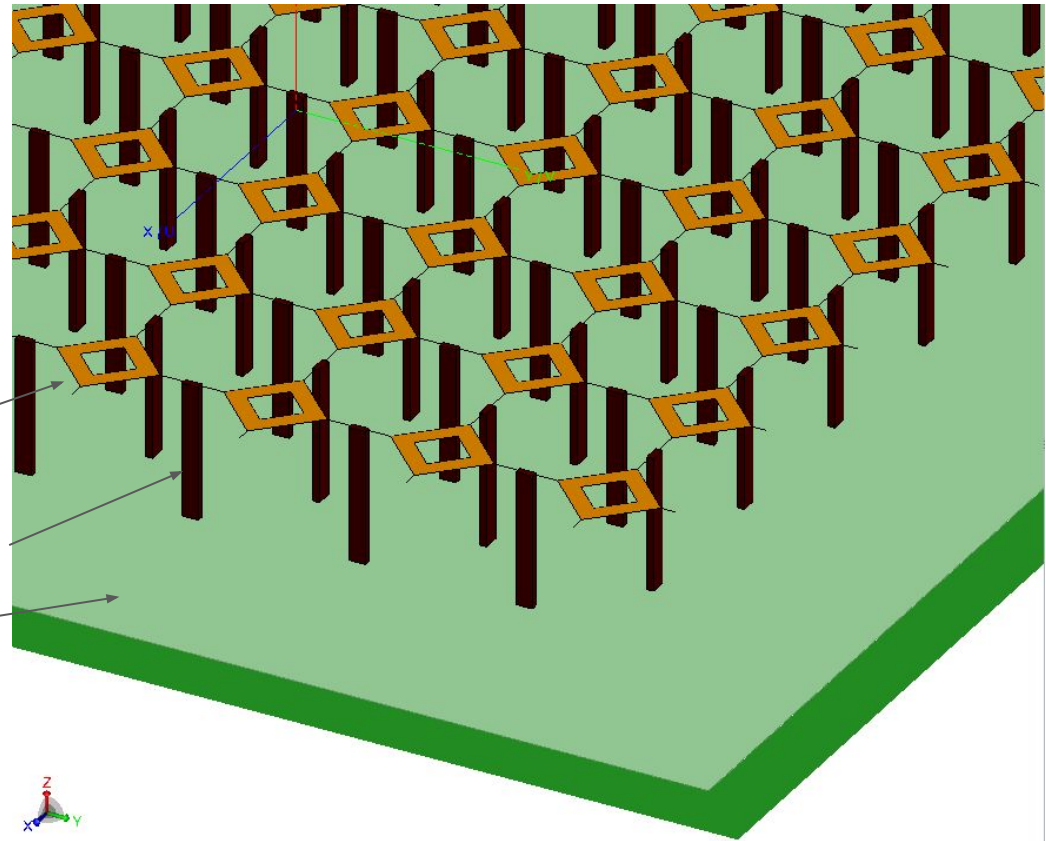
Reduce power consumption

Simpler

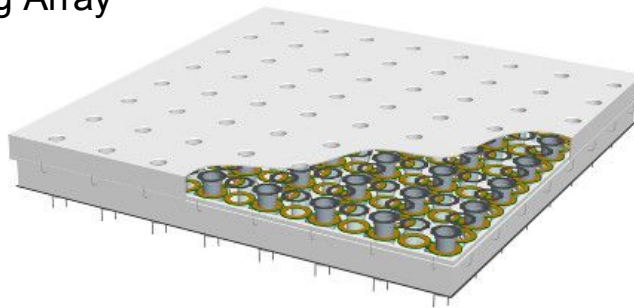


Aperture array example

- Construction
 - Copper wires and pads
 - Thin PCBs with LNA
 - Ground plane

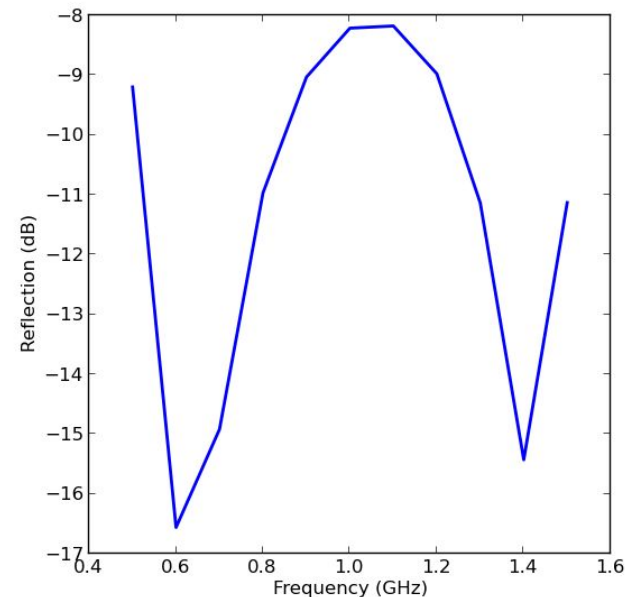
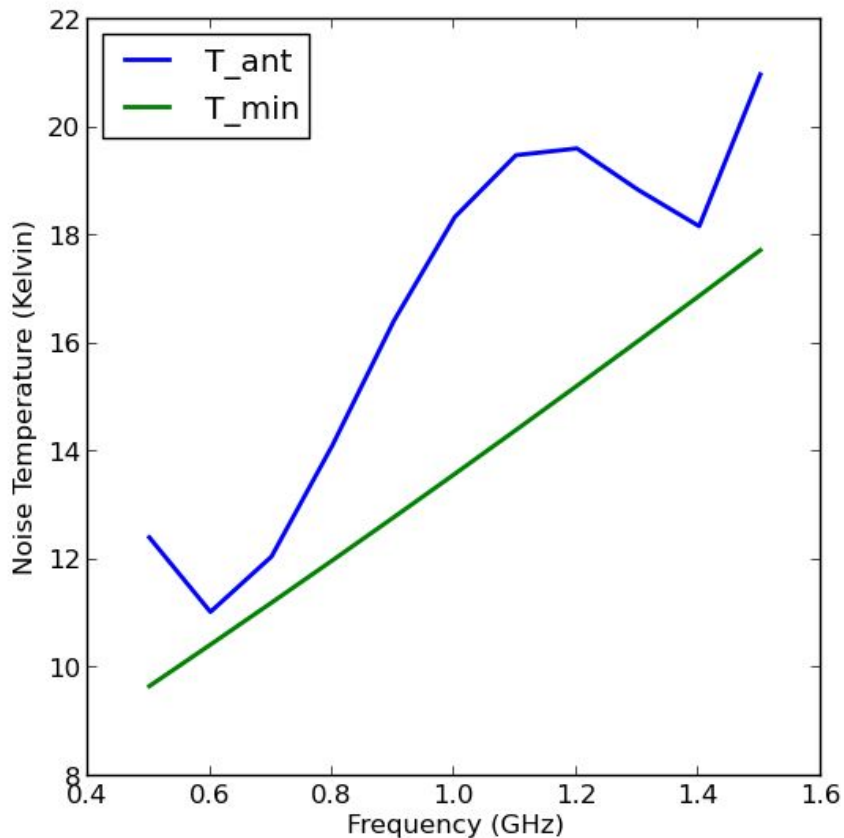
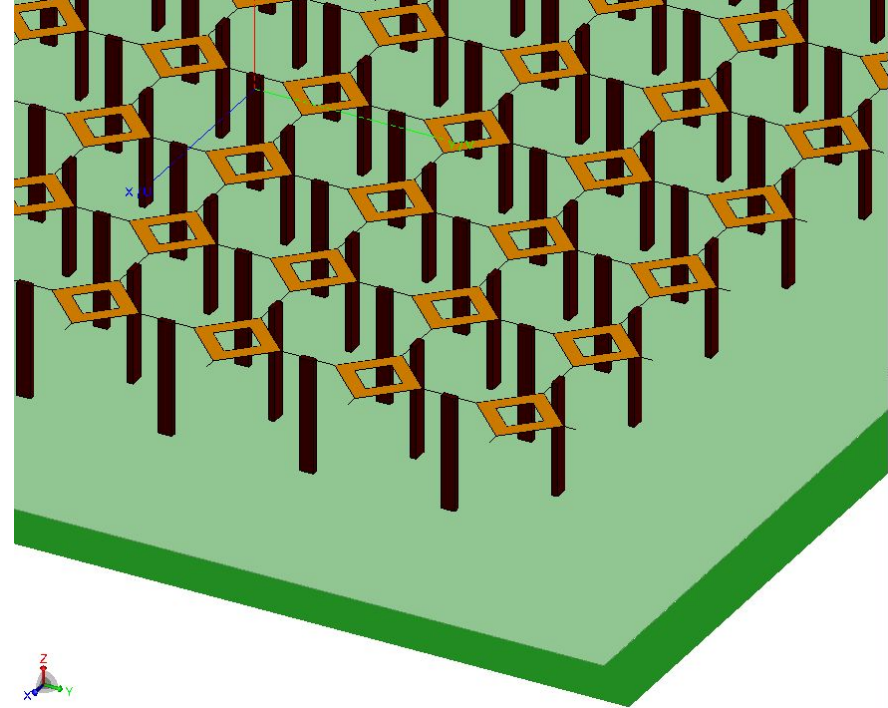


Based on Manchester's
Octal Ring Array



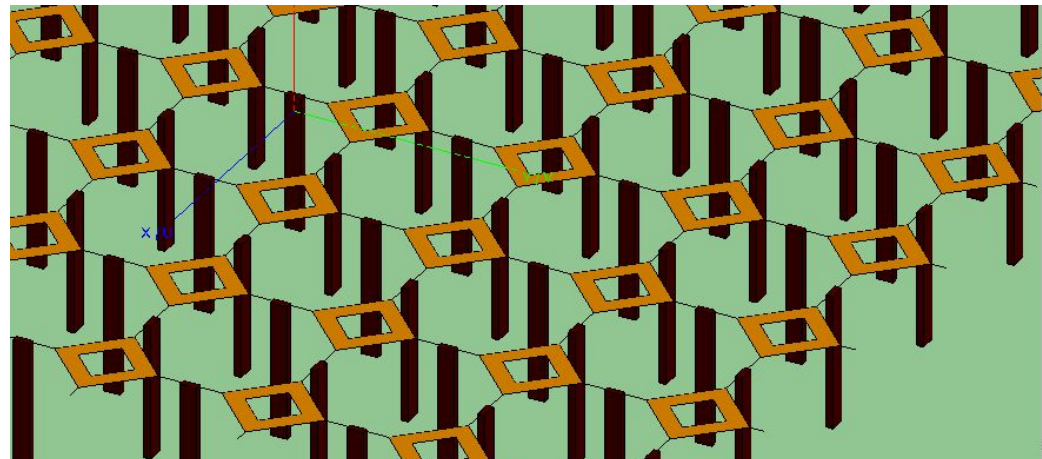
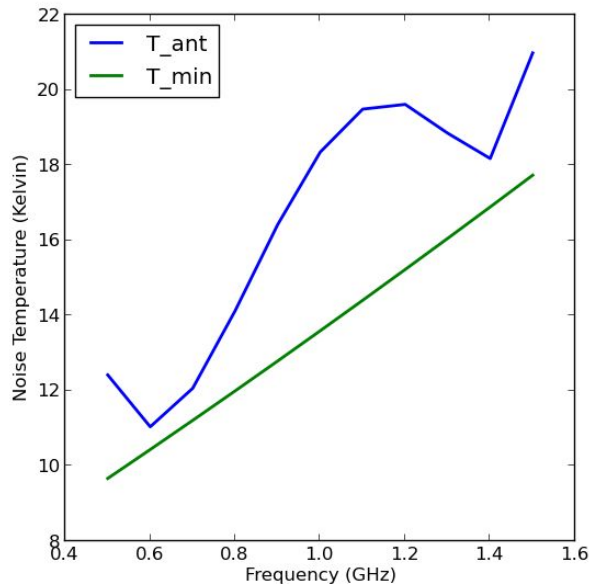
Apperture array example 1

- 12 cm between elements
(Grating effect >45 degrees @ 1.5GHz)



Aperture array example 1

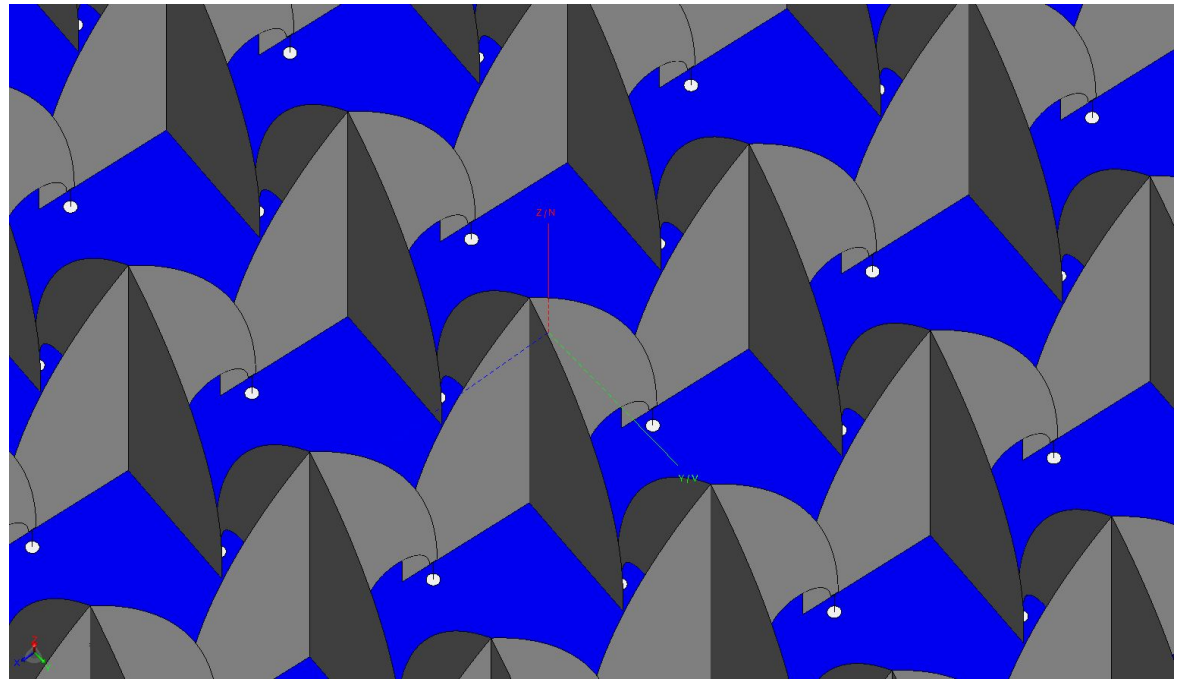
- 0.5 - 1.5 GHz
- Noise Temp <30 K (LNA & antenna loss)
- 100mW (for 30 dB gain) / LNA
- Dual-polarised



Apperture array example 2

12 cm between
elements

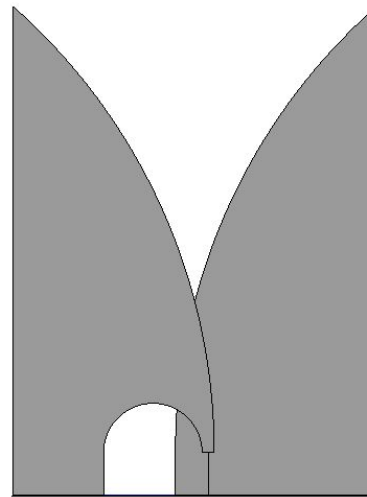
(Grating effect >45
degrees @ 1.5GHz)



Based on ASTRON's design

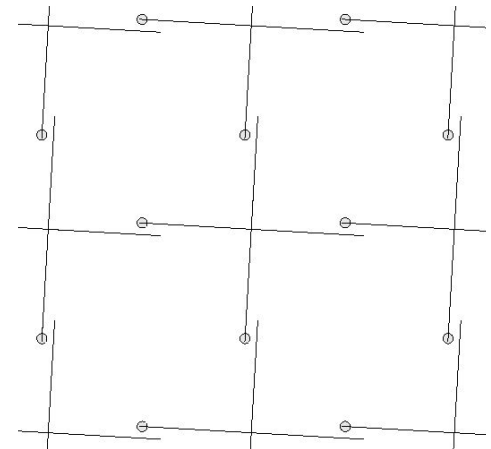


Side view



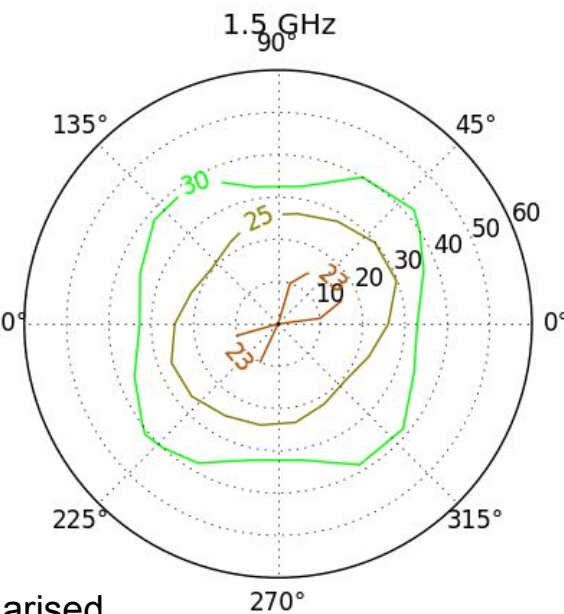
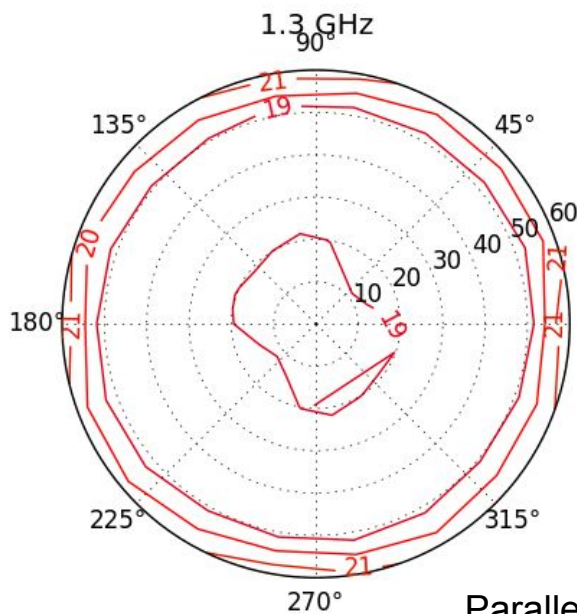
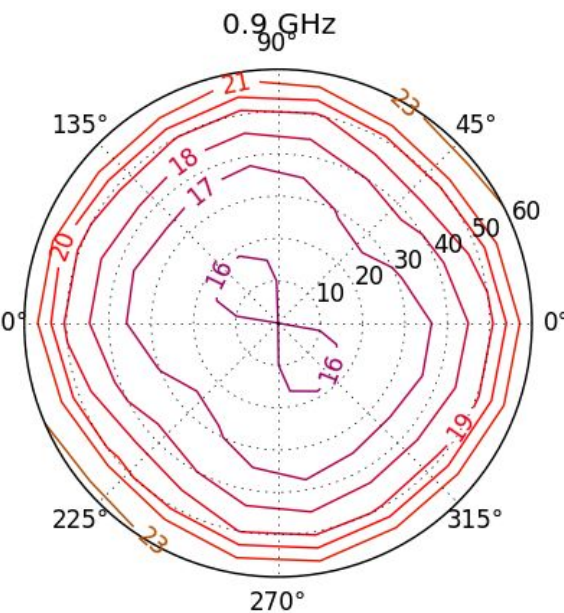
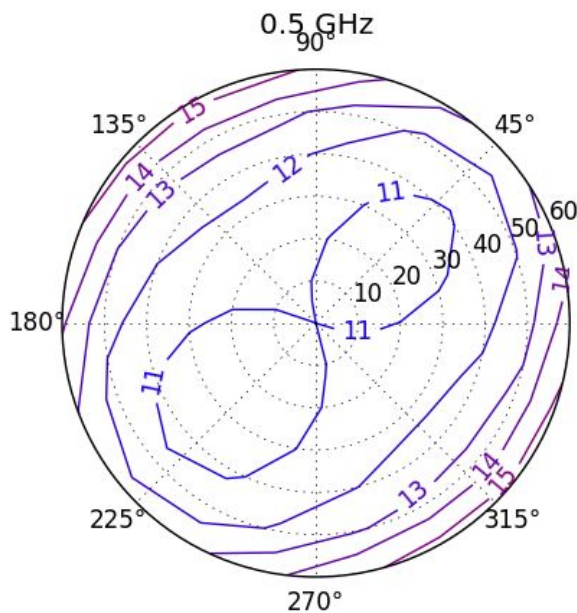
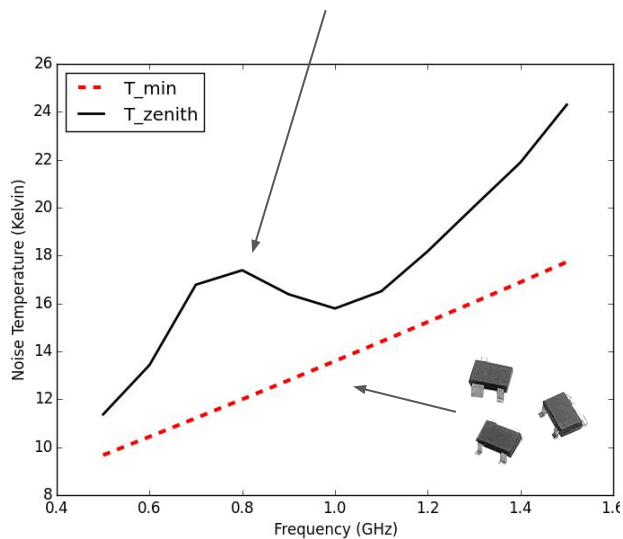
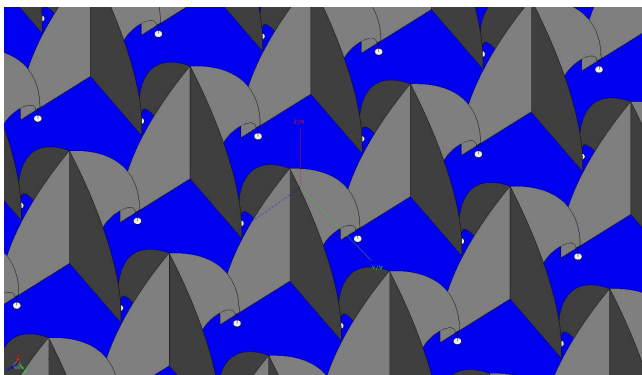
LNA

Top view



Apperture array example 2

Receiver Noise Temperature (Kelvin)

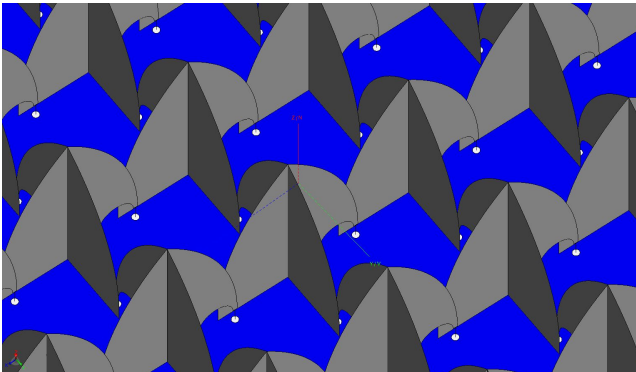
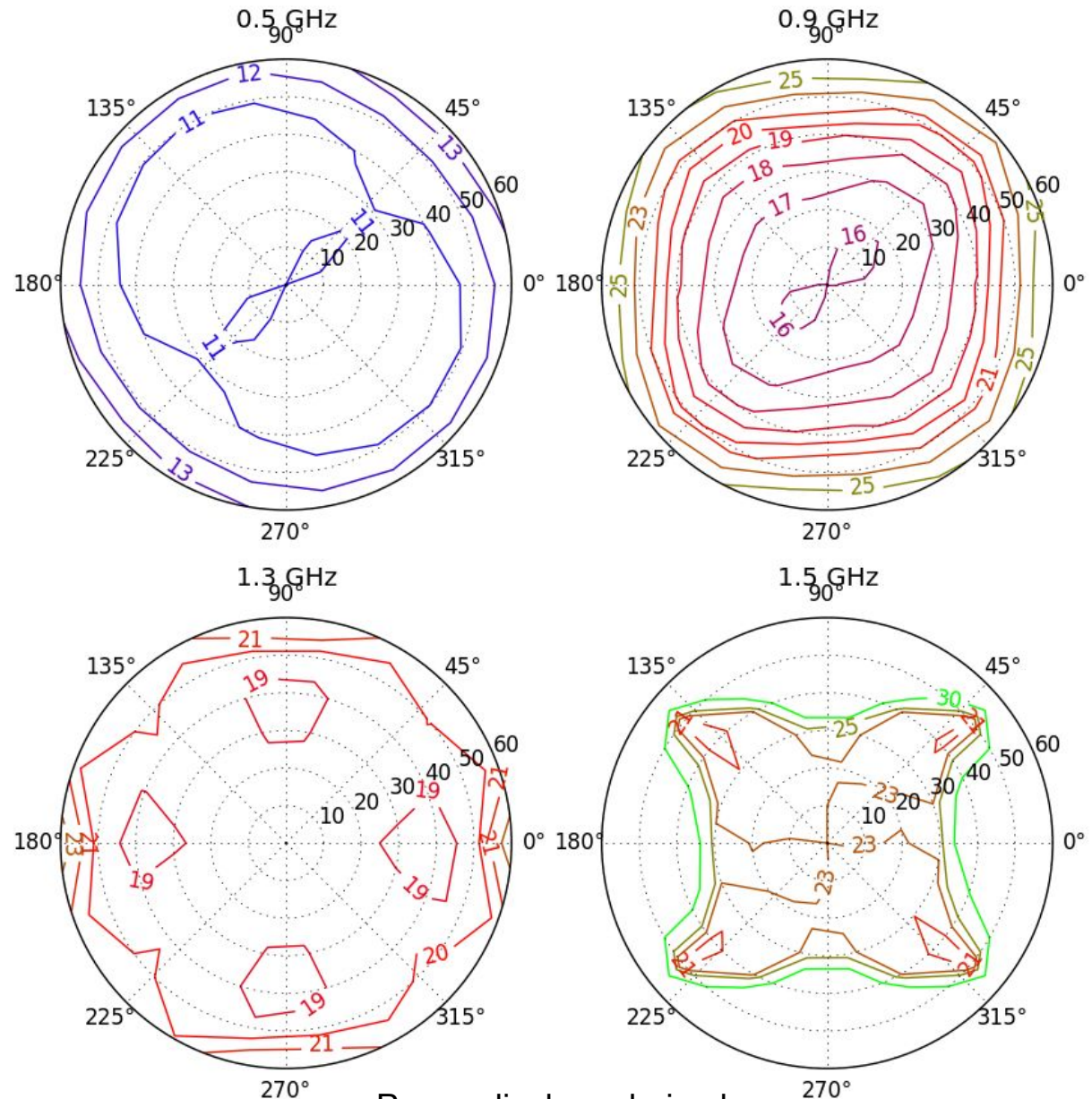


Parallel polarised

Apperture array example 2

- $T_{\text{rec}} < 30 \text{ K}$
(Mostly $< 20 \text{ K}$)
 - 0.5 - 1.5 GHz
 - 0-45° Zenith
- $< 50 \text{ mW}$ / LNA
(for 25 dB gain)
- Single ended

Receiver Noise Temperature (Kelvin)



What about S_{11} ?

Not important, but ...

Quite easy to design HEMT LNA to have Z_{in} close to Z_{opt}^* .

A good match to Z_{opt}^* then also give a good match to Z_{in}

Conclusion

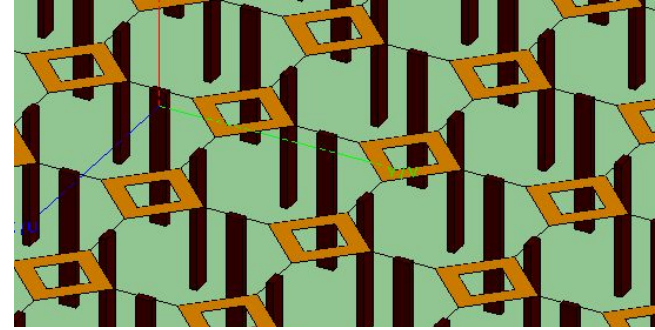
Using commercial HEMTs for LNAs may be a good and easy solution for MFAA.

Requires impedance matching:

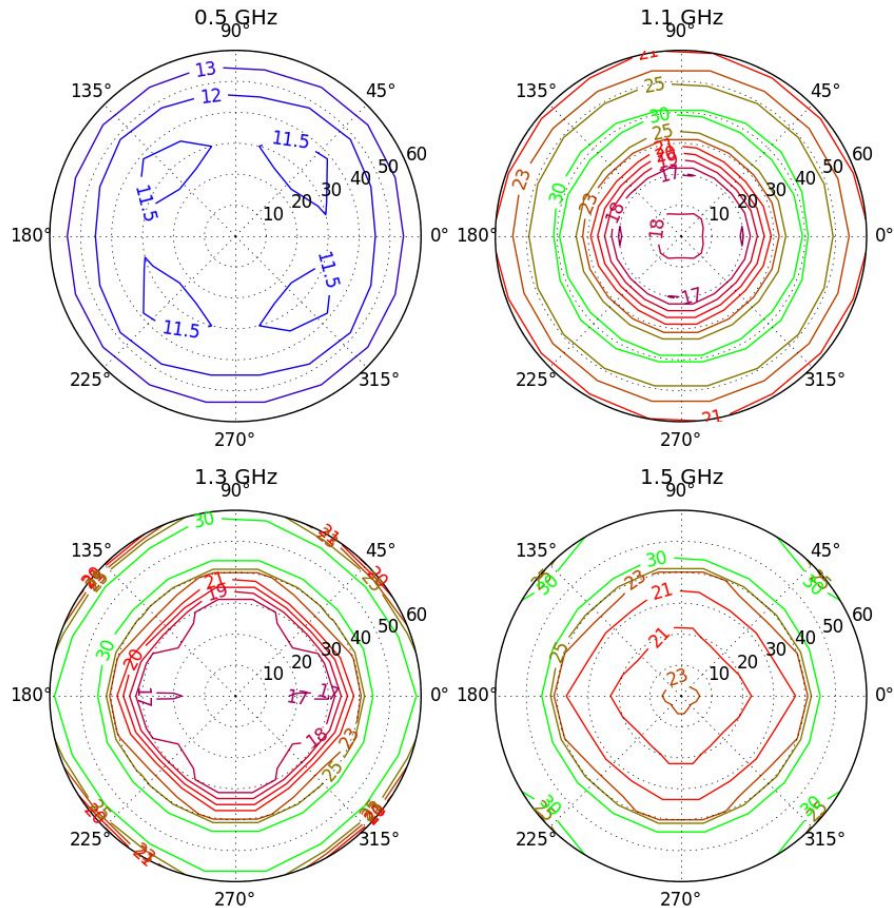
- By using a multipath amplifier
 - Advantage: Seperate LNA & antenna design
- By matching the antenna to the HEMT
 - Advantage:
 - Lower semiconductor cost
 - Lower power consumption
 - Simpler

Apperture array example 1

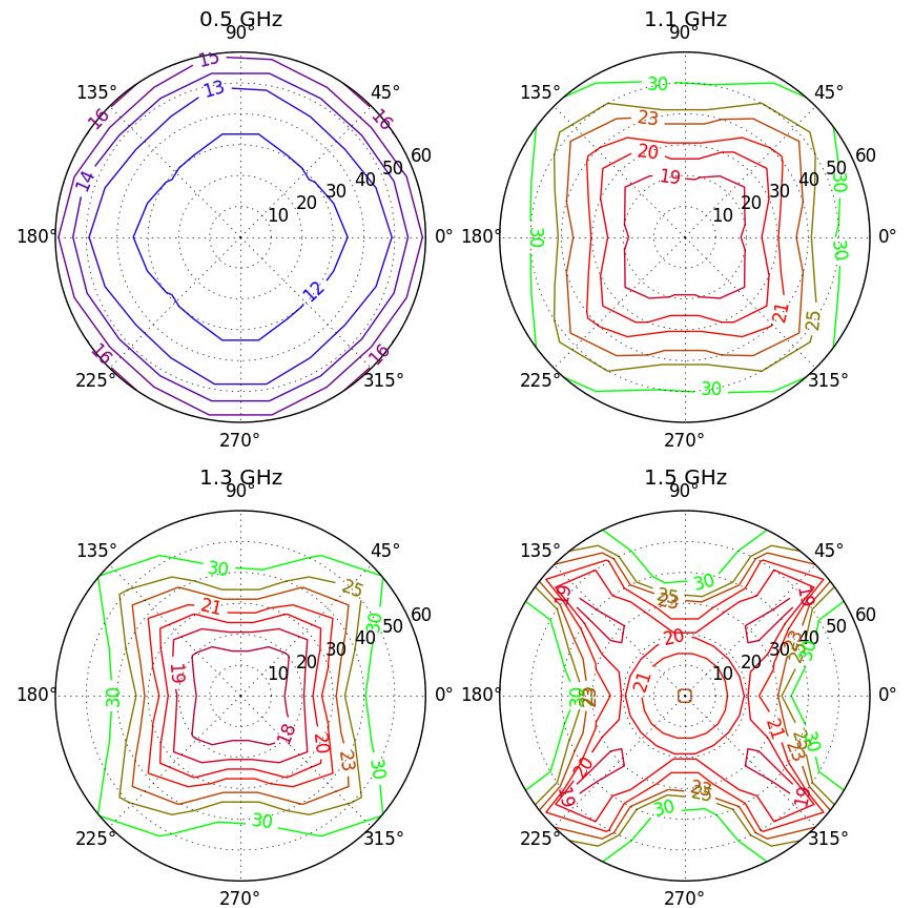
Receiver Noise Temperature (Kelvin)



Parallel polarised



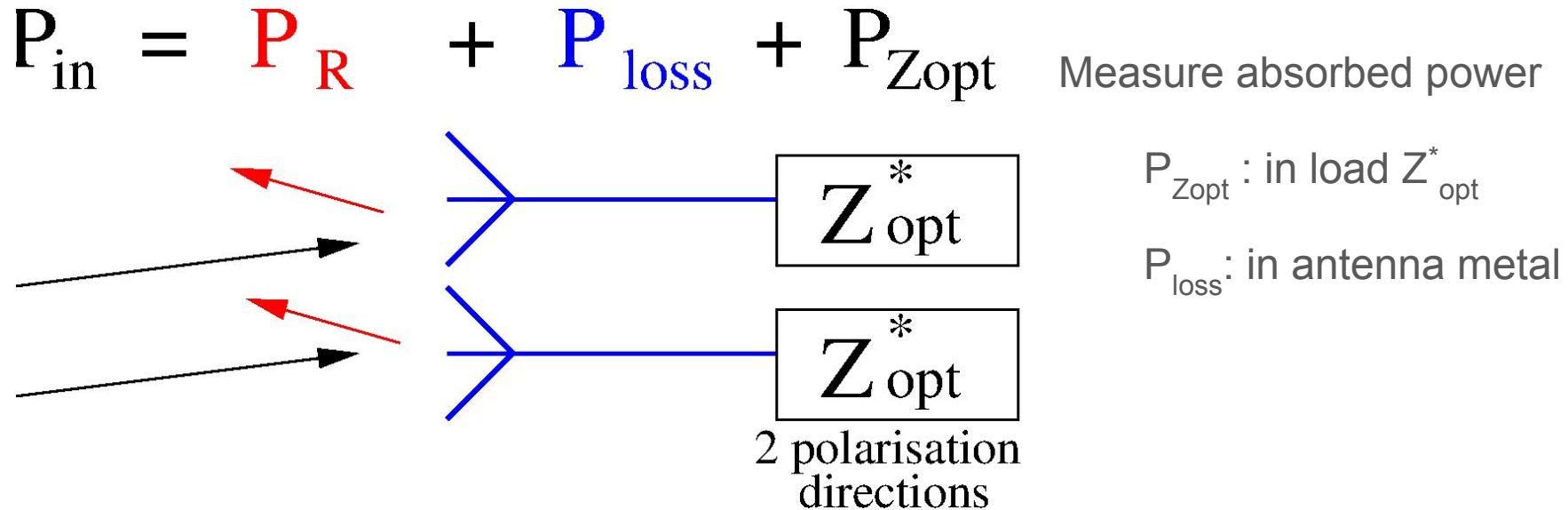
Perpendicular polarised



Receiver temperature

Load antenna with Z_{opt}^* (at each LNA)

Simulated incoming wave: P_{in} = power per array element



Receiver temperature

Load antenna with Z_{opt}^* (at each LNA)

$$T_{\text{rec}} = T_{\text{min}} + 2T_n(P_{\text{in}}/P_{Z_{\text{opt}}} - 1) + T_0 P_{\text{loss}}/P_{\text{in}}$$

T_0 - ambient temperature

T_{min} - amplifier minimum noise temperature

T_n - amplifier mismatch noise temperature ($\approx T_{\text{min}}$)

$T_n = 2 R_n G_{\text{opt}} T_0$ (G_{opt} - optimum conductance, R_n noise resistance)

$$T_{\text{rec}} / A_{\text{element}} = T_{\text{LNA+antenna}} / A_{\text{eff}}$$

Amplifier noise measurement: Procedure

3 Output power measurements

- P_s : Input shorted
- P_o : Input open
- P_t : Input terminated at room temperature T_R

Y-factor: $Y \equiv 2P_t / (P_s + P_o)$

Correction factor: $\alpha = 1 + S_{11}^2 + 2\text{Re}(S_{11}) (P_s - P_o) / (P_s + P_o)$

$\alpha \approx 1$ when S_{11} small.

LNA Noise temperature: $T_{\text{lna}} \approx T_R / (2\alpha Y - 1)$

Amplifier noise measurement: Simulated example

